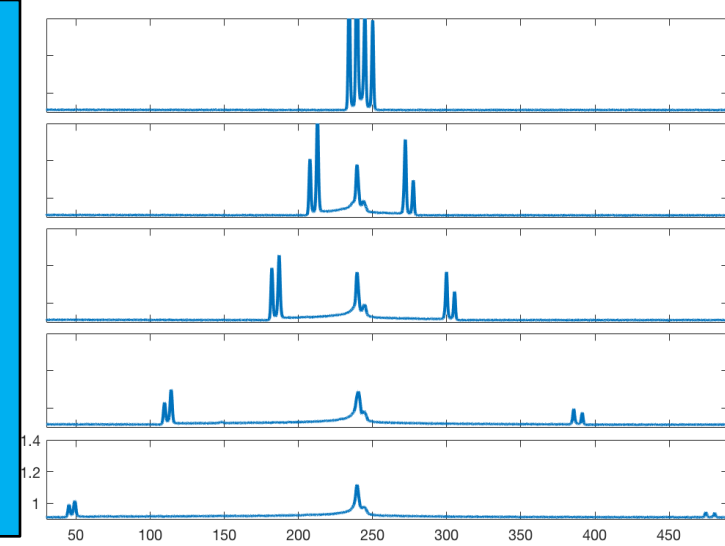


A measurement of the fine structure constant as test of the standard model



Richard H. Parker, Chenghui Yu, Weicheng Zhong, Brian Estey, and H. Müller

U.C. Berkeley



the David &
Lucile Packard
FOUNDATION

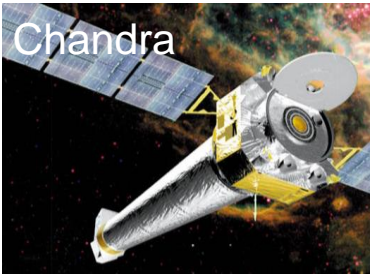


The Era of precision uncertainty

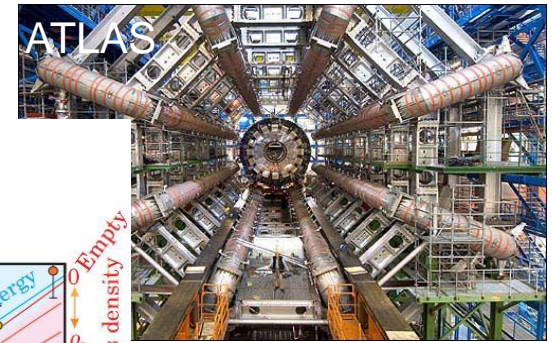
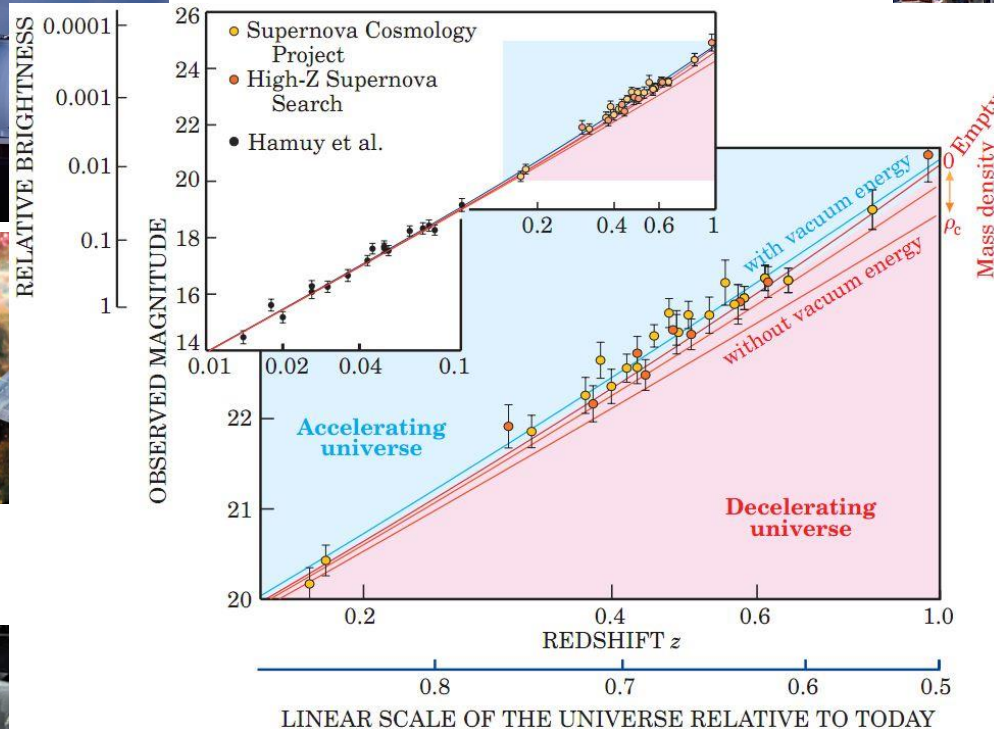
Loud and clear signals from the skies....

...but silence in our detectors

W.M. Keck Observatory



Planck



Finishing my postdoc

2004-2018

2003: “We can write 3 PRLs in the first year”

The plan

- PRL 1: Multiphoton Bragg diffraction
- PRL 2: Simultaneous interferometers
- PRL 3: Fine structure constant

Reality

- Measurement of the fine-structure constant as a test of the Standard Model. Richard H. Parker, Chenghui Yu, Weicheng Zhong, Brian Estey, and Holger Müller, *Science* **360**, 191-195 (2018).
- Controlling the Multiphoton Nature of Bragg Diffraction in Atom Interferometry. Richard H. Parker, Chenghui Yu, Brian Estey, Weicheng Zhong, Eric Huang, and Holger Müller, *Phys. Rev. A* **94**, 053618 (2016).
- High resolution atom interferometers with suppressed diffraction phases. Brian Estey, Chenghui Yu, Holger Müller, Pei-Chen Kuan, and Shau-Yu Lan, *Phys. Rev. Lett.* **115**, 083002 (2015)
- A clock directly linking time to a particle’s mass. Shau-Yu Lan, Pei-Chen Kuan, Brian Estey, Damon English, Justin Brown, Michael Hohensee, and Holger Müller, *Science* **339**, 554 (2013)
- Influence of the Coriolis force in atom interferometry. Shau-Yu Lan, Pei-Chen Kuan, Brian Estey, Philipp Haslinger, and Holger Müller, *Phys. Rev. Lett.* **108**, 090402 (2012).
- A precision measurement of the gravitational redshift by the interference of matter waves. Holger Müller, Achim Peters, and Steve Chu, *Nature* **463**, 926 (2010).
- Atom interferometers with scalable enclosed area, Holger Mueller, Sheng-wei Chiow, Sven Herrmann, and Steven Chu, *Phys. Rev. Lett.* **102**, 240403 (2009).
- 6 W, 1 kHz linewidth, tunable continuous-wave near-infrared laser, Sheng-wei Chiow, Sven Herrmann, Holger Müller, and Steven Chu, *Optics Express* **17**, 5246 (2009).
- Noise-Immune Conjugate Large-Area Atom Interferometers, Sheng-wei Chiow, Sven Herrmann, Steven Chu, and Holger Müller, *Phys. Rev. Lett.* **103**, 050402 (2009).
- Atom Interferometry with up to 24-Photon-Momentum-Transfer Beam Splitters, Holger Müller, Sheng-wei Chiow, Quan Long, Sven Herrmann, and Steven Chu, *Phys. Rev. Lett.* **100**, 180405 (2008)
- Diffraction between the Raman-Nath and the Bragg regime: Effective Rabi frequency, losses, and phase shifts. Holger Müller, Sheng-wei Chiow, and Steven Chu, *Phys. Rev. A* **77**, 023609 (2008).
- Multiphoton- and simultaneous conjugate Ramsey-Borde atom interferometers. Holger Müller, Sheng-wei Chiow, S. Herrmann, and S. Chu, *AIP Conf. Proc.* **977**, 291 (2008).
- Coherent Control of Ultracold Matter: Fractional Quantum Hall Physics and Large-Area Atom Interferometry. Edina Sarajlic, Nathan Gemelke, Sheng-wei Chiow, Sven Herrman, Holger Müller, and Steven Chu, *Proc. 21st ICAP* (2008).
- Extended cavity diode lasers with tracked resonances, Holger Müller, Sheng-wei Chiow, Quan Long, Christoph Vo, and Steven Chu, *Appl. Opt.* **46**, 7997-8001 (2007).
- Nanosecond electro-optical switching with a repetition rate above 20MHz, Holger Müller, Sheng-wei Chiow, Sven Herrmann, Steven Chu, *Rev. Sci. Instrum.* **78**, 124702 (2007).
- A new photon recoil experiment: towards a determination of the fine structure constant. Holger Müller, Sheng-wei Chiow, Quan Long, Christoph Vo, and Steven Chu, *Appl. Phys. B* **84**, 633-642 (2006).



“Will you ever leave Stanford?”

The plan

- No

Reality



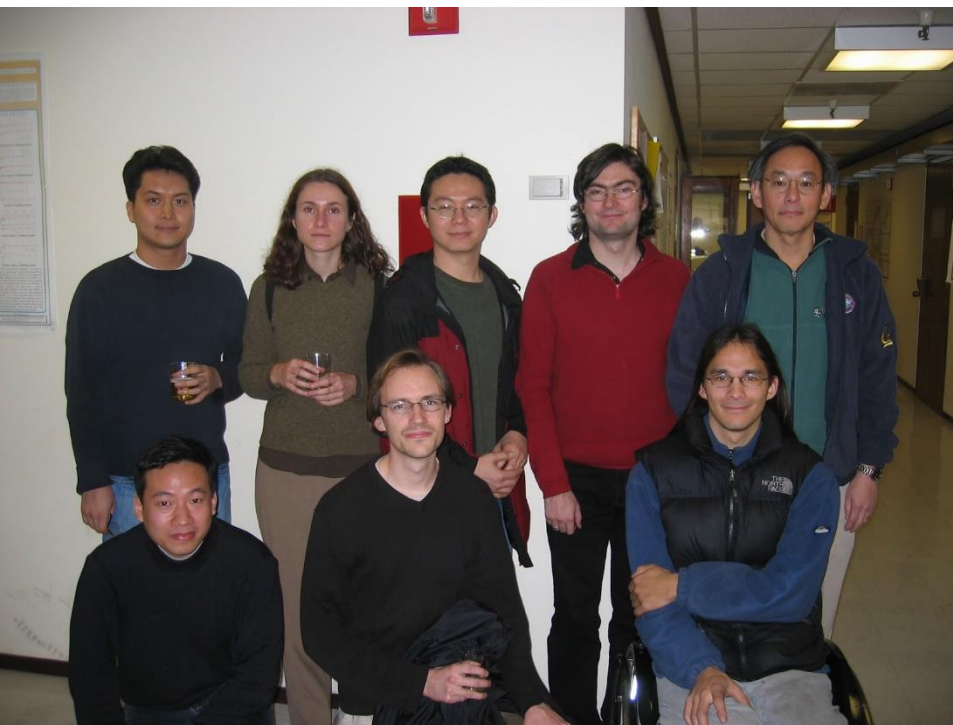
NEWS IN BRIEF

Hungover Energy Secretary Wakes Up Next To Solar Panel

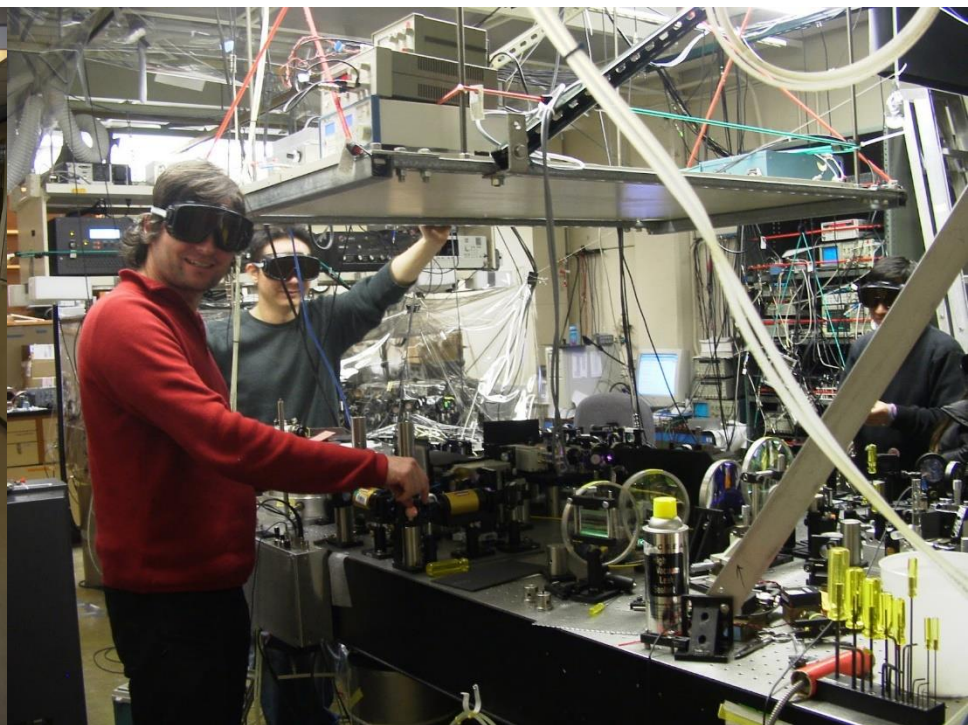
2/07/13 11:52am · SEE MORE: SCIENCE & TECHNOLOGY



WASHINGTON—Sources have reported that following a long night of carousing at a series of D.C. watering holes, Energy Secretary Steven Chu



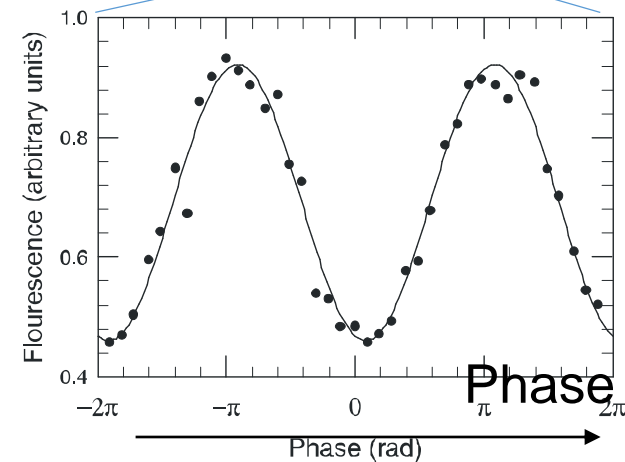
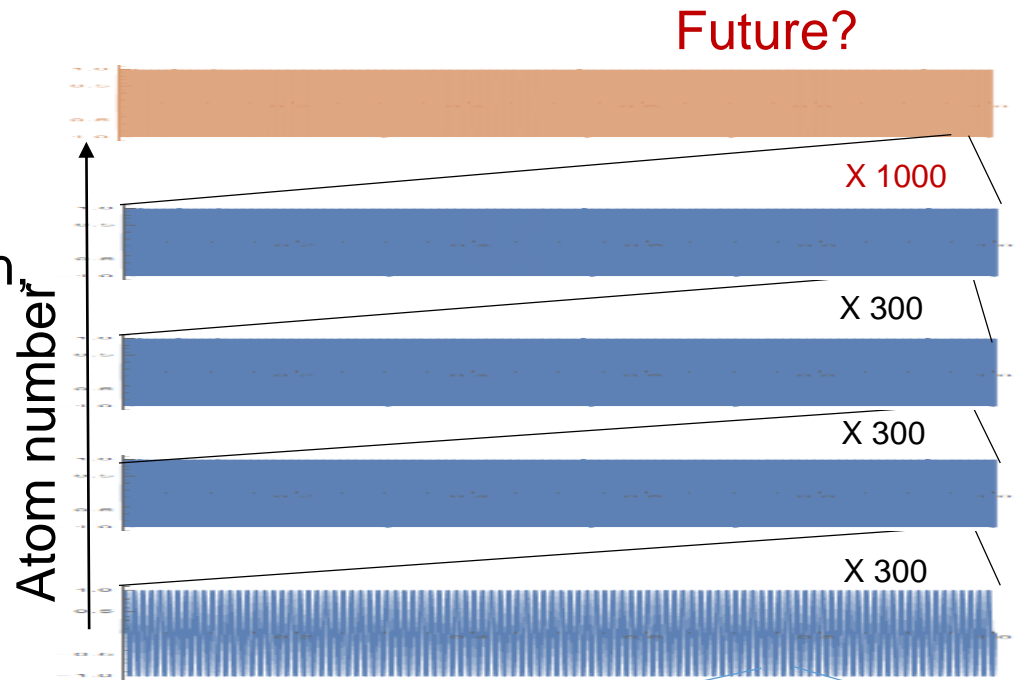
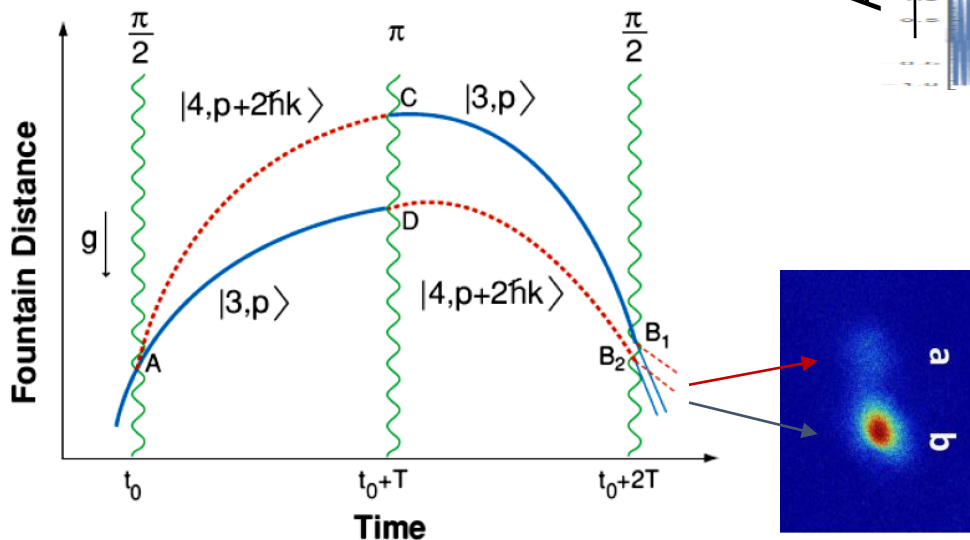
Feb 2005



