



# Neutron – antineutron oscillation

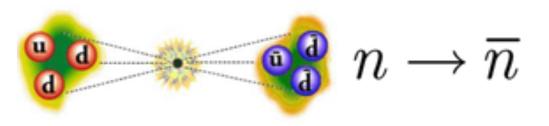
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#### What is neutron – antineutron oscillation?

• Neutron – antineutron oscillation is exactly what the name would insinuates, a neutron turning into an antineutron

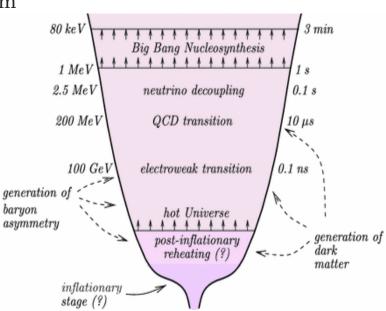


- Observation of electrically neutral particles into other species are no longer a surprising phenomenon in particle physics
  - Ex. Neutrino and neutral meson oscillations
- Only conservation of baryon number forbids a neutron from transforming to an anti-neutron
  - This conservation law does not follow from any known physical principle but is inferred from experiment.

# Why is it important?

- Simple answer: baryon asymmetry
  - Observe more matter than anti-matter in the universe
- To produce matter and antimatter at different rates, a set of three necessary conditions that a baryon-generating interaction must satisfy (Sakharov conditions):
  - Baryon number *B* violation
  - C-symmetry and CP-symmetry violation
  - Interactions out of thermal equilibrium

- However there are a number of other reasons to study  $n \bar{n}$  oscillations including:
  - Majorana neutrinos
  - Baryogenesis mechanisms
  - Theory independent







# $n-\bar{n}$ transition probability

• The time evolution of a system in which a neutron state can transform into an antineutron follows:

$$i\hbarrac{\partial}{\partial t} \begin{pmatrix} n\\ ar{n} \end{pmatrix} = \begin{pmatrix} E_n & \delta m\\ \delta m & E_{ar{n}} \end{pmatrix} \begin{pmatrix} n\\ ar{n} \end{pmatrix}$$

- where  $\delta m$  is the  $n \bar{n}$  mixing rate of the process that violates baryon number conservation
- From this you can extract the antineutron transition probability using the boundary conditions that at time zero  $P_n(0) = 1$  and and  $P_{\bar{n}}(0) = 0$

$$P_{\overline{n}}(t) = \left(\frac{t}{\tau_{n-\overline{n}}}\right)^2$$
 where  $\tau_{n-\overline{n}} = \frac{1}{|\delta m|}$  oscillation lifetime

# $n-\bar{n}$ transition probability (II)

- $n \bar{n}$  oscillation can occur in two different settings:
  - Free: neutron propagates while not being bound within and a nucleus
  - **Bound**: Neutron propagates while bound in a nucleus
- These two settings have different transition probabilities because there is a suppression factor within the nucleus which affects the oscillation lifetime
  - Situation complicated by neighboring nucleons
- The suppressed oscillation lifetime in the nuclei and the free oscillation lifetime are then related by:

$$\tau_{nucl} = R \times (\tau_{n-\bar{n}})^2$$

- where R is a nucleus dependent suppression factor (order of  $10^{22\text{-}23}$  sec)



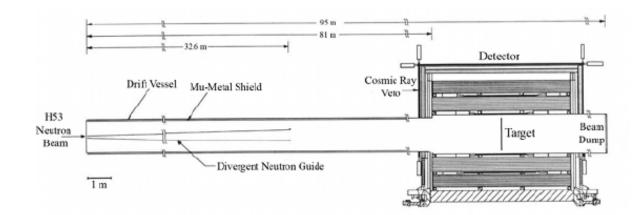
# Experimental methods

- Two methods for observing  $n \overline{n}$  oscillations:
  - Oscillations of a beam of neutrons against an annihilation target
  - Oscillations of neutrons to antineutrons in bound nuclei



