

Neutrino oscillation

NOvA + T2K

Peter Madigan

Outline

- Neutrino oscillation basics
- Design principles
- T2K and NOvA

Neutrino oscillation basics

Neutrinos interact weakly and have mass

Mass eigenstates are not the same as flavor eigenstates

Neutrinos are produced in weak eigenstates, but evolve as mass eigenstates

Non-zero probability of observing other weak eigenstates at future times (distances)

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

$$|\nu_i(t)\rangle \sim e^{-ip \cdot x} |\nu_i(0)\rangle$$

$$|\nu_i(L)\rangle \sim e^{-im_i^2 L/2E} |\nu_i(0)\rangle$$

Neutrino oscillation basics

PMNS mixing matrix:

- 3x3 unitary matrix \rightarrow 9 dof
- Split into 3 Euler angles, and 6 complex phases
- At least 3 of these phases can be absorbed into the definitions of the neutrino states and are not physical
- If neutrinos are Dirac particles \rightarrow 2 more phases can be absorbed
- In total, it is commonly expressed as...

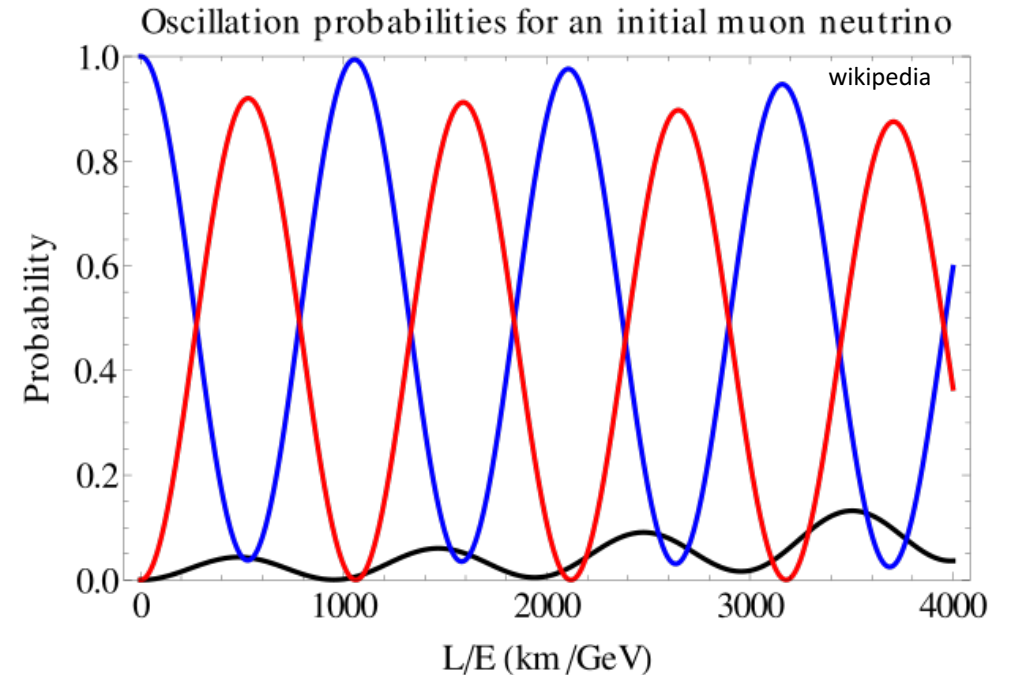
$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{bmatrix} \begin{bmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{-i\delta} & 0 & \cos \theta_{13} \end{bmatrix} \begin{bmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Neutrino oscillation basics

Oscillation probability can be found by taking the square of the inner product

$$\langle \nu_\beta(L) | \nu_\alpha(0) \rangle$$

- Probability depends on L/E , mixing angles, mass splitting, and complex phases



$$P_{\alpha \rightarrow \beta}(L) = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2(\delta m_{ij}^2 L / 4E) + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin(\delta m_{ij}^2 L / 2E)$$

Requirements

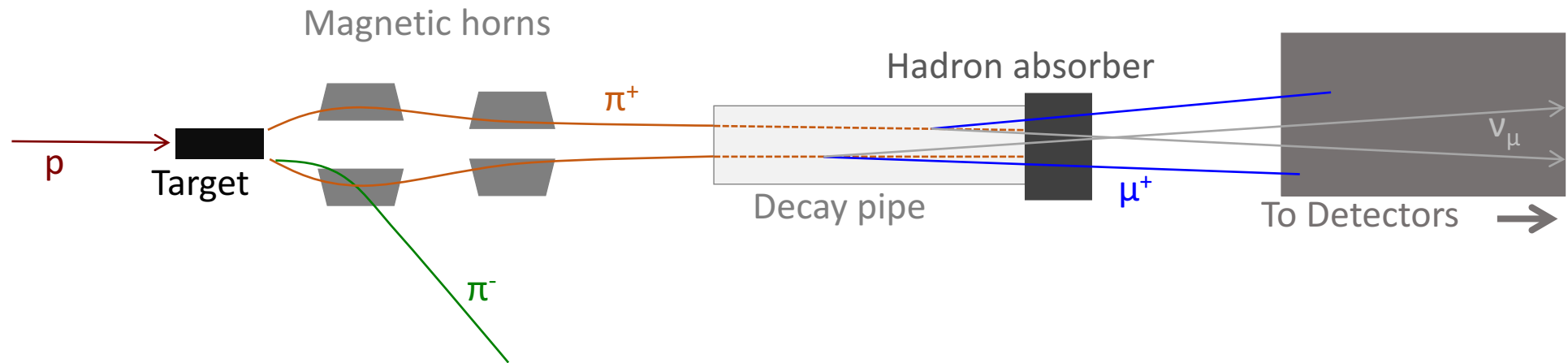
What goes into making a good oscillation measurement?

- Distinguish between flavors
- Large target to improve statistics
- Well understood energy spectrum (L/E dependence)
- Low backgrounds

Focus on accelerator-produced muon neutrinos (T2K + NOvA)

General design: Beam

Beam production

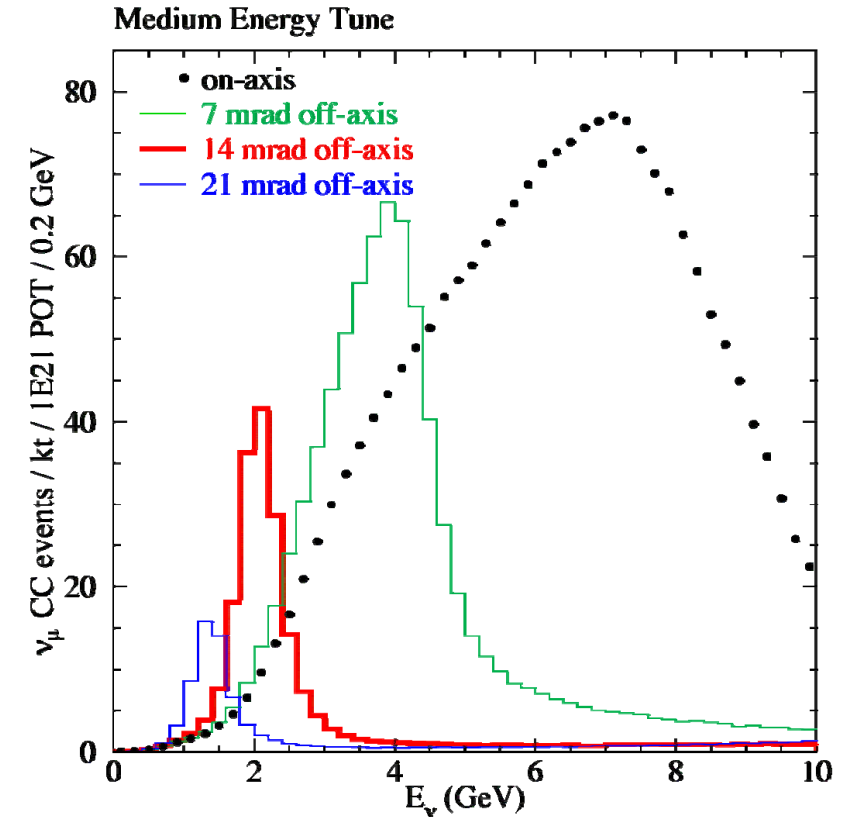


General design: Beam

Off-axis beam

2-body decay in flight kinematics produce a strong relationship between parent energy and neutrino energy

- produces an enhanced neutrino energy peak at off-axis angles
- peak energy depends on off-axis angle



$$E_\nu = \frac{0.43 E_\pi}{1 + \gamma^2 \theta^2}$$

General design: Detector technology

Specific requirements for neutrinos (compared to Atlas / etc.)