Apr 2005

Light pulse atom interferometer

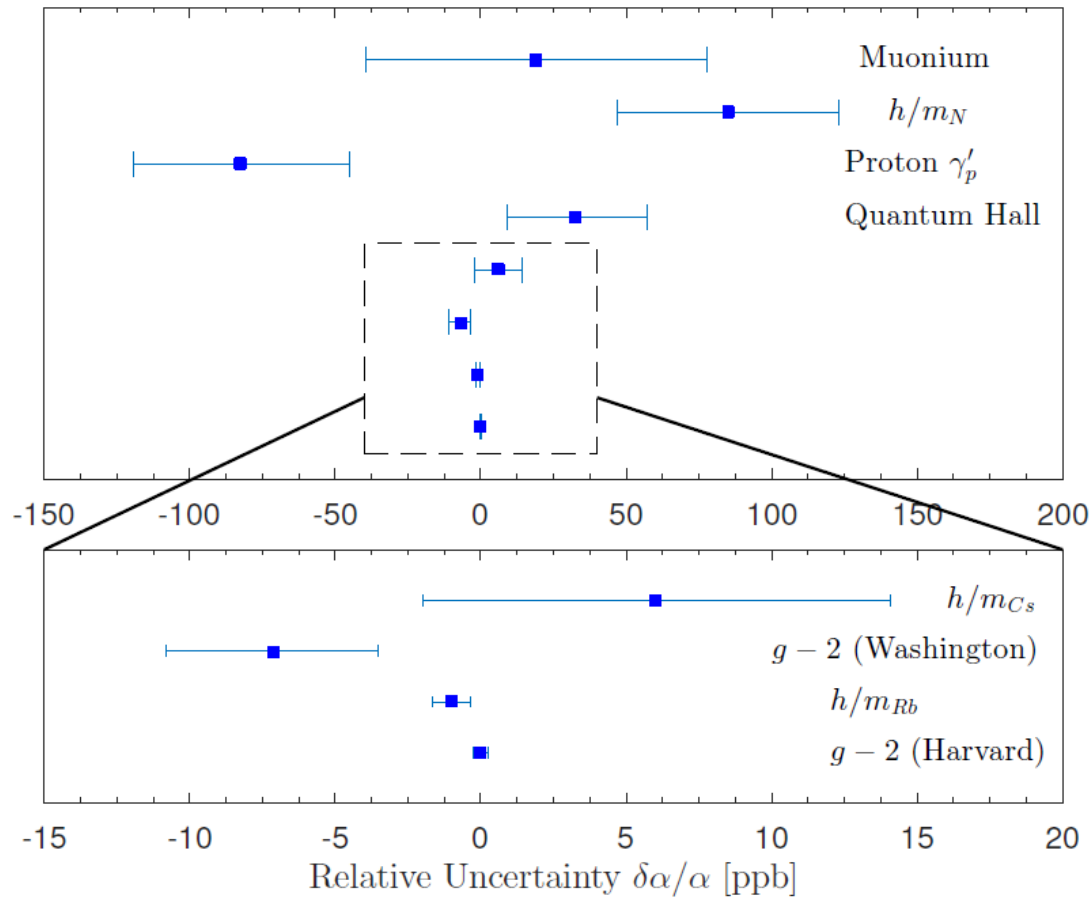
- Particles / waves
- Phase determined by Lagrangian,
 $L = E_{\text{kin}} - E_{\text{pot}}$
- Laser wavelength as a ruler



The fine structure constant

2014 CODATA

$$\alpha = \frac{1}{4\pi\epsilon_0} \frac{e^2}{\hbar c} = \frac{1}{137.035999139(31)} \quad (0.23\text{ppb})$$



The Fine Structure Constant

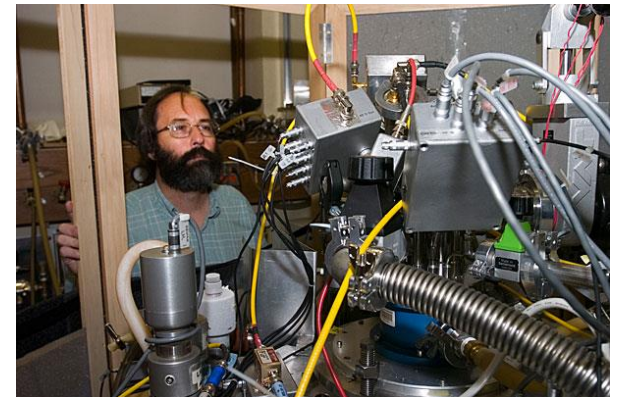
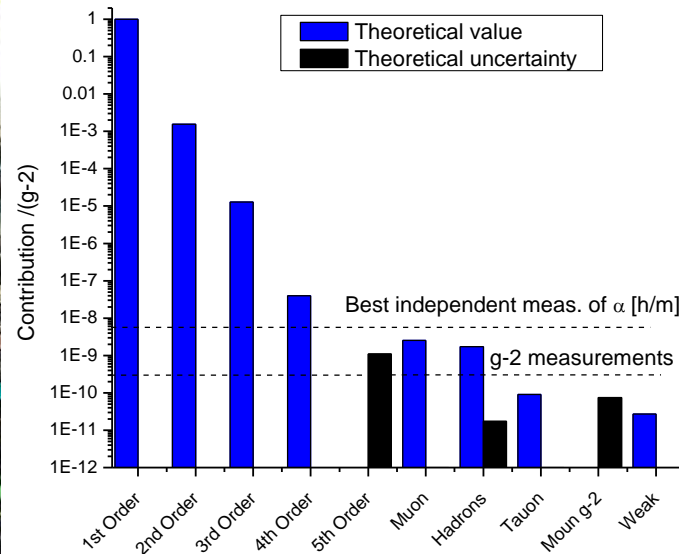
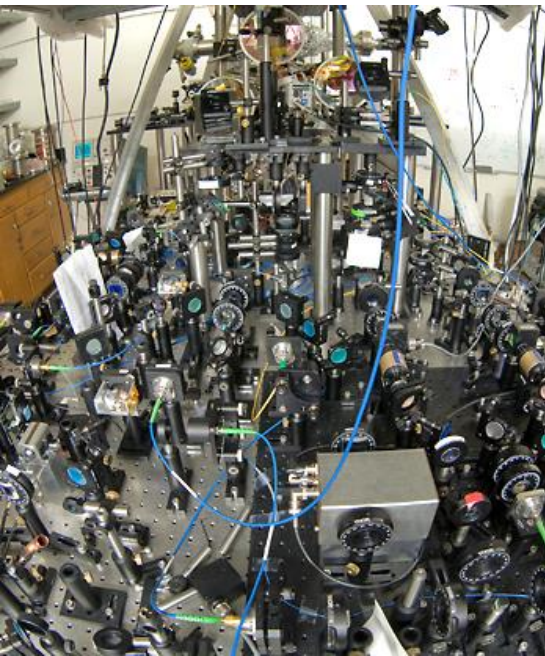
137 is...(<https://primes.utm.edu/curios/page.php/137.html>)

- The numerical value of “Kaballah” (קַבְּלָה)
- Genesis 25:17: “And these are the years of the life of Ishmael, an **hundred and thirty and seven** years. ...” It is also the age of Moses father and Levi
- The day before his inauguration, President Obama made a 137-mile train trip from Philadelphia to Washington, DC.
- There are 137 Hawaiian islands, islets, and shoals
- W. Pauli died in room 137 of Rotkreuz Hospital, Zurich.
- WMAP’s age of the universe is 13.7 billion years
- 33rd prime number
- The largest prime factor of 123456787654321
- Karpov and Kasparov played 137 draws in chess
- Chlorophyll ($C_{55}H_{72}MgN_4O_5$) consists of 137 atoms

The most precise theory/experiment comparison in science

Fine structure constant

Electron gyromagnetic moment



Dark photons shift magnetic moment versus fine structure constant

Measures how strong a magnet an electron is

Alpha in the atom recoil frequency

$$\alpha = \left[2 \frac{R_\infty}{c} \frac{u}{m_e} \frac{M}{u} \frac{h}{M} \right]^{1/2}$$

Rydberg Constant

0.007 ppb P. J. Mohr *et al.*,
Rev. Mod. Phys. **80**, 633 (2008).

Cs mass in u

0.18ppb M. P. Bradley *et al.*, Phys. Rev.
Lett. **83**, 4510 (1999).

Electron mass in atomic mass units u

0.43 ppb P. J. Mohr *et al.*, Rev. Mod. Phys. **80**, 633 (2008).

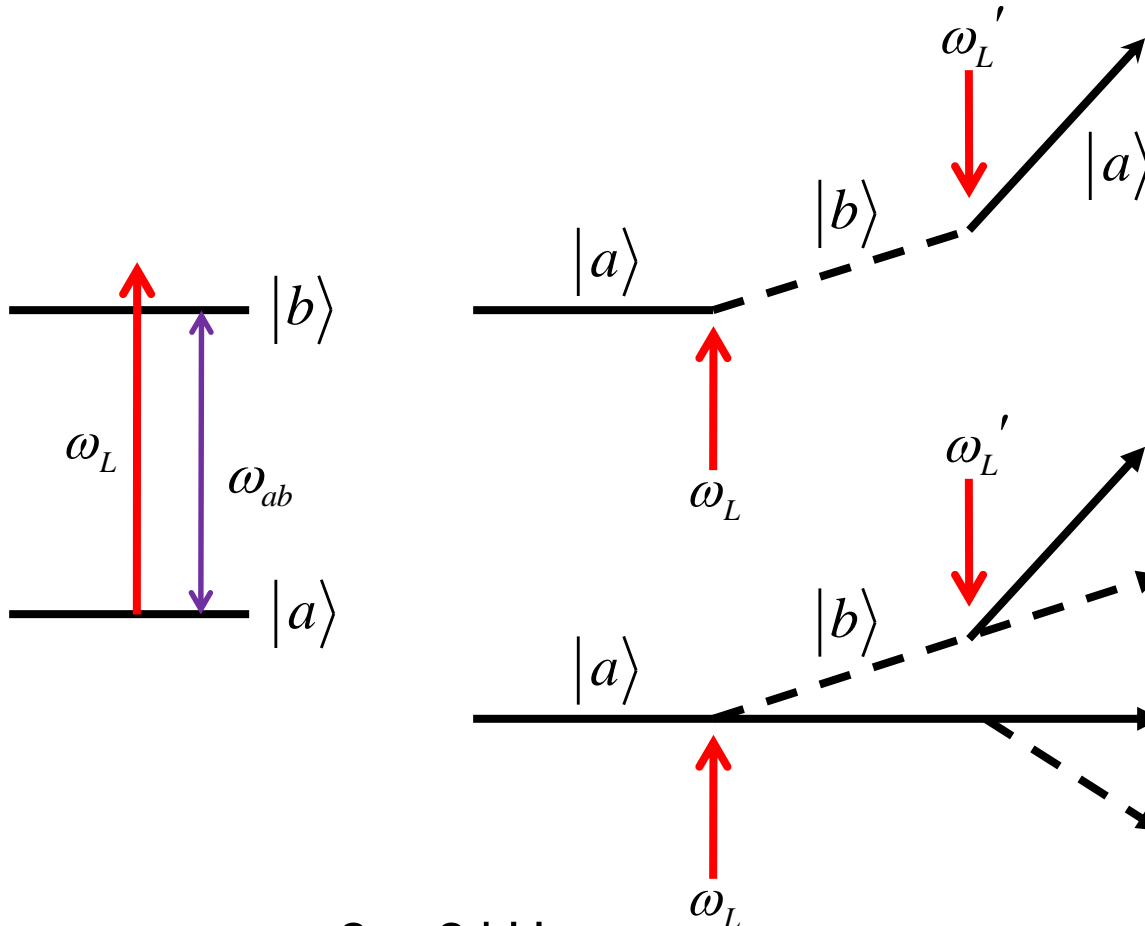
Determined by the atom recoil
frequency

$$\frac{h}{M} = \frac{4\pi c^2 \omega_r}{\omega^2}$$

Lowest : 0.23ppb

Cs D2 Transition
0.015ppb

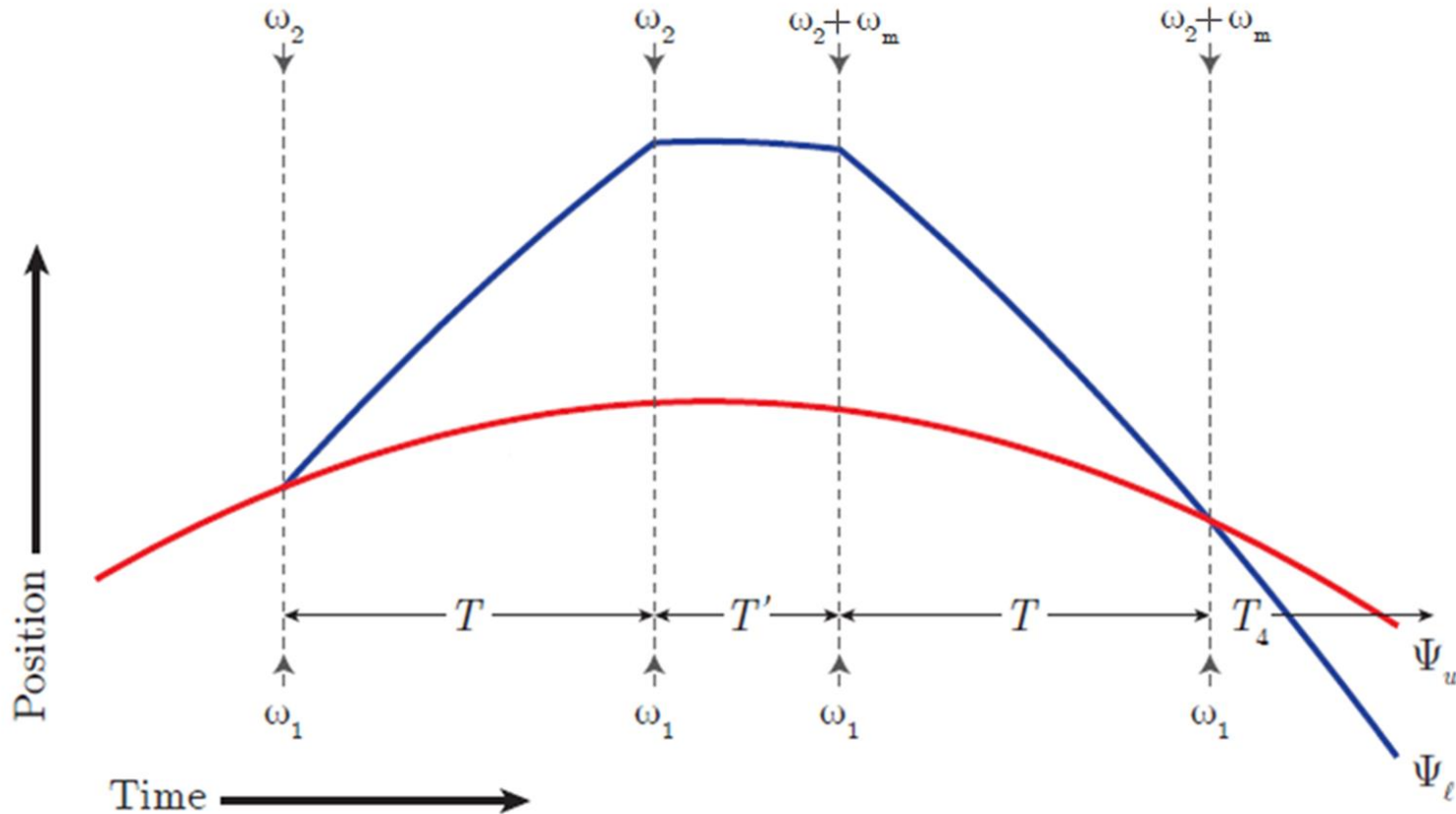
Recoil measurement



$$\omega'_L(|a\rangle) - \omega'_L(|b\rangle) = 2\omega_{rec}$$

- $\omega r \sim 2\pi \times 2$ kHz,
- Accuracy 10^{-10}
- Need to pinpoint resonance to $0.2 \mu\text{Hz}$ or 6×10^{-22}
- 10,000 times better accuracy than precision of best clocks

