

Unravelling the Particle World and the Cosmos at Berkeley

Sep. 28th, 2024

Higgs and Z_2 symmetry

Keisuke Harigaya (U Chicago)

Hitoshi, particle, and cosmos

“Mystery of missing antimatter”



Oct. 25th, 2008
at the University of Tokyo

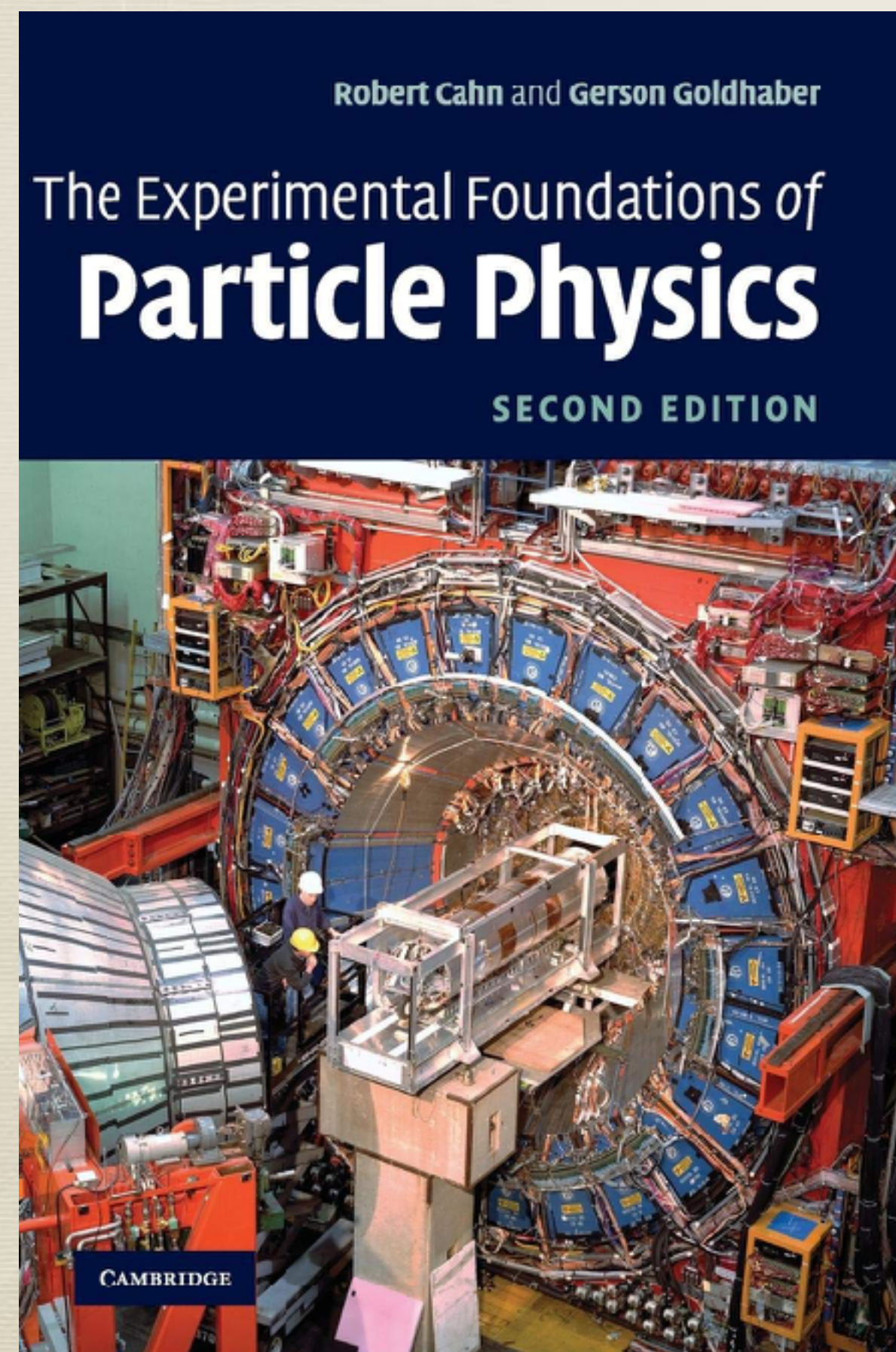
<https://www.ipmu.jp/ja/story/7534>

Kavli IPMU 4th anniversary



Reading club

Hitoshi organized a weekly reading club for first-year Ph.D students



We learned the use of symmetry, the Standard Model (except for Higgs), etc.

