Impacts and Imprints of Axion Dynamics

UC Berkeley Unraveling the Particle World and the Cosmos at Berkeley Workshop in Honor of Lawrence Hall and Hitoshi Murayama

September 28th, 2024

Raymond Co

Indiana University

For the record, Co is my last name and I do not represent a company.

Raymond Co

For the record, Co is my last name and I do not represent a company.

I still remember being so excited to see my full name appear in a paper when I was a grad student at Berkeley. However, I quickly said what Lawrence likes to say a lot: "wait a minute!" Why is there a period after Co?

Raymond Co

The paper was written by Stuart.

2016 Dec ∞ [hep-ph]

Reheating and Leptogenesis after Pati-Salam F-term Subcritical Hybrid Inflation

B. Charles Bryant, Zijie Poh, and Stuart Raby

Department of Physics The Ohio State University 191 W. Woodruff Ave, Columbus, OH 43210, USA

June 15, 2021

Raymond Co

The paper was written by Stuart.

It's too bad that Staurt had to leave this workshop early. When I took this opportunity and sent him an email to introduce myself, he kindly invited me to give a seminar at OSU. However, to this date, I still have not visited OSU.

Reheating and Leptogenesis after Pati-Salam F-term Subcritical Hybrid Inflation

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June 15, 2021

Raymond Co

 10 A recent paper by Raymond Co. et. al. [38], points out that for models of gravitino or axino dark matter with high reheat temperature, such as our model, the decay of the saxions can produce 3 orders of magnitude of entropy in the case that the saxion vev, s_I , equals the PQ breaking vev, V_{PQ} and $s_I = V_{PQ} \approx 10^{14}$ GeV (see Eqn. 2.9, Ref. [38]). This would make the plateau of our baryon-to-entropy ratio of order the observed experimental value.

2016

Dec

 ∞

 $\overline{}$

[hep-ph]

That email also reminded of what Yasunori said to me at the time.

"You should be happy because Lawrence suddenly becomes the employee of your company now!"

2016 Dec ∞ $[hep-ph]$

Reheating and Leptogenesis after Pati-Salam F-term Subcritical Hybrid Inflation

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June 15, 2021

Raymond Co

Myths about Hitoshi I have heard during grad school

7

1. Hitoshi's average speed is 40 mph throughout the year.

Myths about Hitoshi I have heard during grad school

8

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Myths about Hitoshi I have heard during grad school

9

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$D = 40$ mph \times 365 days \times 16 hours/day \simeq 233k miles USA-JPN trip distance \simeq 5k miles

assuming domestic flight is 2k miles each way

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Data collected: I was in Hitoshi's QFT class, we had to reschedule classes every two weeks due to his trips to Japan

10

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11

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2. Hitoshi's airline miles exceed the distance from the Earth to the Moon.

Myths about Hitoshi I have heard during grad school

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 $D = 40$ mph \times 365 days \times 16 hours/day \simeq 233k miles

 $D_{\text{earth,moon}} \simeq 239k$ miles

Myths about Hitoshi I have heard during grad school

2. Hitoshi's airline miles exceed the distance from the Earth to the Moon.

233k miles

SaxiGUTs and their Predictions

Raymond T. $Co^{1,2}$, Francesco D'Eramo^{3,4} and Lawrence J. Hall^{1,2}

¹ Berkeley Center for Theoretical Physics, Department of Physics, University of California, Berkeley, CA 94720, USA ² Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA ³Department of Physics, University of California Santa Cruz, Santa Cruz, CA 95064, USA ⁴Santa Cruz Institute for Particle Physics, Santa Cruz, CA 95064, USA

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On 2/29/2016, I gave a 4D seminar. Hitoshi soon raised a serious concern that would basically kill our entire project. After the talk, he explained the issue in detail, but then

 \circ **Oct 201** $\boxed{1}$ ph

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On 2/29/2016, I gave a 4D seminar. Hitoshi soon raised a serious concern that would basically kill our entire project. After the talk, he explained the issue in detail, but then he immediately offered a solution..