- Target / detector combo
- High mass = scalability, must be cheap!

~kton scale is required

muon / electron id

cheap
...but good! ✓

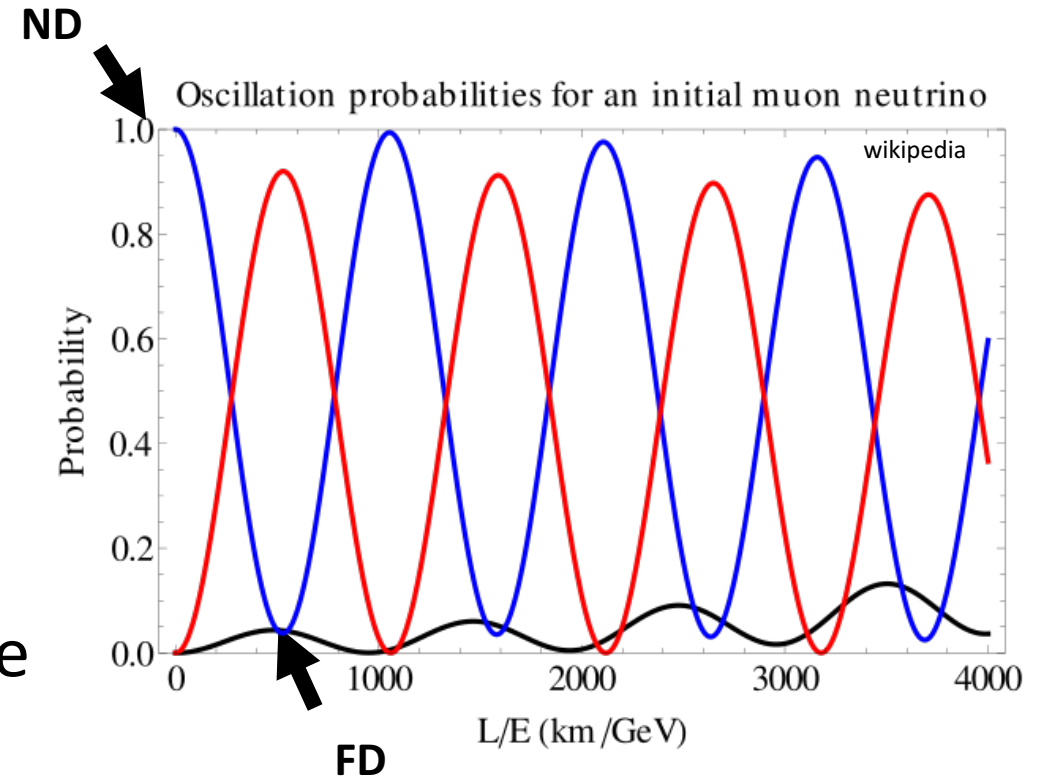
General design: Near / far detector

Near / far detector goals:

- Measure flavor content of beam
- Measure energy spectrum of beam
- Constrain backgrounds

First oscillation maximum

- Build near / far detector to maximize expected signal



The two experiments

T2K

NOvA

recent neutrino experiments with goals of

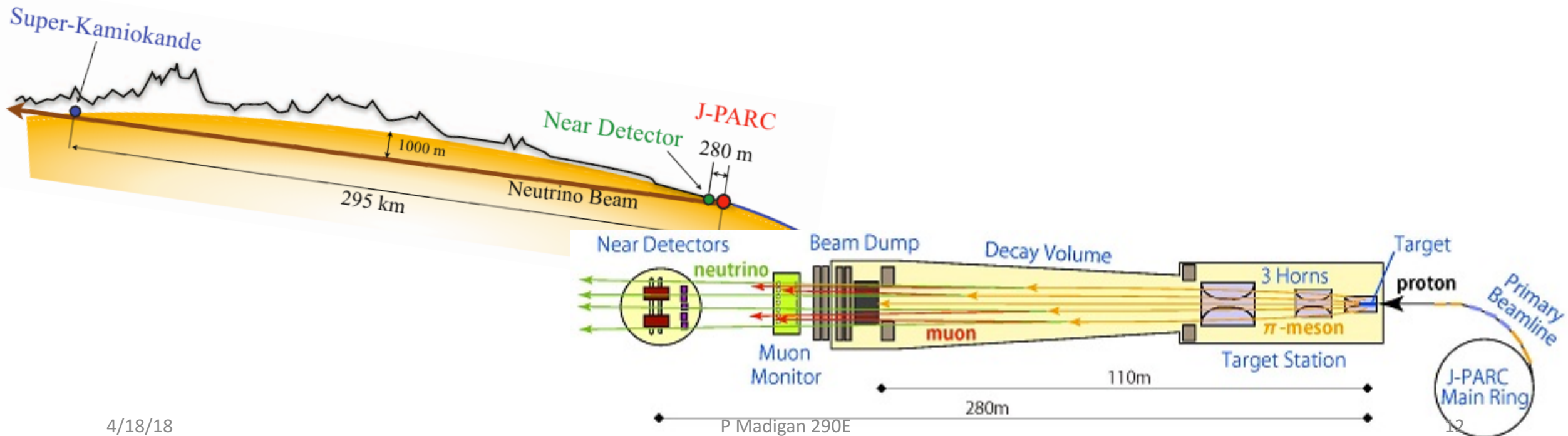
- measuring θ_{23} octant (precision of ~ 0.01)
- measurement of Δm_{23}^2 (precision of $\sim 10^{-4} \text{eV}^2$)
- search for sterile oscillations