Berkeley and Kavli IPMU



2013? at Kyoto



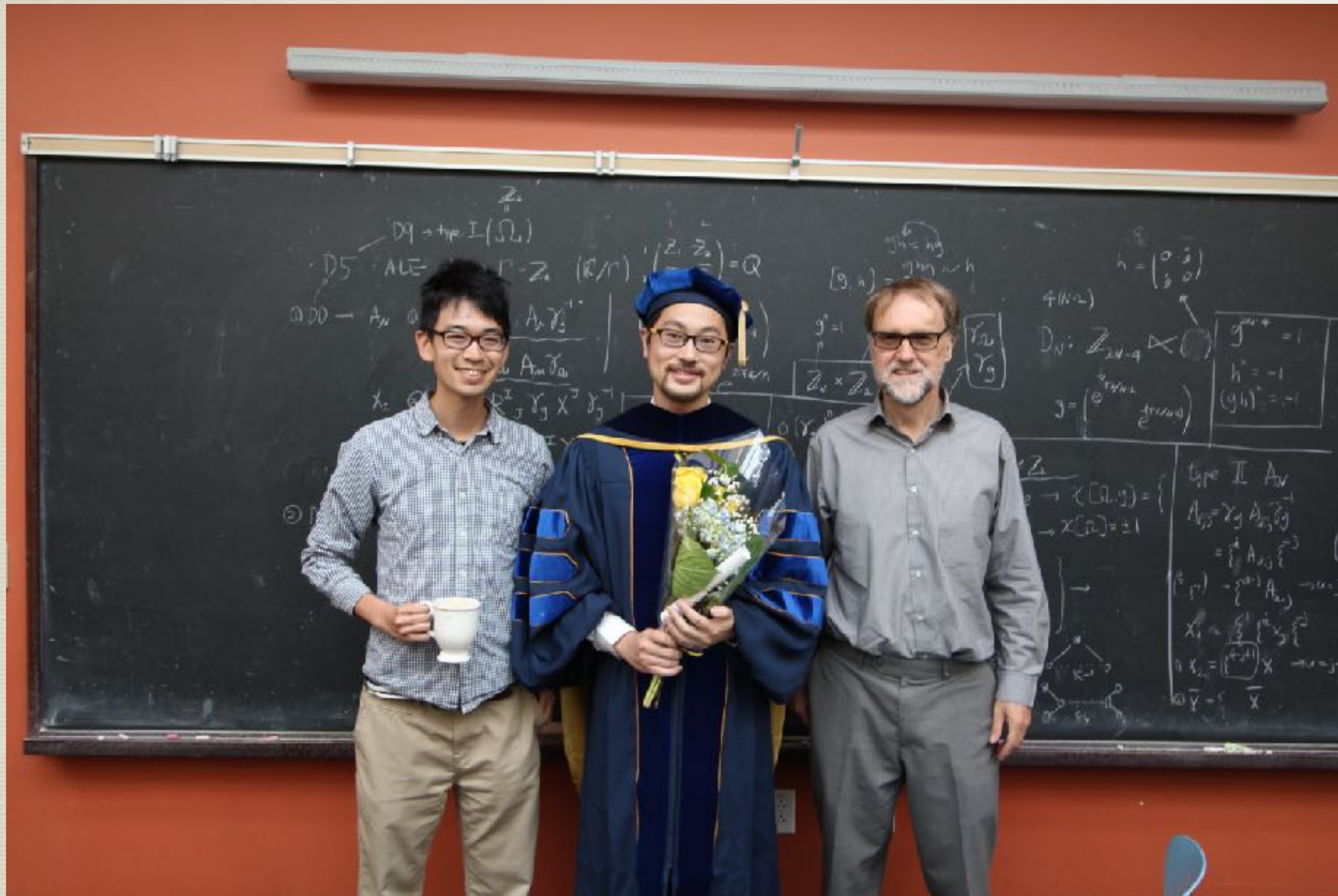
2016 Berkeley-IPMU week

Berkeley and Kavli IPMU



2018 Berkeley-IPMU week

Working with Lawrence



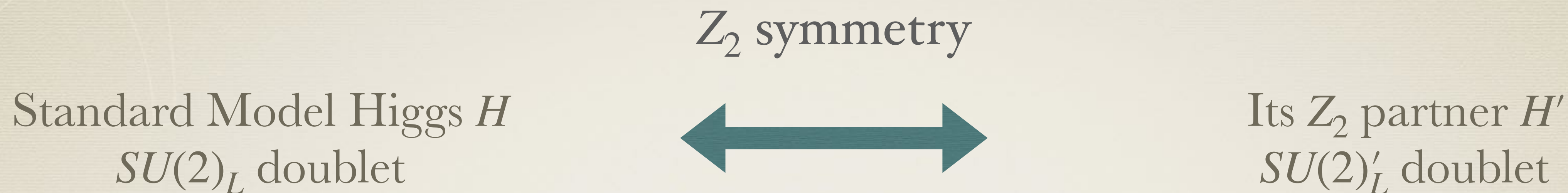
Development in QCD axion cosmology with Lawrence, Raymond Co, Aaron Pierce, ...