 \circ **Oct 201** $\boxed{1}$ ph

SaxiGUTs and their Predictions

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In principle the PQ symmetry could be an R symmetry, as in the case of both the Nilles-Raby $SU(5)$ theory [6] and the Hall-Raby $SO(10)$ theory [21]. However, it has been shown quite generally that in flat space supersymmetric theories with a continuous R symmetry broken at scale V_{PQ} the vacuum value of the superpotential is bounded by $|\langle W \rangle| \leq F V_{PQ}/2$ where F is the scale of supersymmetry breaking [22]. Such values of $|\langle W \rangle|$ are insufficient to cancel the cosmological constant in supergravity unless V_{PQ} is of order the reduced Planck mass. Hence we restrict our attention to non- R symmetries.

Acknowledgments

We thank Michael Dine, Keisuke Harigaya and Hitoshi Murayama for useful discussions. This work was supported in part by the Director, Office of Science, Office of High Energy and Nuclear

Whenever I had some results after a long calculation in Mathematica to show Lawrence. He always pulled out his notes and checked my results using this hand-drawn figures. He was often able to catch my mistakes in O(1) factors, line slopes, and crossing points.

Mathematica calculation

Lawrence hand-drawn Mathematica calculation

21

One time we encountered a serious problem in our project. I found a solution in the literature and explained to Lawrence. He was immediately amazed by the existing solution and asked who wrote the paper.

One time we encountered a serious problem in our project. I found a solution in the literature and explained to Lawrence. He was immediately amazed by the existing solution and asked who wrote the paper.

It was him.

9/12/2015 Tahoe Summit 2015

9/27/2024 Lawrence/Hitoshi Fest

24

As a grad student, I am not allowed to give you any presents that are worth a lot of money. Therefore, instead of spending a lot of money, I spent a lot of time making this present. And I am sure everyone here will agree that a grad student's time is worth no money. α

9/12/2015 Tahoe Summit 2015

As a grad student, I am not allowed to give you any presents that are worth a lot of money. Therefore, instead of spending a lot of money, I spent a lot of time making this present. And I am sure everyone here will agree that a grad student's time is worth no money. "
"

(Actually, after 9 years and having just gotten my own grant, I now take it back. Grad students' time is expensive.)

9/12/2015 Tahoe Summit 2015

As a grad student, I am not allowed to give you any presents that are worth a lot of money. Therefore, instead of spending a lot of money, I spent a lot of time making this present. And I am sure everyone here will agree that a grad student's time is worth no money. "
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This is a picture of Lawrence, made of 999 dice. This is not a printed picture but literally 999 dice. We know God does not play dice with the Universe. Neither does Lawrence. However, I do. I simply put all of the dice in the box and kept
shaking it until the face of Lawrence emerged shaking it until the face of Lawrence emerged. "

made for Lawrence's birthday celebration during Tahoe Summit 2015

Raymond

made for Lawrence's birthday celebration during Tahoe Summit 2015

Raymond

28

As a grad student, I am not allowed to give you any presents that are worth a lot of money. Therefore, instead of spending a lot of money, I spent a lot of time making this present. And I am sure everyone here will agree that a grad student's time is worth no money. "
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shaking it until the face of Lawrence emerged shaking it until the face of Lawrence emerged. "

(Lawrence, I hope you have not shaken it since. I have to remind you that I am now faculty and do not have the time to shake it back. Just kidding. For you, I will do a lifetime warranty.)

For those lucky few who have used it, it can truly be called the *Michael Jordan* of chalk,
the *Relle Reserve* for the Unit**y** the *Rolls Royce* of chalk. "

https://math.williams.edu/dream-chalk/

Disclaimer: I do not own any shares of the company.

For those lucky few who have used it, it can truly be called the *Michael Jordan* of chalk,
the *Relle Reserve* for the Unit**y** the *Rolls Royce* of chalk. ϵ

https://math.williams.edu/dream-chalk/

Lawrence, I would like to offer you an unlimited supply for life. Here is the first box. Just shoot me an email when you run out of it.

Disclaimer: I do not own any shares of the company.