T2K beam

J-PARC beam in Tokai, Japan to SuperK in Kamioka

50GeV proton beam, off-axis by 43mrad = peak neutrino energy of 0.6GeV

Beamline of 295km



T2K detectors

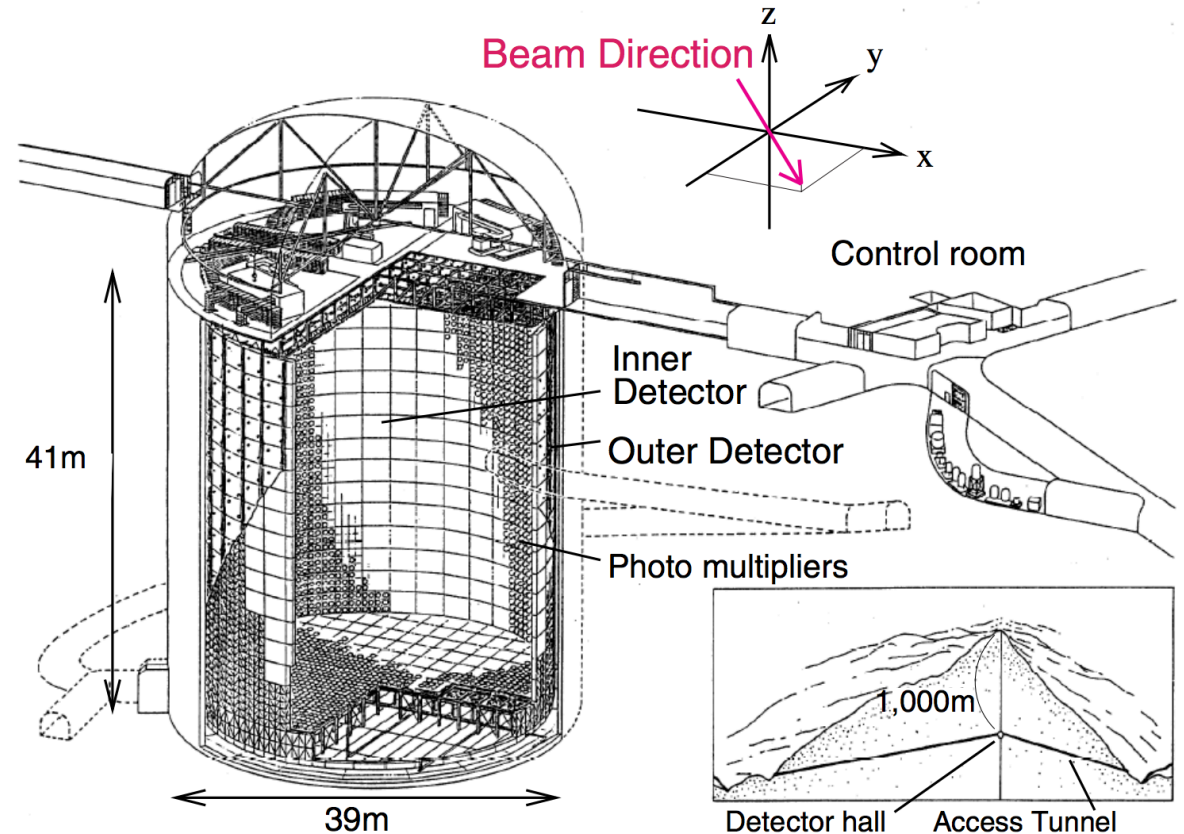
Super Kamiokande – 1km underground

22.5kt water cherenkov detector

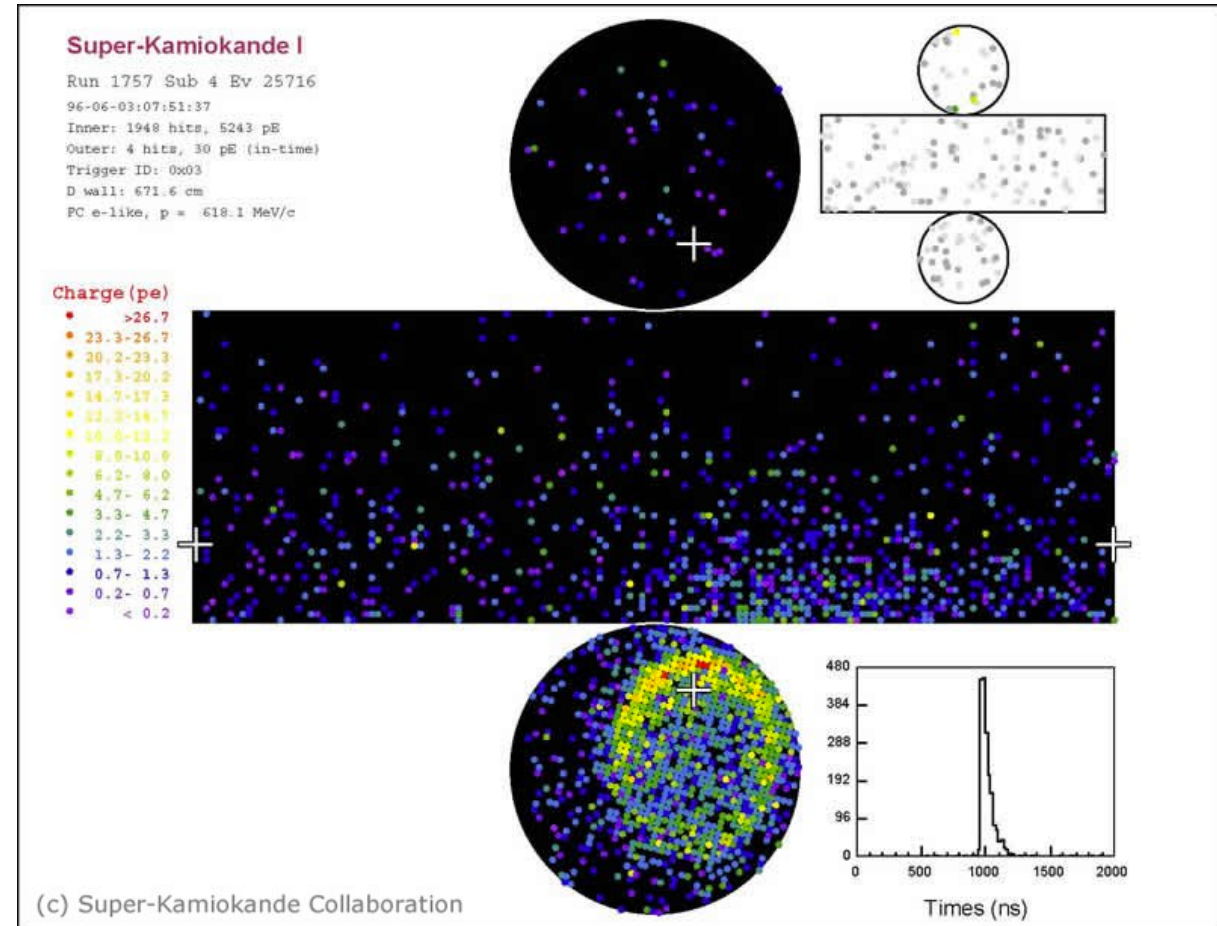
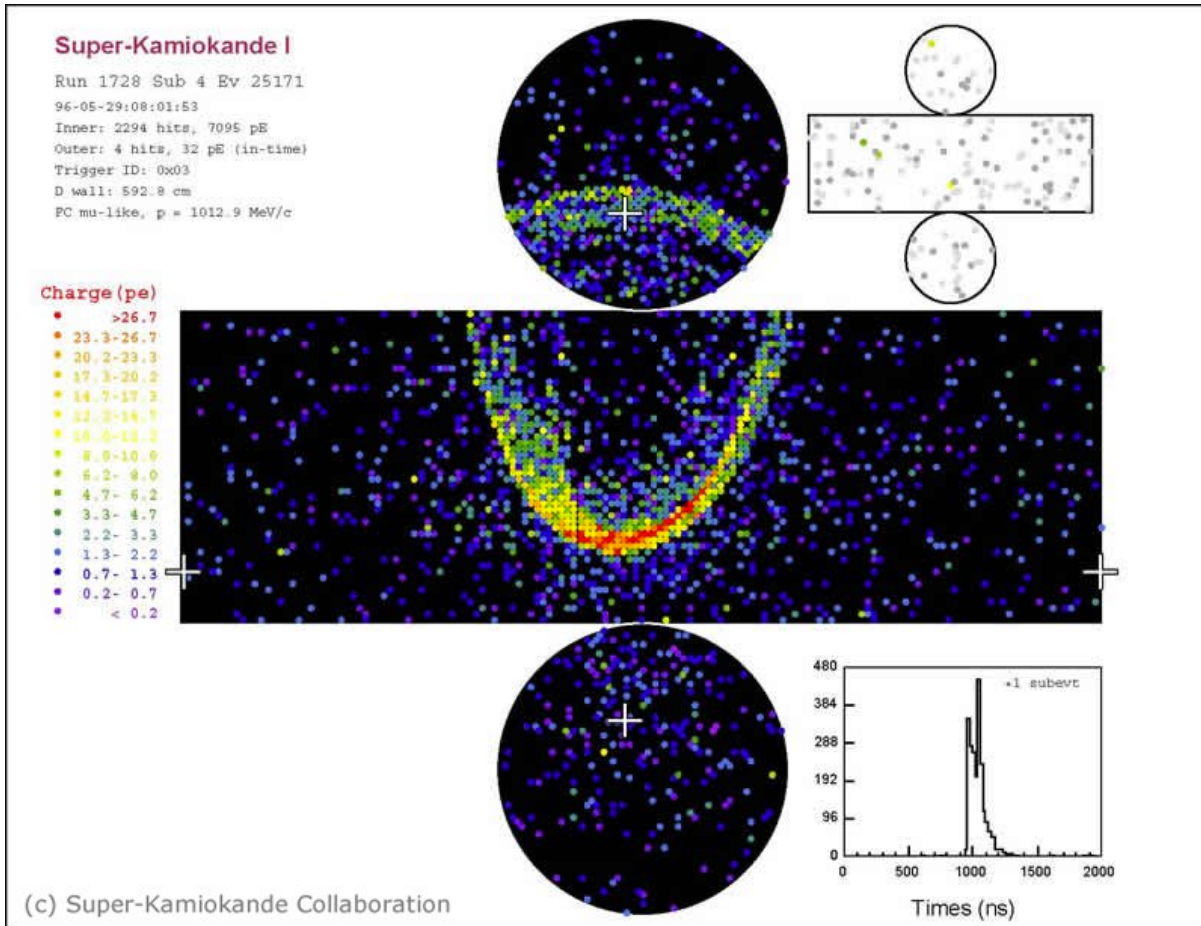
ID 10k PMTs