Raymond Co's talk

Outline

- * Z_2 symmetry of Higgs and its spontaneous breaking, how it is correlated with Standard Model parameters
- * Parity, $SO(10)$, and dark matter

Z_2 symmetry of Higgs



- * If Z_2 introduces the mirror copy of the SM, dark matter candidates

Mirror baryons, electrons, atoms, ...

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times (SU(3)_c \times SU(2)_L \times U(1)_Y)'$$

Goldberg and Hall (1986)

- * If Z_2 involves space-time parity, the strong CP problem can be solved

Babu and Mohapatra (1989), Barr, Chang and Senjanovic (1991),

Hall and KH (2018, 2019)

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times SU(2)'_L \times U(1)'_Y$$

~~$\theta G\tilde{G}$~~

Z_2 needs to be broken

- * For mirror Z_2 , there would be too much extra radiation in the universe. Also, mirror atom dark matter would be too strongly self-interacting

$$\sigma_{\text{atom}} \sim \frac{1}{m_e^2 \alpha^2}$$

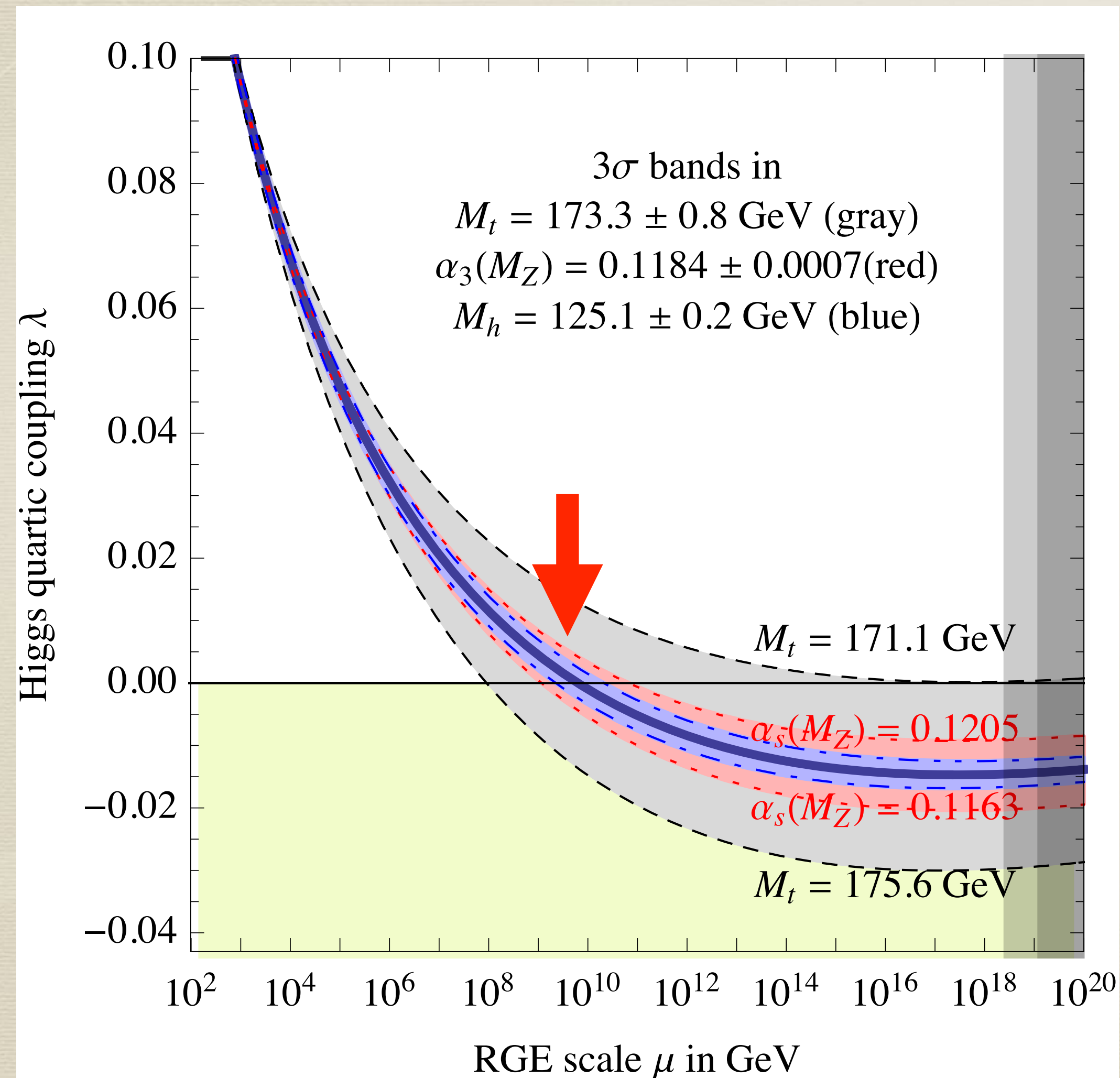
- * For space-time parity, there would too light particles that has $O(1)$ coupling to SM particles