Impacts and Imprints of Axion Dynamics

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September 28th, 2024

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Indiana University

Outline of the Talk

strong CP problem

Source: Planck

matter (fermions) gauge bosons electro-
magnetic \mathbf{v} ก quarks photon up quark charm quark top quark g $\mathbf d$ strong $\overline{1}$ $\overline{9}$ down quark bottom quark qluon strange quark Z, W е weak (11) $\binom{1}{3}$ $\frac{1}{\sqrt{2}}$ leptons electron tau muon weak bosons V_{e} Н H higgs boson electron neutrino muon neutrino tau neutrino **Higgs bosons**

matter-antimatter dark matter asymmetry

Review of the QCD Axion

Strong CP Problem

u d

neutron

34

-1/3 e

gluon

d

-1/3 e

"Charge-Parity" symmetry

Quantum field theory:
 $\mathcal{L} \supset \bar{\theta} \frac{\alpha_s}{8\pi} \overset{\sim}{G_{b\mu\nu}} \overset{\text{gluon}}{\widetilde{G}_{b\mu\nu}}$

Crewther, Di Vecchia, Veneziano, Witten 1979, Pospelov, Ritz 2000

Experiments:

 $|d_n| \lesssim 1.8 \times 10^{-26}$ e·cm

PRL 97, 131801 2006, PRL 124, 081803 2020

Strong CP problem:

 $\overline{\theta} \lesssim 10^{-10}$ Why exceedingly small?

Strong CP Problem solution

Peccei and the axion for the axion of the axion of the axiom of the axiom of the axiom of the axion of the axion of $\vec{\theta}$ and $\vec{\theta}$ a Peccei, Quinn 1977 Weinberg 1978 Wilczek 1978 Promoted to a dynamical field: 2017 Countamagazine **2017 COD** axion Peccei Quinn

decay constant

QCD effects automatically generate an axion potential $\ V(a)=m_a^2f_a^2\ \left[1-\cos\left(\bar{\theta}+\frac{a}{f_a}\right)\right]$
Status of Axion Dark Matter

Experimental Searches

Experimental progress

shaded regions: excluded broken lines: future sensitivities

damped simple harmonic oscillator

$\left(\partial_t^2 + 3H\partial_t + m_a^2\right)a = 0$
Hubble "friction" Equation of motion:

(from cosmic expansion)

damped simple harmonic oscillator

Equation of motion:

$$
\partial_t^2 + 3H\partial_t + m_a^2\big)\,a = 0
$$

Misalignment Mechanism

Hubble "friction" (from cosmic expansion)

Abbott, Sikivie 1983 Dine, Fischler 1983

Preskill, Wise, Wilczek 1983 **by Callier Controller Manual Article Controller Start when**

 $H \lesssim m_a$

 $\theta \equiv a/f_a$

We assume $\theta_i = \mathcal{O}(1)$ here.

Axion Dynamics

Axion UV Completion

 $\mathcal{L} \supset \frac{\alpha_s}{8\pi} \frac{a}{f_a} G^{\mu\nu}_b \widetilde{G}_{b\mu\nu}$

(radial direction) spontaneously breaks Peccei-Quinn symmetry

S

axion (angular direction)

pseudo Nambu-Goldstone boson of Peccei-Quinn symmetry breaking

Axion Dynamics

✓ Parametric Resonance ✓ Axion Rotations

radial

mode "axion"

 $S+i\chi$

Pedagogical example:

Quartic potential:

 $V=\lambda^2|P|^4$

 $=\frac{\lambda ^2}{4}\left(S^2+ \right.$

Equation of motions:

$$
\ddot{\chi} - \nabla^2 \chi + V''(\chi) \chi = 0
$$

(Non-expanding Universe)

PHYSICAL REVIEW I FTTERS

Highlights Collections Authors Referees Recen Accented

EDITORS' SUGGESTION

QCD Axion Dark Matter with a Small Decay Constant

A proposed new cosmological production mechanism for QCD axion dark matter that involves parametric resonance in field oscillation predicts larger axion masses than the conventional misalignment mechanism.

Raymond T. Co, Lawrence J. Hall, and Keisuke Harigaya Phys. Rev. Lett. 120, 211602 (2018)

$$
V''(\chi) = \lambda^2 S^2 = \lambda^2 S_0^2 \cos^2(\lambda S_0 t)
$$

$$
S \sim S_0 \cos(\lambda S_0 t)
$$

PRL 120, 211602 (2018) RC, L. Hall, and K. Harigaya

Oscillation frequency of the field in absence of the driving force

Oscillation frequency of the driving force

$$
\ddot{\chi} - \frac{\mathbf{v}}{\mathbf{X}^2} \chi + V''(\chi) \chi = 0 \qquad V''(\chi) = \frac{\chi^2 S^2}{\mathbf{A}^2} = \lambda^2 S_0^2 \cos^2(\lambda S_0 t)
$$

Resonance occurs for some specific frequencies: warm dark matter.

