

# Dark Matter Detection with Kinetic Inductance Detectors

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# Overview

Review: SuperCDMS SNOLAB low-mass reach with iZIPs and CDMS-HV

Motivation for kinetic inductance detectors in DM detection

## Multiplexability

- improve fiducial volume for ionization-based surface-event rejection
- phonon-only surface-event rejection and fiducial volume definition

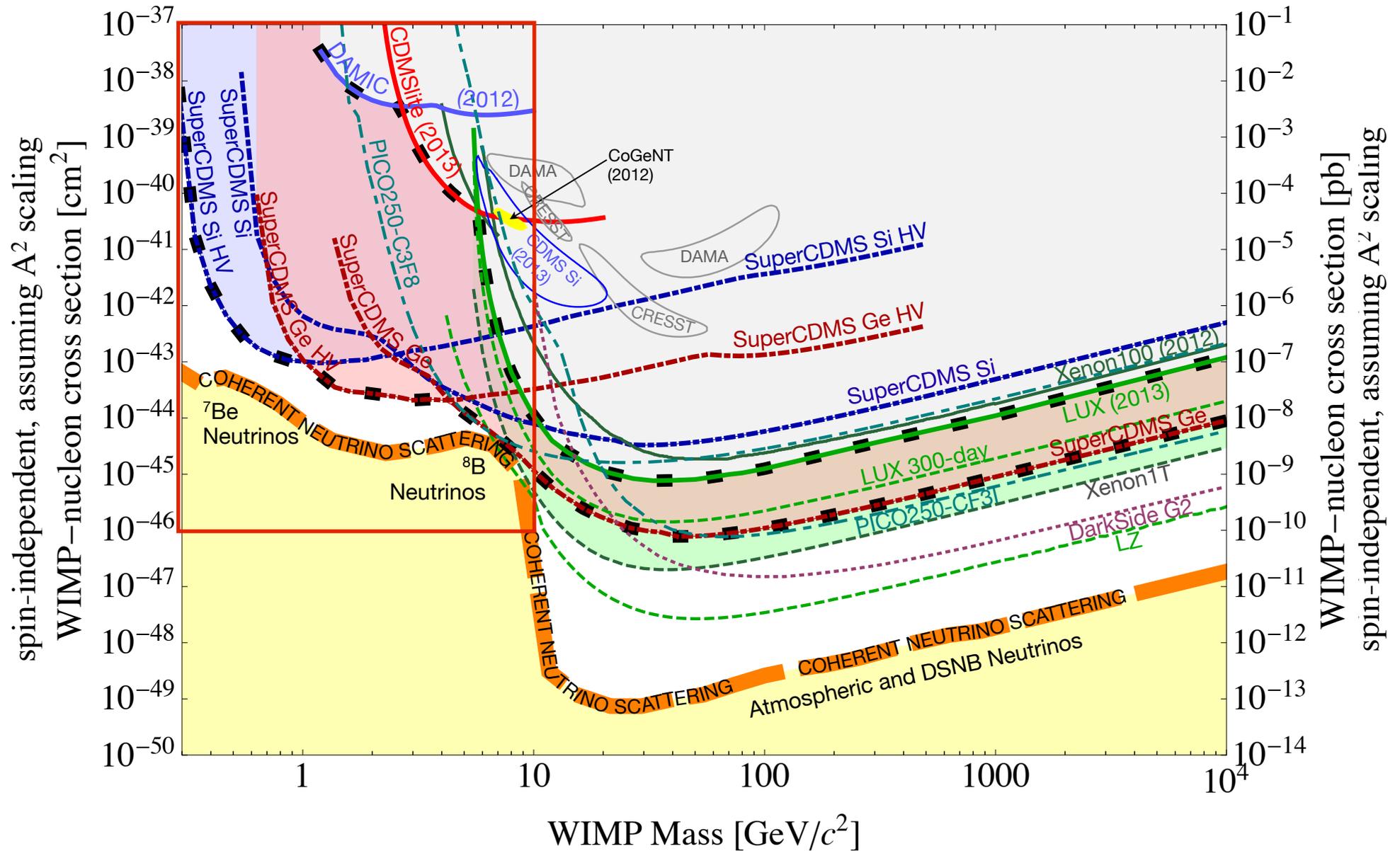
## Sensitivity

- reduced threshold, fiducial volume definition for CDMS-HV operation

Expected sensitivity

Progress to date

# Searching for Low Mass DM with SuperCDMS SNOLAB



# Searching for Low Mass DM with SuperCDMS SNOLAB

## Better energy resolution

Recently developed HEMTs + modified amplifier design:

→  $\sigma_q = 300 \text{ eV}_{ee}$  will improve to  $\sigma_q = 100 \text{ eV}_{ee}$

Reduced  $T_c$  (60 mK) phonon sensors are baseline

→  $\sigma_p = 200 \text{ eV}_p$  will improve to  $\sigma_p = 50 \text{ eV}_p$

Ge low-mass dominates down to  $\sim 5 \text{ GeV}$

Ionization asymmetry-based rejection to  $\sim 10 \text{ GeV}$

Phonon asymmetry-based rejection to  $\sim 5 \text{ GeV}$

HV search at 100V bias extends down to 1 GeV with Ge, to 0.5 GeV with Si

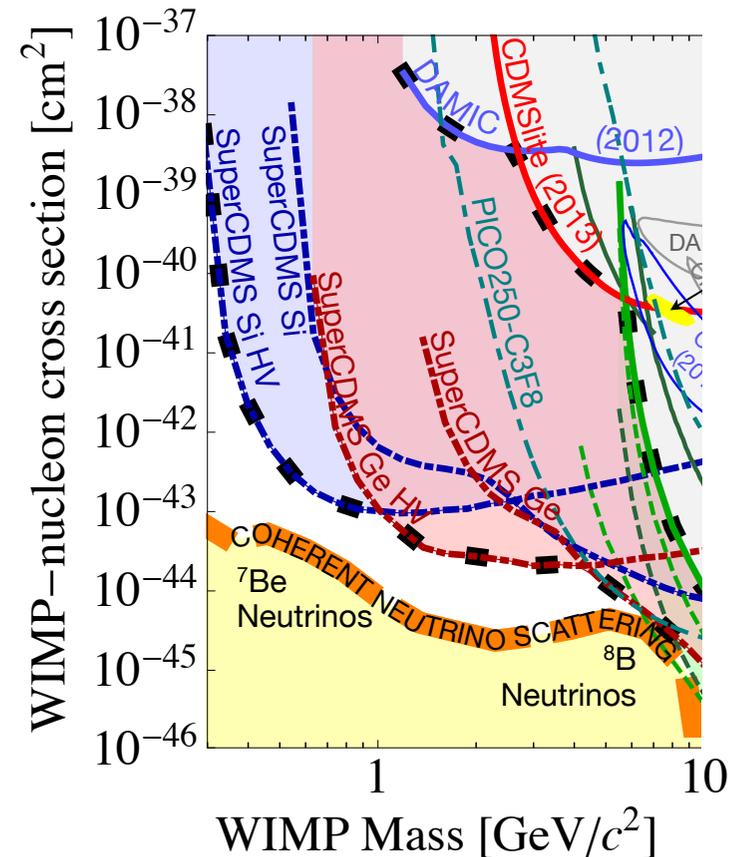
Also, need to reduce backgrounds  
(lower cross sections)

Compton background reduced by improved materials selection, shielding (200x)

$^{210}\text{Pb}$  background from Cu will be reduced to levels observed on Ge (20x)

Should enable reach approaching coherent solar neutrino scattering background

$^{32}\text{Si}$  (225 keV endpoint, 150 yr half-life) contamination in detector-grade Si unknown



# Beyond SuperCDMS SNOLAB Baseline

Better  $\sigma_p$  (10 eV goal,  $T_c = 40$  mK) will extend phonon asymmetry-based rejection

Higher voltage operation

Initial tests of p-type point contact (PPC) Ge detectors, complementary amorphous Ge and Si blocking layers suggests  $V_b \sim 400$  V

→ lower threshold, stretch ER background

Single e-h pair detection? (Pyle talk)

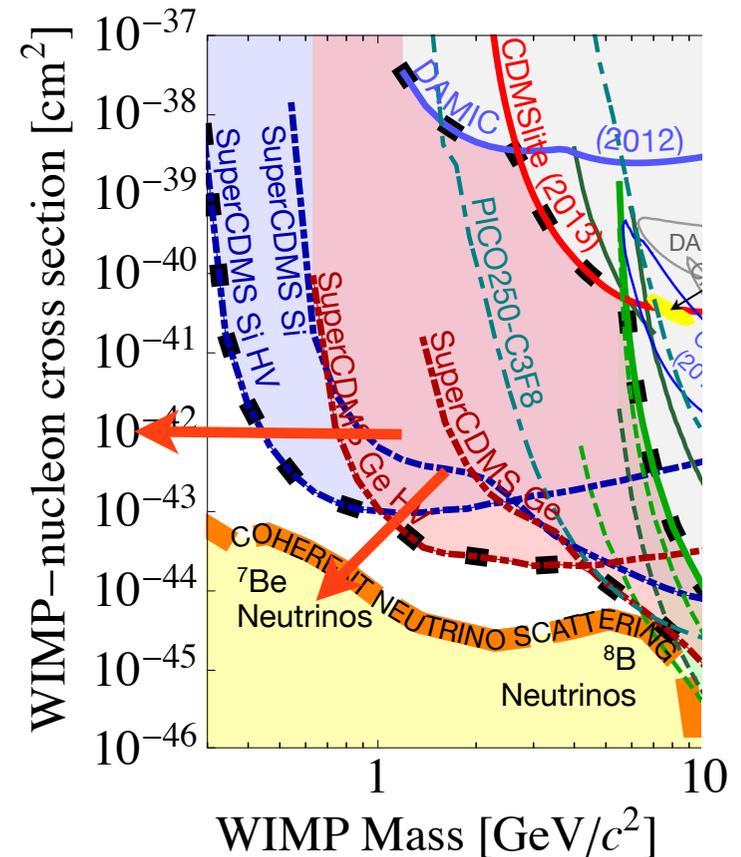
At large  $V_b$  and excellent  $\sigma_p$ , expect peaks in phonon spectrum due to integral e-h pairs

$V_b = 100$  eV, 3 eV to create e-h pair → peak at 103 eV

At  $\sigma_p \sim 10$  eV<sub>p</sub> and  $V_b = 100$  V, single e-h pair peaks become resolvable

At  $\sigma_p \sim 3$  eV<sub>p</sub> and  $V_b = 100$  V, single e-h pair peaks are separated and NRs can occupy empty space (because more recoil energy per e-h pair)

Sub-GeV dark matter at CNS limit accessible with reasonable extrapolations of current technology