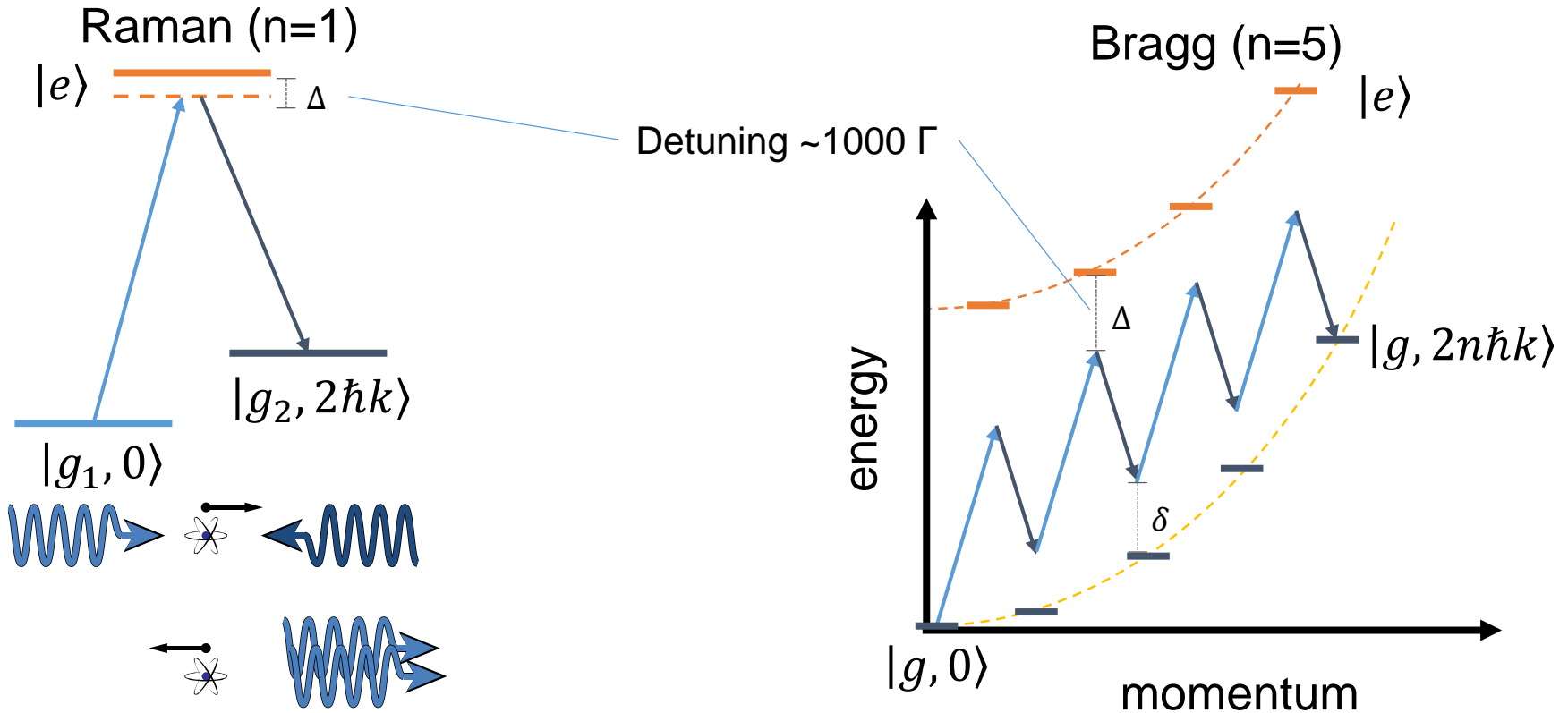
Ramsey-Borde interferometer



$$\Phi_{RB} = 8n^2 \omega_r T - 2nk g (T + T') T - n\omega_m T$$

$$\frac{1}{2} m v_r^2 = \hbar \left(\frac{\hbar k^2}{2m} \right) = \hbar \omega_r \quad \boxed{\frac{\omega_r}{k}} \Rightarrow \hbar/m \rightarrow \alpha$$

Multiphoton Bragg diffraction

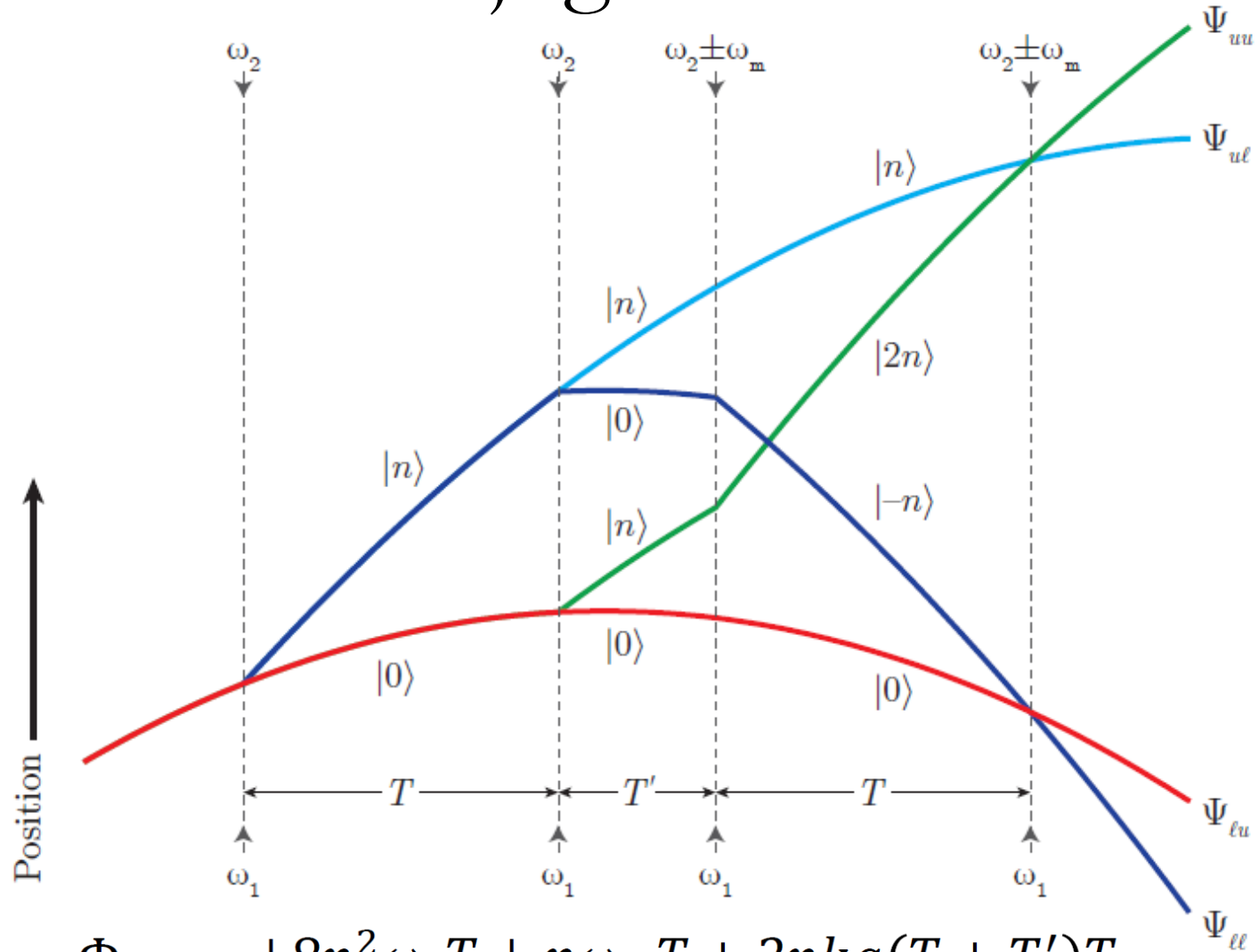


Bragg gives you:

- More photons transferred per pulse (higher sensitivity)
- Atoms stay in same internal state (Zeeman, AC Stark systematics suppressed)

$$\Phi_{RB,Diff} = 16n^2 \omega_r T - 2n\omega_m T$$

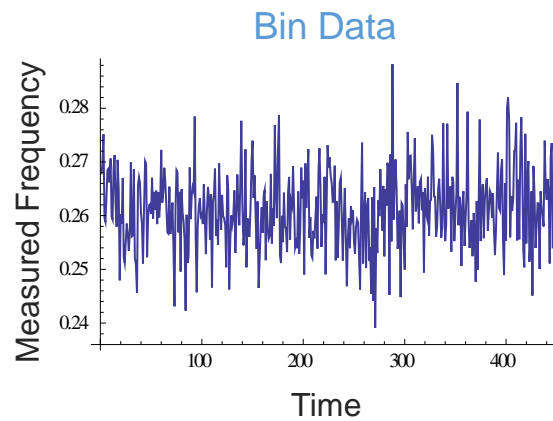
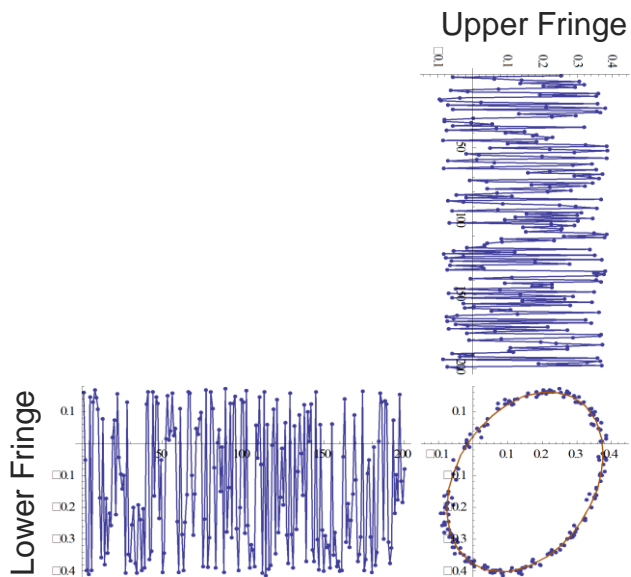
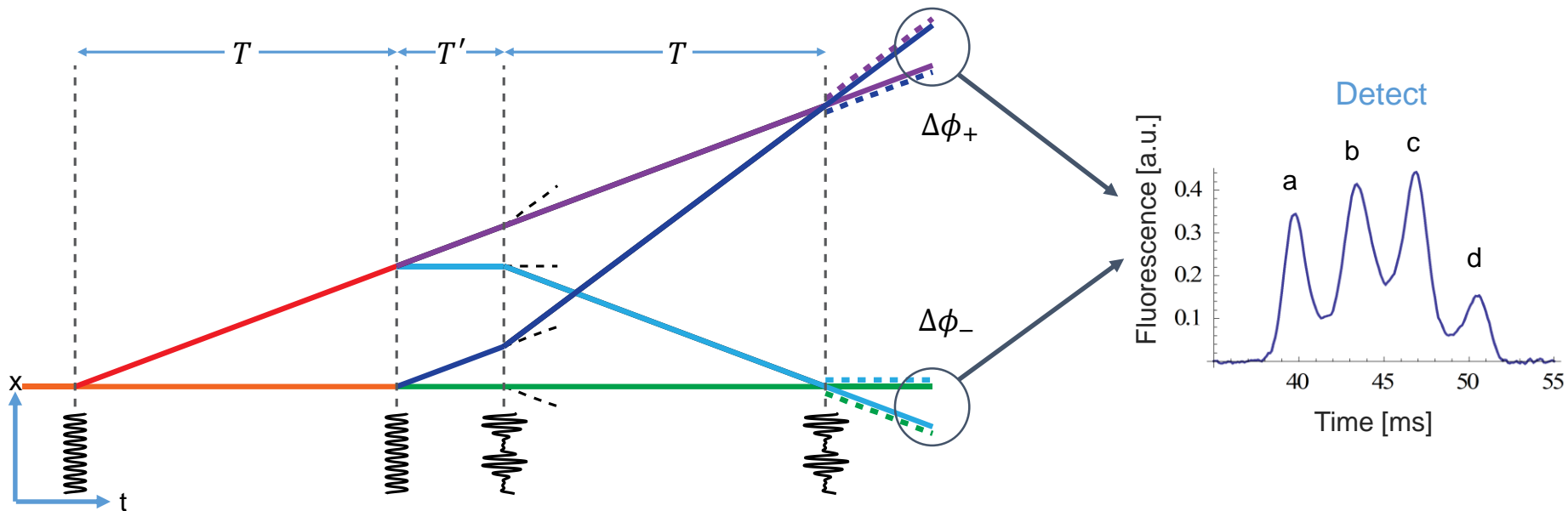
Simultaneous conjugate interferometers

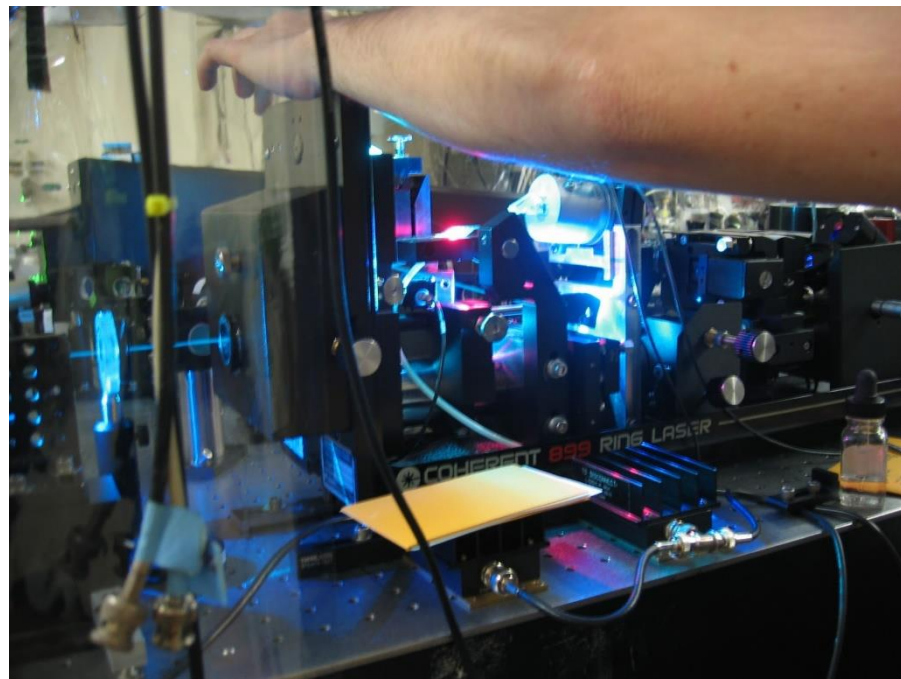
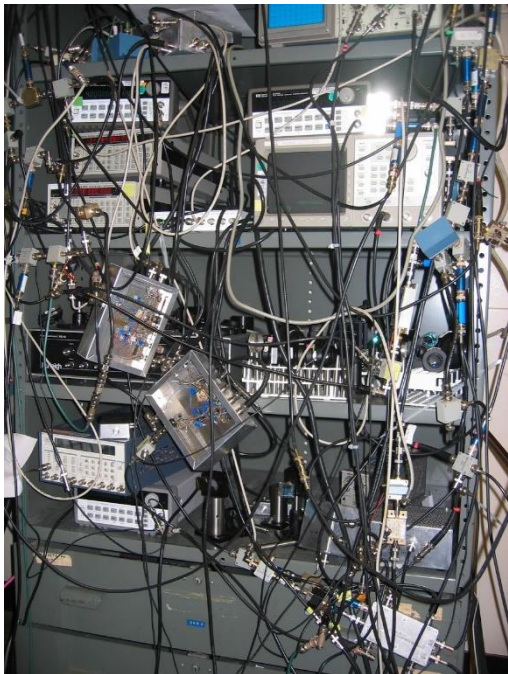