OD 1800 PMTs

Reconstruction via ring imaging, intensity, and timing



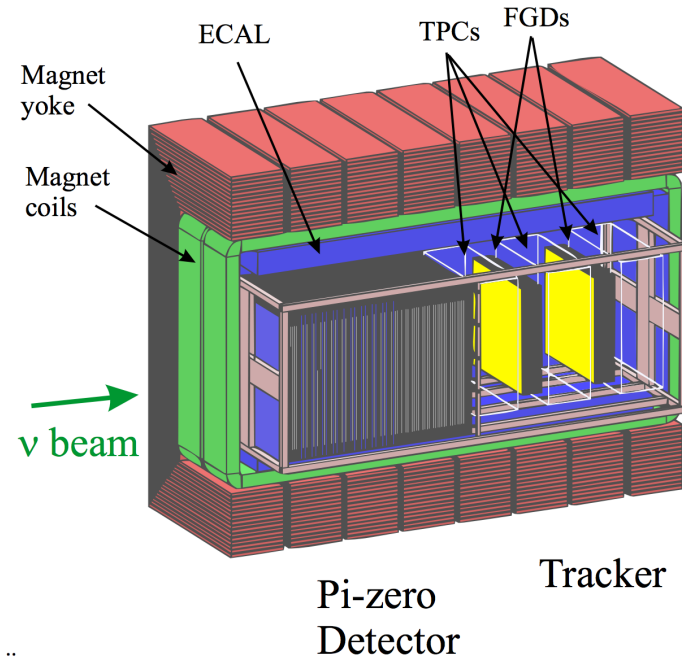
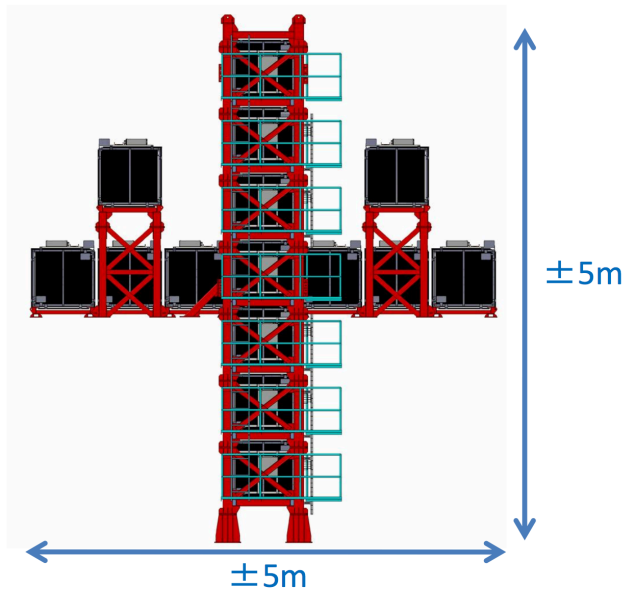
T2K detectors



T2K detectors

Near detector system consists of 2 main detectors

- INGRID (on axis)
- ND280 (off axis)



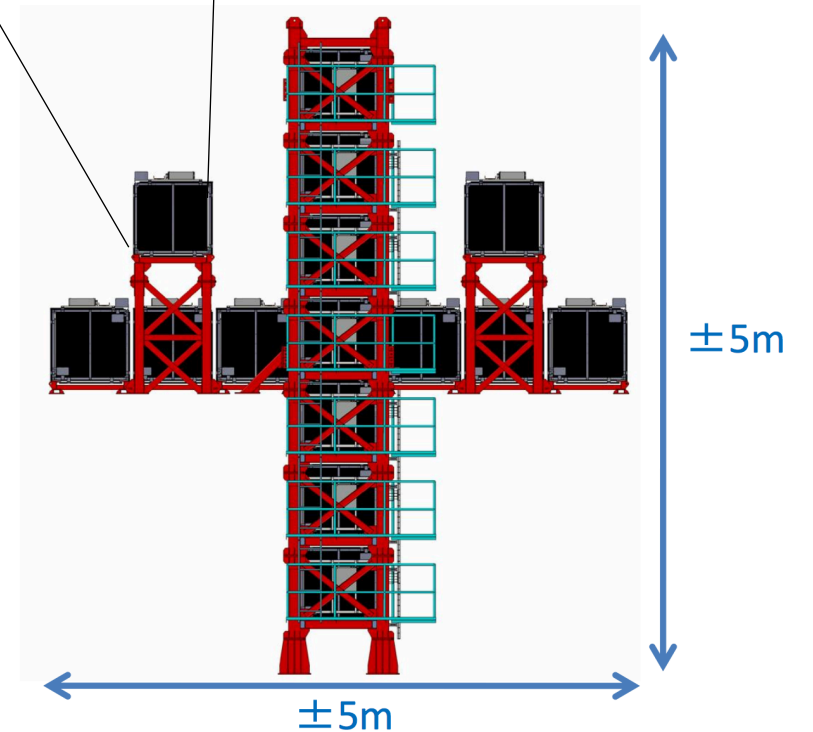
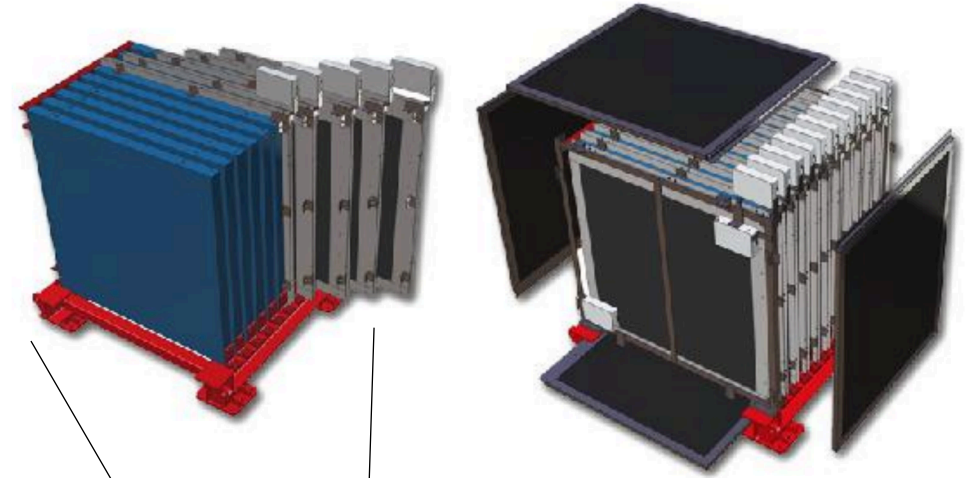
T2K detectors