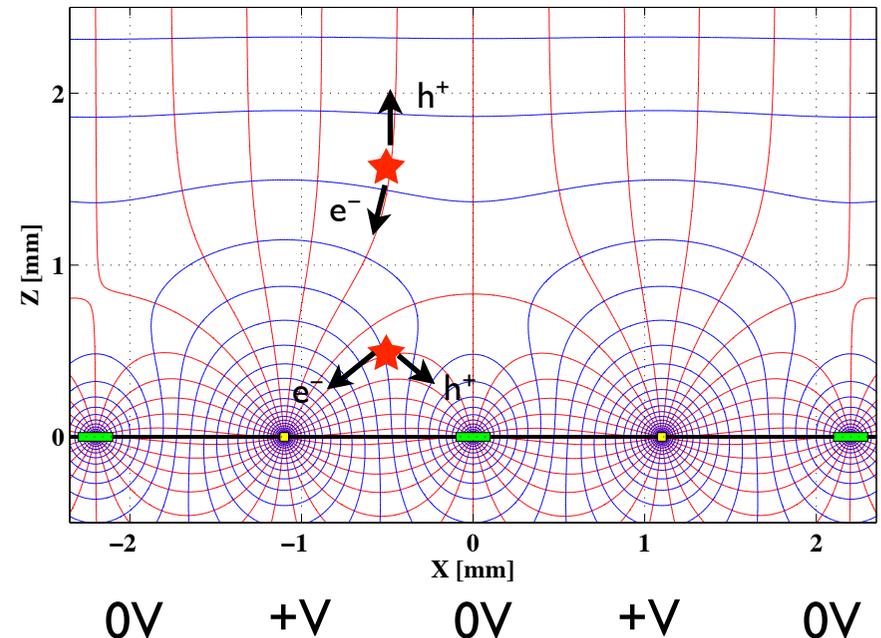
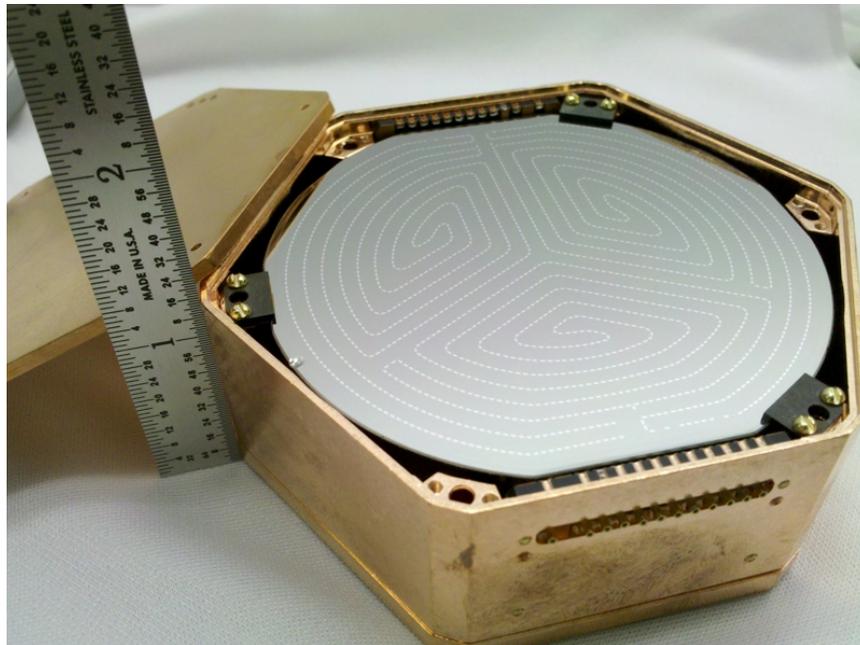
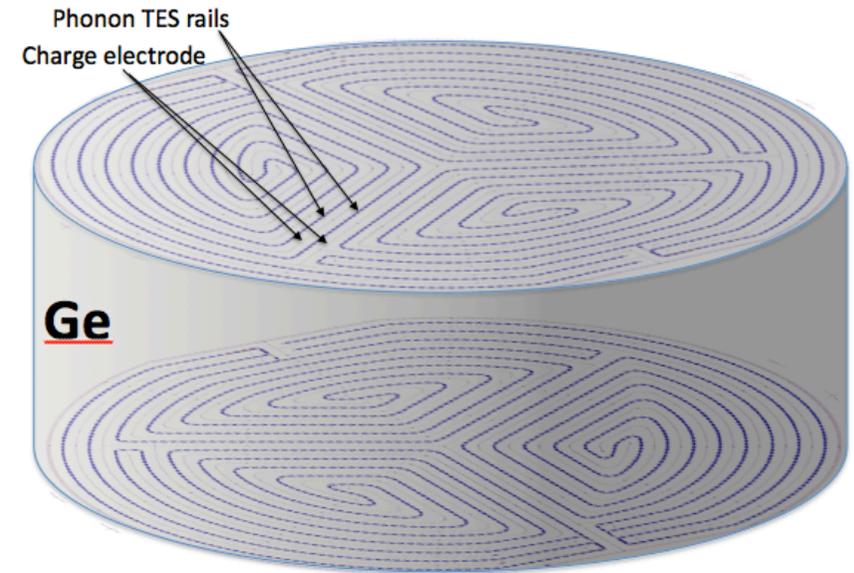


# SuperCDMS iZIP Ionization-Based Surface Event Rejection

Alternating ground and biased electrodes yield rejection via sophisticated field configuration:

Bulk events have z-symmetric hole/  
electron collection

Surface ERs are z-asymmetric



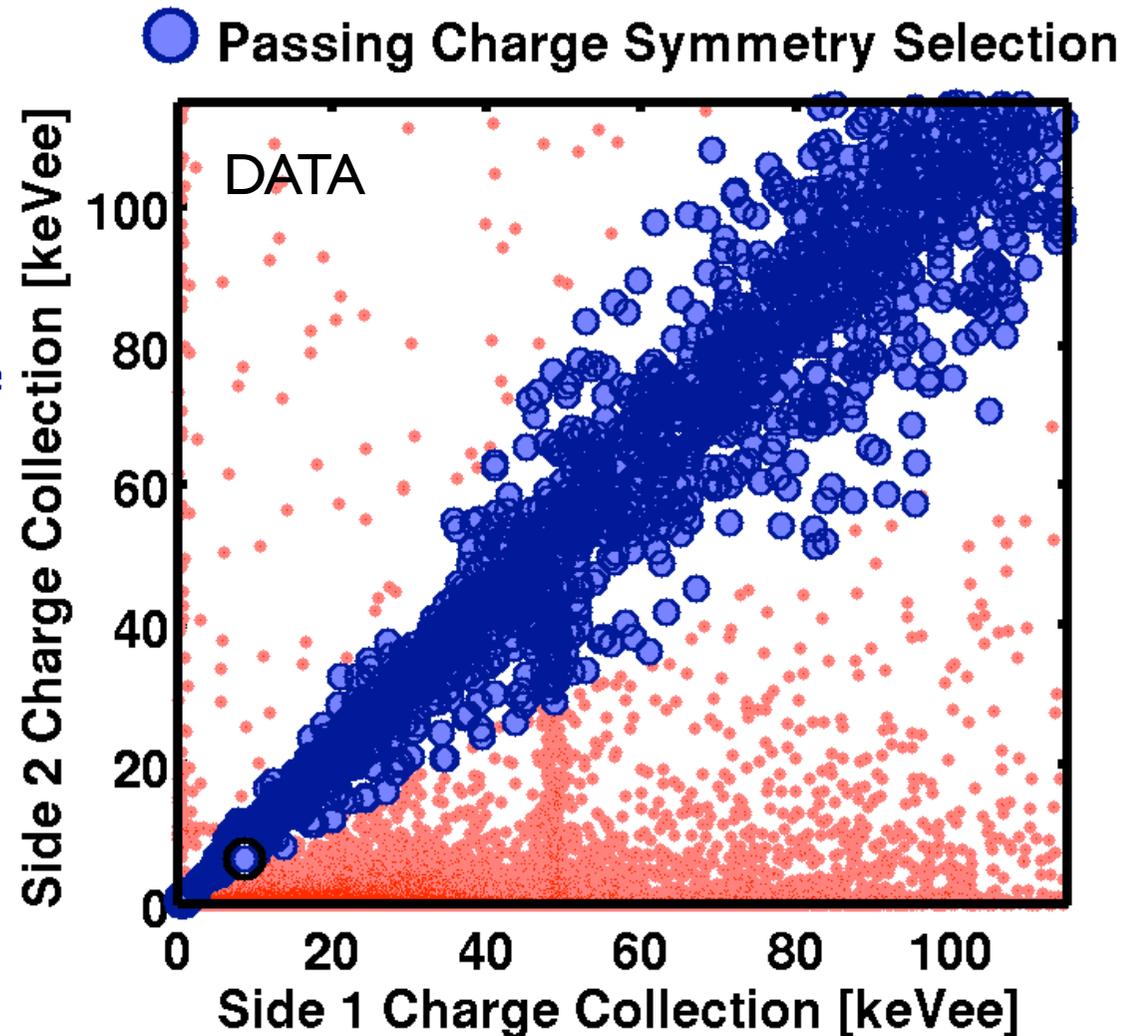
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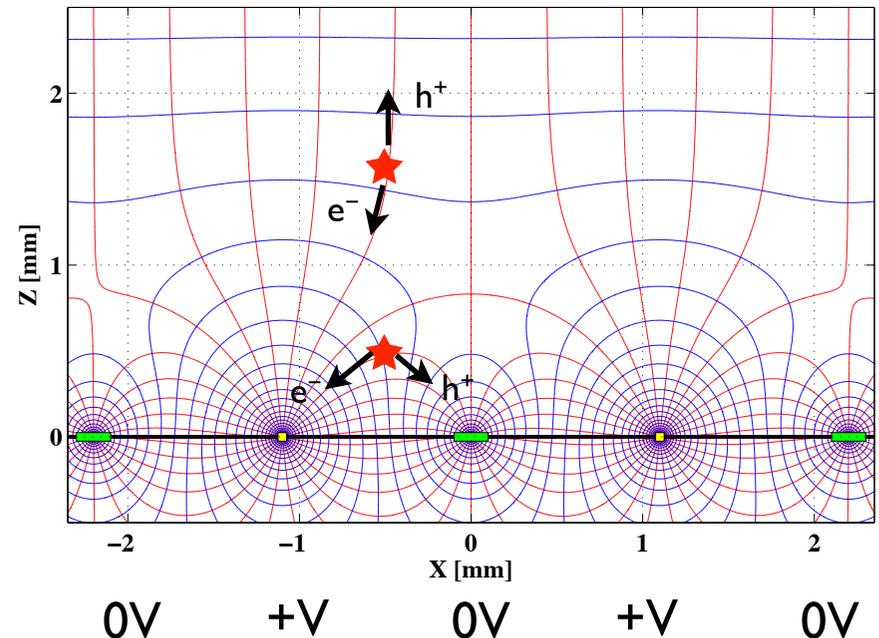
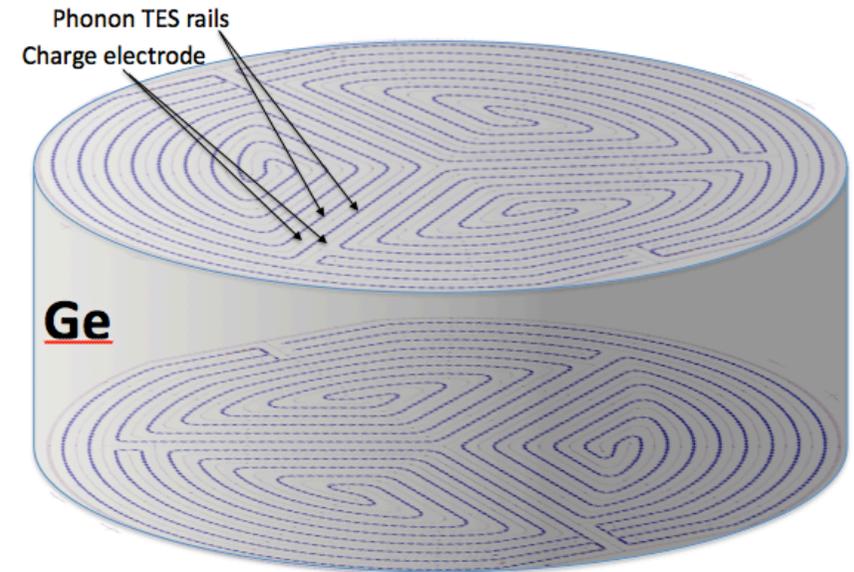
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Field strength