$$\Phi_{RB} = \pm 8n^2 \omega_r T \pm n \omega_m T + 2nkg(T + T')T$$

$$\Phi_{RB,Diff} = 16n^2 \omega_r T - 2n \omega_m T$$

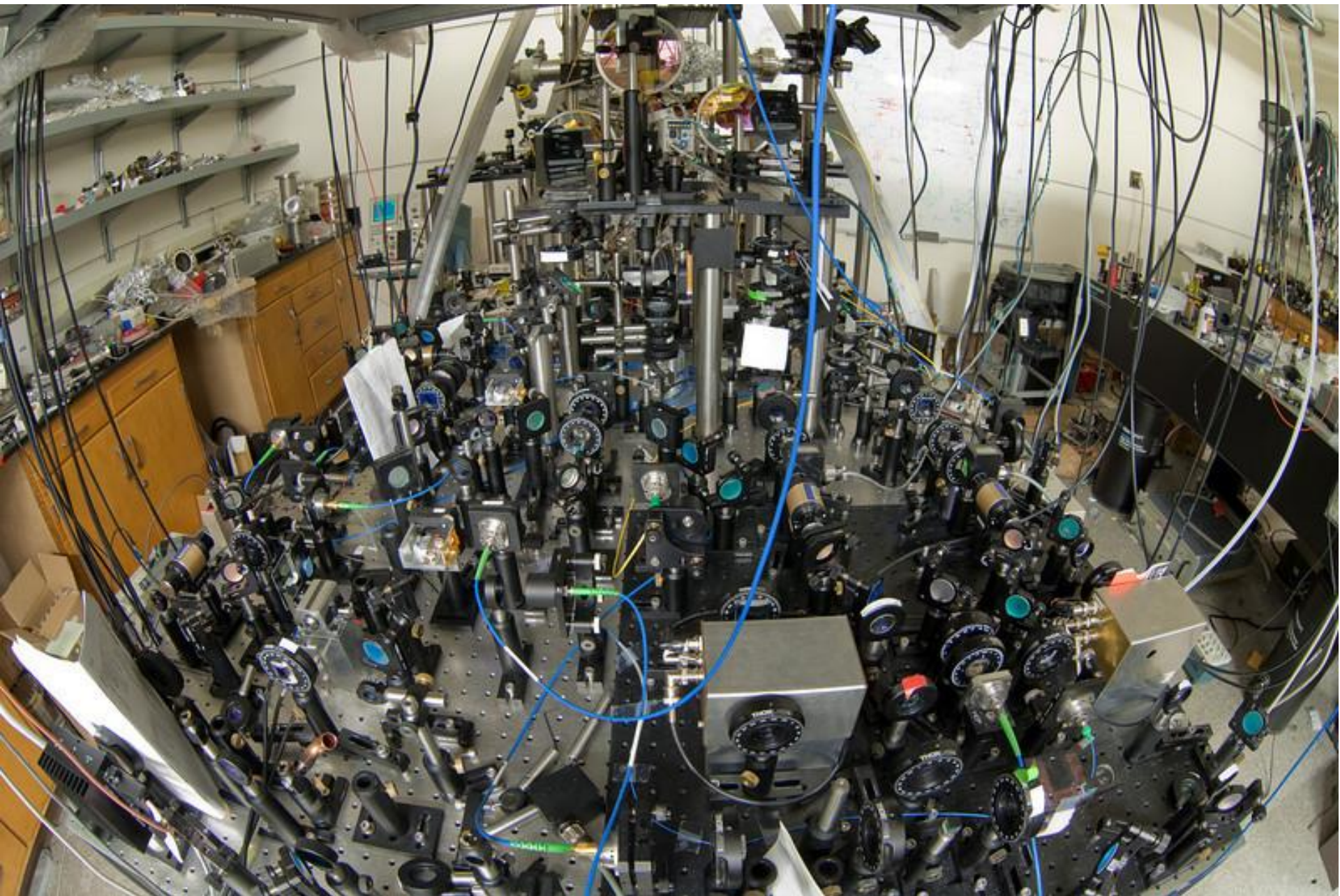
Phase extraction



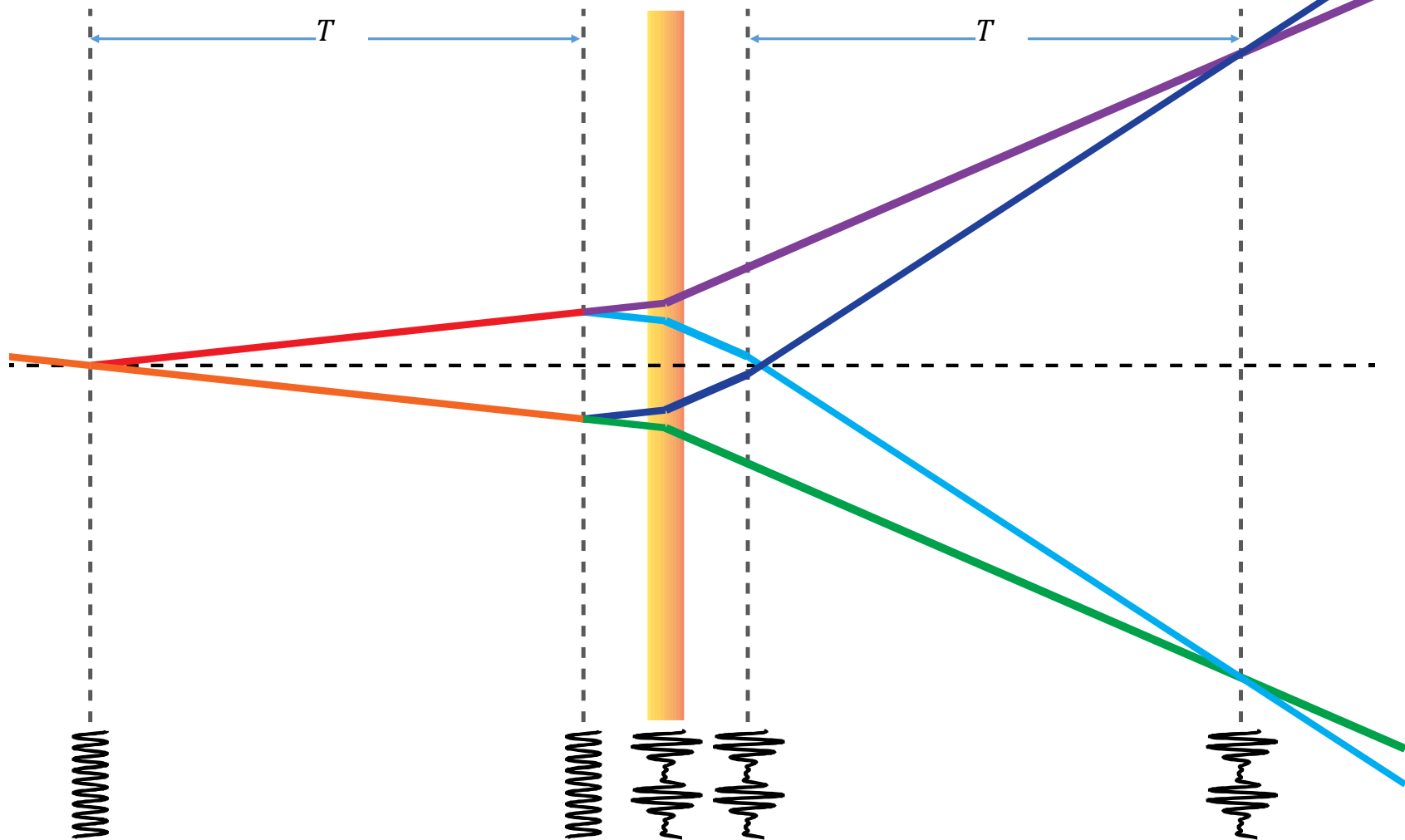


Now it's my turn to leave Stanford





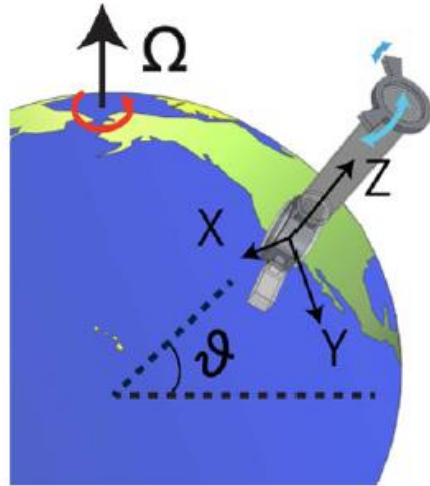
Bloch oscillations



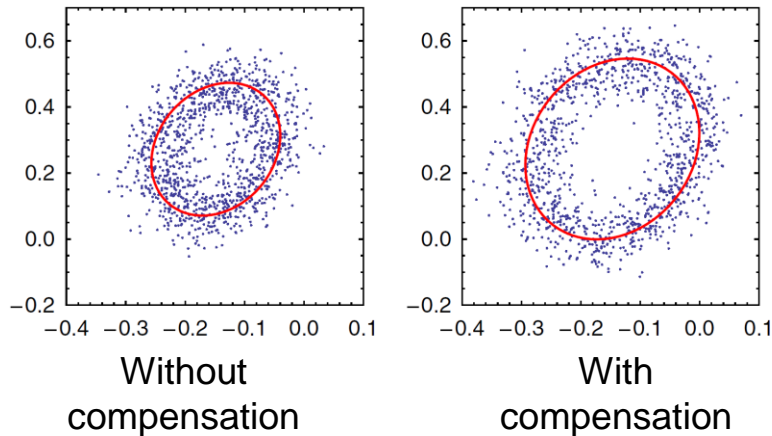
$$\Delta\Phi_{RB+Bloch} = 16n(n + N)\omega_r T - 2n\omega_m T$$

Tricks for Increased Sensitivity

Coriolis Compensation



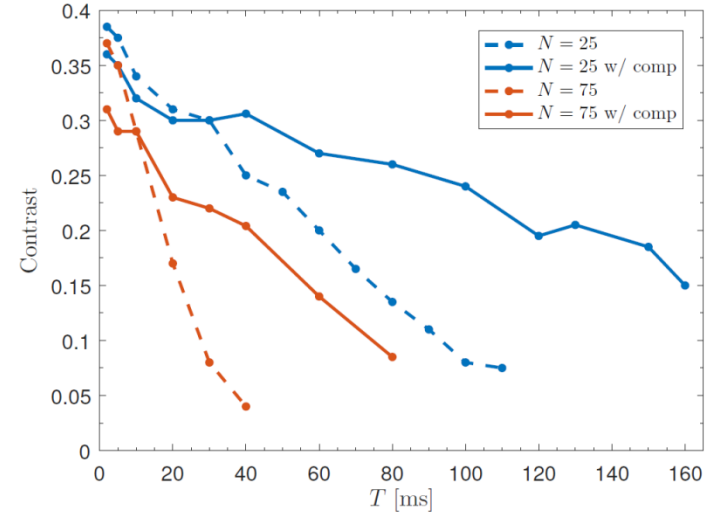
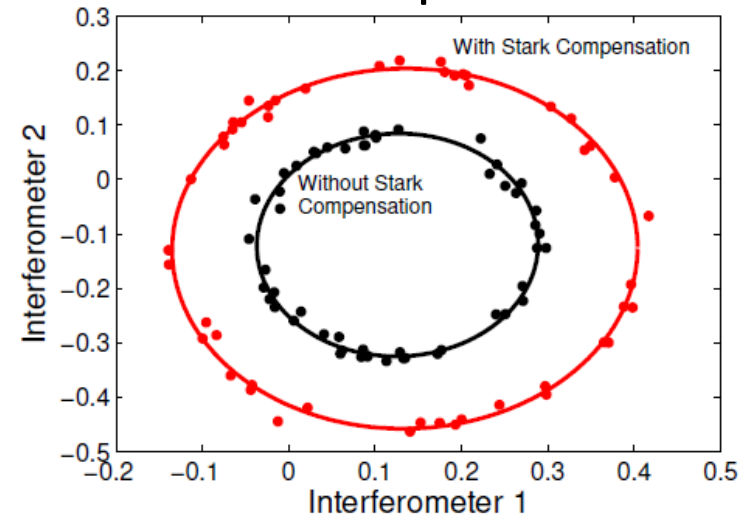
$10\hbar k, T = 180\text{ms}$



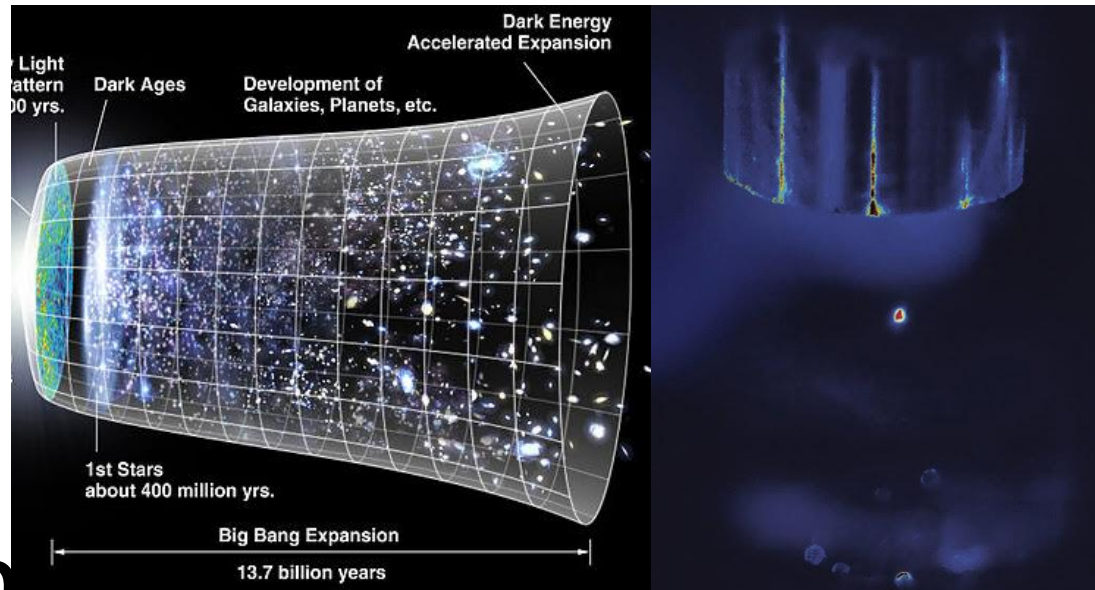
x3.5 contrast gain

>12Mrad phase diff. measurable!

Stark Compensation



Up to $N=200$



Dark energy scalar fields

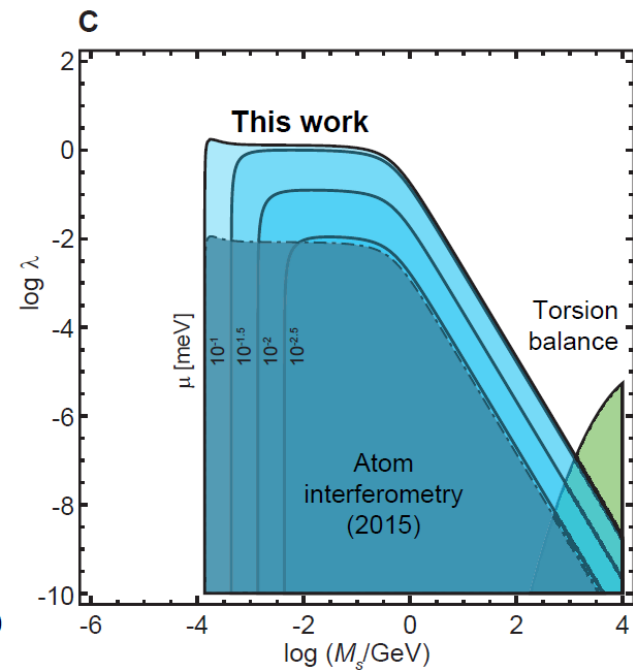
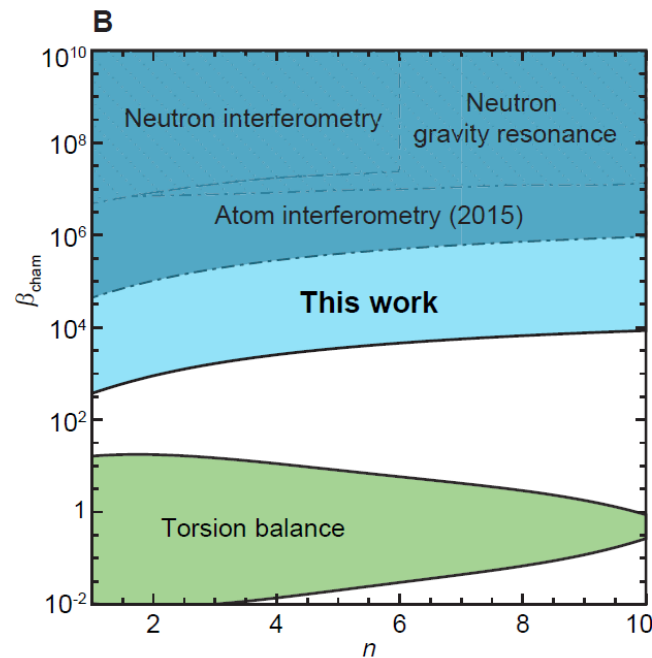
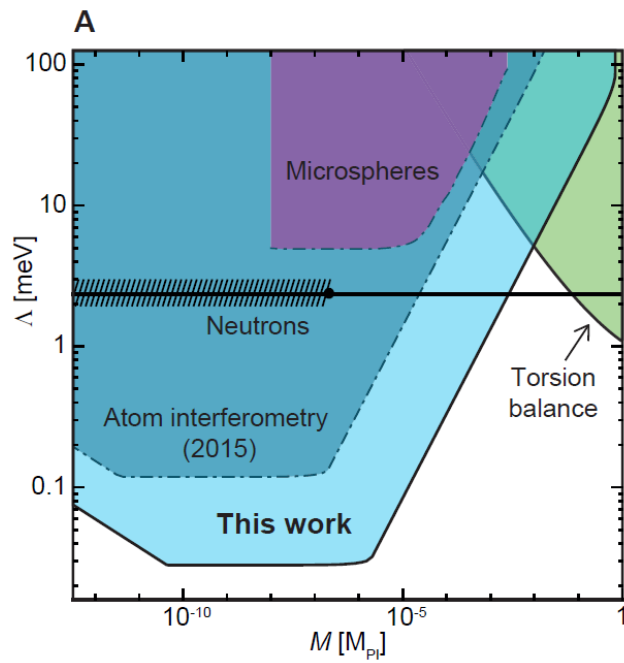
M. Jaffe, P. Haslinger, V. Xu, J. Khoury, B. Elder, M. Upadhye and H. Muller



Limits on dark-energy scalars

Chameleons

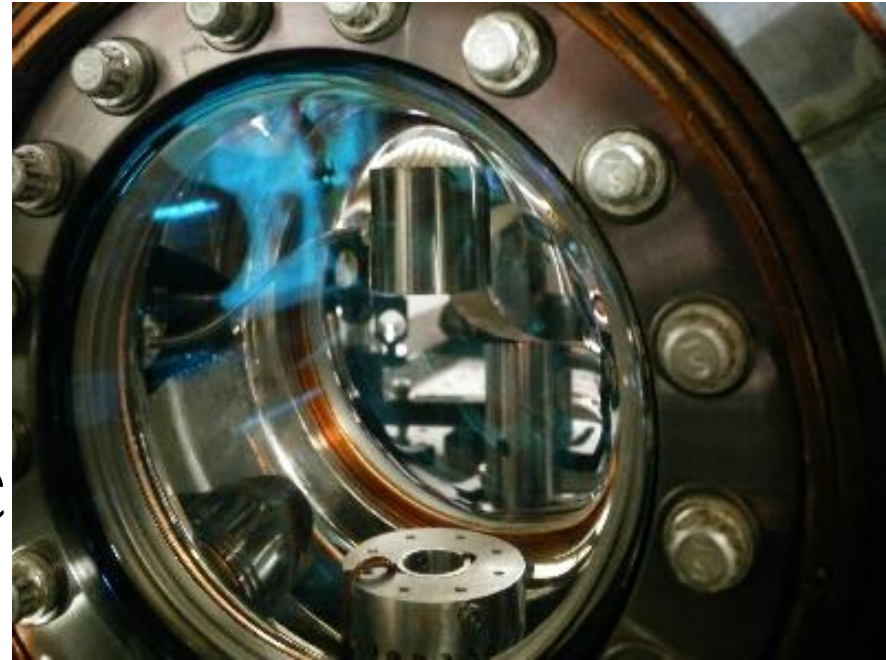
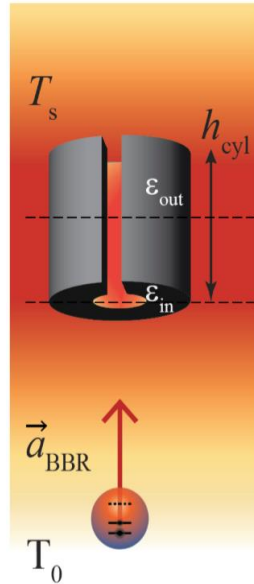
Symmetrons



Interaction strength

M. Jaffe, P. Haslinger, V. Xu, P. Hamilton, A. Upadhye, B. Elder, J. Khoury, and HM, [Nature Physics](https://doi.org/10.1038/nphys4189), doi: 10.1038/nphys4189