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

$$m_{W_R} = m_{W_L}$$

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times SU(2)'_L \times U(1)'_Y \quad m_{q'} = m_q$$

Z_2 and vanishing SM Higgs quartic



In 2017, we found a relation between the vanishing quartic coupling and Z_2 symmetry

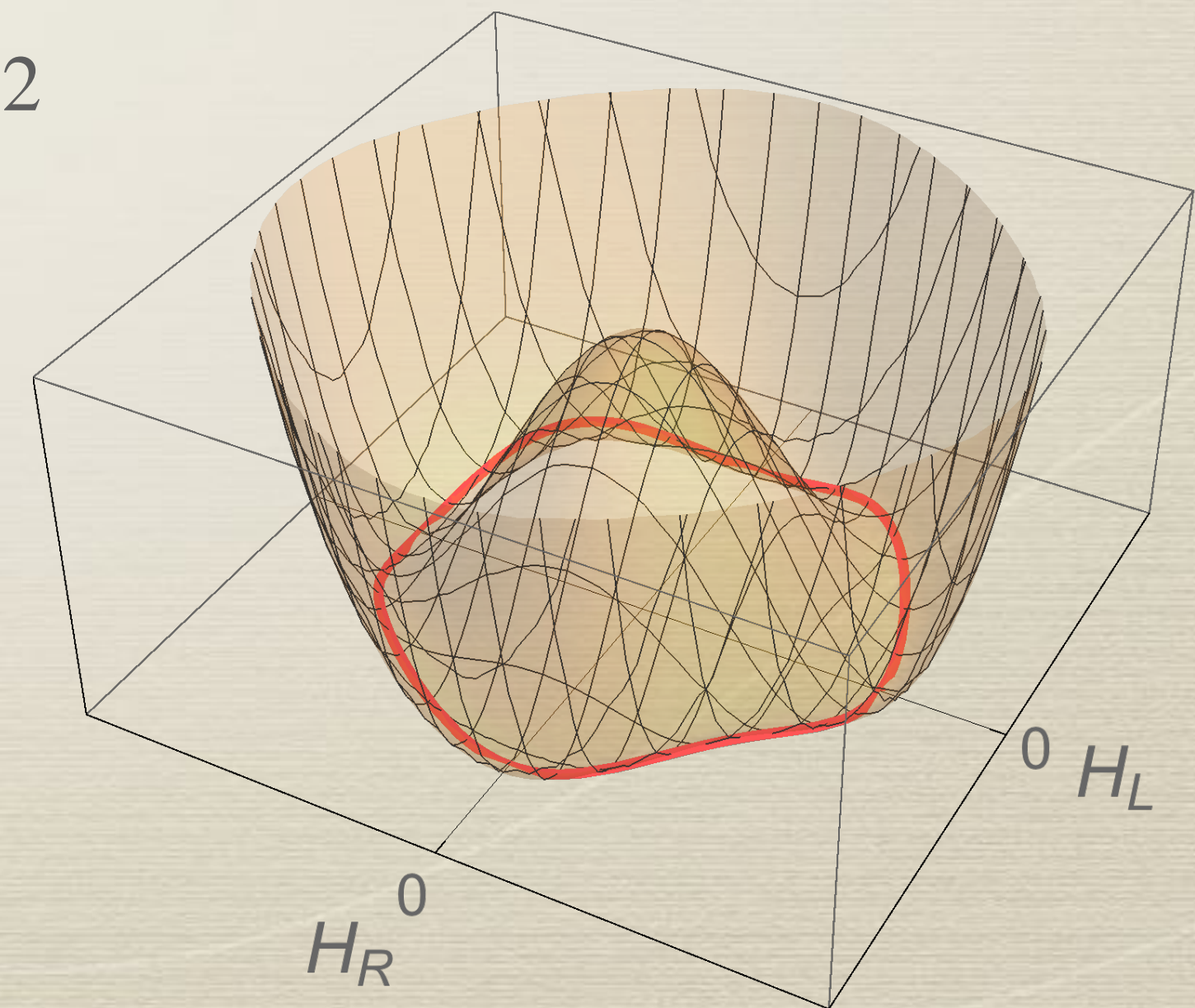
Higgs potential



$$\begin{aligned}
 V &= \left(\lambda |H|^4 - m^2 |H|^2 \right) + \left(\lambda |H'|^4 - m^2 |H'|^2 \right) + \tilde{y} |H|^2 |H'|^2 \\
 &= \lambda \left(|H|^2 + |H'|^2 - v'^2 \right)^2 + y |H|^2 |H'|^2
 \end{aligned}$$