High field near surface raises ionization collection for surface electron recoils



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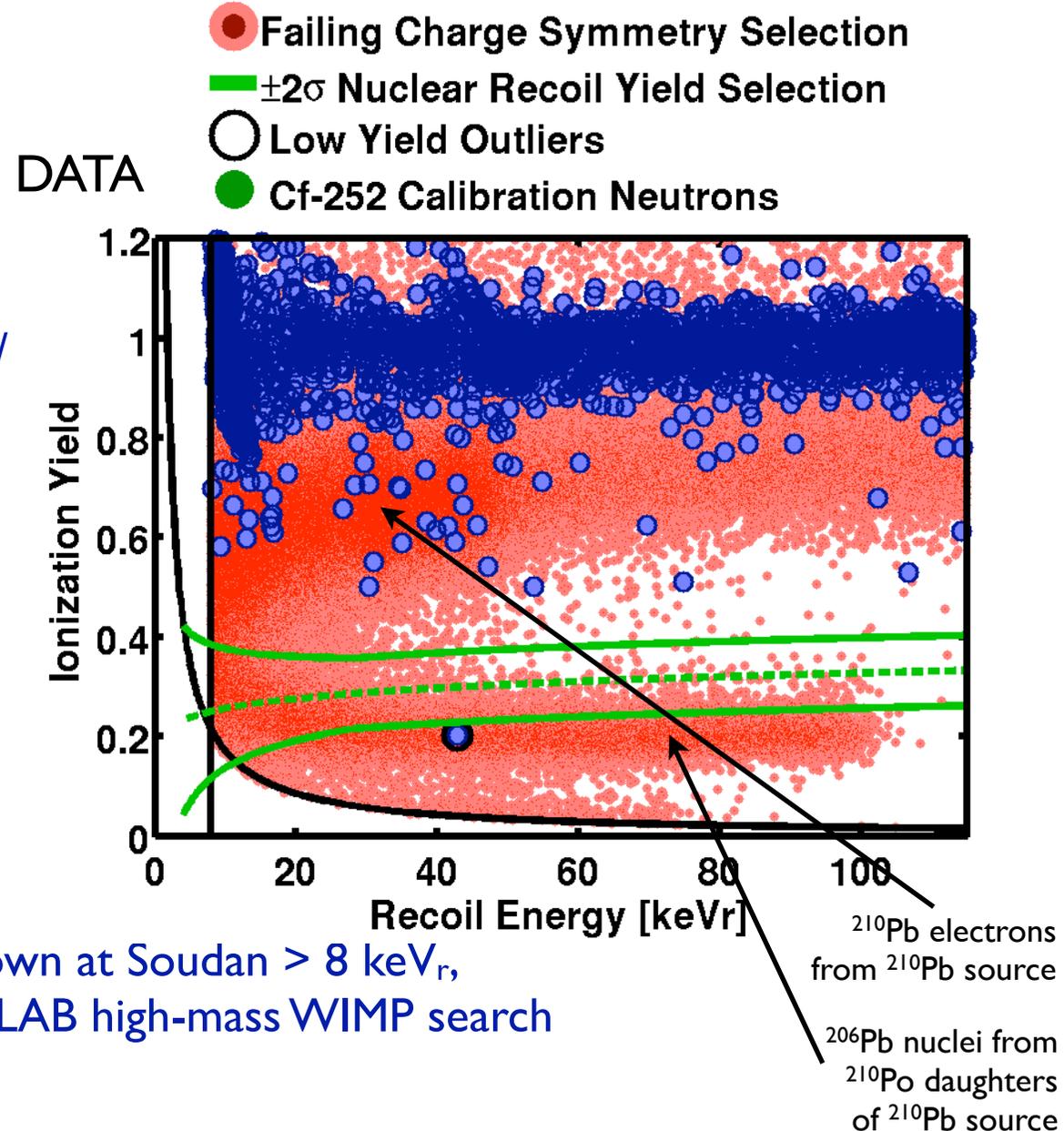
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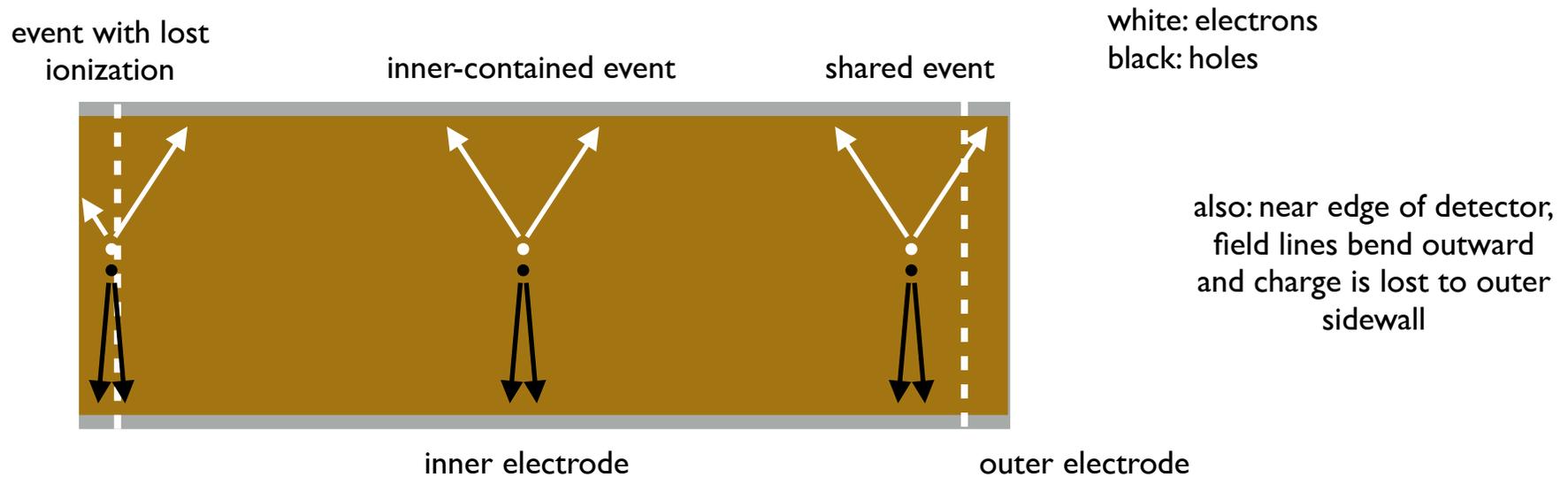
$^{206}\text{Pb}$  nuclear recoils visible

Achieves excellent rejection

Surface ER rejection  $\sim 1 \times 10^{-5}$  shown at Soudan  $> 8 \text{ keV}_r$ , sufficient for SuperCDMS SNOLAB high-mass WIMP search



# SuperCDMS iZIP Ionization-Based Surface Event Rejection



## Fiducialization: inner-electrode containment

$e^-$  propagate obliquely at low biases:  $33^\circ$  from  $z$

need full containment to trust both  $e$  and  $h$  side measurements to measure asym

fiducial volume fraction for inner electrode = 55%

## Implications

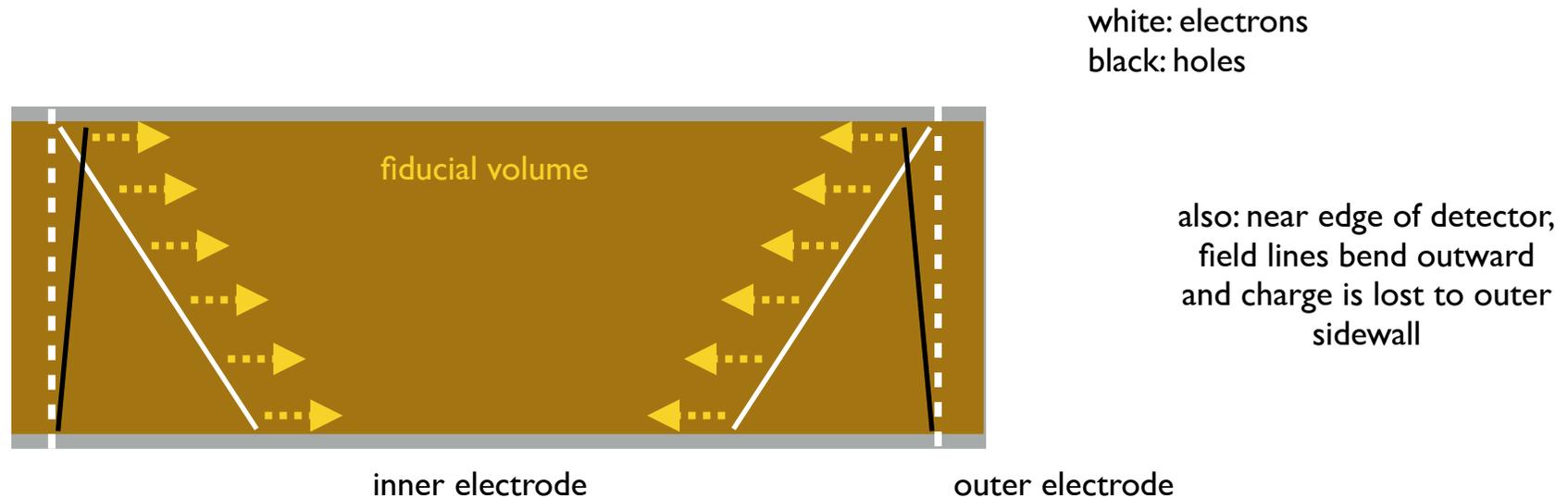
Efficiency, obviously

Threshold for surface-event rejection set by ionization resolution:

doesn't benefit from excellent phonon energy resolution

Would like to be able to rely on holes alone

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Phonon energy resolution much better than ionization