INGRID

280m downstream of absorber

Directly measures neutrino beam profile and direction

Provides a better monitor of the neutrino beam direction than muon monitors

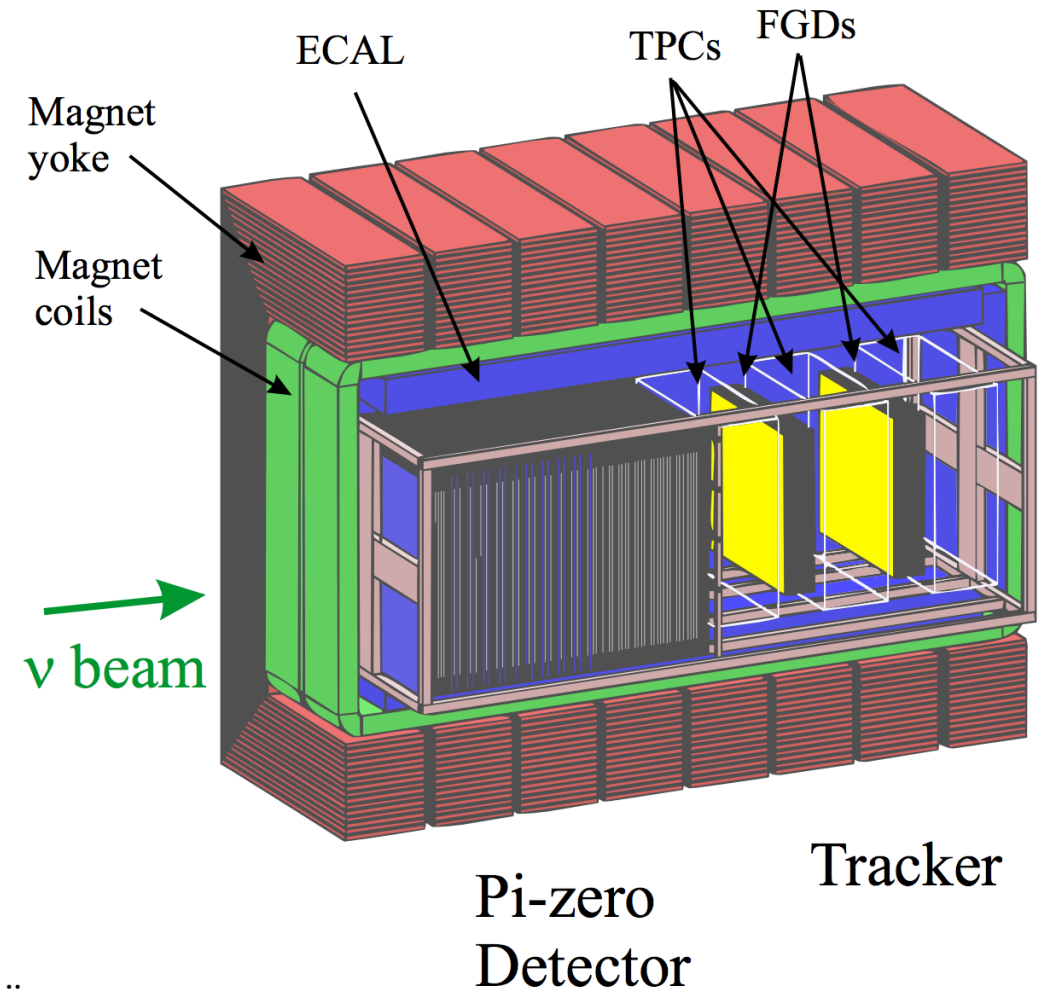


T2K detectors

ND280 – 280m off axis by 43mrad

5 submodules:

- POD: specialized π^0 detector
- TPCs: muon momentum and PID
- 1t FGDs: neutrino target mass
- Ecals: π^0 decays
- SMRD: energy of escaping muons and cosmic ray trigger

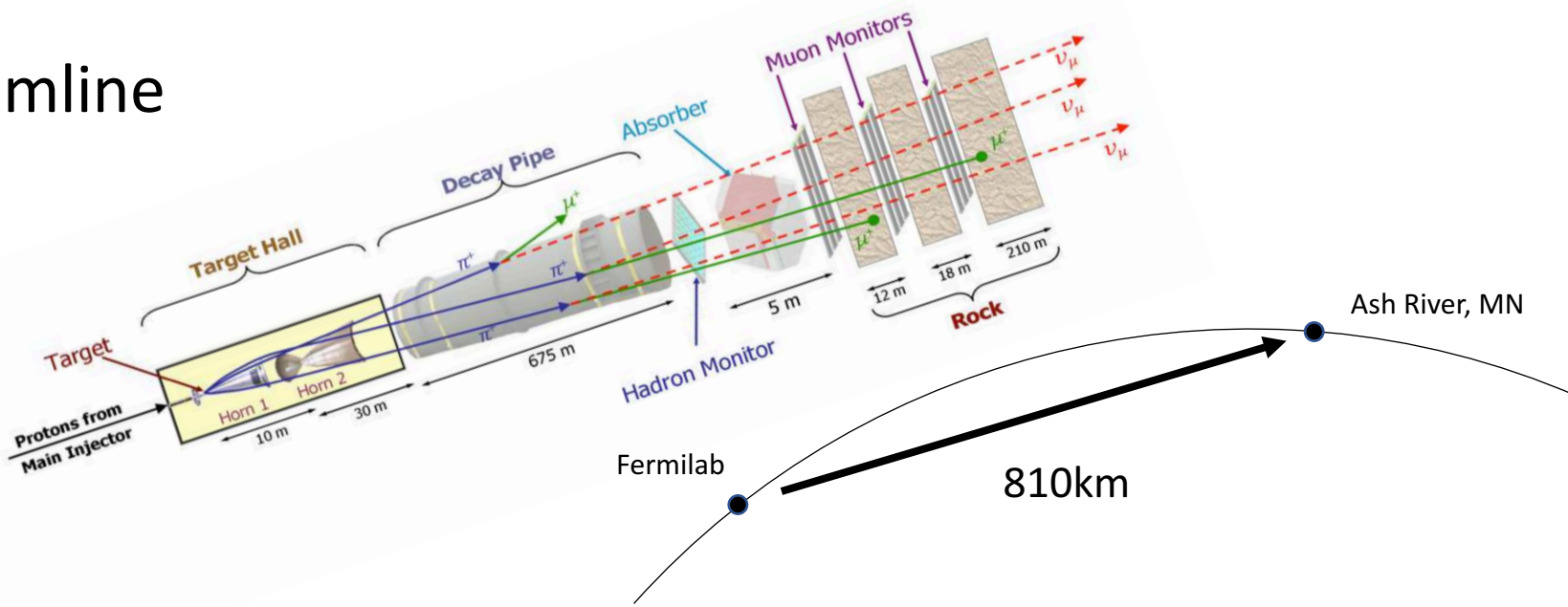
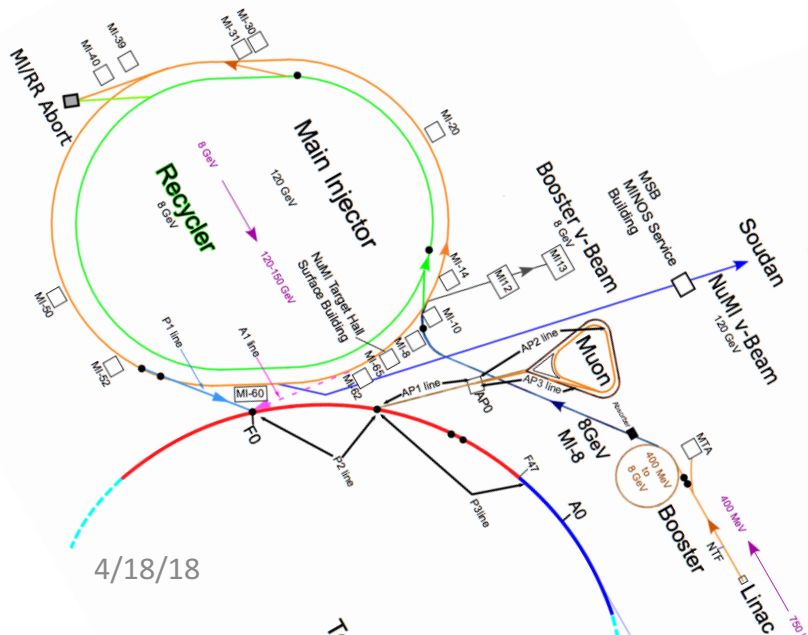