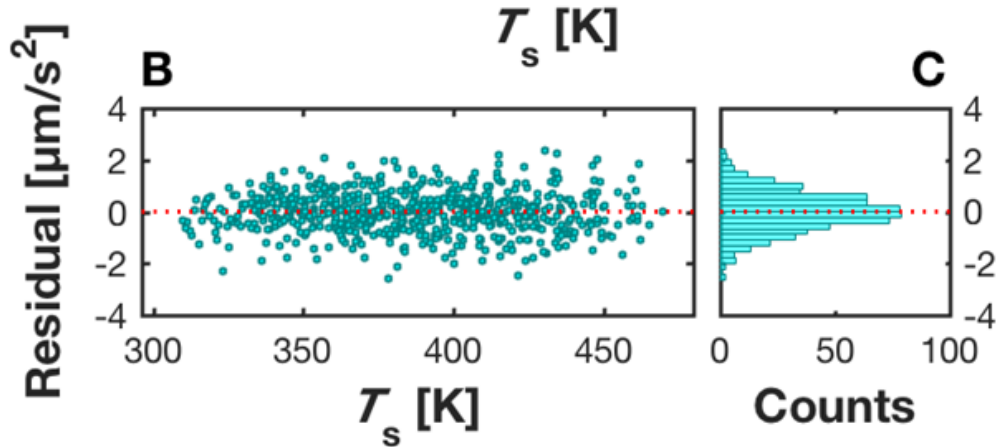
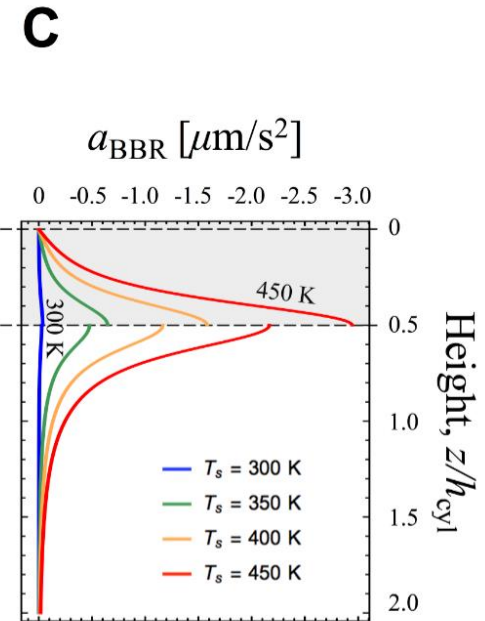
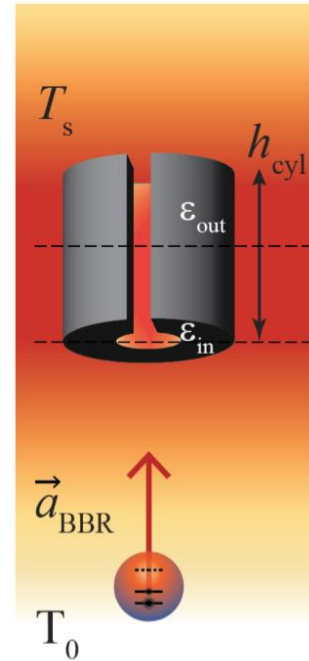
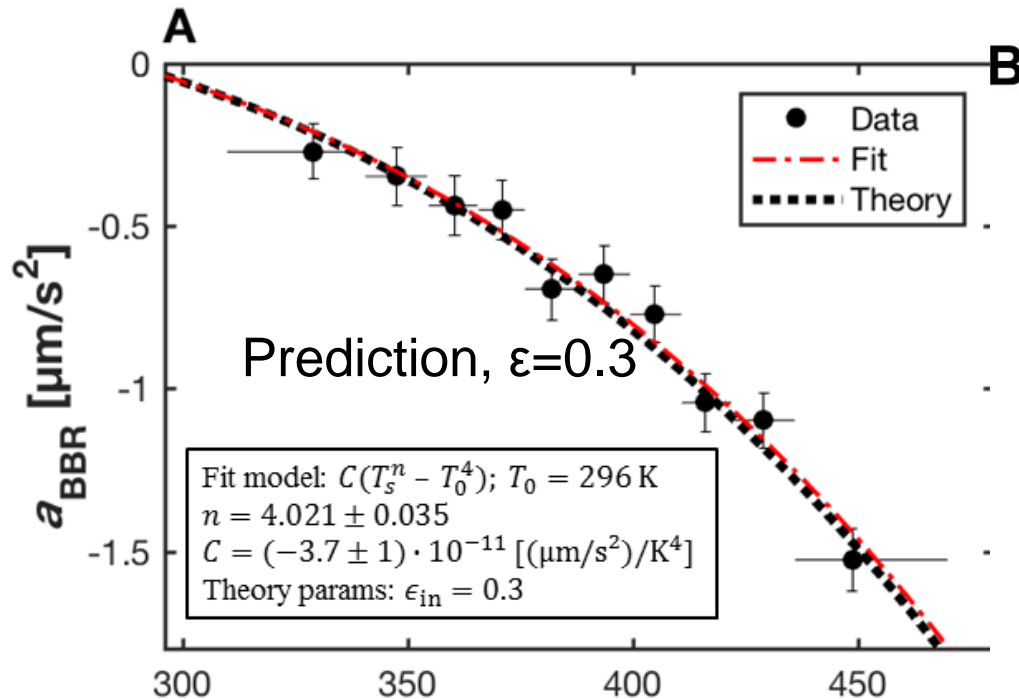
Attractive blackbody radiation



P. Haslinger, M. Jaffe, V. Xu, M. Sonnleithner, H. Ritsch, M. Ritsch & HM

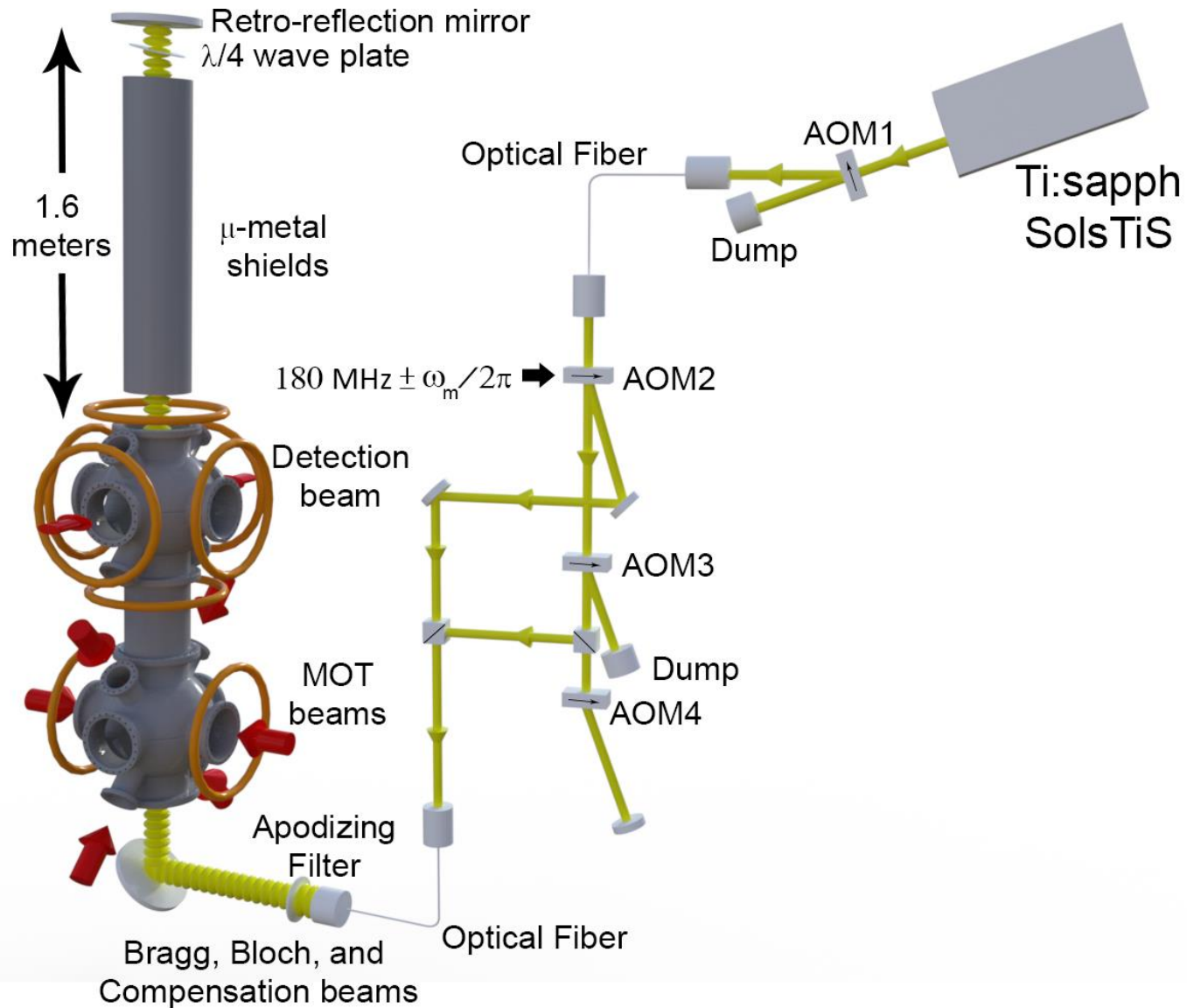
Only one room temp. blackbody photon absorbed every 10^5 years...