200 eV<sub>p</sub> now vs. 300 eV<sub>ee</sub>

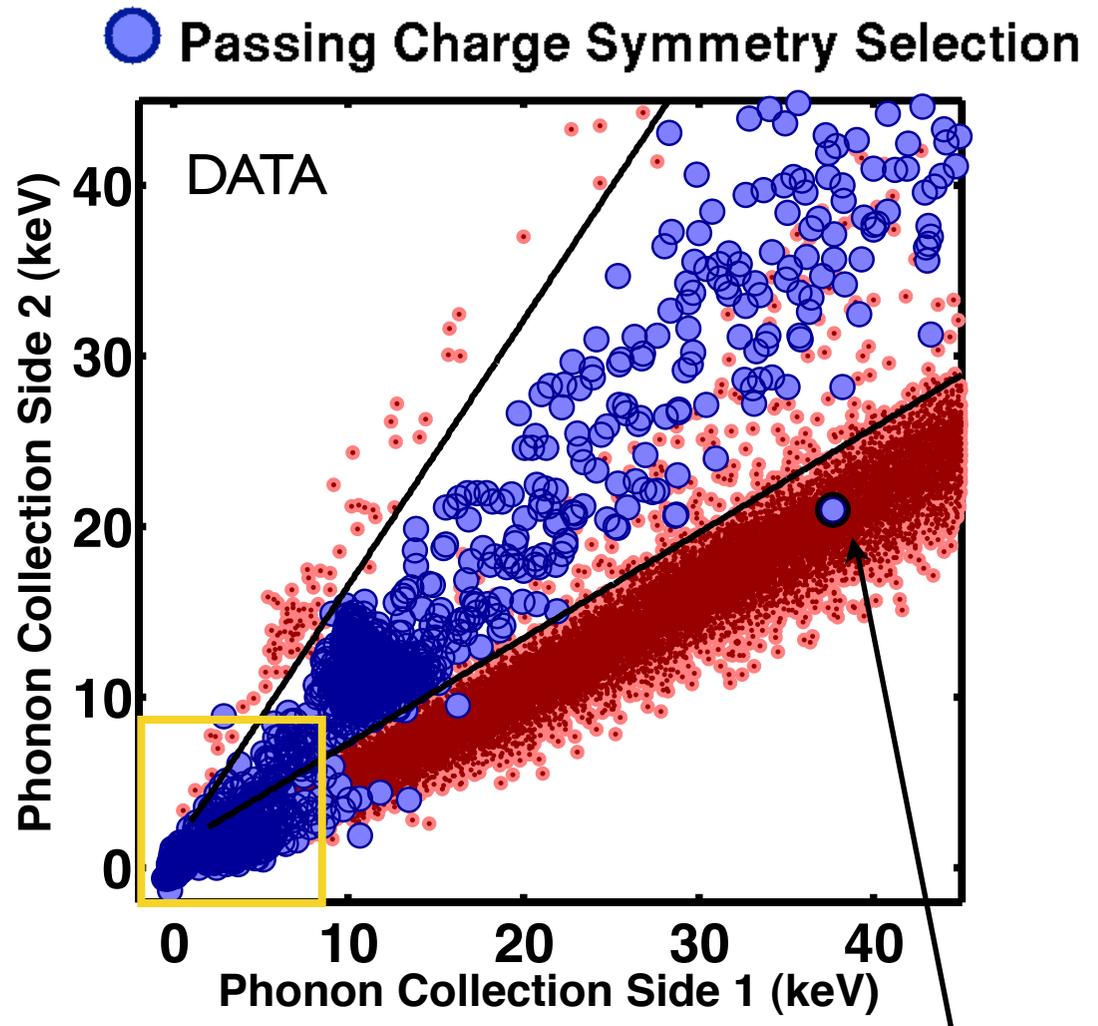
Ionization asymmetry only useful  
 $\geq 8$  keV<sub>r</sub>,  $\geq 10$  GeV WIMP mass

Will improve for SNOLAB:

→ 50 eV<sub>p</sub> vs. 100 eV<sub>q</sub>,

$\geq 3$  keV<sub>r</sub>,  $\geq 3$  GeV WIMP mass

At low mass, define asymmetry using phonons only



outlier event at  
42 keV<sub>r</sub>, 8 keV<sub>ee</sub>  
in prior slide; easily  
rejected using  
phonon asymmetry

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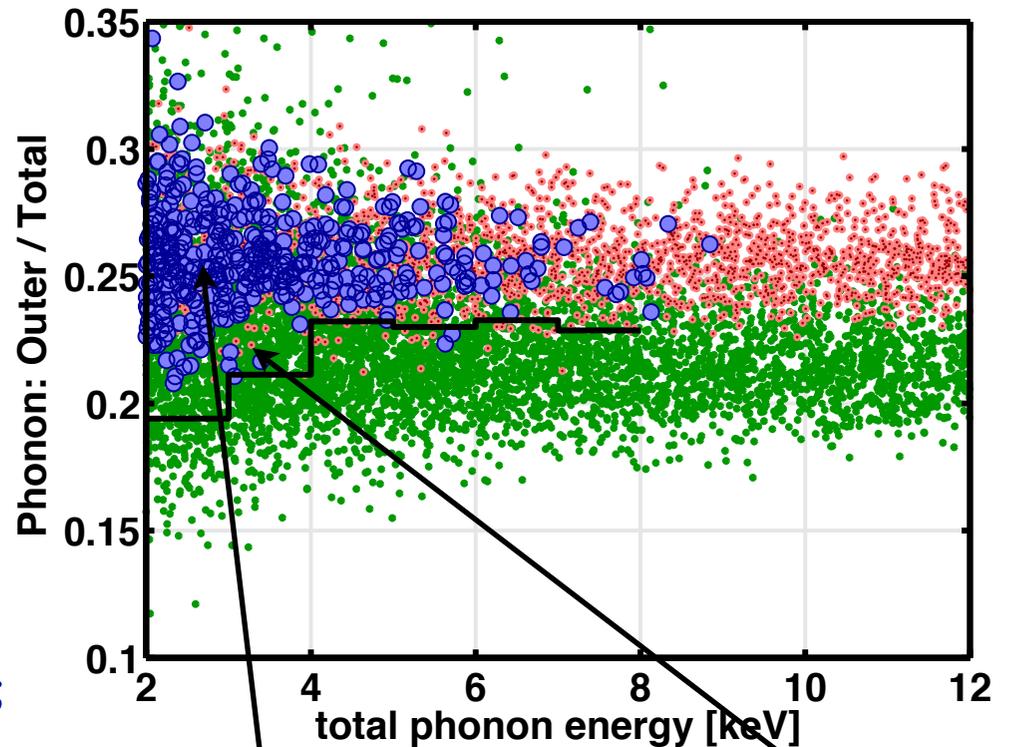
At low mass, define asymmetry using phonons only

But can't use ionization to define fiducial volume (radial)

Use phonon radial partition instead; but phonon radial rejection of outer wall events not as good, yields analysis limited by radially misid'd <sup>210</sup>Pb; 30% fiducial volume

data-driven simulation

- Failing charge fiducial and NR band
- Passing charge fiducial and NR band
- Cf-252 Neutrons



events that would have been accepted as bulk nuclear recoils by ionization asymmetry: rejected by phonon fiducialization cut

phonon radial rejection cut limited by radial misid of <sup>210</sup>Pb

# SuperCDMS iZIP Phonon-Based Surface Event Rejection

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200 eV<sub>p</sub> now vs. 300 eV<sub>ee</sub>

Ionization asymmetry only useful  
 $\gtrsim 8$  keV<sub>r</sub>,  $\gtrsim 10$  GeV WIMP mass

Will improve for SNOLAB:

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Ideally:

Use phonons to define radial fiducial volume with high efficiency

Use phonons for surface-event rejection to accept hole-only collection

Use ionization signal only for defining bulk electron-recoil rejection (using holes only)

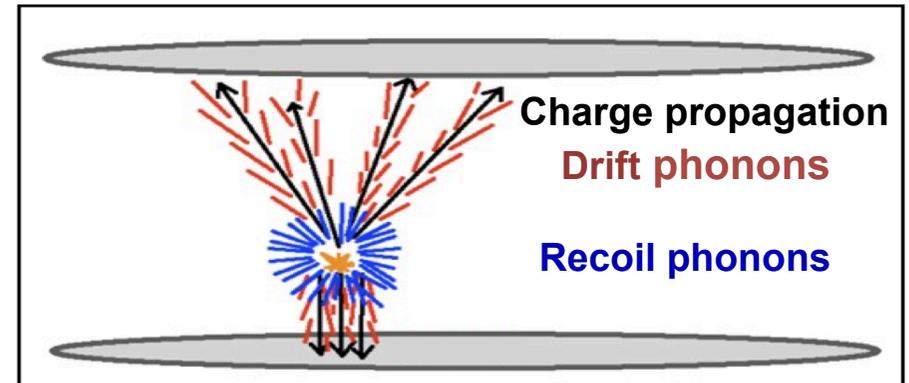
Requires much better position fidelity from phonon signal

# SuperCDMS-HV: Technique and Challenges

Use phonons to measure e-h production via Luke-Neganov phonon production

$$E_p = E_r + N_Q e V_b = E_r + E_{ee} (e V_b / E_{Fano})$$

$$e V_b \gg E_{Fano} = 3 \text{ eV} \rightarrow E_p \gg E_r, E_{ee}$$



Gain =  $(e V_b / E_{Fano})$  requires full drift of e and h through bias voltage

At high bias, phonon emission isotropizes  $e^-$  drift  $\rightarrow$  higher fiducial volume than iZIP

Fiducialization, background rejection critical to sensitivity

For continuum operation (not single e-h mode)

Lost ionization smears ER and NR spectra: not a big problem

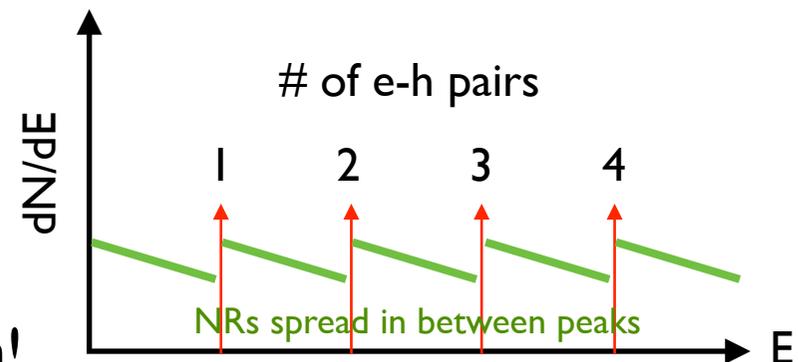
But backgrounds are larger at outer walls, so want to reject to reduce bgnd because continuum operation is a bgnd-limited experiment

For single e-h operation

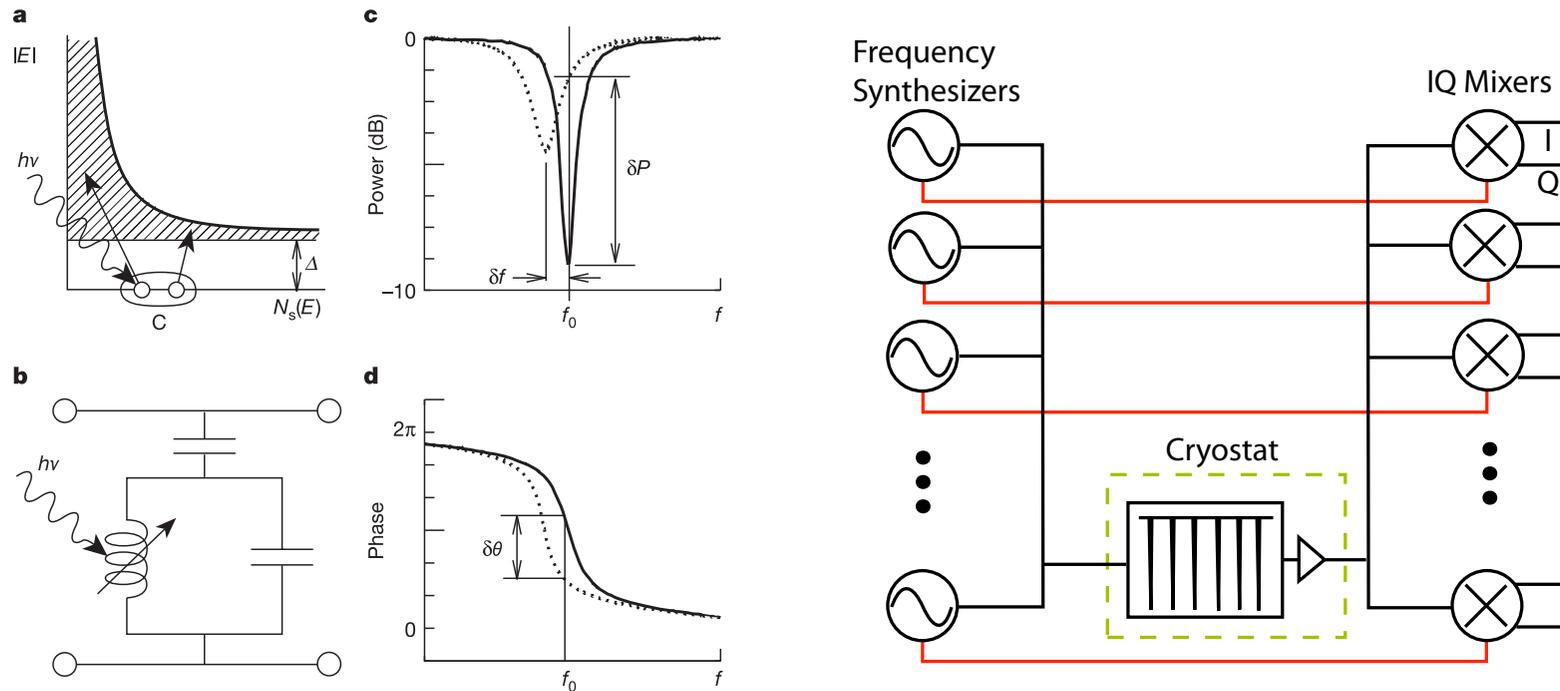
Lost ionization smears ER peak:  
degrades ER rejection

Still need to worry about bgnd rejection:  
 $^{210}\text{Pb}$  nuclei from sidewalls, housing

Need phonon position reconstruction!