# Experimental methods

- Two methods for observing  $n \overline{n}$  oscillations:
  - Oscillations of a beam of neutrons against an annihilation target
    - Beam of slow neutron which propagate freely from the exit of a neutron optical guild to a distant antineutron annihilation target



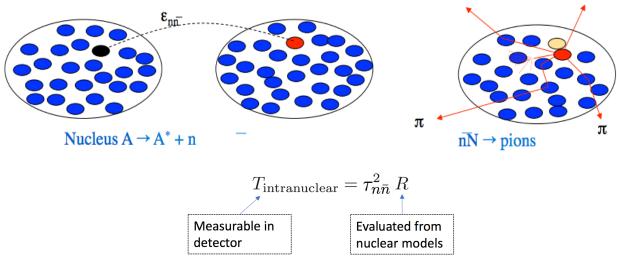
- Pros:
  - Robust background suppression
  - Control signal with magnetic field
  - Can achieve orders of magnitude improvement in sensitivity with updated technology
- Cons:
  - Difficult to find source of neutrons tailored for this search
  - Small overlap with other neutron experiments





# Experimental methods

- Two methods for observing  $n \overline{n}$  oscillations:
  - Oscillations of neutrons to antineutrons in bound nuclei
    - Look for spontaneous transition of neutron to antineutron in nuclei in large underground detectors which are mainly built for proton decay and neutrino oscillation studies
      - SNO, DUNE, etc.



- Pros:
  - Overlap with other proton and neutrino experiments
  - Large number of neutrons to help overcome suppression factor
- Cons:
  - Considerable backgrounds to account for (10-50%)



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### Previous search with free neutrons

 Current best limit performed at the Institut Laue-Langevin in Grenoble in the early 1990's

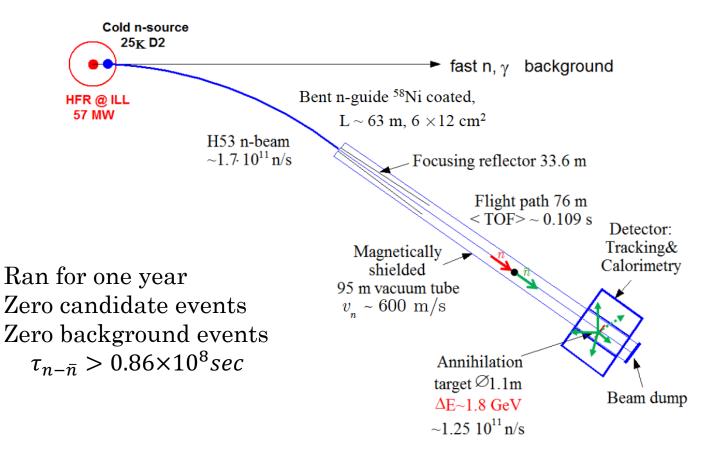
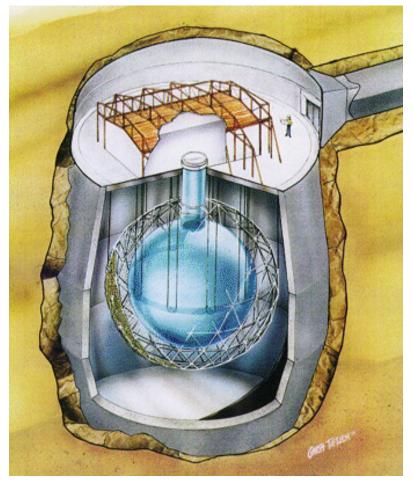


Figure 2: Configuration of the horizontal  $n - \bar{n}$  search experiment at ILL/Grenoble [63].