Force from blackbody radiation

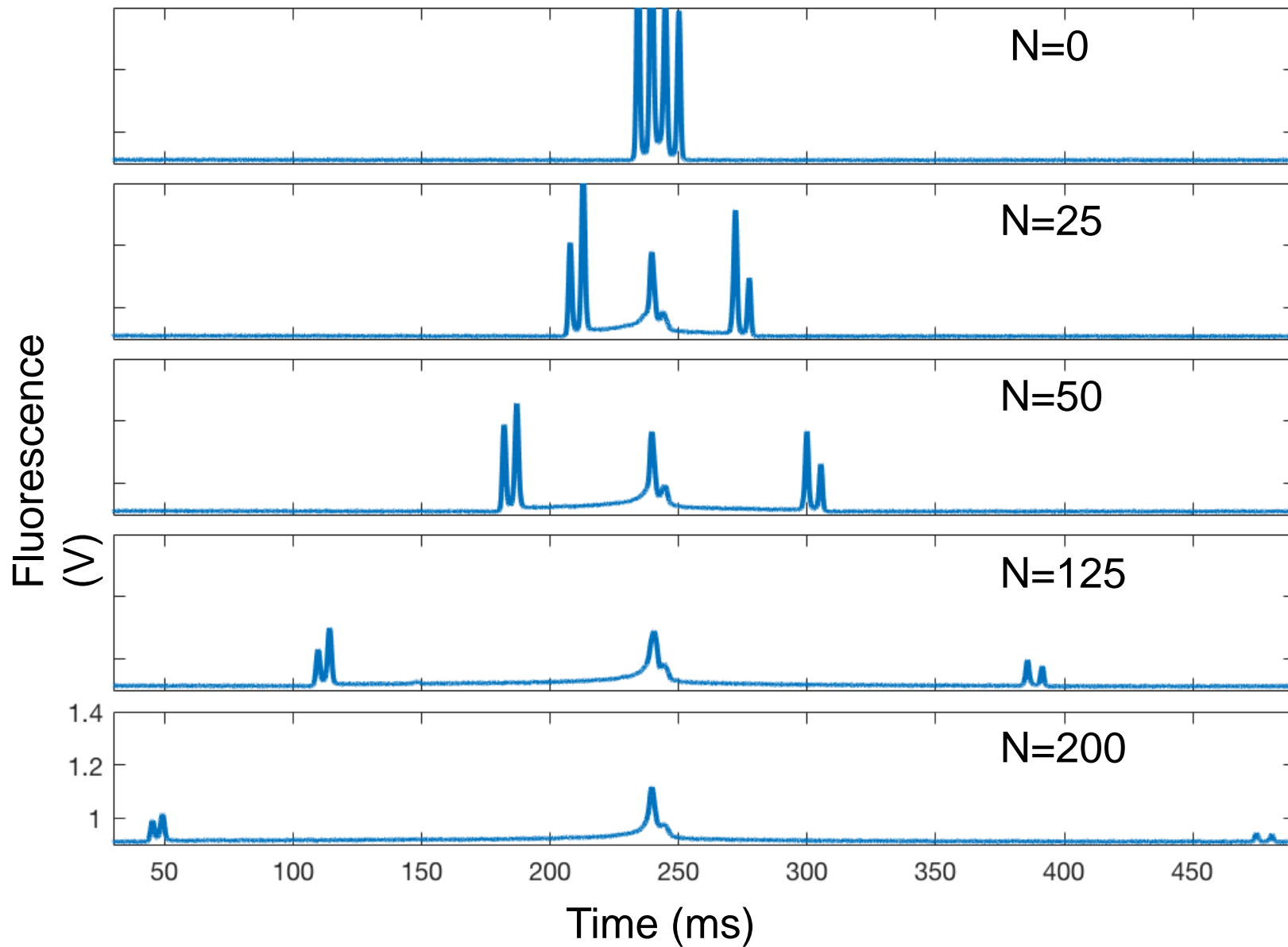


arXiv:1704.03577

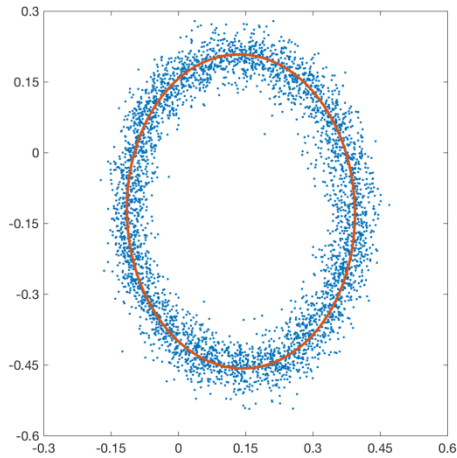
Setup



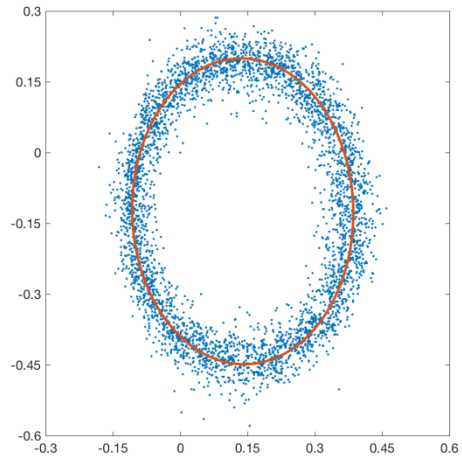
Fluorescence Traces



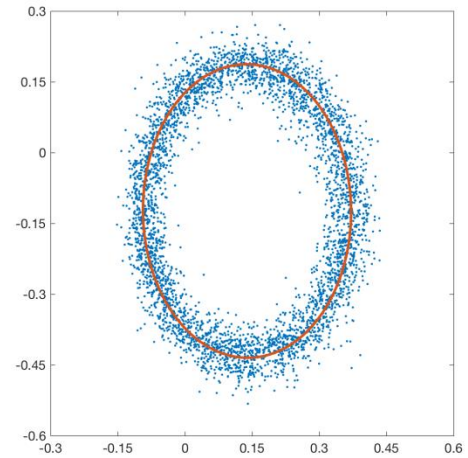
$n=5, N=125$ Ellipses



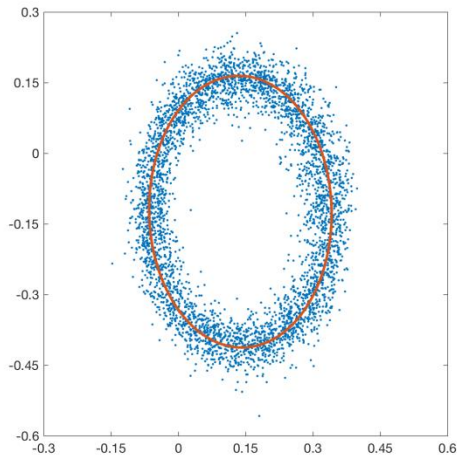
$T=5$ ms



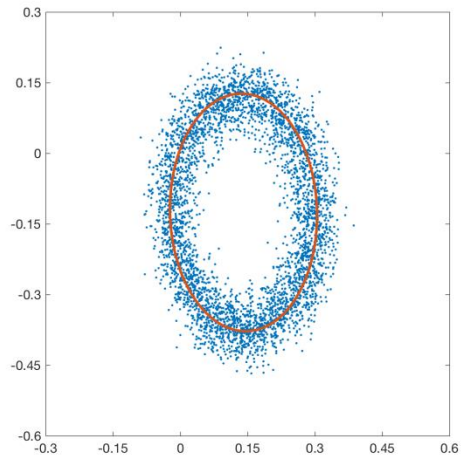
$T=10$ ms



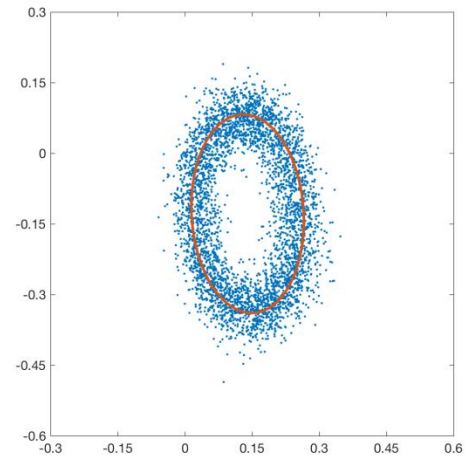
$T=20$ ms



$T=40$ ms

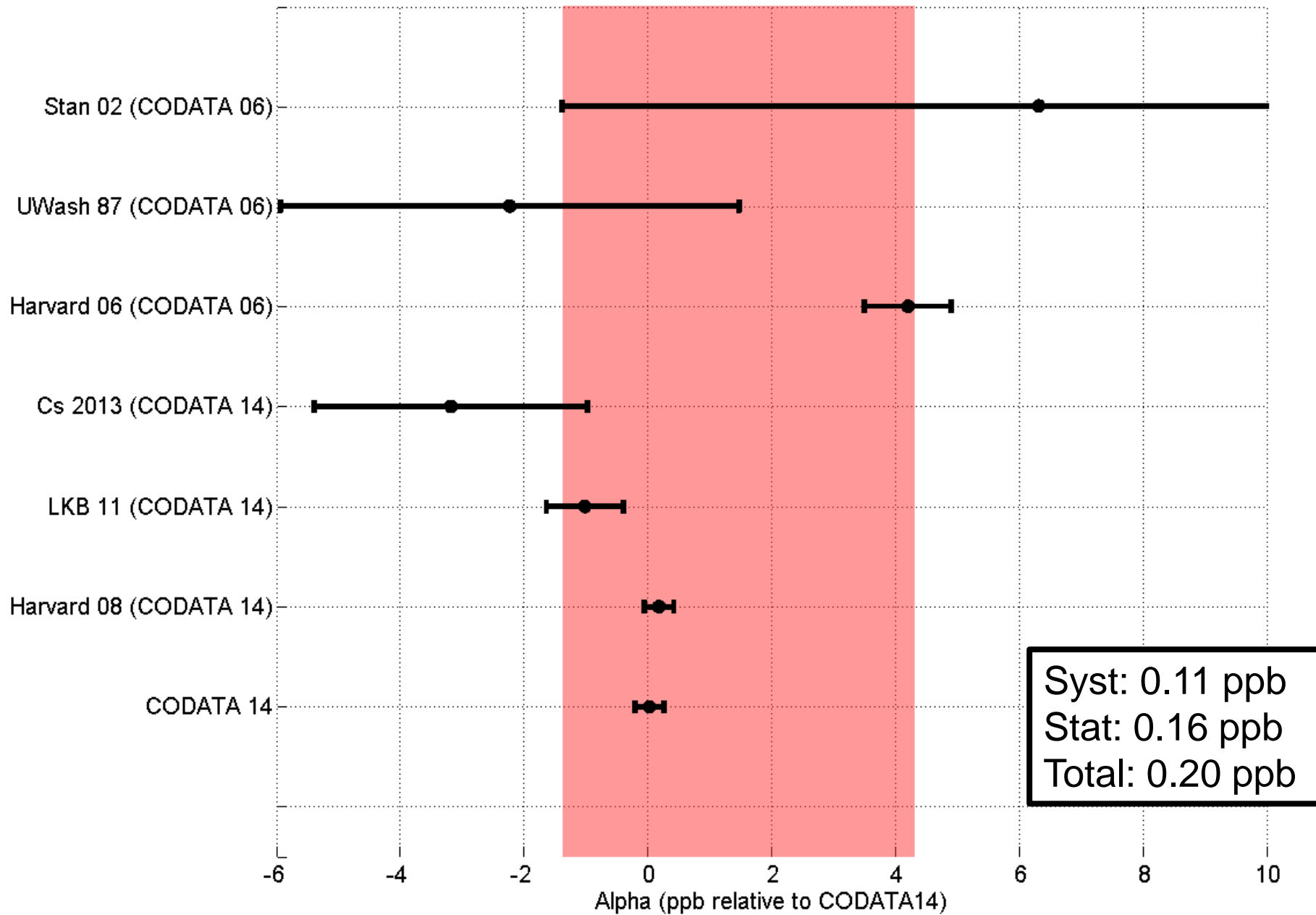


$T=60$ ms



$T=80$ ms

Blind Analysis



0.16 ppb systematic errors

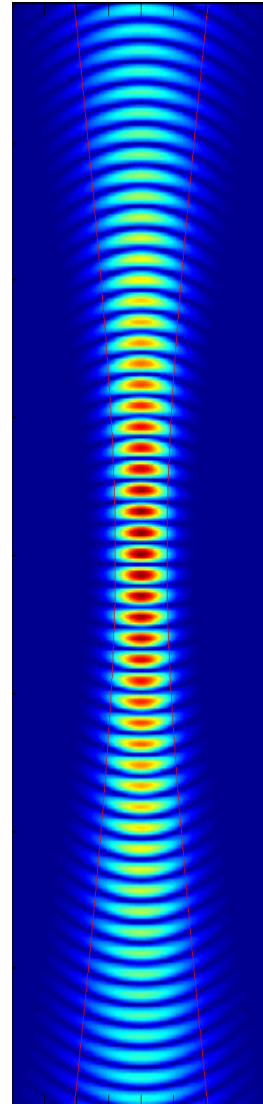
Effect	Sect.	Value	$\delta\alpha/\alpha$ (ppb)
Laser Frequency	1	N/A	-0.24 ± 0.03
Acceleration Gradient	4A	$\square=(2.13 \pm 0.01)\times 10^{-6}/s^2$	-1.69 ± 0.02
Gouy phase	3	$w_0=3.21\pm 0.008$ mm, $z_0=0.5\pm 1.0$ m	-3.60 ± 0.03
Wavefront Curvature	12	$\langle r^2 \rangle^{1/2}=0.58$ mm	0.15 ± 0.03
Beam Alignment	5	N/A	0.05 ± 0.03
BO Light Shift	6	N/A	0 ± 0.004
Density Shift	7	$\rho=10^6$ atoms/cm ³	0 ± 0.003
Index of Refraction	8	$n_{\text{cloud}}-1=30\times 10^{-12}$	0 ± 0.03
Speckle Phase Shift	4B	N/A	0 ± 0.04
Sagnac Effect	9	N/A	0 ± 0.001
Mod. Frequency Wavenumber	10	N/A	0 ± 0.001
Thermal Motion of Atoms	11	N/A	0 ± 0.08
Non-Gaussian Waveform	13	N/A	0 ± 0.03
Parasitic Interferometers	14	N/A	0 ± 0.03
Total Systematic Error			-5.33 ± 0.12
Total Statistical Error			± 0.16
Electron Mass (18)		$5.48579909067\times 10^{-4}$ u	± 0.02
Cesium Mass (4,17)		132.9054519615 u	± 0.03



Big

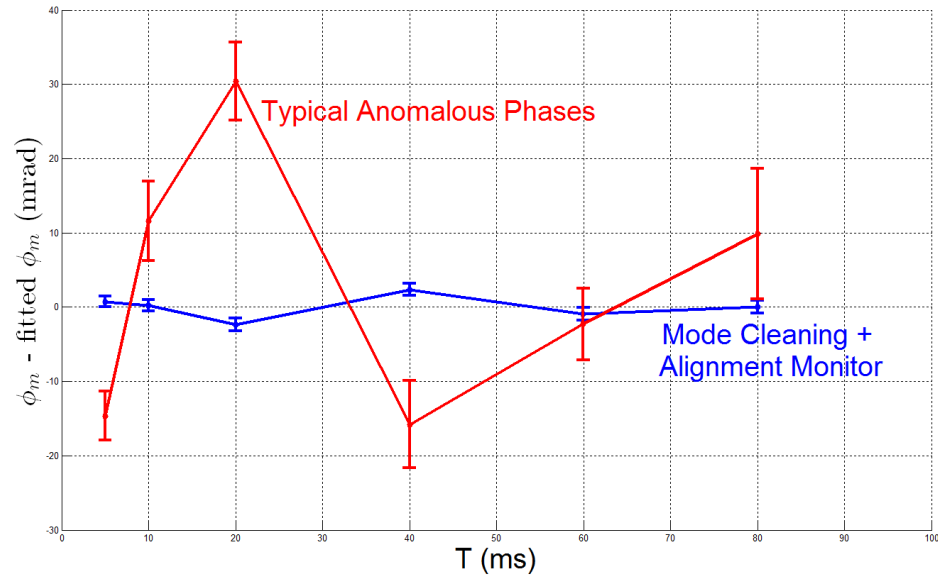
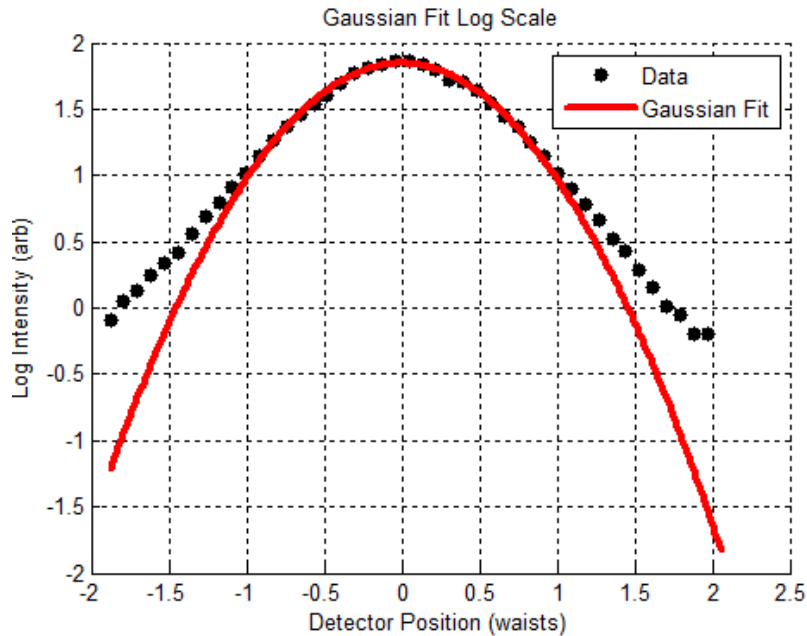


'New'



Speckle Phase

- 30 mrad anomaly \rightarrow 8 ppb at $N=0$
- $>10x$ suppression by mode filtering
- 25x suppression by $N=125$



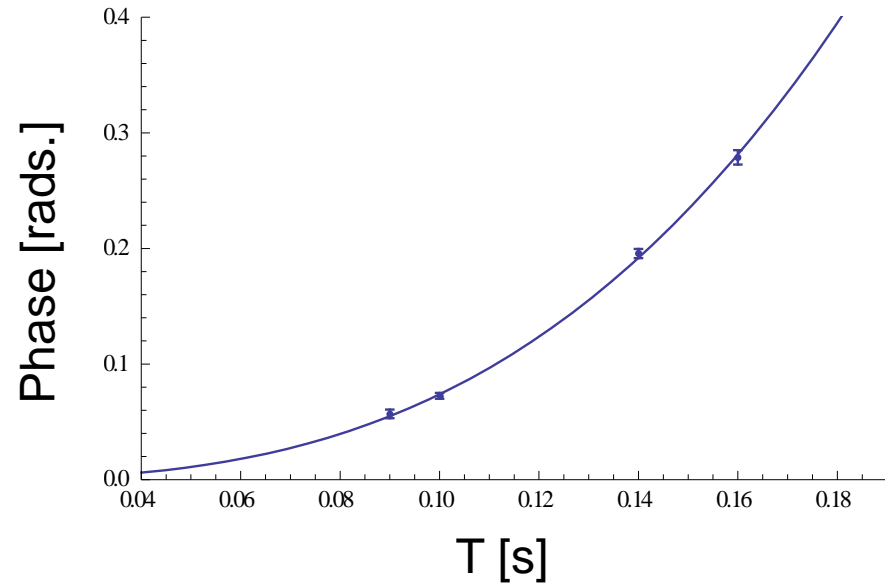
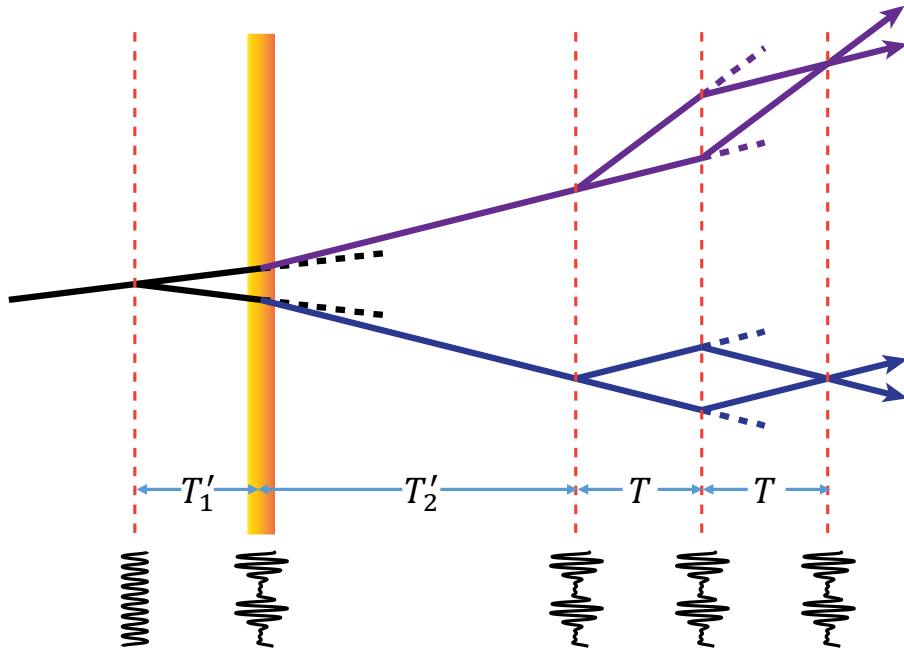
\$200 Apodizing Filter



- Fiber doesn't make Gaussian beams
- Spatial Filtering via Apodizer + Fountain Alignment Monitor

Gravity Gradient

$$\Delta\Phi = 16n(n + N)\omega_r T - 2n\omega_m T + \frac{4}{3}n\omega_r\gamma T \left[n(2T^2 + 3T(T'_1 + T'_2) + 3(T'_1 + T'_2)^2) + N(2T^2 + 6TT'_2 + 6T'^2_2) \right]$$



$$\Delta\phi = 2\gamma n\omega_r T^2 (2N(2T + T'_2) + n(2T + T'_1 + T'_2))$$

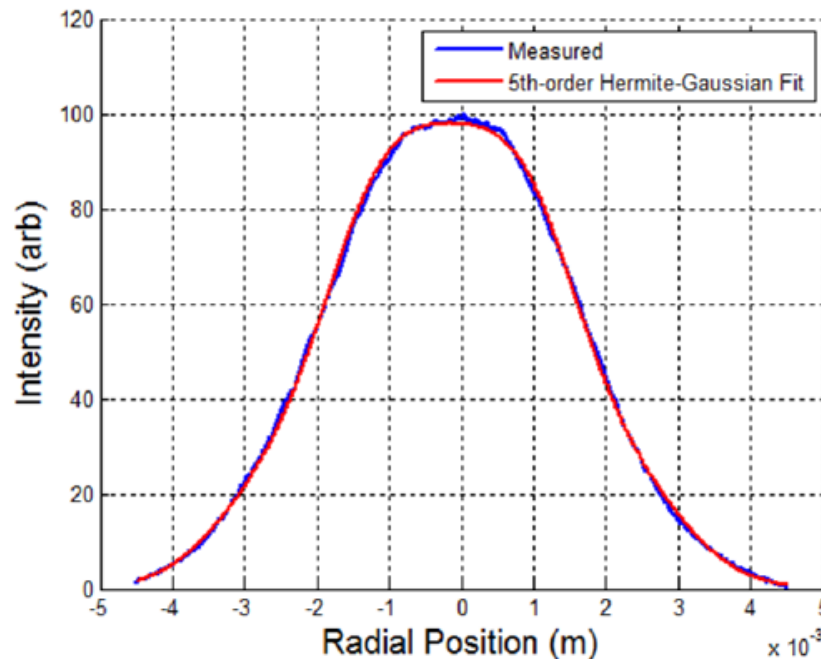
$$\gamma = 1.295(32) \times 10^{-6} \frac{m}{s^2} \frac{1}{m}$$

Shift in alpha = -1.41 +/- 0.02 ppb

Revised Gouy Phase

$$\frac{\delta k_{\text{eff}}}{k_{\text{eff}}} = -\frac{\lambda^2}{2\pi^2 w_0^2} \left(1 - \frac{z_0^2}{z_R^2} - \frac{r^2}{w_0^2} \right)$$

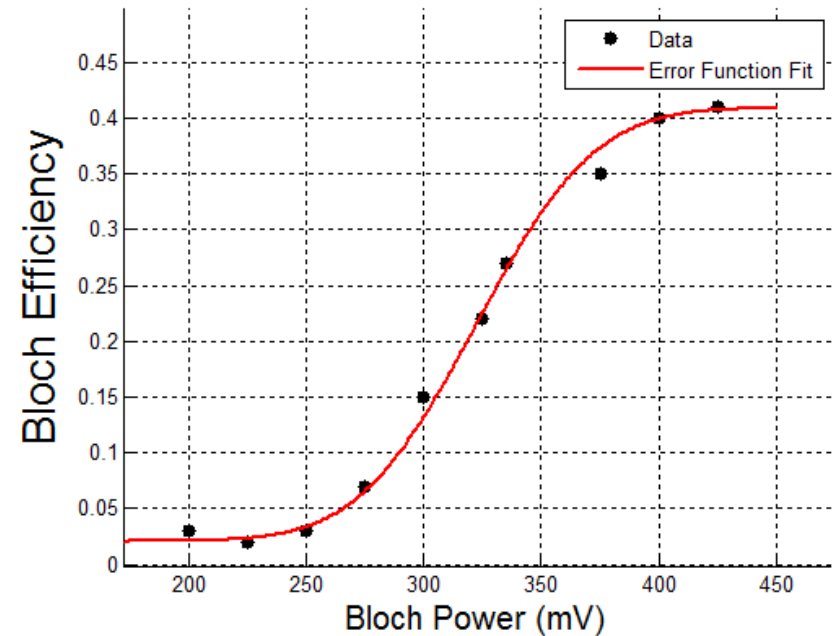
- Previously used knife-edge measurements to verify beam was Gaussian (within error)
- Suspected not Gaussian, based on 3D propagation
- With Scanning-slit/CCD, determined not Gaussian
- Use Monte Carlo to determine on-axis and wavefront-curvature corrections



French Effect

$$\frac{\delta k}{k} = \frac{1}{2k^2} \frac{\nabla_{\perp}^2 E}{E}$$

- Small-scale intensity variations can lead to dramatic changes in Gouy phase
- Doesn't average out!
- Can be >10ppb for LKB
- Use 3D Monte Carlo, CCD images, and Bloch Efficiency data to estimate effect
- <0.1 ppb for us



Systematic Checks

Variations of alpha w.r.t.:

- Bloch order
- Bloch power
- Contrast
- Detection region
- Pulse intensity: overall and pulse/pulse ratio
- Speckle phase
- ω_m mixing (RF)
- ω_m mixing (optics)
- Delay of interferometer sequence
- Bias B-field
- Single-photon detuning
- Data Analysis parameters (cuts, fitting, etc.)
- Fountain alignment (launch direction, no spatial filtering)

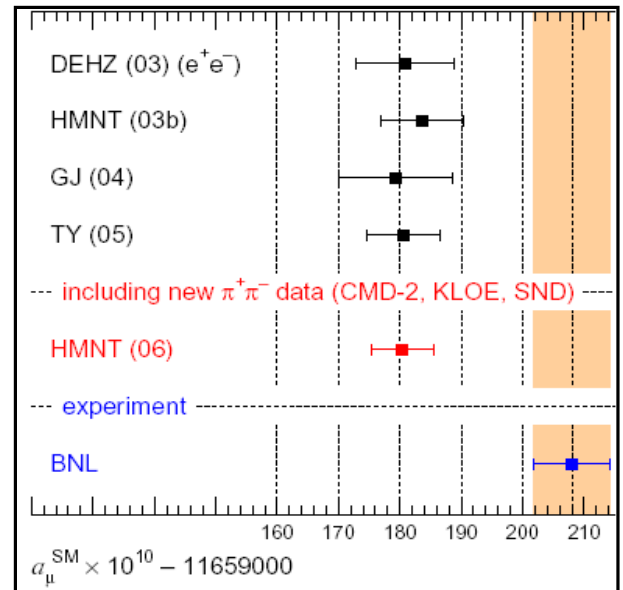
Dark photons

Whatever the dark sector is made of, only three interactions are allowed by standard model symmetries