# Basics of Kinetic Inductance Detectors



Superconductors have an AC inductance due to inertia of Cooper pairs  
alternately, due to magnetic energy stored in screening supercurrent

Changes when Cooper pairs broken by energy, creating quasiparticles (qps)  
Sense the change by monitoring a resonant circuit

Key point: superconductors provide very high Q ( $Q_i > 10^7$  achieved), so  
thousands of such resonators can be monitored with a single feedline  
enormous cryogenic multiplex technology relative to existing ones  
very simple cryogenic readout components

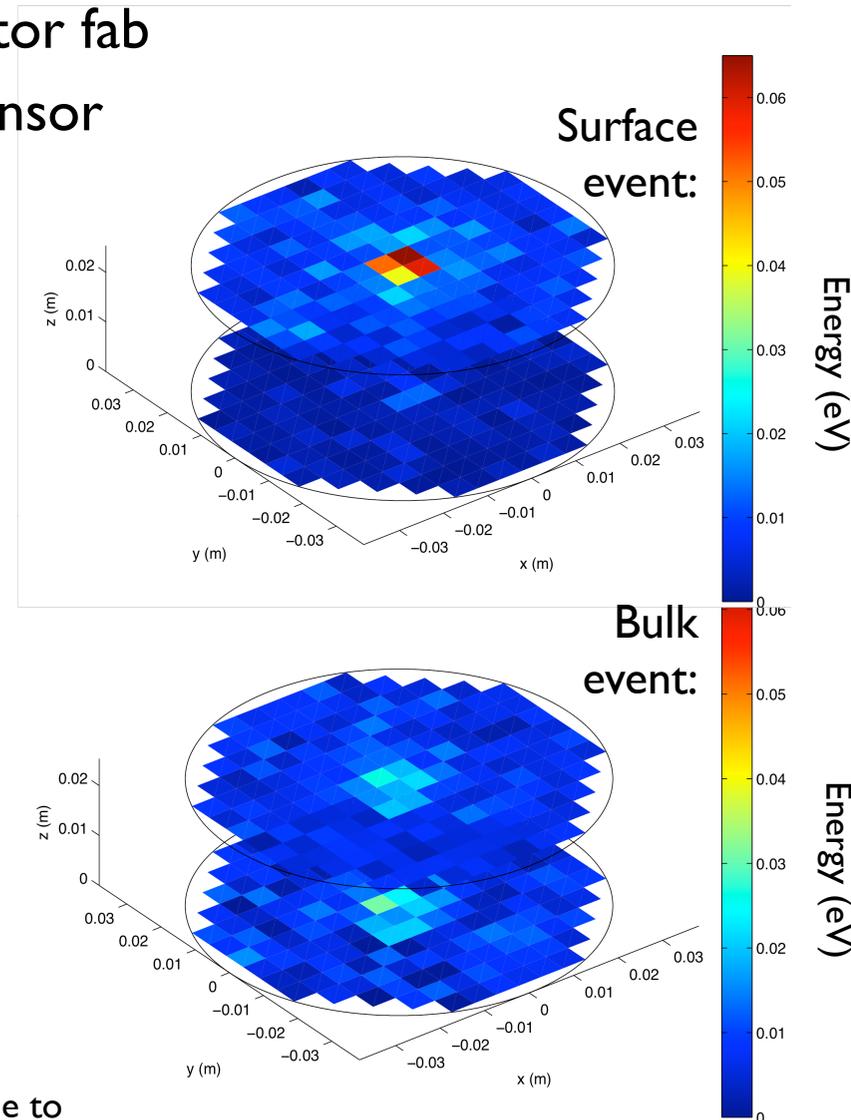
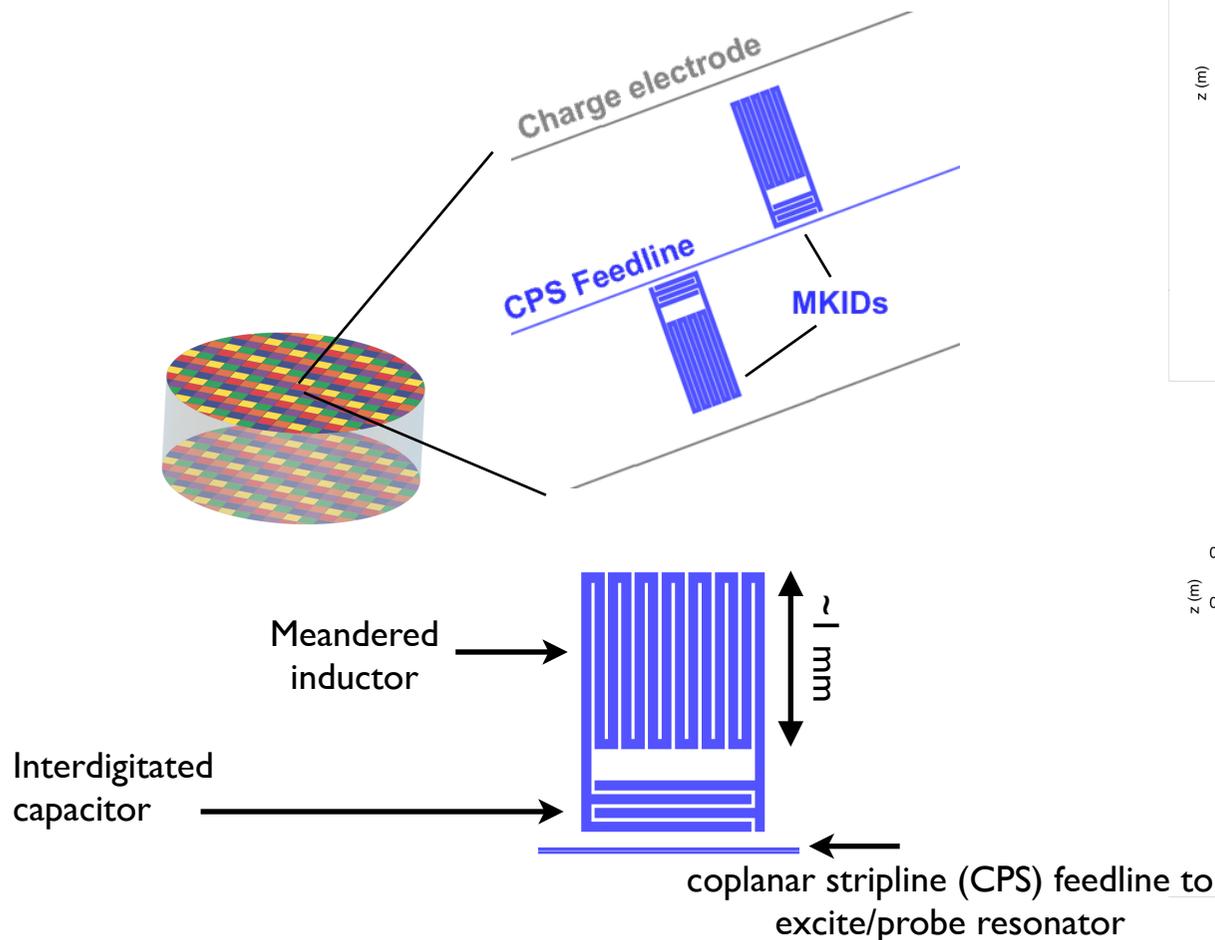
# Phonon-Mediated DM Detection Using KIDs

Lumped-element designs enables large-area resonators for phonon sensing

Single film, 10  $\mu\text{m}$  features would simplify detector fab

Multiplexing enables finely pixellated phonon sensor

→ better surface event rejection



Figures by D. Moore

# Energy Resolution

## Assume:

readout noise dominates ( $T \ll T_c$  ensures g-r noise subdominant)

qp density determined by readout power generation

resonator is coupling dominated ( $Q_c \ll Q_i$ )

quasiparticle lifetime in KID  $\ll$  phonon absorption timescale ( $\tau_{qp} \ll \tau_{abs}$ )

$$\begin{aligned} \Delta E &= 2.355 \frac{\Delta}{\eta_{ph}} \sqrt{\frac{\eta_{read}}{2\pi\alpha p t}} \sqrt{\frac{A_{sub} \lambda_{pb}}{f_r \tau_{qp}}} \sqrt{\frac{k T_N N_0}{\gamma_s S_1(f_r, T_{qp}, \Delta)}} \\ &= (0.052 \text{ keV}) \left( \frac{1}{\eta_{ph}} \sqrt{\frac{1}{\alpha p t}} \right) \left( \frac{\Delta}{200 \text{ } \mu\text{eV}} \right) \sqrt{\frac{A_{sub}}{100 \text{ cm}^2} \frac{\lambda_{pb}}{1 \text{ } \mu\text{m}} \frac{5 \text{ GHz}}{f_r} \frac{100 \text{ } \mu\text{s}}{\tau_{qp}} \frac{T_N}{3 \text{ K}} \frac{1.18}{S_1(f_r, T_{qp}, \Delta)}} \end{aligned}$$

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gap energy

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area of substrate (including sidewalls)

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Annotations for the equation:

- superconducting gap energy  $\Delta$
- efficiency of qp creation by readout power (assume = 1 to be conservative)  $\eta_{read}$
- area of substrate (including sidewalls)  $A_{sub}$
- pair-breaking length in KID film  $\lambda_{pb}$
- efficiency for converting phonons to qps  $\eta_{ph}$
- fraction of inductance due to KI ( $\sim 1$ )  $\alpha p t$
- probability for phonon to enter KID per try  $\frac{1}{\alpha p t}$

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efficiency for converting phonons to qps  $\eta_{ph}$   
 fraction of inductance due to KI ( $\sim 1$ )  $\alpha$   
 probability for phonon to enter KID per try  $p$   
 KID resonant frequency  $f_r$

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 pair-breaking length in KID film  $\lambda_{pb}$

$$\Delta E = 2.355 \frac{\Delta}{\eta_{ph}} \sqrt{\frac{\eta_{read}}{2\pi\alpha p t}} \sqrt{\frac{A_{sub} \lambda_{pb}}{f_r \tau_{qp}}} \sqrt{\frac{k T_N N_0}{\gamma_s S_1(f_r, T_{qp}, \Delta)}}$$

$$= (0.052 \text{ keV}) \left( \frac{1}{\eta_{ph}} \sqrt{\frac{1}{\alpha p t}} \right) \left( \frac{\Delta}{200 \mu\text{eV}} \right) \sqrt{\frac{A_{sub}}{100 \text{ cm}^2} \frac{\lambda_{pb}}{1 \mu\text{m}} \frac{5 \text{ GHz}}{f_r} \frac{100 \mu\text{s}}{\tau_{qp}} \frac{T_N}{3 \text{ K}} \frac{1.18}{S_1(f_r, T_{qp}, \Delta)}}$$

efficiency for converting phonons to qps  $\eta_{ph}$   
 fraction of inductance due to KI ( $\sim 1$ )  $\alpha$   
 probability for phonon to enter KID per try  $p$   
 KID resonant frequency  $f_r$   
 quasiparticle lifetime  $\tau_{qp}$