NOvA beam

Fermilab NuMI beam

120GeV proton beam of 120GeV, off-axis by 14.6mrad = neutrino peak at 2GeV

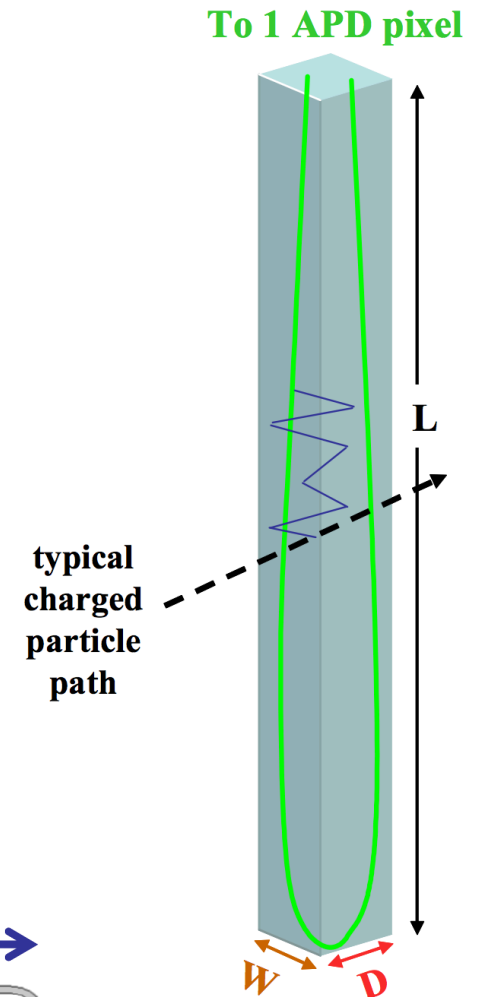
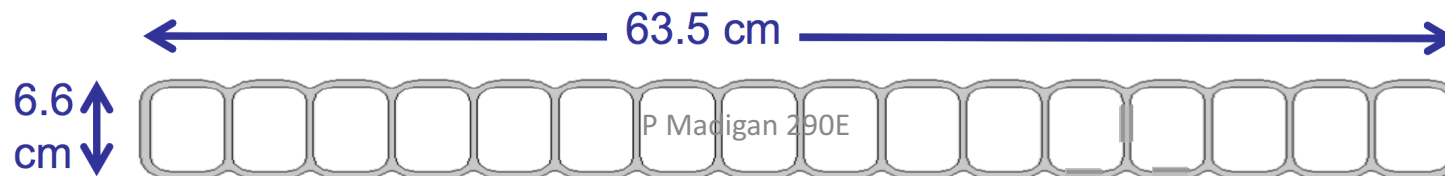
Beamline of 810km beamline



NOvA detectors

NOvA basic building block

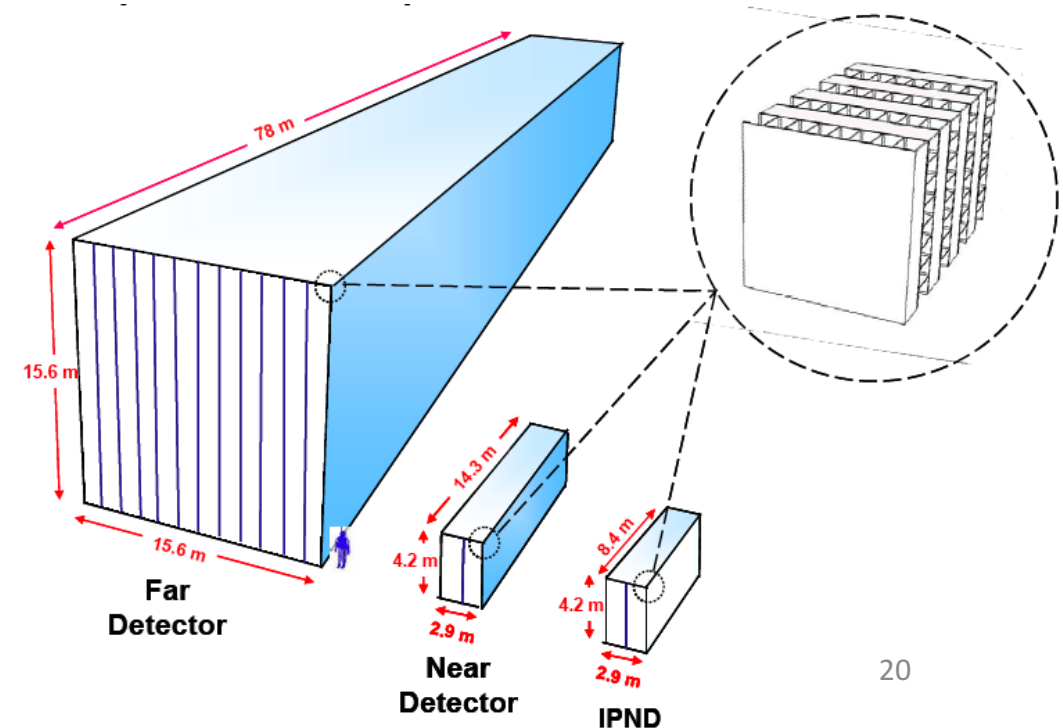
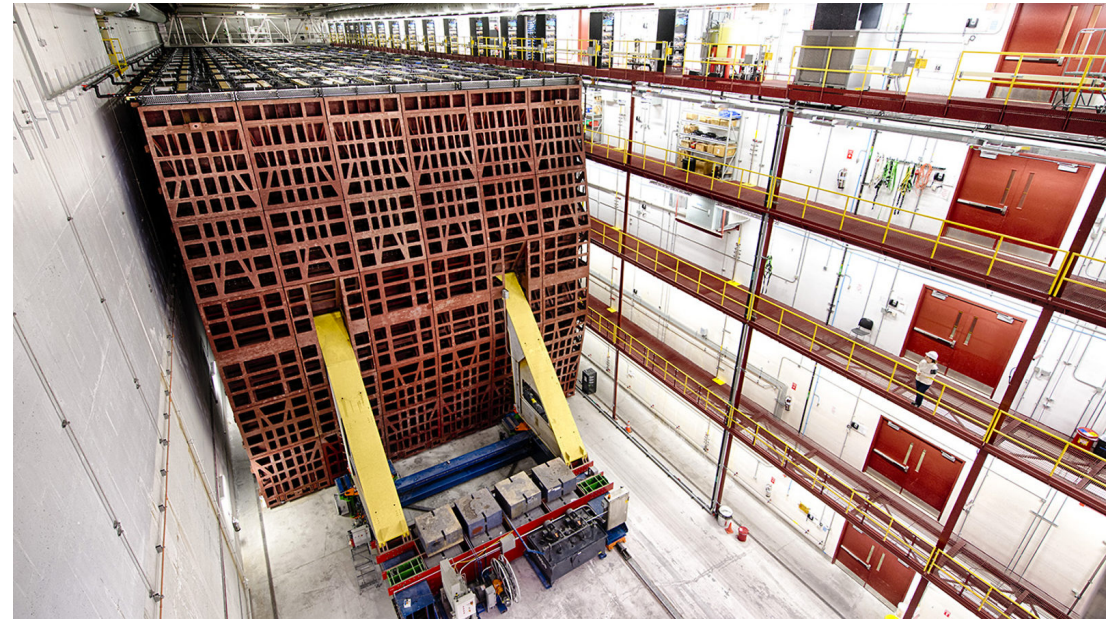
- Liquid scintillator bars with wavelength shifting (WLS) fibers
- Extruded plastic cells 4cm x 6cm x 15.5m sized for 54-foot trucks
- Scintillator produces in 360-390nm range
 - Additives downshift to 400-450nm
 - WLS downshifts to 490-550nm for higher PMT efficiency
- 13000km of fiber, layered to provide total internal reflection of light