- **Vector portal** “massive photon”
- Higgs portal
- Neutrino portal

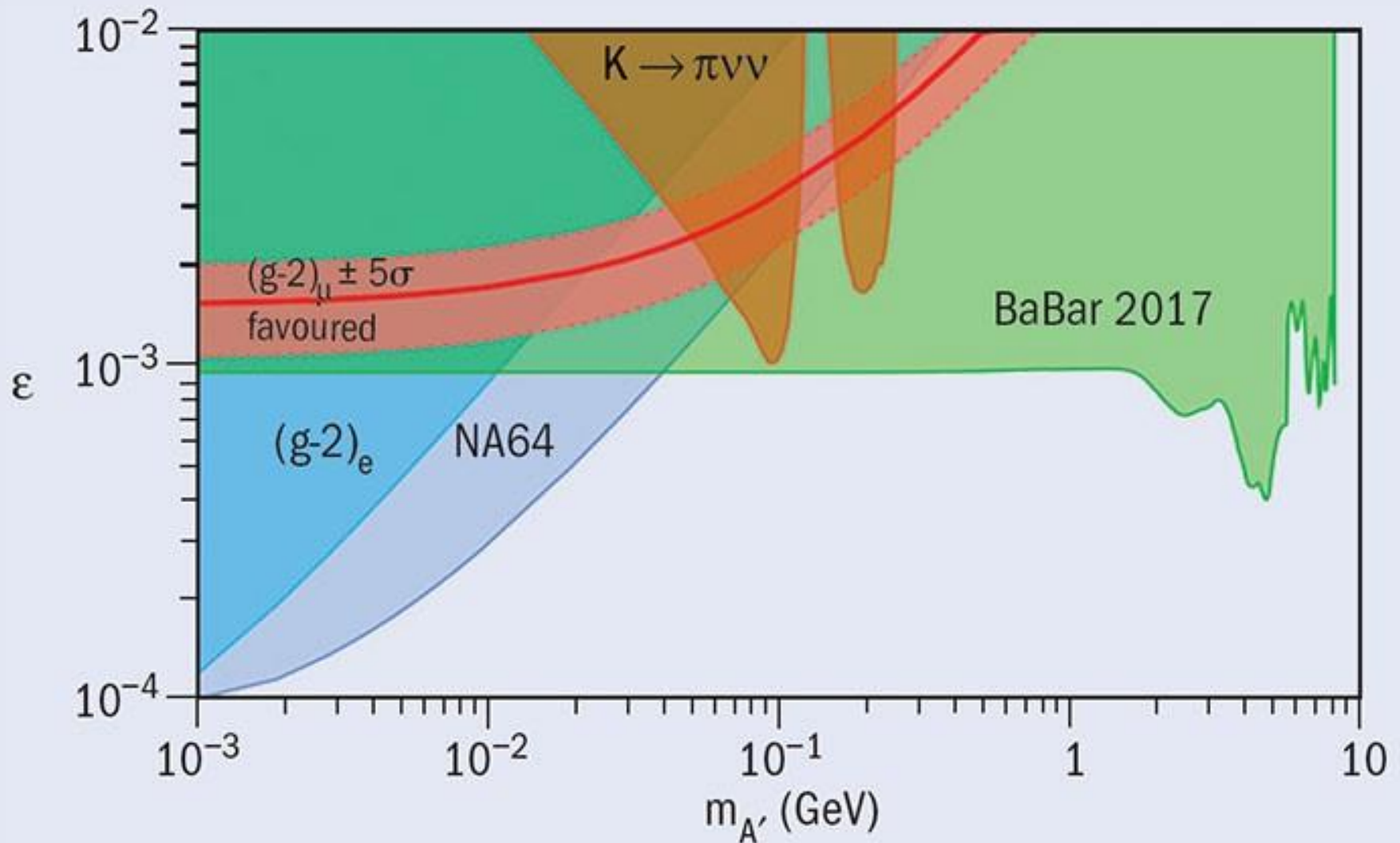
Hints

- Muon $g-2$
- Proton radius puzzle?
- ^8Be decay
- Astrophysical hints?
 - 511 keV line
 - keV gamma-ray excess
 - Galactic center excess



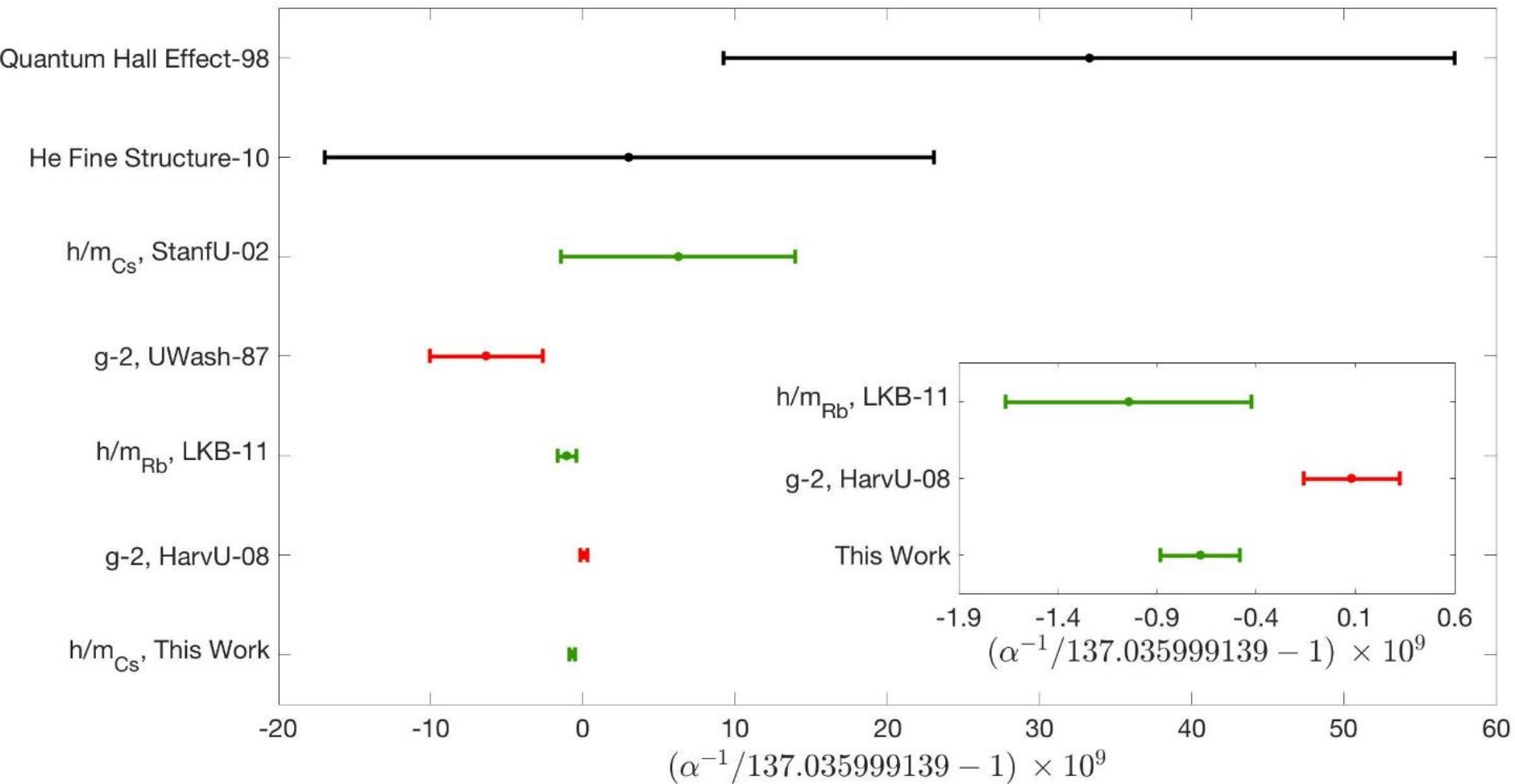
“Arguably, the strongest experimental evidence for physics beyond the standard model” (David Hertzog)

Ongoing dark photon searches



Results

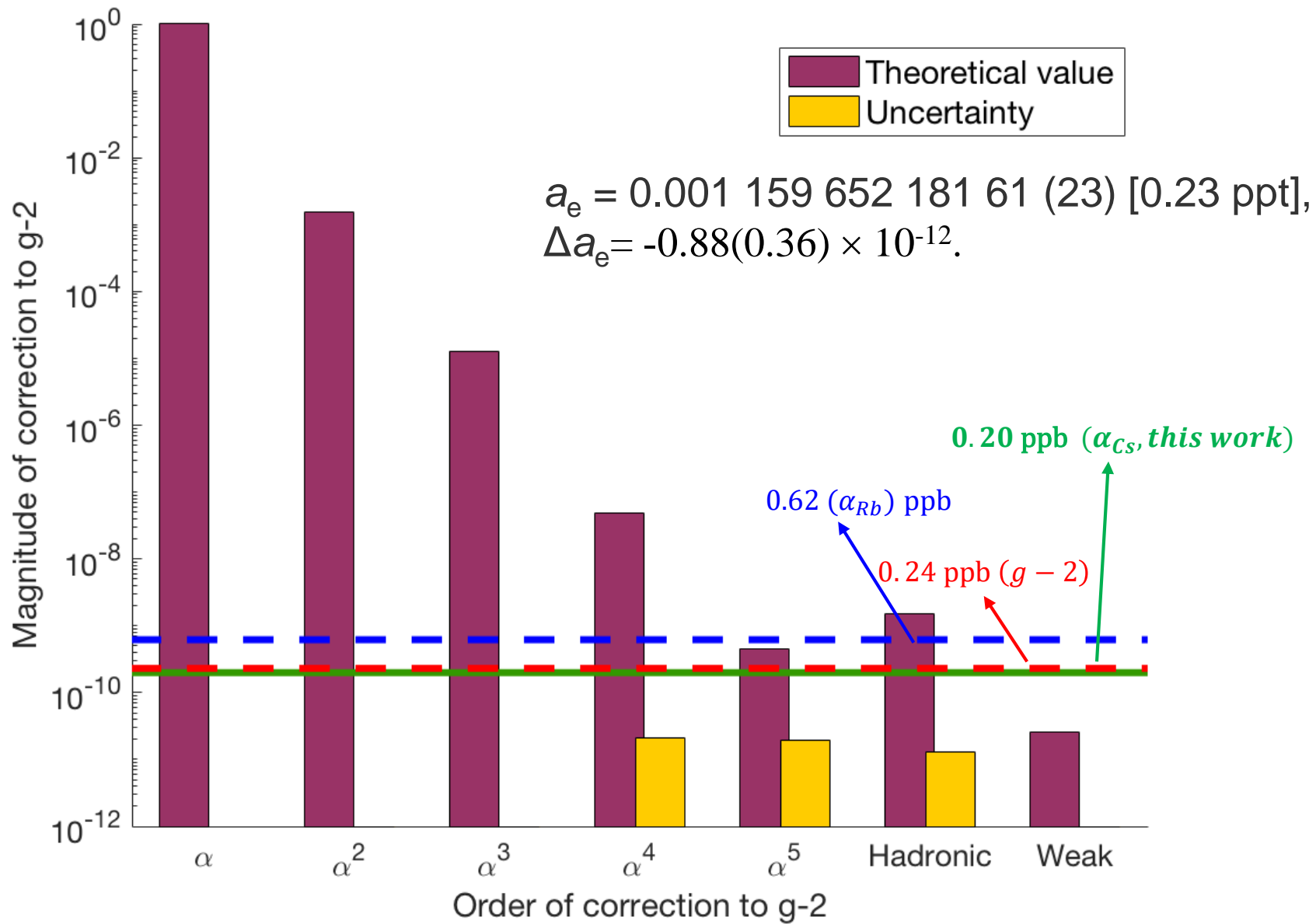
Results



$\alpha = 1/137.035\ 999\ 046\ (27)\ [0.20\ \text{ppb}]$

1.3 million sigma from 1/137

Results



Two $g-2$ anomalies

A Tale of Two Anomalies

Hooman Davoudiasl^{*1} and William J. Marciano^{†1}

¹Department of Physics, Brookhaven National Laboratory, Upton, NY 11973, USA

The most recent determination of the fine structure constant α seems to point to a $\sim 2.4\sigma$ negative deviation in the measured electron anomalous magnetic moment $g_e - 2$. The corresponding experimental value for the muon, $g_\mu - 2$, has long had a $\sim 3.7\sigma$ deviation, in the positive direction. In this short letter, we point out that one real scalar, with a mass of $\sim 250 - 1000$ MeV, could explain the deviations in $g_\mu - 2$ and $g_e - 2$, through one- and two-loop processes, respectively. We briefly discuss potential implications of this simple scenario for low and high energy phenomena.

and muon anomalous magnetic moments see Ref. [19], where the authors discuss the relative contributions of one- and two-loop diagrams, but focus on the case of a pseudoscalar boson. Here, we focus on the effect of a light scalar where the Barr-Zee contribution represents an extension of earlier work in Ref. [10]. Work on the contribution of Barr-Zee type diagrams to $g_\mu - 2$ in the context of two Higgs doublet models and supersymmetry can be found in Ref. [20].

Let us consider the following Lagrangian for the real scalar ϕ of mass m_ϕ

$$\mathcal{L}_\phi = -\frac{1}{2}m_\phi^2\phi^2 - \sum_f \lambda_f \phi \bar{f}f - \frac{\kappa_\gamma}{4} \phi F_{\mu\nu}F^{\mu\nu}, \quad (4)$$

where we only include an explicit coupling to a fermion f with strength λ_f and have omitted various kinetic terms and fermion masses. In this work, we allow f to correspond to known quarks and leptons, as well as other potential more massive charged fermions. The λ_f are constrained by phenomenology, as will be discussed later. We assume that the coupling to photons, through the

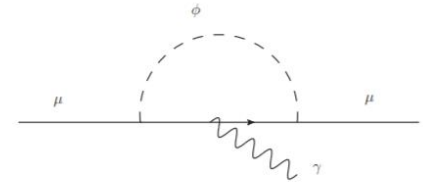


FIG. 1: One-loop ϕ contribution to $g_\mu - 2$.

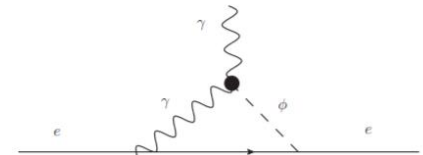
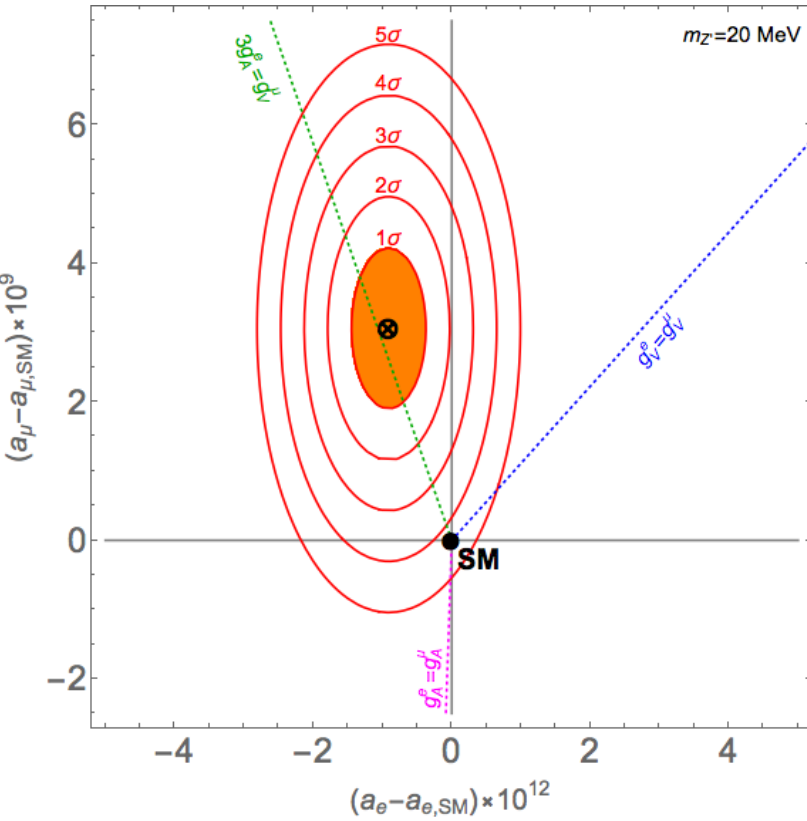


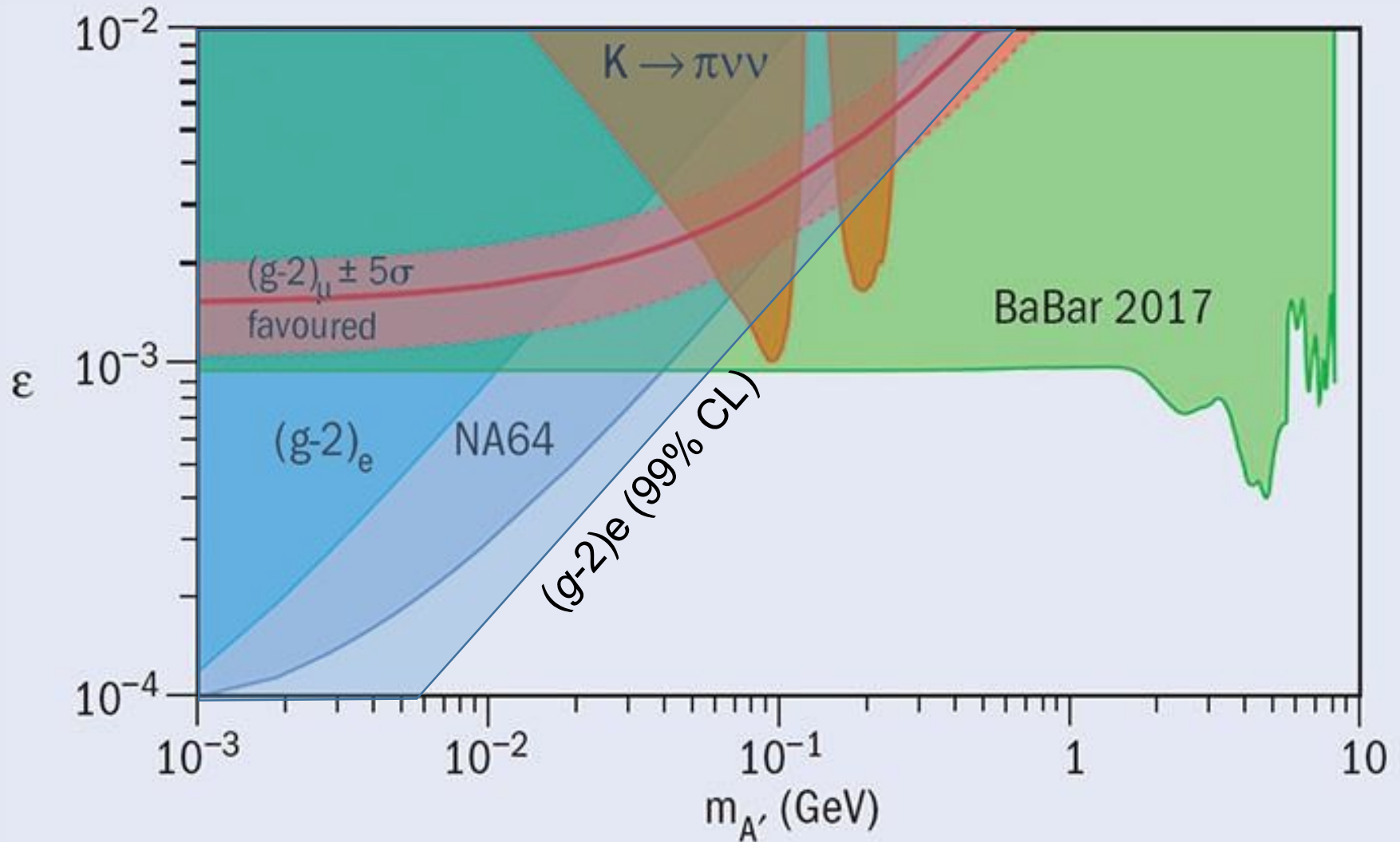
FIG. 2: Effective two-loop Barr-Zee diagram contribution to $g_e - 2$, with heavy fermion loops integrated out.