# SNO results with bound nuclei

- SNO (Sudbury Neutrino Observatory) was a heavy water (2H2O or D2O) Cherenkov imaging detector that was in operation from November 2, 1999 to November 28, 2006
- Deuteron is an intra-nuclear source for neutron – antineutron oscillations
  - Has lower suppression factor by factor of 4 compared with Oxygen
- Looking for signature of multiprong events with multiple charged and neutral pions from nbar-p or nbar-n interactions





## SNO results with bound nuclei (II)

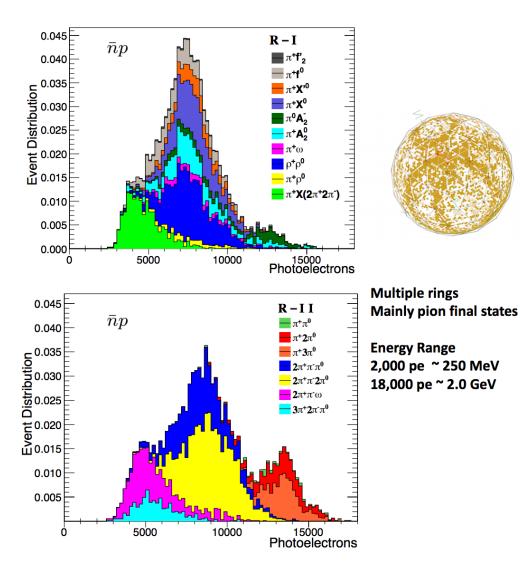
• Momentum regime I (at rest)

A study of channels of an antiproton colliding with a neutron near rest showed a majority of 2-body intermediate states\*. These intermediate states can then decay into channels including multiple pions.

 Momentum regime II (~ 250 MeV) : Alternative interaction channels [\*\*] for n p annihilation have also been modeled using beam data of p n collisions at momenta comparable to the <sup>16</sup>O

Both models are investigated and **give similar results.** Weighted average of both is used in analysis

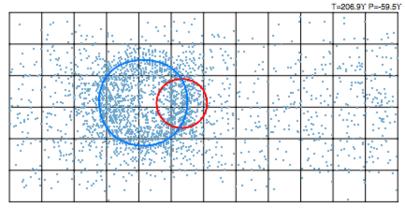
\*R. Bridges et al., Phys. Rev. Lett. 56, 215 (1986) \*\*K. Abe et al., Phys. Rev. D 91, 072006 (2015)



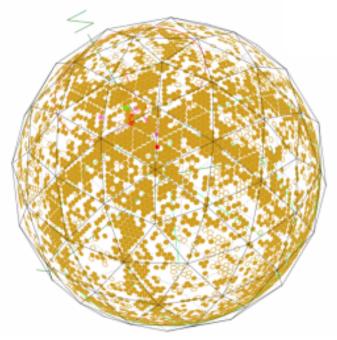


# SNO results with bound nuclei (III)

• They look for evidence of the pion decays by reconstructing rings left on detector

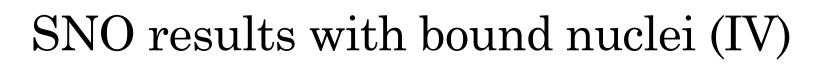


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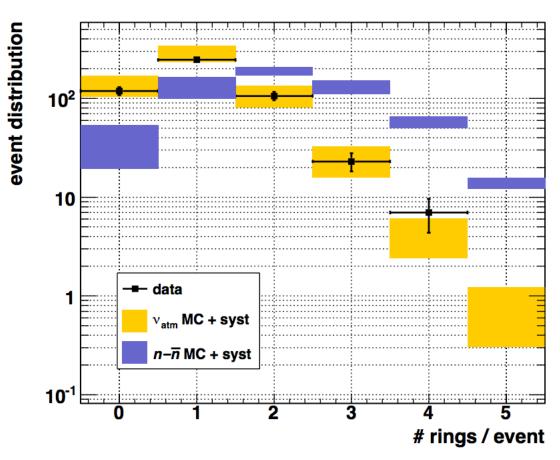


Simulation of a neutron-antineutron oscillation event in SNO





• Data events compared with MC simulations for signal and background are consistent with the background



#### **Ring distribution of contained events**

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### Bound nuclei limits on oscillation lifetime

• A number of different experiments have set limits on the free neutron – antineutron oscillation lifetime using bound neutrons in nuclei