46 PRL 120, 211602 (2018) RC, L. Hall, and K. Harigaya

Graphical understanding

No Enhancement **Parametric Resonance**

PRL 120, 211602 (2018) RC, L. Hall, and K. Harigaya

48 PRL 120, 211602 (2018) RC, L. Hall, and K. Harigaya

Non-thermal PQ Symmetry Restoration

parametric resonance when *P* oscillates

fluctuations create effective mass term

end of inflation \mathcal{L}

|

|

 $\mathsf{T}_{\mathrm{osc}}$ $\mathsf{-}$

 $\overline{\mathbf{A}}$

E

PQ symmetry restored cosmic strings created

kinetic misalignment

RC, K. Harigaya, T. Lee, A. Pierce Grad student: Taegyu Lee

Animation for illustration purposes only. Caution: string core \sim 0.5 lattice spacing

Axion Parametric Resonance

Parametric resonance from hilltop/trapped misalignment

PRD 101 (2020) 8, 083014 A. Arvanitaki, S. Dimopoulos, M. Galanis, L. Lehner, J. Thompson, K. Van Tilburg JCAP 10 (2021) 001 L. Luzio, B. Gavela, P. Quilez, A. Ringwald JHEP 09 (2024) 145 RC, T. Gherghetta, Z. Liu, K. Lyu 2408.04623 L. Luzio, P. Sørensen

Parametric resonance from kinetic misalignment

JCAP 01 (2017) 036 J. Jaeckel, V. Mehta, L. Witkowski JCAP 08 (2019) 020 J. Berges, A. Chatrchyan, J.Jaeckel JHEP 04 (2020) 010 N. Fonseca, E. Morgante, R. Sato, G. Servant JHEP 12 (2021) 099 RC, K. Harigaya, A. Pierce JCAP 10 (2022) 053 C. Eroncel, R. Sato, G. Servant, P. Sørensen JCAP 01 (2023) 009 C. Eroncel, G. Servant

end of inflation $\frac{1}{2}$

|

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 OSC

 $\overline{\mathbf{A}}$

E

$|\mathcal{C}{\rm osm}o\mathcal{L}{\rm attice}|$

2102.01031

Work in Progress

High schooler: Ardashir Kocer Undergrad: Owen Leonard Undergrad: Bohao Wang Grad student: Taegyu Lee

kinetic misalignment

Axion Rotation

 $\mathcal{L} \supset \frac{\alpha_s}{8\pi} \frac{a}{f_a} G^{\mu\nu}_b \widetilde{G}_{b\mu\nu}$

(radial direction)

spontaneously breaks Peccei-Quinn symmetry axion (angular direction)

pseudo Nambu-Goldstone boson of Peccei-Quinn symmetry breaking

matter-antimatter example and dark matter asymmetry

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PHYS BORG

Paper sheds light on infant universe and origin of matter

10 March 2020

The rotation of the QCD axion (black ball) produces an excess of matter (colored balls) over antimatter, allowing "The versatility of the QCD axion in solving the galaxies and human beings to exist. Credit: Graphic: Harigaya and Co; Photo: NASA

A new study, conducted to better understand the origin of the universe, has provided insight into some of the most enduring questions in fundamental physics: How can the Standard Model Harigaya and Co have reasoned that the QCD of particle physics be extended to explain the cosmological excess of matter over antimatter? What is dark matter? And what is the theoretical origin of an unexpected but observed symmetry in the force that binds protons and neutrons together?

In the paper "Axiogenesis," scheduled to be 2020, researchers Keisuke Harigaya, Member in

the School of Natural Sciences at the Institute for Advanced Study, and Raymond T. Co of the University of Michigan, have presented a compelling case in which the quantum chromodynamics (QCD) axion, first theorized in 1977, provides several important answers to these questions

> "We revealed that the rotation of the QCD axion can account for the excess of matter found in the universe," stated Harigaya. "We named this mechanism axiogenesis."