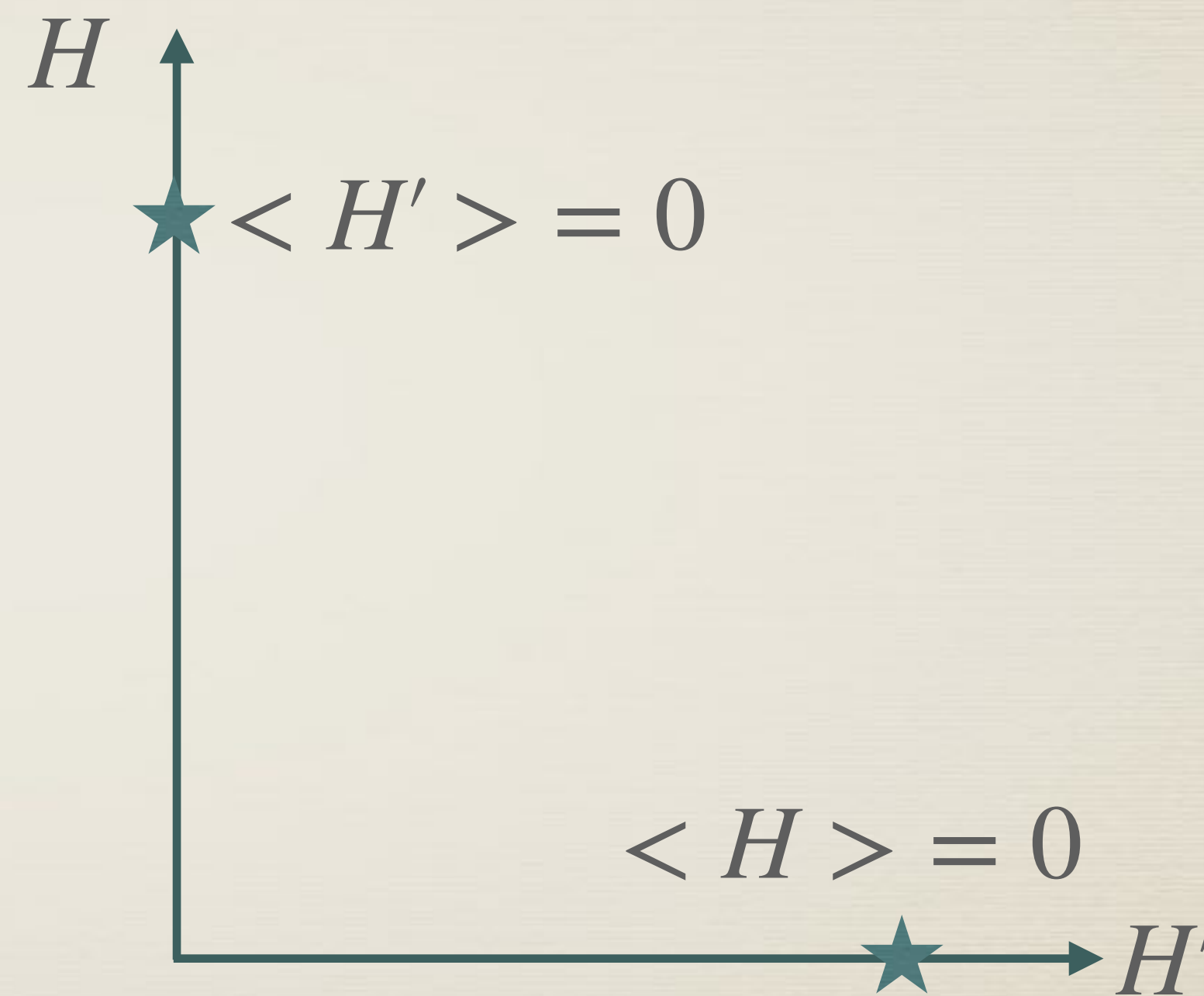
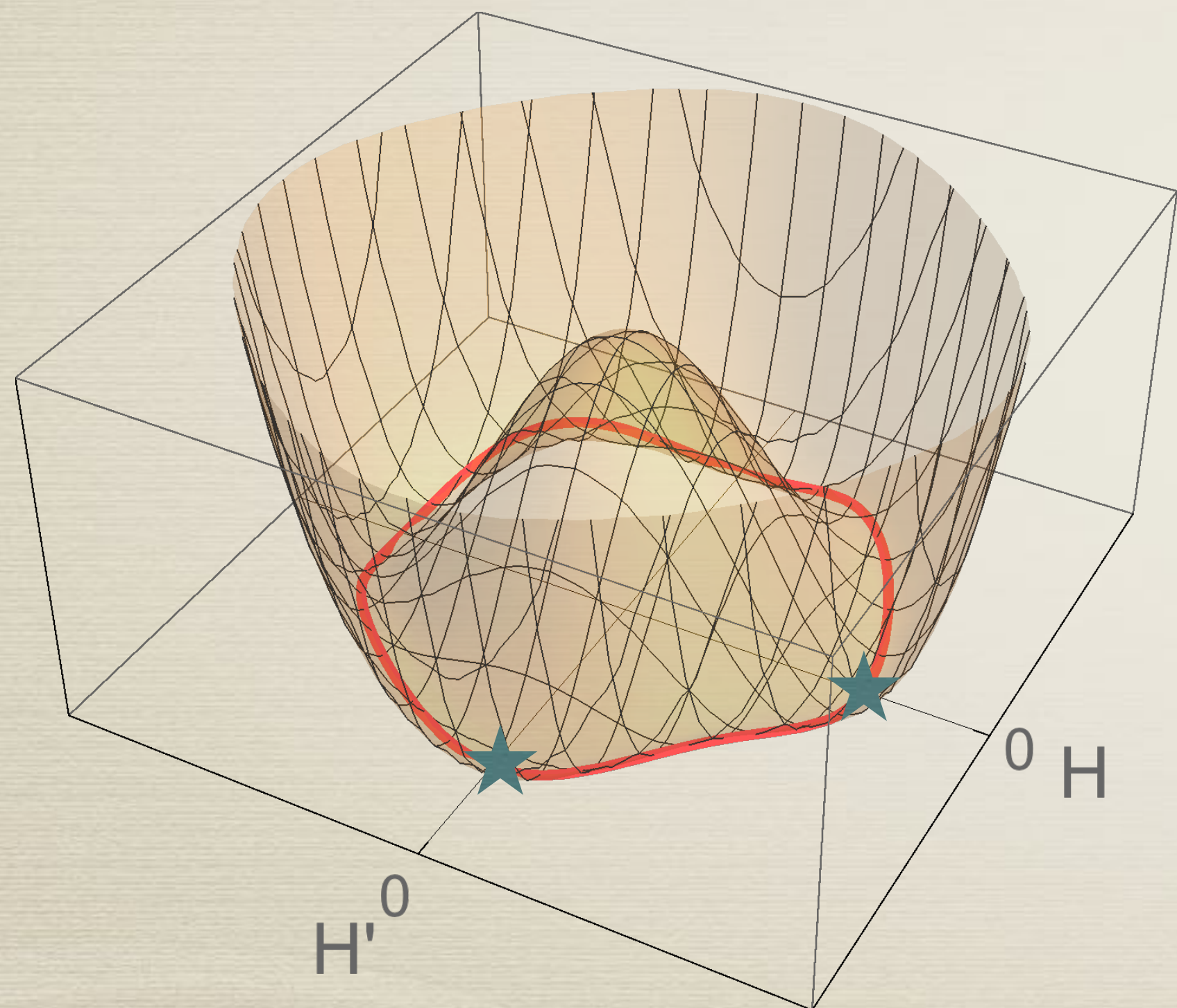
Can we find the minimum with