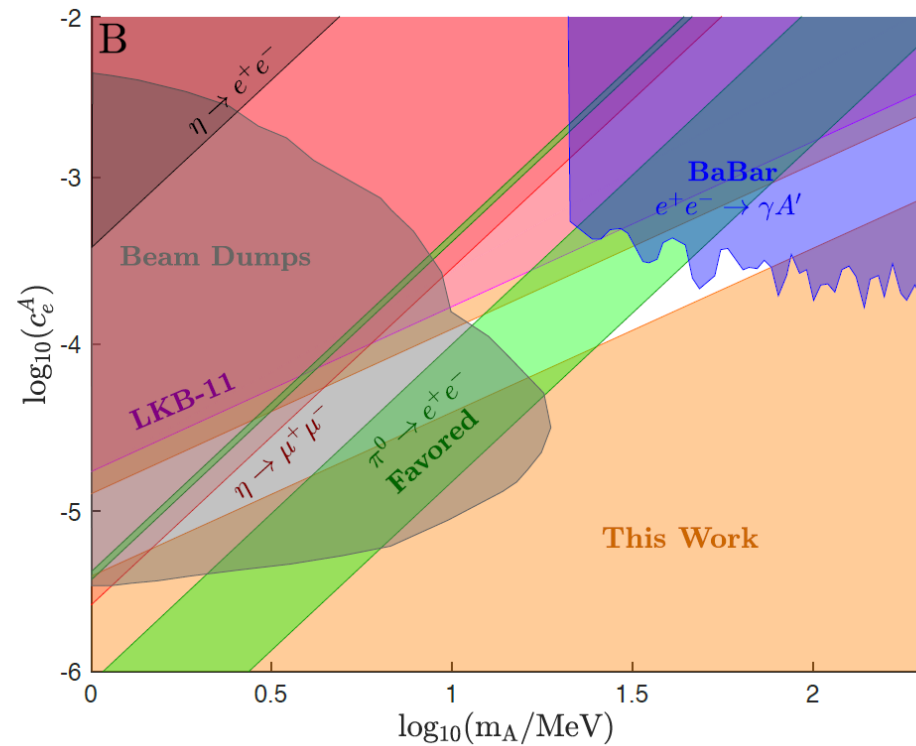
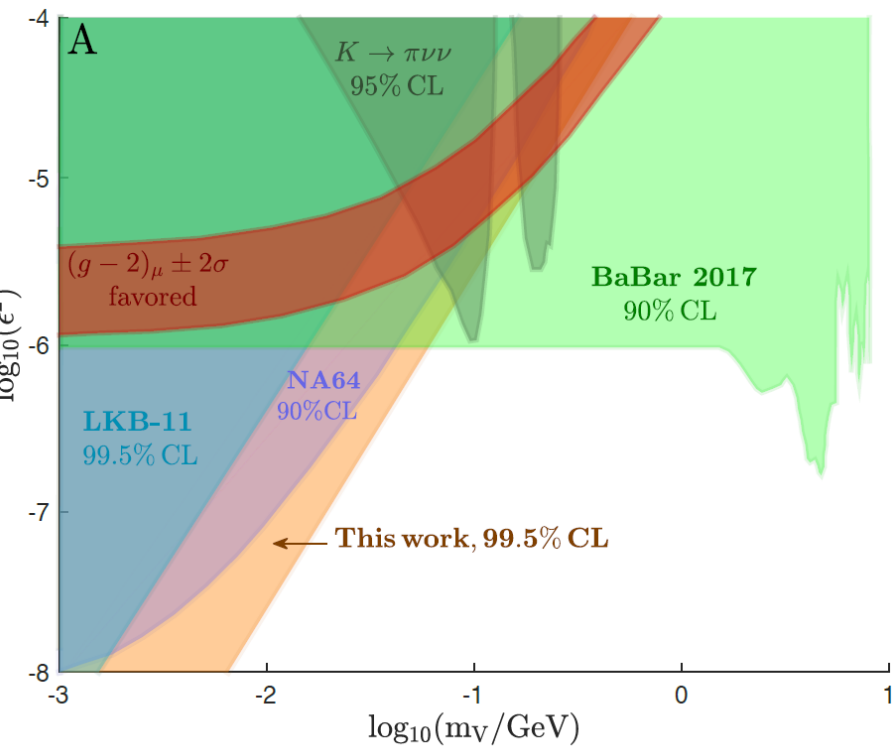
<http://resonances.blogspot.com/2018/06/alpha-and-g-minus-two.html>

Davoudiasl & Marciano,
arXiv:1806.10252

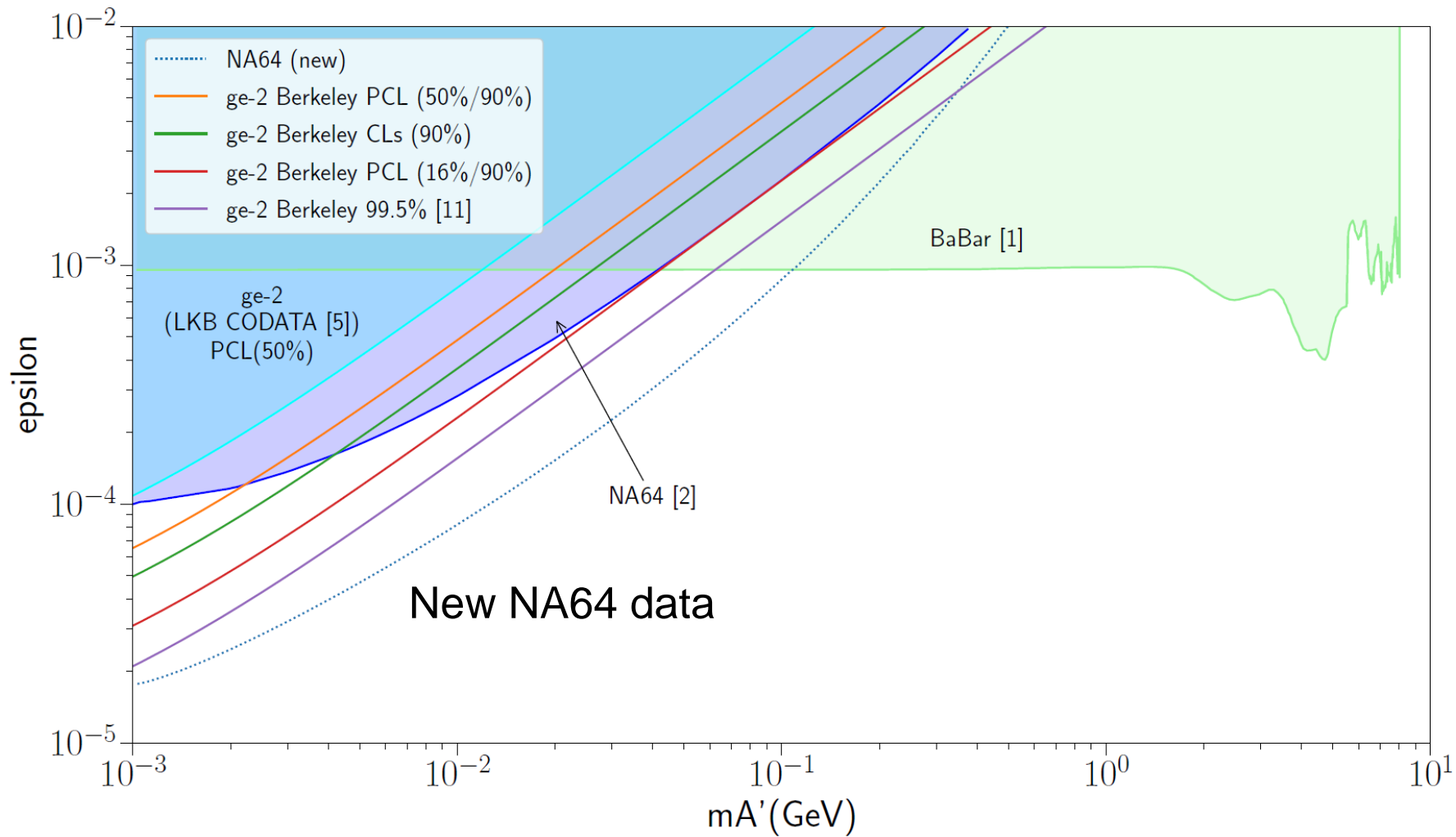
Dark photon limits



Dark photons & axial vectors



Dark photon update



Other interpretations

+3.7 σ positive deviation in the muon, -2.4 σ for the electron

- “One real scalar, $M \sim 250 - 1000$ MeV, could explain [both] deviations through 1- and 2-loop processes, respectively.” [1]
- “Requires that the muon and electron sectors effectively decouple to avoid constraints from $\mu \rightarrow e\gamma$ Improved measurements of a_e , a_μ , as well as α are not only set to provide exciting precision tests of the SM, but, in combination with EDMs, to reveal crucial insights into the flavor structure of physics beyond the SM.” [2]

[1] Davoudiasl and Marciano, [arXiv:1806.10252](https://arxiv.org/abs/1806.10252) (2018)

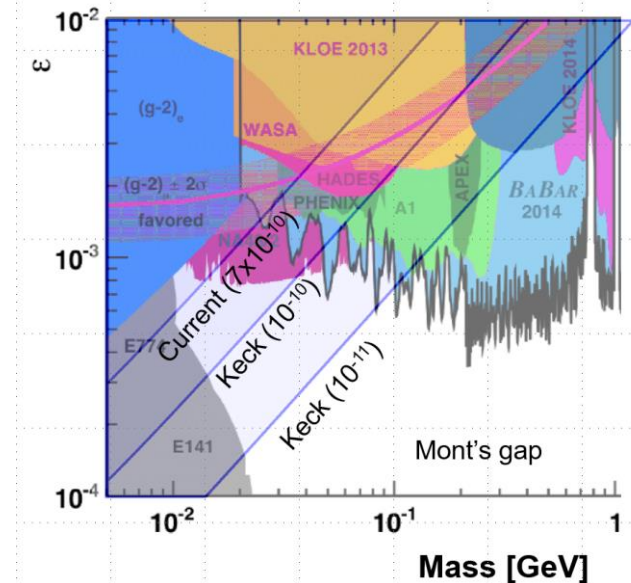
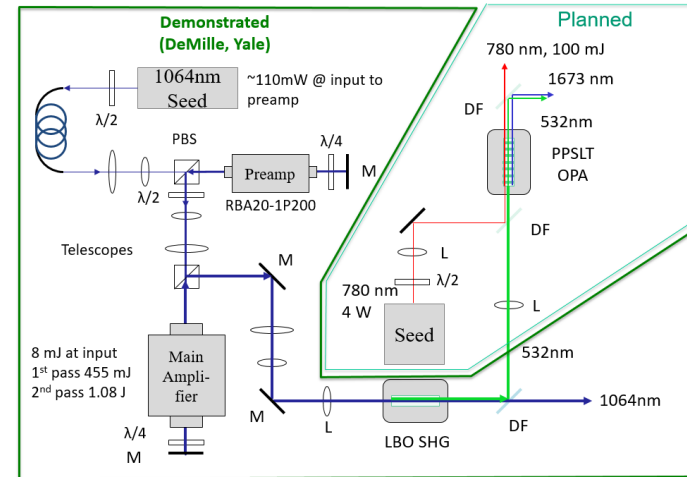
[2] Crivellin, Hoferichter, and Schmidt-Wellenburg, [arXiv:1807.1148](https://arxiv.org/abs/1807.1148) (2018)





Future Upgrades

- Broad beam
 - x20 waist \rightarrow 1/400 beam-related systematics
 - x1000 eff. power
- Acoustic Shielded Room
 - Controls gravity anomalies by keeping heavy objects out
- Science Chamber
 - Dark Matter studies



Thank you!

Fine Structure Constant

Postdoc:

Richard Parker

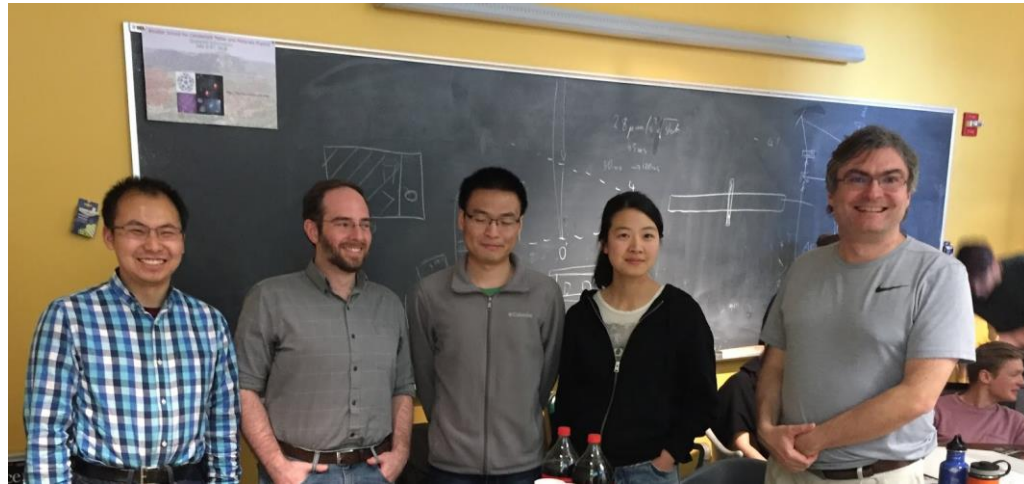
Grads:

Brian Estey,

Chenghui Yu

Weicheng Zhong

Undergrad: Joyce Kwan



Atom interferometry

Postdocs:

Philipp Haslinger,

Xuejian Wu

Grads:

Kayleigh Cassella,

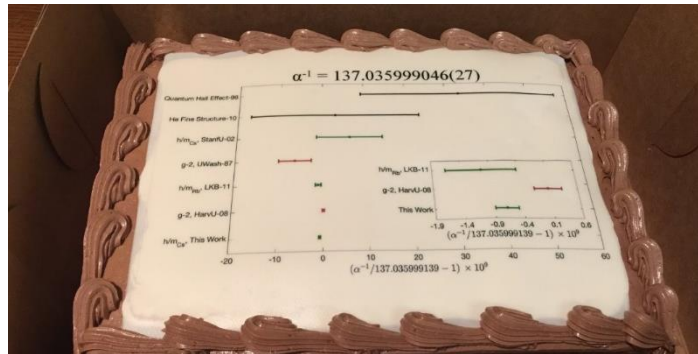
Eric Copenhaver,

Jordan Dudley,

Matt Jaffe,

Zachary Pagel

Victoria Xu



Phase-Contrast TEM

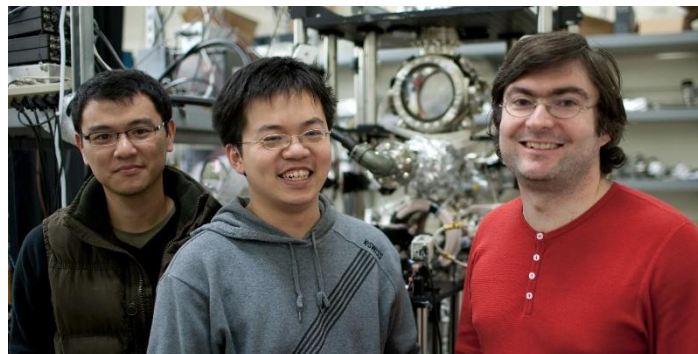
Postdocs:

Sara Campbell

Osip Schwartz

Grad:

Jeremy Axelrod,



Faculty Alumni

Paul Hamilton (UCLA)

Mike Hohensee (LLNL)

Geena Kim (Regis)

Pei-Chen Kuan (NCKU)

Shau-Yu Lan (NTU)