Infinitesimally light, the QCD axion-at least one billion times lighter than a proton-is nearly ghostlike. Millions of these particles pass through ordinary matter every second without notice. However, the subatomic level interaction of the QCD axion can still leave detectable signals in experiments with unprecedented sensitivities. While the QCD axion has never been directly detected, this study provides added fuel for experimentalists to hunt down the elusive particle.

mysteries of fundamental physics is truly amazing, stated Co. "We are thrilled about the unexplored theoretical possibilities that this new aspect of the QCD axion can bring. More importantly, experiments may soon tell us whether the mysteries of nature truly hint towards the QCD axion."

axion is capable of filling three missing pieces of the physics jigsaw puzzle simultaneously. First, the QCD axion was originally proposed to explain the so-called strong CP problem-why the strong force, which binds protons and neutrons together, unexpectedly preserves a symmetry called the Charge Parity (CP) symmetry. The CP symmetry is inferred from the observation that a neutron does published in Physical Review Letters on March 17, not react with an electric field despite its charged constituents. Second, the QCD axion was found to

 $\overline{1/3}$

Articles published week ending

PHYSICAL

REVIEW

LETTERS

C: Quantamagazine

Physics **Mathematics**

ABSTRACTIONS BLOG **Axions Would Solve Another Major Problem in Physics**

 \bullet \bullet \blacksquare In a new paper, physicists arque that hypothetical particles called axions could explain why the universe isn't empty.

Published by **American Physical Society**

Volume 124, Number 11

Dynamics analogous to that in Affleck-Dine baryogenesis

I. Affleck and M. Dine 1991

PRL 124, 111602 (2020) RC and K. Harigaya PRL 92, 011301 (2004) T. Chiba, F. Takahashi, M. Yamaguchi

S

 \mathbf{z}

Wiggles :

$$
V(P) \sim \frac{P^n}{M^{n-4}} e^{i\varphi} + \text{h.c.} \sim \frac{|P|^n}{M^{n-4}} \cos\left(n\frac{a}{f_a} + \varphi\right)
$$

Explicit PQ breaking

expected from quantum gravity or PQ as an accidental symmetry S. Giddings et al. 1988 S. Coleman 1988 G. Gilbert 1988

Dynamics analogous to that in Affleck-Dine baryogenesis

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axion

 $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$

S

D

 $V(P)$

Wiggles :

$$
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$$

Explicit PQ breaking

expected from quantum gravity or PQ as an accidental symmetry S. Giddings et al. 1988 S. Coleman 1988 G. Gilbert 1988

Large field value : Flat potential

For example, as an initial condition or set dynamically by inflationary dynamics

$$
V(|P|)\sim -H_I^2|P|^2+\frac{|I|}{M}
$$

Hubble-induced mass

M. Dine, L. Randall, and S. D. Thomas 1991

Dynamics analogous to that in Affleck-Dine baryogenesis

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axion

 $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{q}{f_a}}$

Wiggles :

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$$

Explicit PQ breaking

expected from quantum gravity or PQ as an accidental symmetry S. Giddings et al. 1988 S. Coleman 1988 G. Gilbert 1988

Work in Progress: Axion Non-Gaussianity

PQ field and inflaton interaction JHEP 12 (2023) 197 X. Chen, J. Fan, L. Li Higher dimensional $U(1)_{PQ}$ breaking

Postdoc: Sai Chaitanya Tadepalli Grad student: Taegyu Lee

S

Dynamics analogous to that in Affleck-Dine baryogenesis

I. Affleck and M. Dine 1991

PRL 124, 111602 (2020) RC and K. Harigaya PRL 92, 011301 (2004) T. Chiba, F. Takahashi, M. Yamaguchi

axion

 $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{q}{f_a}}$

Asymmetry of PQ Charge

Noether charge associated with the shift symmetry

$$
n_{\theta} = S^2 \dot{\theta}
$$

this is nothing but "angular momentum" $r^2\omega$

S

 $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$

axion

PQ asymmetry PQ charge density PQ Fotation of PQ field =

This is conserved soon after the initial kick.

Asymmetry of PQ Charge

Noether charge associated with the shift symmetry

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this is nothing but "angular momentum" $r^2\omega$

 $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$

PQ asymmetry =

 PQ charge density PQ Fotation of PQ field

This is conserved soon after the initial kick. *^S*

 $F_c = ma_c$ What determines θ ? Centripetal force! $V'(S) = S\dot{\theta}^2$ $V(S) \simeq \frac{1}{2} m_S^2 S^2$ $m_S^2 \hat{S} = S \dot{\theta}^2$ from supersymmetry $\theta \equiv m_S^-$ which is in turn set by supersymmetry scale.