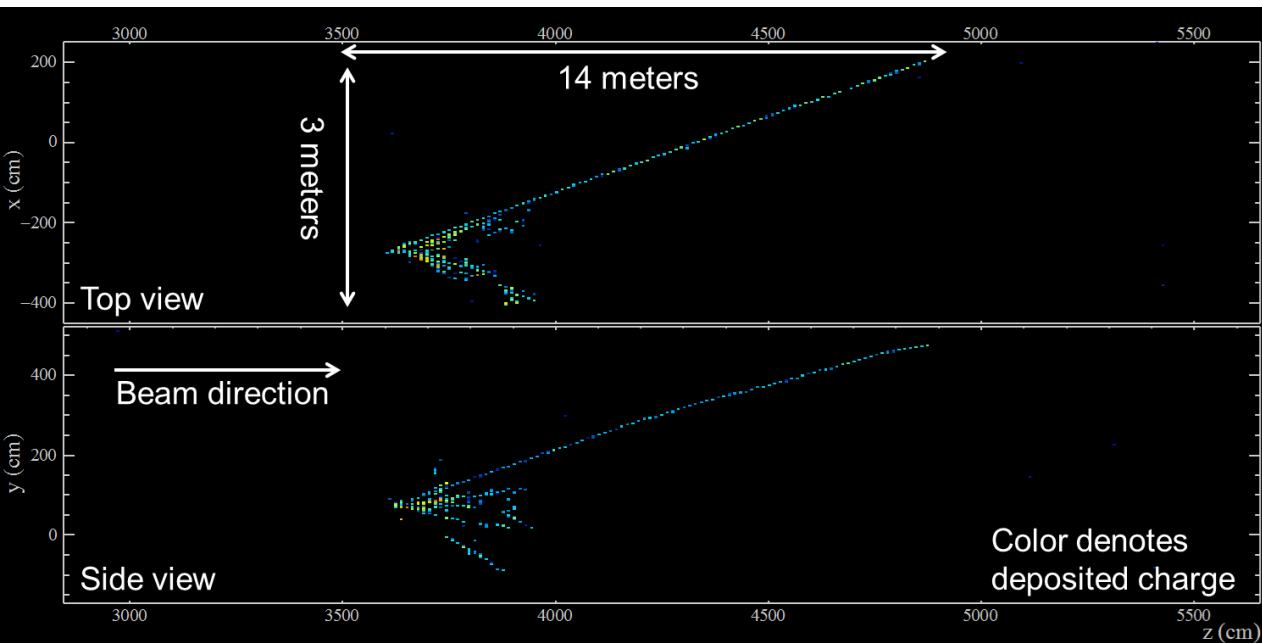
NOvA detectors

Far Detector

- Surface detector, 810km beamline, 15kt, 10kt active
- 385,000 scintillator cells
- flavor id via track topology (muon-like or electron-like)



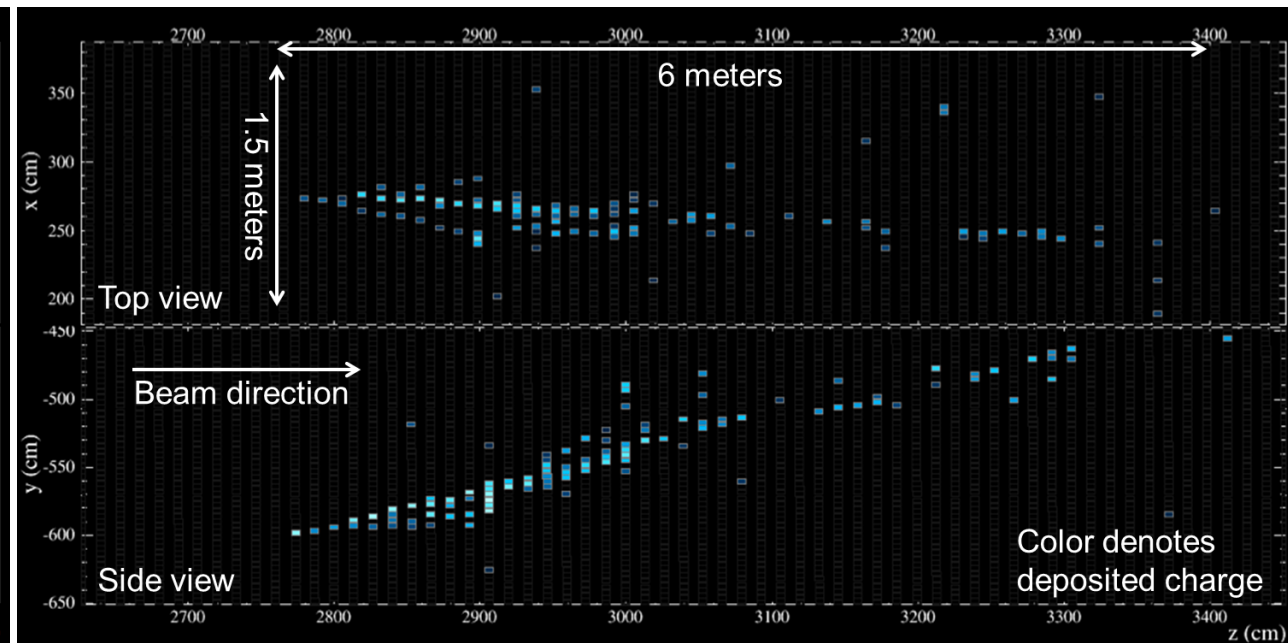
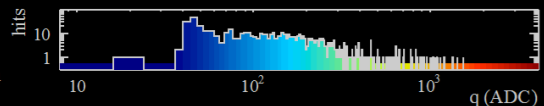
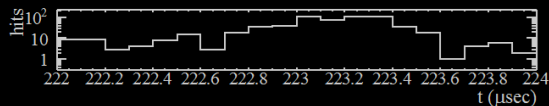
NOvA detectors