$$\langle H \rangle \ll \langle H' \rangle ?$$



$$y > 0$$

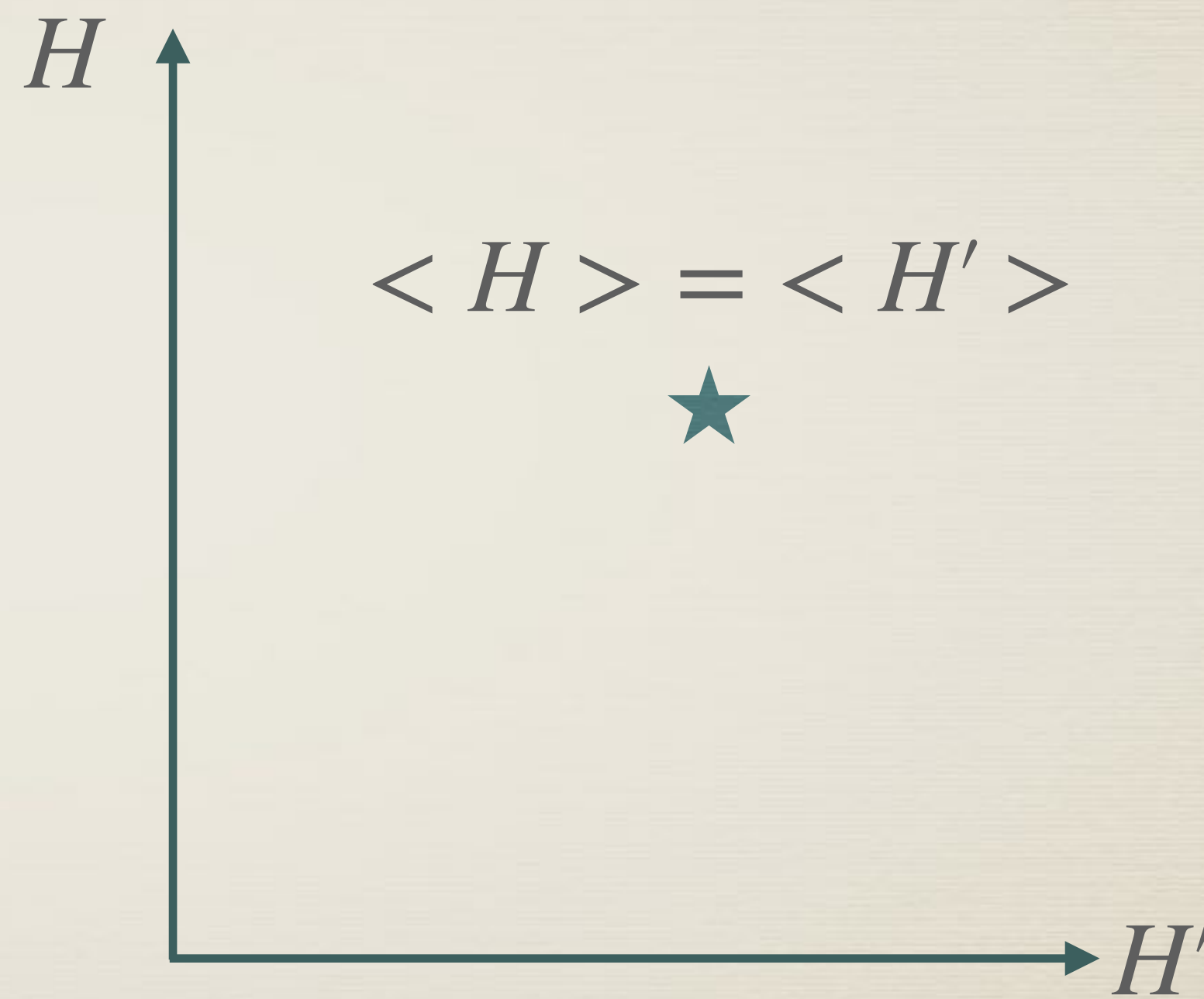