axion

PQ Charge Evolution

Charge conservation:

scale factor of the universe

 $n_{\theta} = S^2 \dot{\theta} \propto \dot{R}^{-3}$

dilution due to cosmic expansion

PRL 124, 111602 (2020) RC and K. Harigaya

S

axion

 $\overline{f_a}$

 θ

PQ Charge Evolution

Charge conservation:

scale factor of the universe

$$
n_\theta = S^2 \dot{\theta} \propto \mathring{R}^{-3}
$$

Large field $(S \gg f_a)$:

$$
S^2 \propto R^{-3}
$$

for quadratic potential $V(S) \simeq \frac{1}{2} m_S^2 S^2$

dilution due to cosmic expansion

 $\theta = constant$ $\rho_\theta = \dot{\theta}^2 S^2 \propto R^{-3}$ matter

S

axion

 θ

At the minimum:

$$
S^{2} = f_{a}^{2} \qquad \dot{\theta} \propto R^{-3}
$$

$$
\rho_{\theta} = \dot{\theta}^{2} f_{a}^{2} \propto R^{-6}
$$

kination

PRL 124, 111602 (2020) RC and K. Harigaya

Gravitational Wave Signatures

The energy content determines the universe's expansion rate.

64 **Conserverse Cools and temperature drops.** Over time, the universe cools and temperature drops.

The energy content determines the universe's expansion rate.

Over time, the universe cools and temperature drops.

The energy content determines the universe's expansion rate.

Smoking Gun: Triangular Peak in Gravitational Wave Spectra

2108.09299 JHEP 09 (2022) 116 RC, D. Dunsky, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya, J. Shelton 2108.10328, 2111.01150 Y. Gouttenoire, G. Servant, P. Simakachorn

Smoking Gun: Triangular Peak in Gravitational Wave Spectra

2108.09299 JHEP 09 (2022) 116 RC, D. Dunsky, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya, J. Shelton 2108.10328, 2111.01150 Y. Gouttenoire, G. Servant, P. Simakachorn

Smoking Gun: Triangular Peak in Gravitational Wave Spectra

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The energy content determines the universe's expansion rate.

2108.09299 RC, N. Fernandez, A. Ghalsasi, K. Harigaya, J. Shelton

Axion Dark Matter
a novel scenario where the axion field has a nonzero initial velocity, e.g., from axion rotations.

axion cosine potential

PRL 124, 251802 (2020) RC, L. Hall, K. Harigaya

a novel scenario where the axion field has a nonzero initial velocity, e.g., from axion rotations.

kinetic energy > potential energy enhancing axion abundance

axion cosine potential

PRL 124, 251802 (2020) RC, L. Hall, K. Harigaya

giving a strong motivation for axion dark matter experiment

Baryon Asymmetry

Axiogenesis

Baryogenesis is automatic, thanks to Standard Model processes exercine that the matter-antimatter asymmetry (production of baryon asymmetry)

$$
Y_B \equiv \frac{n_B}{s} = c_B Y_\theta \left(\frac{T_{\rm EW}}{f_a}\right)^2
$$

produced by axion rotations

 $Y_B^{\rm obs} \simeq 8.7 \times 10^{-11}$

experimentally measured value (*Planck* 2018)

PRL 124, 111602 (2020) RC and K. Harigaya

Axiogenesis

Baryogenesis is automatic, thanks to Standard Model processes exercine that the matter-antimatter asymmetry (production of baryon asymmetry)

$$
Y_B \equiv \frac{n_B}{s} = c_B Y_\theta \left(\frac{T_{\rm EW}}{f_a}\right)^2 \qquad Y_B^{\rm obs} \simeq 8.7 \times 10^{-11} \tag{1}
$$

 $2\,$

 $\stackrel{Y_B}{\longleftrightarrow} Y_\theta$

Namely, the baryon asymmetry relates

experimentally measured value (*Planck* 2018)

PRL 124, 111602 (2020) RC and K. Harigaya

 $T_{\rm EW}$

Axion/ALP Cogenesis

Kinetic Misalignment + Axiogenesis

 $\overline{2}$ $T_{\rm EW}$ $m_a \xleftrightarrow{DM} Y_\theta \xleftrightarrow{Y_B}$

Kinetic Misalignment + Axiogenesis

 $\overline{2}$

EW

Prediction:

$$
m_a \overset{\mathrm{DM}}{\longleftrightarrow} Y_\theta \overset{Y_B}{\longleftrightarrow} \Big(\overset{\cdot}{\neg}
$$

 $\int a$ ALP cogenesis : $ALP = axion-like particle$ (no gluon coupling) cogenesis = production of both dark matter & matter-antimatter asymmetry We assume T_{EW} = 130 GeV.