NOvA - FNAL E929

Run: 18620 / 13
Event: 178402 / -

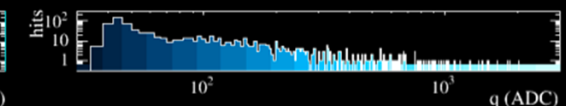
UTC Fri Jan 9, 2015
00:13:53.087341608



NOvA - FNAL E929

Run: 15392 / 55
Event: 125664 / NuMI

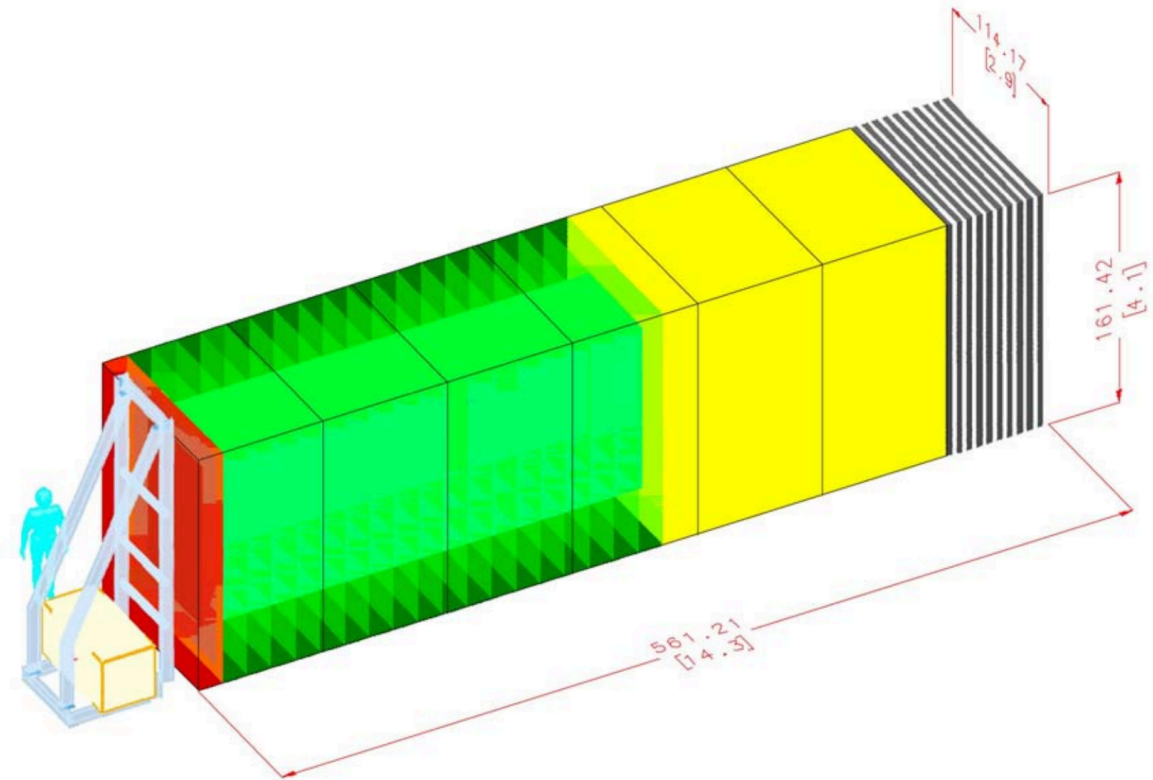
UTC Wed May 28, 2014
04:55:46.939251776



NOvA detectors

Near Detector

- Similar design as FD – scintillator cells
- Consists of an upstream veto, fiducial + containment, and a muon catcher
- Muon catcher is steel interspersed with active scintillator



Analysis

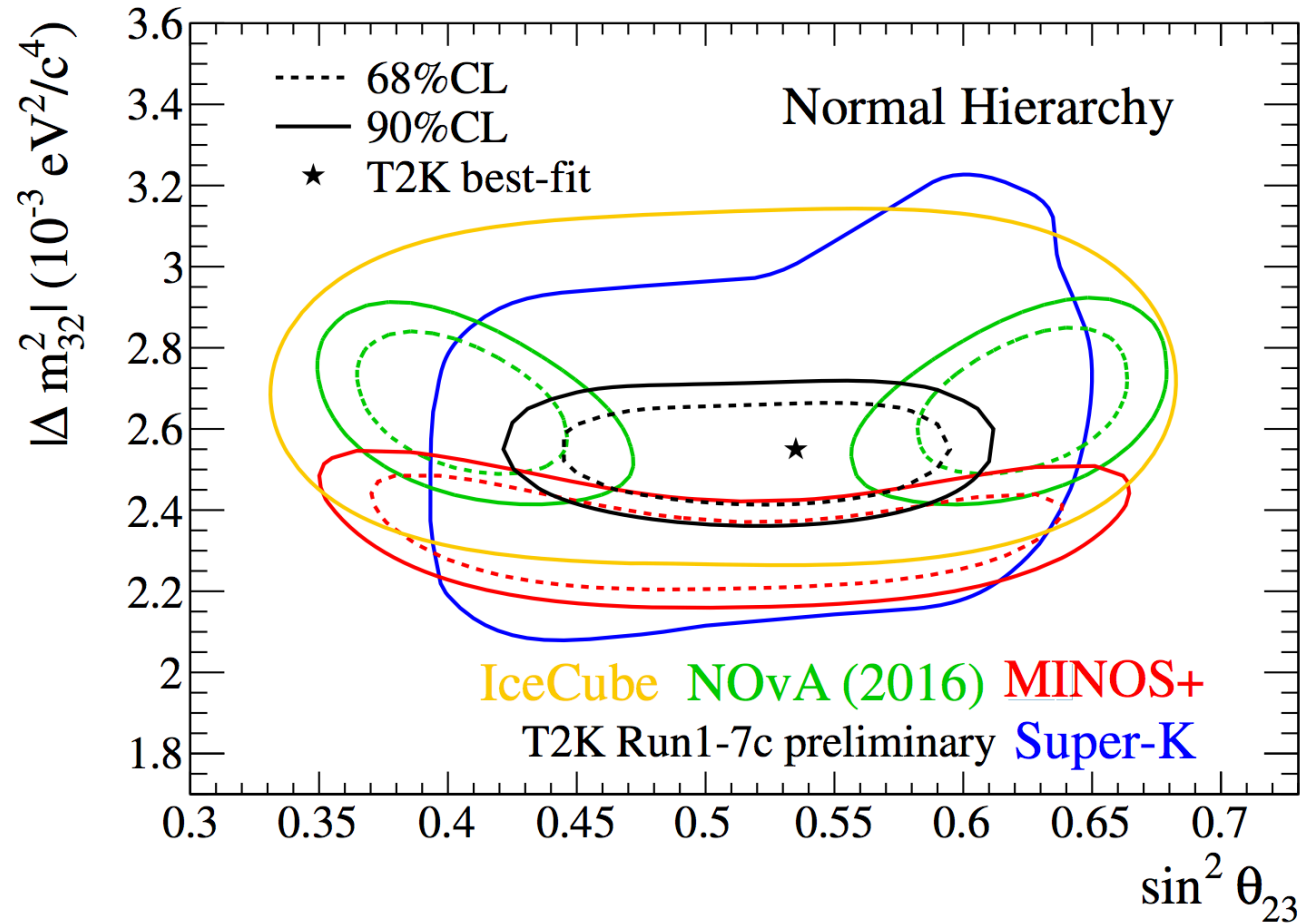
Generally, analysis is performed by comparing the expected far detector flux (as predicted by near detector) to the observed far detector flux

Use extensive detector and beam modelling to determine the translation from near detector to far detector

- NOvA: FLUKA + Geant4 + GENIE
- T2K: Constrained FLUKA + Geant3 based models (NA61/SHINE)

Results

Tension!



Summary

General qualities for oscillation experiments

- Large-mass active-target detectors
- Well-understood neutrino spectra

T2K and NOvA

- Two oscillation experiments observing muon neutrino disappearance and electron neutrino appearance
- Tension between the two experiments

References

- T2K TDR + NIM papers on submodules
- NOvA TDR
- [arXiv:hep-ex/1409.7469](https://arxiv.org/abs/hep-ex/1409.7469)
- [arXiv:hep-ex/1701.00432](https://arxiv.org/abs/hep-ex/1701.00432)

Neutrino oscillation

circa 2012

Parameter	best-fit ($\pm 1\sigma$)	3σ
Δm_{\odot}^2 [10^{-5} eV ²]	$7.58^{+0.22}_{-0.26}$	6.99 – 8.18
$ \Delta m_A^2 $ [10^{-3} eV ²]	$2.35^{+0.12}_{-0.09}$	2.06 – 2.67
$\sin^2 \theta_{12}$	0.306 (0.312) $^{+0.018}_{-0.015}$	0.259 (0.265) – 0.359 (0.364)
$\sin^2 \theta_{23}$	$0.42^{+0.08}_{-0.03}$	0.34 – 0.64
$\sin^2 \theta_{13}$ [140]	0.021 (0.025) $^{+0.007}_{-0.008}$	0.001 (0.005) – 0.044 (0.050)
$\sin^2 \theta_{13}$ [142]	0.0251 ± 0.0034	0.015 – 0.036