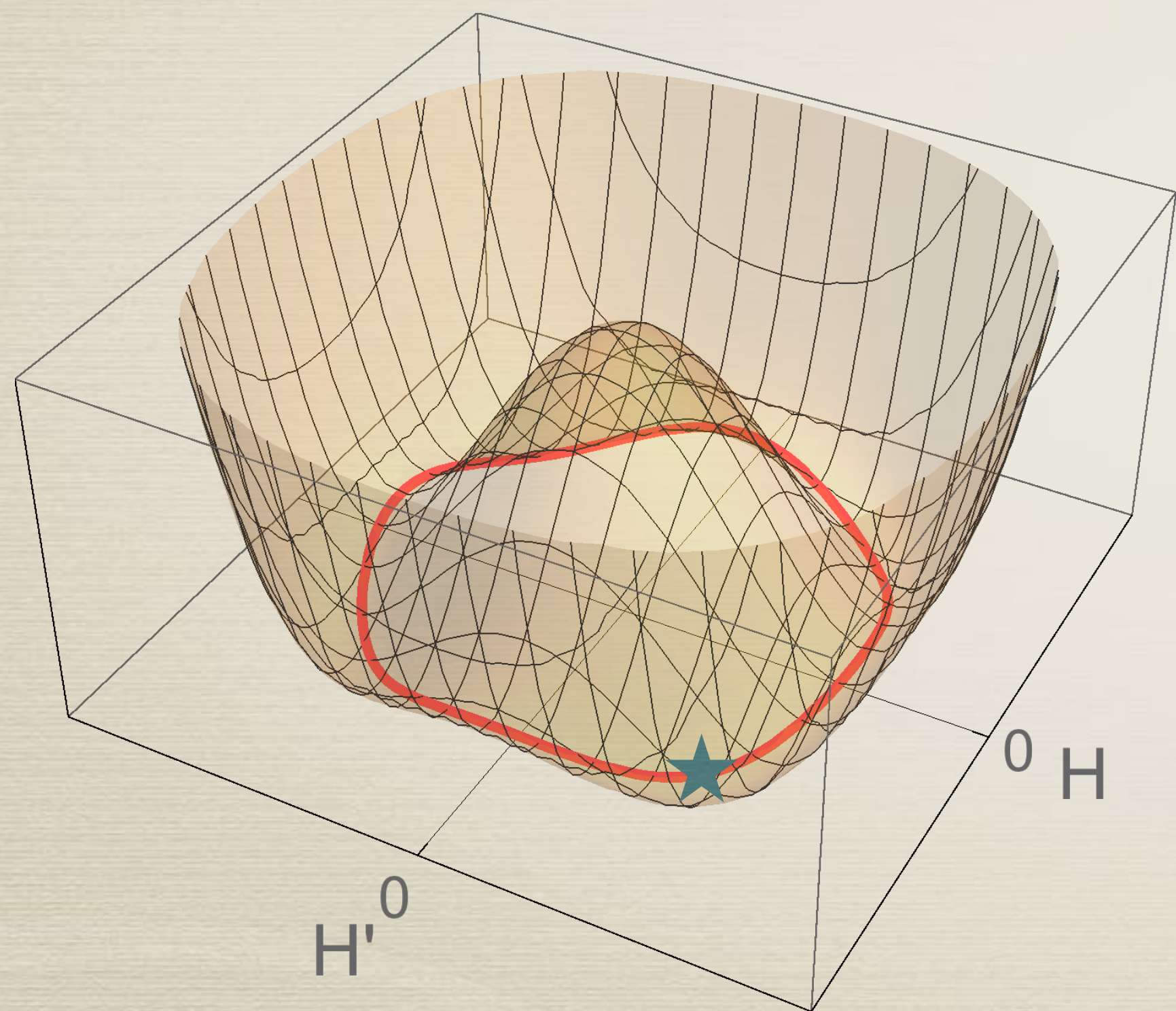
$$V = \lambda(|H|^2 + |H'|^2 - v'^2)^2 + y|H|^2|H'|^2$$