# Energy Resolution

Assume:

readout noise dominates ( $T \ll T_c$  ensures g-r noise subdominant)

qp density determined by readout power generation

resonator is coupling dominated ( $Q_c \ll Q_i$ )

quasiparticle lifetime in KID  $\ll$  phonon absorption timescale ( $\tau_{qp} \ll \tau_{abs}$ )

Annotations for the equation:

- superconducting gap energy  $\Delta$
- efficiency of qp creation by readout power (assume = 1 to be conservative)  $\eta_{read}$
- area of substrate (including sidewalls)  $A_{sub}$
- pair-breaking length in KID film  $\lambda_{pb}$
- amplifier noise temperature  $T_N$
- efficiency for converting phonons to qps  $\eta_{ph}$
- fraction of inductance due to KI ( $\sim 1$ )  $\alpha_{pt}$
- probability for phonon to enter KID per try  $S_1(f_r, T_{qp}, \Delta)$
- KID resonant frequency  $f_r$
- quasiparticle lifetime  $\tau_{qp}$

$$\Delta E = 2.355 \frac{\Delta}{\eta_{ph}} \sqrt{\frac{\eta_{read}}{2\pi\alpha_{pt}}} \sqrt{\frac{A_{sub}\lambda_{pb}}{f_r\tau_{qp}}} \sqrt{\frac{kT_N N_0}{\gamma_s S_1(f_r, T_{qp}, \Delta)}}$$

$$= (0.052 \text{ keV}) \left( \frac{1}{\eta_{ph}} \sqrt{\frac{1}{\alpha_{pt}}} \right) \left( \frac{\Delta}{200 \mu\text{eV}} \right) \sqrt{\frac{A_{sub}}{100 \text{ cm}^2} \frac{\lambda_{pb}}{1 \mu\text{m}} \frac{5 \text{ GHz}}{f_r} \frac{100 \mu\text{s}}{\tau_{qp}} \frac{T_N}{3 \text{ K}} \frac{1.18}{S_1(f_r, T_{qp}, \Delta)}}$$

# Energy Resolution

Assume:

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quasiparticle lifetime in KID  $\ll$  phonon absorption timescale ( $\tau_{qp} \ll \tau_{abs}$ )

efficiency of qp creation by readout power (assume = 1 to be conservative)  
 area of substrate (including sidewalls)  
 pair-breaking length in KID film  
 amplifier noise temperature  
 density of states  
 superconducting gap energy

$$\Delta E = 2.355 \frac{\Delta}{\eta_{ph}} \sqrt{\frac{\eta_{read}}{2\pi\alpha p t}} \sqrt{\frac{A_{sub} \lambda_{pb}}{f_r \tau_{qp}}} \sqrt{\frac{k T_N N_0}{\gamma_s S_1(f_r, T_{qp}, \Delta)}}$$


---


$$= (0.052 \text{ keV}) \left( \frac{1}{\eta_{ph}} \sqrt{\frac{1}{\alpha p t}} \right) \left( \frac{\Delta}{200 \mu\text{eV}} \right) \sqrt{\frac{A_{sub}}{100 \text{ cm}^2} \frac{\lambda_{pb}}{1 \mu\text{m}} \frac{5 \text{ GHz}}{f_r} \frac{100 \mu\text{s}}{\tau_{qp}} \frac{T_N}{3 \text{ K}} \frac{1.18}{S_1(f_r, T_{qp}, \Delta)}}$$

efficiency for converting phonons to qps  
 fraction of inductance due to KI (~1)  
 probability for phonon to enter KID per try  
 KID resonant frequency  
 quasiparticle lifetime

# Energy Resolution

Assume:

readout noise dominates ( $T \ll T_c$  ensures g-r noise subdominant)

qp density determined by readout power generation

resonator is coupling dominated ( $Q_c \ll Q_i$ )

quasiparticle lifetime in KID  $\ll$  phonon absorption timescale ( $\tau_{qp} \ll \tau_{abs}$ )

efficiency of qp creation by readout power (assume = 1 to be conservative)  
 area of substrate (including sidewalls)  
 pair-breaking length in KID film  
 amplifier noise temperature  
 density of states  
 superconductivity factors,  $\sim 1$

$$\Delta E = 2.355 \frac{\Delta}{\eta_{ph}} \sqrt{\frac{\eta_{read}}{2\pi\alpha p t}} \sqrt{\frac{A_{sub} \lambda_{pb}}{f_r \tau_{qp}}} \sqrt{\frac{k T_N N_0}{\gamma_s S_1(f_r, T_{qp}, \Delta)}}$$

superconducting gap energy

$$= (0.052 \text{ keV}) \left( \frac{1}{\eta_{ph}} \sqrt{\frac{1}{\alpha p t}} \right) \left( \frac{\Delta}{200 \mu\text{eV}} \right) \sqrt{\frac{A_{sub}}{100 \text{ cm}^2} \frac{\lambda_{pb}}{1 \mu\text{m}} \frac{5 \text{ GHz}}{f_r} \frac{100 \mu\text{s}}{\tau_{qp}} \frac{T_N}{3 \text{ K}} \frac{1.18}{S_1(f_r, T_{qp}, \Delta)}}$$

efficiency for converting phonons to qps  
 fraction of inductance due to KI ( $\sim 1$ )  
 probability for phonon to enter KID per try  
 KID resonant frequency  
 quasiparticle lifetime

# Energy Resolution

Assume:

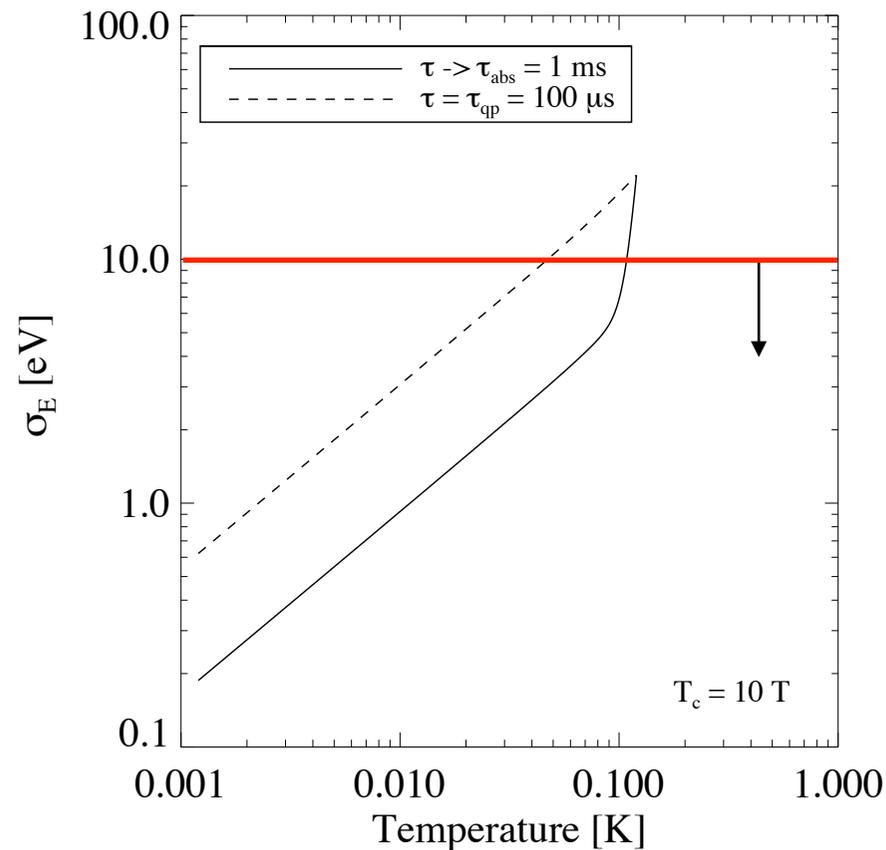
readout noise dominates ( $T \ll T_c$  ensures g-r noise subdominant)

qp density determined by readout power generation

resonator is coupling dominated ( $Q_c \ll Q_i$ )

quasiparticle lifetime in KID  $\ll$  phonon absorption timescale ( $\tau_{qp} \ll \tau_{abs}$ )

very competitive and  
interesting resolution  
possible if most  
phonons can  
be collected!



$\eta_{ph} = 1$  assumed

single e-h pair  
detection  
possible

# Small Prototype

12 mm x 16 mm arrays of 20 resonators on 22 mm x 22 mm x 1 mm silicon substrates show excellent performance

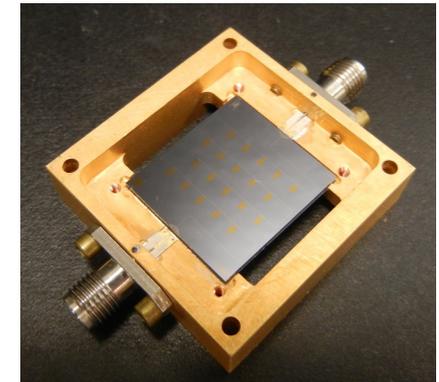
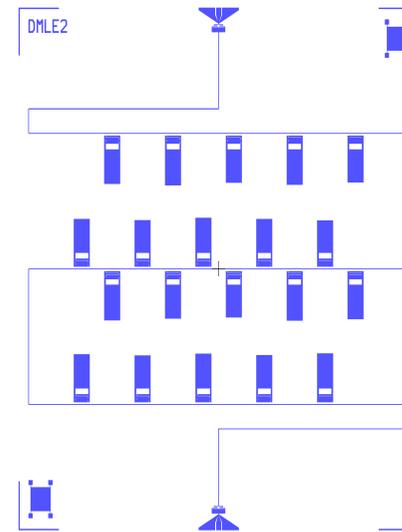
position correction recovers close to baseline performance; energy resolution ~ existing phonon-mediated detectors

$\tau_{qp} = 13 \mu\text{s}$  measured (low!)