Experimental probes are happening! ALP Cogenesis

Kinetic Misalignment + Axiogenesis

 $2\,$

EW

Prediction:

$$
m_a \stackrel{\text{DM}}{\longleftrightarrow} Y_\theta \stackrel{Y_B}{\longleftrightarrow} \Big(\stackrel{\cdot}{\cdot}
$$

 $\int a$ ALP cogenesis : $ALP = axion-like particle$ (no gluon coupling) cogenesis = production of both dark matter & matter-antimatter asymmetry We assume T_{EW} = 130 GeV. QCD axion cogenesis? Can the QCD axion be compatible with cogenesis from axion rotations?

Yes. This is a great opportunity to bring other open questions into the picture!

89 JHEP 01 (2021) 172 RC, L. Hall, K. Harigaya

Extensions of Axiogenesis

any additional *lepton* or *baryon* number-violating processes

Extensions of Axiogenesis

any additional *lepton* or *baryon* number-violating processes

Lepto-Axiogenesis other extensions

$$
\mathcal{L} \supset \frac{m_\nu}{2 v_{\rm EW}^2} \ell \ell \, H^\dagger H^\dagger
$$

91

This Weinberg operator gives Majorana neutrino masses, breaks lepton number, and thus affects the charge transfer.

JHEP 03 (2021) 017 RC, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya

Lepto-Axiogenesis

Producing *L* at high temperatures Converting to *B* at T_{EW}

Lepto-Axiogenesis

simultaneous production of

- dark matter
- matter-antimatter asymmetry in the framework with
	- QCD axion
	- Majorana neutrinos

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Axion Rotations

Axion Kination

✓ **RC** *et al.* 2108.09299 ✓ Gouttenoire *et al.* 2108.10328 Dark Matter **✓ Gouttenoire** *et al.* 2111.01150

Baryogenesis

✓ **RC** *et al.* 1910.02080

✓ **RC** *et al.* 1910.14152

- ✓ Chang *et al.* 1911.11885
- ✓ **RC** *et al.* 2004.00629
- ✓ Di Luzio *et al.* 2102.01082
- ✓ Rusov *et al.* 2109.01833
- ✓ Barman *et al.* 2111.03677
- ✓ Eröncel *et al.* 2206.14259
- ✓ Eröncel *et al.* 2207.10111
- \checkmark Oikonomou 2208.05544
- ✓ Kozów *et al.* 2212.03518
- ✓ Chatrchyan *et al.* 2305.03756
- ✓ Lee *et al.* 2310.17710
- ✓ **RC** *et al.* 2312.17730

Gravitational Waves **Cosmology**

- ✓ **RC** *et al.* 2104.02077
- ✓ Madge *et al*. 2111.12730
- ✓ Harigaya *et al* 2305.14242

✓ **RC** *et al.* 2006.04809 ✓ Domcke *et al.* 2006.03148 ✓ **RC** *et al.* 2006.05687 ✓ Harigaya *et al.* 2107.09679 ✓ Chakraborty *et al.* 2108.04293 ✓ Kawamura *et al.* 2109.08605

- ✓ **RC** *et al.* 2110.05487
- ✓ **RC** *et al.* 2206.00678
- ✓ Barnes, **RC** *et al.* 2208.07878
- ✓ **RC** *et al.* 2211.12517
- \checkmark Berbig 2307.14121
- ✓ Chao *et al.* 2311.06469
- ✓ Chun *et al.* 2311.09005
- ✓ Wada 2404.10283
- ✓ Datta *et al.*2405.07003

✓ **RC** *et al.* 2202.01785 ✓ **RC** *et al.* 2405.12268

Conclusions

\checkmark New axion dynamics allows the QCD axion to simultaneously explain

- Strong CP problem
- dark matter abundance
- baryon asymmetry
- ✓This paradigm predicts exciting phenomenology
	- \checkmark specific axion mass-coupling relations
	- \checkmark axion kination: unique gravitational wave spectra
- ✓Other possible signatures include
	- ✓ gravitational lensing of axion mini-clusters
	- \checkmark enhanced matter power spectrum
	- warm axion dark matter

✓New model building opportunities

 \checkmark other open questions across disciplines