Experiment	$10^{32} \ n$ -yr	$\tau_m (10^{32} \text{ yr})$	$R(10^{23}/s)$	$\tau_{n-\bar{n}}(10^8 \text{ s})$
ILL (free- $n$ ) [63]	n/a	n/a	n/a	0.86
IMB $(^{16}O)$ <u>96</u>	3.0	0.24	1.0	0.88
Kamiokande $(^{16}O)$ 97	3.0	0.43	1.0	1.2
Frejus (56Fe) [98]	5.0	0.65	1.4	1.2
Soudan-2 ( ${}^{56}$ Fe) [92]	21.9	0.72	1.4	1.3
SNO $(^{2}H)$ [94]	0.54	0.30	0.25	1.96
Super-K $(^{16}O)$ 93	245	1.9	0.517	2.7

Table 1: Neutron-antineutron lifetime lower limits (90% CL).

• Somewhat surprisingly, all results have been consistent with the limit set by the ILL free neutron experiment

# What does this all mean?

- It means that we haven't observed neutron antineutron oscillation yet, and we haven't set a limit that deviates far enough from what is expected to show evidence of new physics
  - The current lower limits set on the neutron antineutron oscillation lifetime are lower than the upper bound set by the seesaw theory (which sets the upper bound at  $10^{10}$  seconds)
- Given that the experiments looking for oscillation in bound nuclei have yet to observe the interaction, it may be worth pursuing free neutron experiments with more seriously.
  - ESS and FNAL have both proposed such experiments









# Extra slides



The oscillation probability of a neutron to an antineutron can be evaluate in analogous ways to neutrino oscillation probability.

A small  $\delta m$  perturbation allows the neutron and antineutron to oscillate between each state

Paris potential is used for deuteron to evaluate suppression factor R [(2.48±0.08)×10<sup>22</sup> s<sup>-1</sup>].

Optical potential used for heavier nuclei; newer calculations for Heavier nuclei showed a decrease of the suppression factor by a factor of 2

If  $|\psi_{ar{n}p}|^2=e^{-\Gamma_R t}$  and

$$i\frac{\partial}{\partial t} \begin{pmatrix} np\\\\ \bar{n}p \end{pmatrix} = \begin{pmatrix} 0 & \delta m\\\\ \delta m & \Delta - i\frac{\Gamma_R}{2} \end{pmatrix} \begin{pmatrix} np\\\\ \bar{n}p \end{pmatrix}$$

$$\Gamma_R \gg \Delta \longrightarrow |\psi_{np}|^2 \approx e^{-4\frac{\delta m^2}{\Gamma_R}t}$$

Since 
$$\delta m = 1/\tau_{n\bar{n}}$$
,  
 $T_{\text{intranuclear}} = \tau_{n\bar{n}}^2 R$   
Measurable in  
detector  
LUNL  
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#### Impact of B-L-conserving SM interaction on B asymmetry

Sphaeleron mechanism in Standard Model lead to violation of lepton and baryon number ('t Hooft, 1976)

• "On anomalous electroweak baryon-number non-conservation in the early universe" (Kuzmin, Rubakov, Shaposhnikov, 1985)

Sphaelerons conserve (B-L) but violate (B+L). Rate of (B+L)-violating processes at T > TeV exceeds the Universe expansion rate. If  $B=L\neq 0$  is set at GUT scale due to B-L conserving process, B asymmetry can be wiped out by (B+L)-violating process

<u>Thus, for the explanation of universe B asymmetry, B-L-violating</u> <u>mechanisms (leptogenesis, NNbar, some nucleon decay modes...) seem</u> <u>to be required</u>

"Proton decay is not a prediction of baryogenesis" [Yanagida'02]



Experimental searches for B Violation: Nucleon Decay and Neutron-Antineutron Oscillations

Mode	Nucleon decay	N-Nbar oscillations
Effect on B and L	$\Delta B=1, \Delta L=1,$ others $\Delta (B-L)=0,2,$	$\Delta B=2, \Delta L=0,$ $\Delta (B-L)=2$
Effective operator	$L = \frac{g}{M^2} Q Q Q L$	$L = \frac{g}{M^5} Q Q Q Q Q Q$
Mass scale probed	~GUT	>~EW