$\eta_{ph} = 0.07$ : large phonon losses

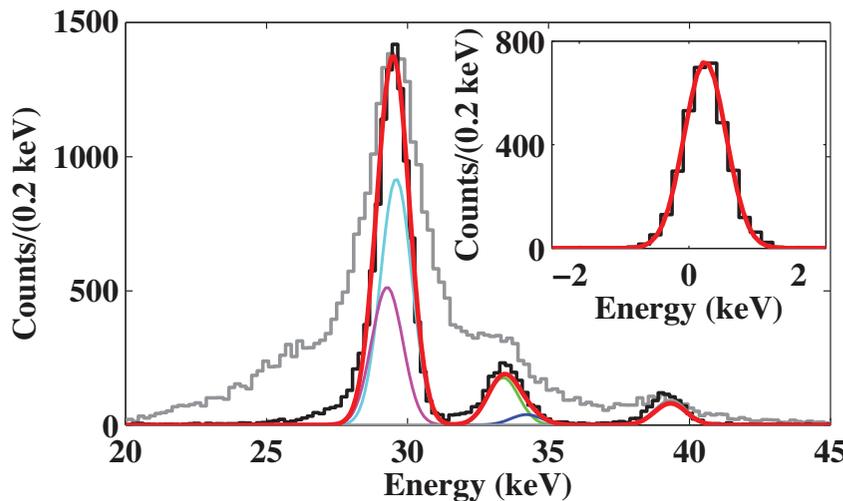
expected  $\sigma_E = 0.48 \text{ keV} \sqrt{(\eta_{read}/p_t)}$ : consistent with measurement

Performance limited by lifetime and phonon losses

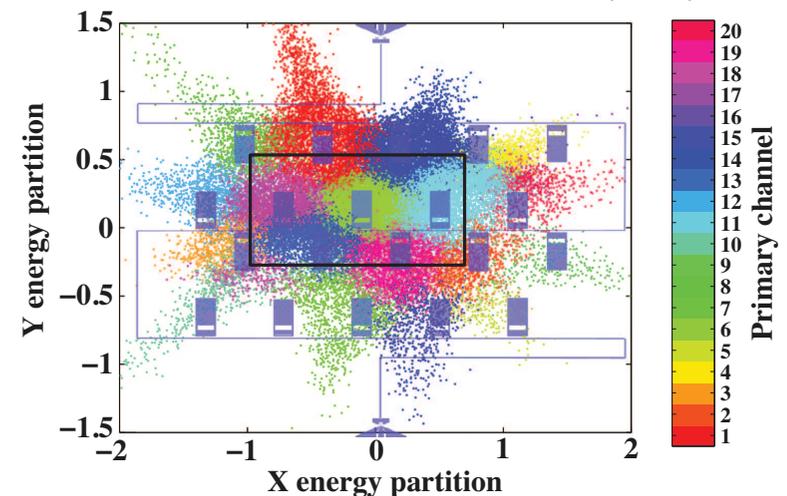


collimated 30 keV  $^{129}\text{I}$  source incident on bottom side of silicon wafer

$\sigma = 0.55 \text{ keV}$   
baseline  $\sigma = 0.38 \text{ keV}$   
grey: before position correction  
black: after position correction



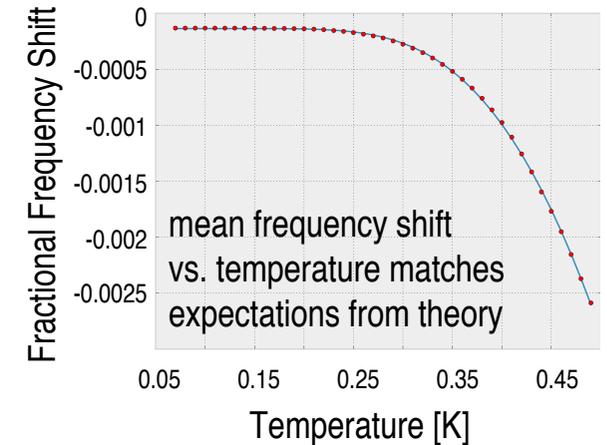
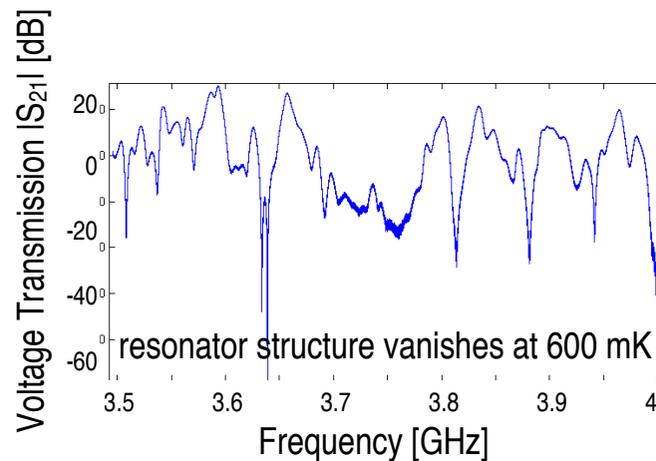
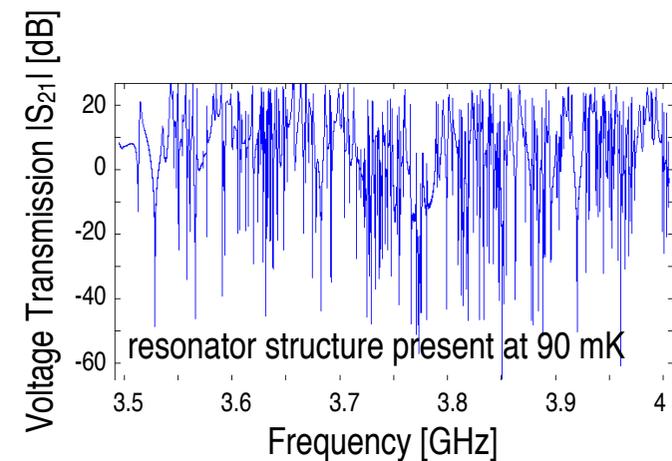
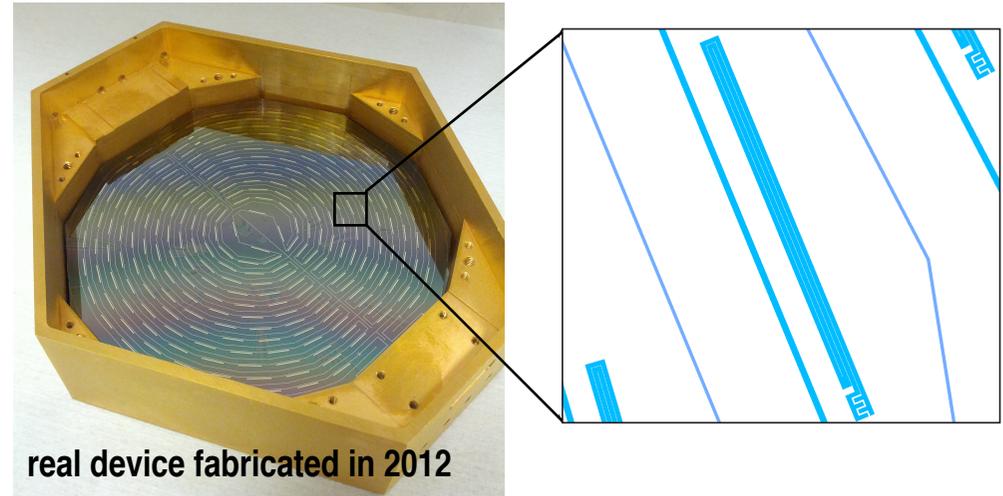
Moore et al, APL 100: 232601 (2012)



# Full Size Prototype

Working on scaling to  
7.5-cm x 1-cm substrates

Have designed and  
fabricated devices and acquired  
characterization data



Working to understand RF structure of device with Nb version (4K testing)

Then pulse data

DOE HEP detector R&D grant awarded

Exploring nuclear nonproliferation funding options (high-res  $\gamma$  spectroscopy)

# Beyond Detection of Individual e-h Pairs: < 100 MeV

Qualitative change in response below ~100 MeV

100 MeV:  $\langle E_R \rangle = 12$  eV in Ge, 22 eV in Si:  
 accessible with  $V_b = 100$  V,  $\sigma_{pt} = 3$  eV

10 MeV:  $\langle E_R \rangle = 1.2$  eV in Ge, 2.2 eV in Si:  
 below e-h pair threshold:  
 there are no states that electrons can scatter into

Except possible relatively low density impurity states  $\rightarrow$  ER-insensitive regime?

$\sigma_p < 1$  eV in principle accessible

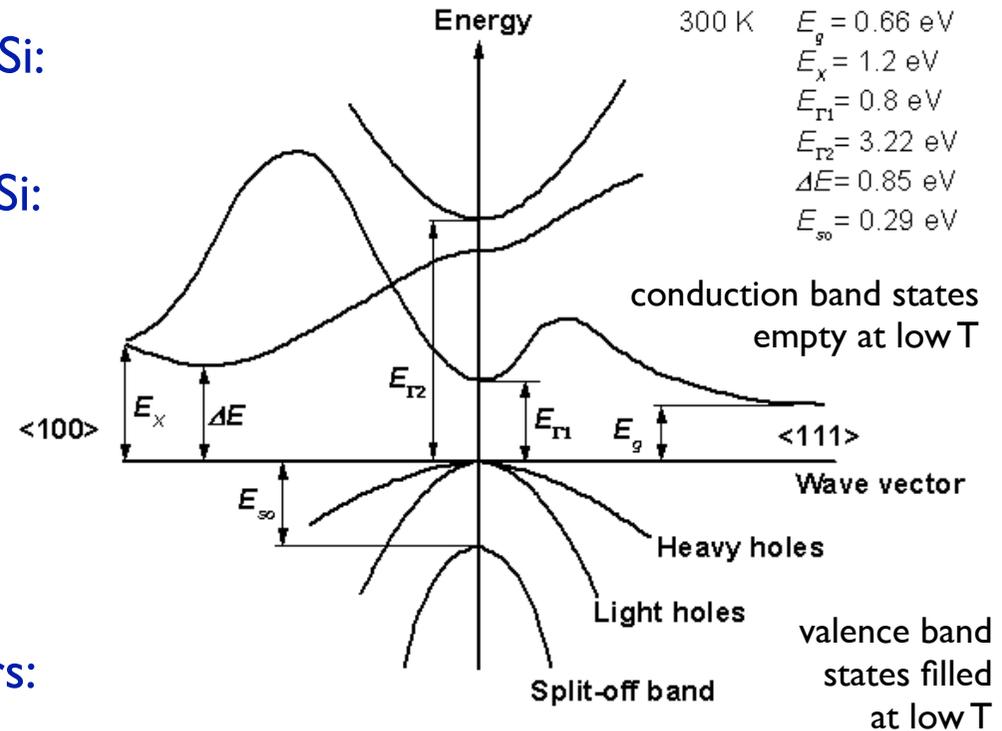
Like IR and FIR photon counting detectors:  
 Resolve depositions less than 1 eV

No quantization effects until reach ~ meV depositions

meV deposition  $\rightarrow$  10 keV DM mass

KID advantage?

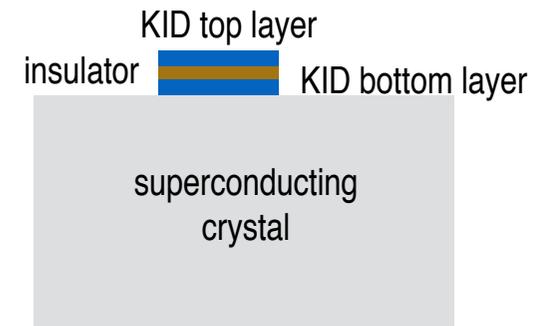
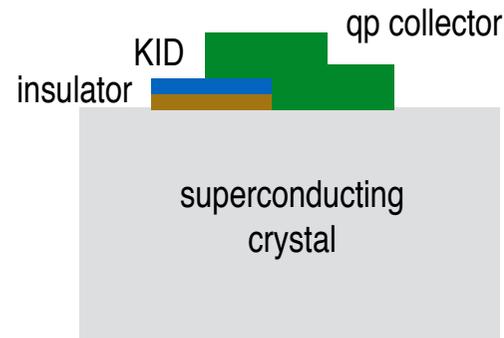
Threshold for qp creation in KIDs perhaps provides some insensitivity to environmental noise, esp. vibrational noise



# What About Direct Quasiparticle Creation?

Long scattering length  
superconducting crystal:  
qp's diffuse

Architectures:



KID on insulator, qps collected via thick superconducting film

Technically straightforward to imagine a design:

Avoids having to deal with operating KID on superconductor

Requires good trapping: qps from crystal into collector film, from collector film into KID

Problem: fast trapping require large  $\Delta$  ratio; large  $\Delta$  ratio  $\rightarrow$  lots of energy lost to phonon emission

Maybe still ok if just interesting in counting substrate qps (still can get meV threshold)

KID on crystal

Need to avoid short-circuiting KID: microstrip structure?

Film needs to be thick to avoid being proximitized by crystal ( $\Delta_{\text{KID}}$  pulled to  $\Delta_{\text{crystal}}$ )

No obvious advantage over phonon mediation *for NR detection*

Phonons already provide sensitivity to meV scale

KIDs are already pair-breaking detectors: insensitive to sub-gap phonons in principle

But definitely interesting for electron scattering

# Conclusions

Improved imaging of phonon signal from substrate can provide significant gains for phonon-mediated detectors

In phonons+ionization mode: fiducialization using phonons, bulk ER rejection using ionization

In HV mode: fiducialization using phonons

Improved energy resolution would reach single e-h pair detection regime

KIDs can provide these improvements

Multiplexability of KIDs promises finer imaging of phonon signal

KID sensitivity promises to reach single e-h pair regime

KID demonstration in hand, working on scaling up to massive substrates



# Solid-State Detectors

## SuperCDMS/EDELWEISS

Semiconducting crystals

Ionization:

Ionization produced in interactions  
drifted w/low electric field

Phonons (thermal and athermal)

Most energy goes into phonons. In Ge: 3.0 eV/e-h pair vs. 0.67 eV bandgap

Energies:

“keV<sub>r</sub>” = recoil energy, energy deposited by particle interaction =  $E_r$

“keV<sub>ee</sub>” = “electron-equivalent” energy =  $N_{e-h} \times 3.0 \text{ eV}$  in Ge =  $E_q$ ;  $E_q = E_r$  for ERs

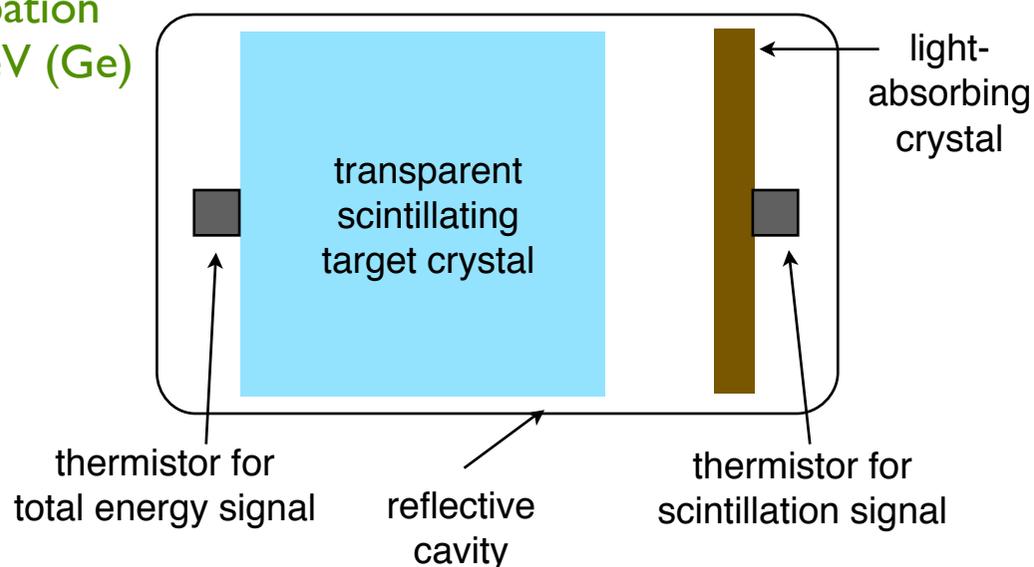
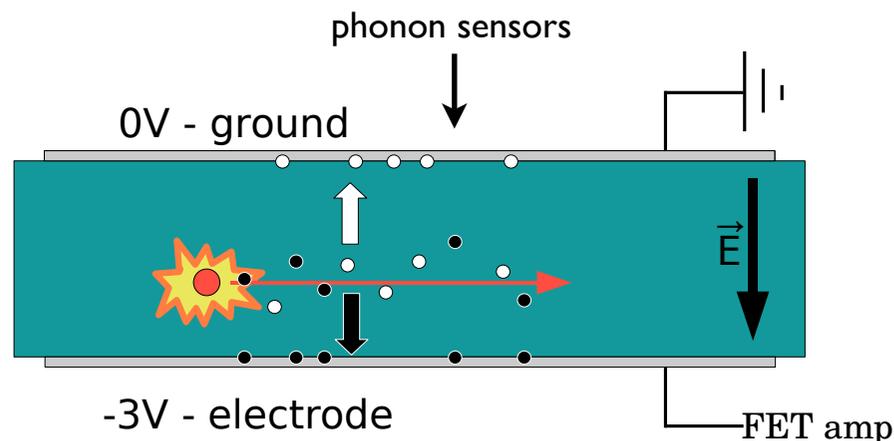
Luke-Neganov energy = drift heating dissipation  
=  $E_{drift} = N_{e-h} \times e \times V_b = E_q \times e \times V_b / 3.0 \text{ eV (Ge)}$

“keV<sub>p</sub>” = phonon energy =  $E_r + E_{drift}$

## CRESST

Photons from scintillating crystals  
instead of ionization (e.g.  $\text{CaWO}_4$ )

Photons detected with separate  
“absorber crystal”



# SuperCDMS: Surface Event Rejection w/Interdig. Electrodes

Alternating ground and biased electrodes further improve rejection

Field configuration:

Bulk events have symmetric hole/electron collection

Surface ERs are asymmetric

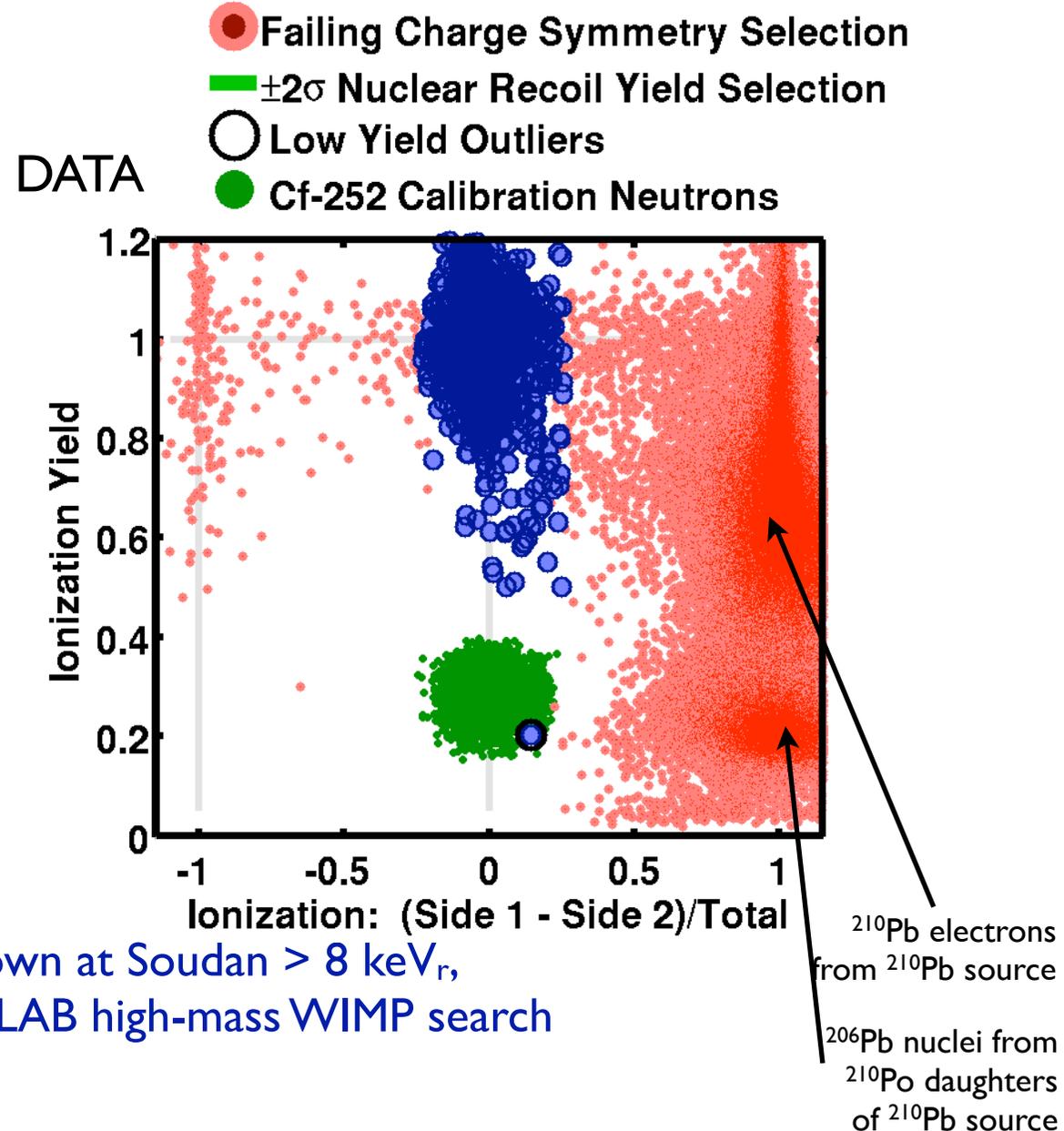
Field strength

High field near surface raises ionization collection for surface electron recoils

$^{206}\text{Pb}$  nuclear recoils visible

Important goals achieved

Surface ER rejection  $\sim 1 \times 10^{-5}$  shown at Soudan  $> 8 \text{ keV}_r$ , sufficient for SuperCDMS SNOLAB high-mass WIMP search

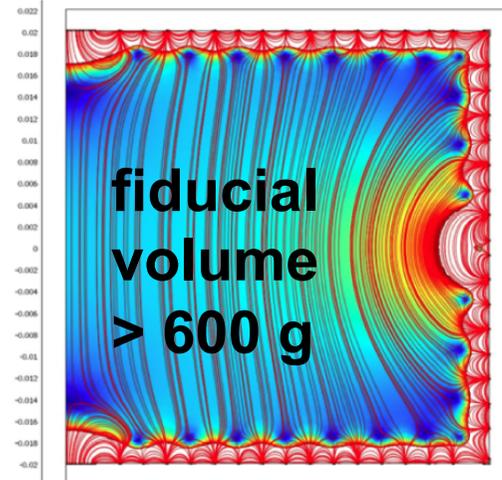
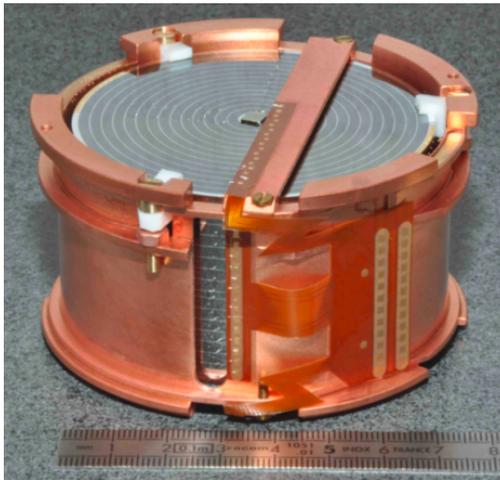


# EDELWEISS: Surface Event Rejection w/Interdig. Electrodes

EDELWEISS also has demonstrated this technology

EDELWEISS III: More mass than SuperCDMS Soudan (30 kg vs. 9 kg),  
but higher threshold  
(15 keV<sub>r</sub> vs. 8 keV<sub>r</sub>)

FID 800g with 40x ~600g fiducial mass



« Full InterDigitised »