$$0 \neq \langle H \rangle \ll \langle H' \rangle$$

$$y < 0$$

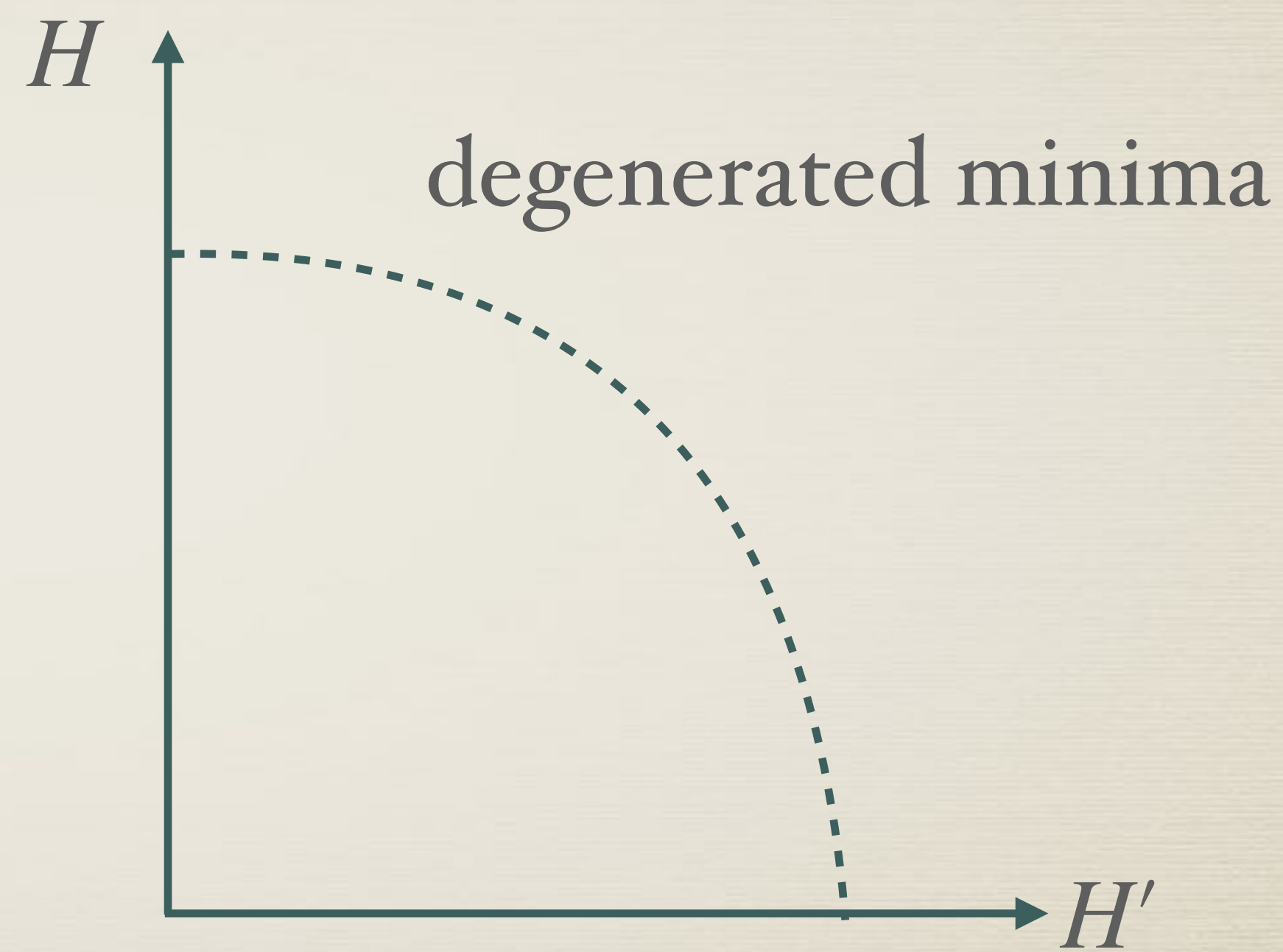
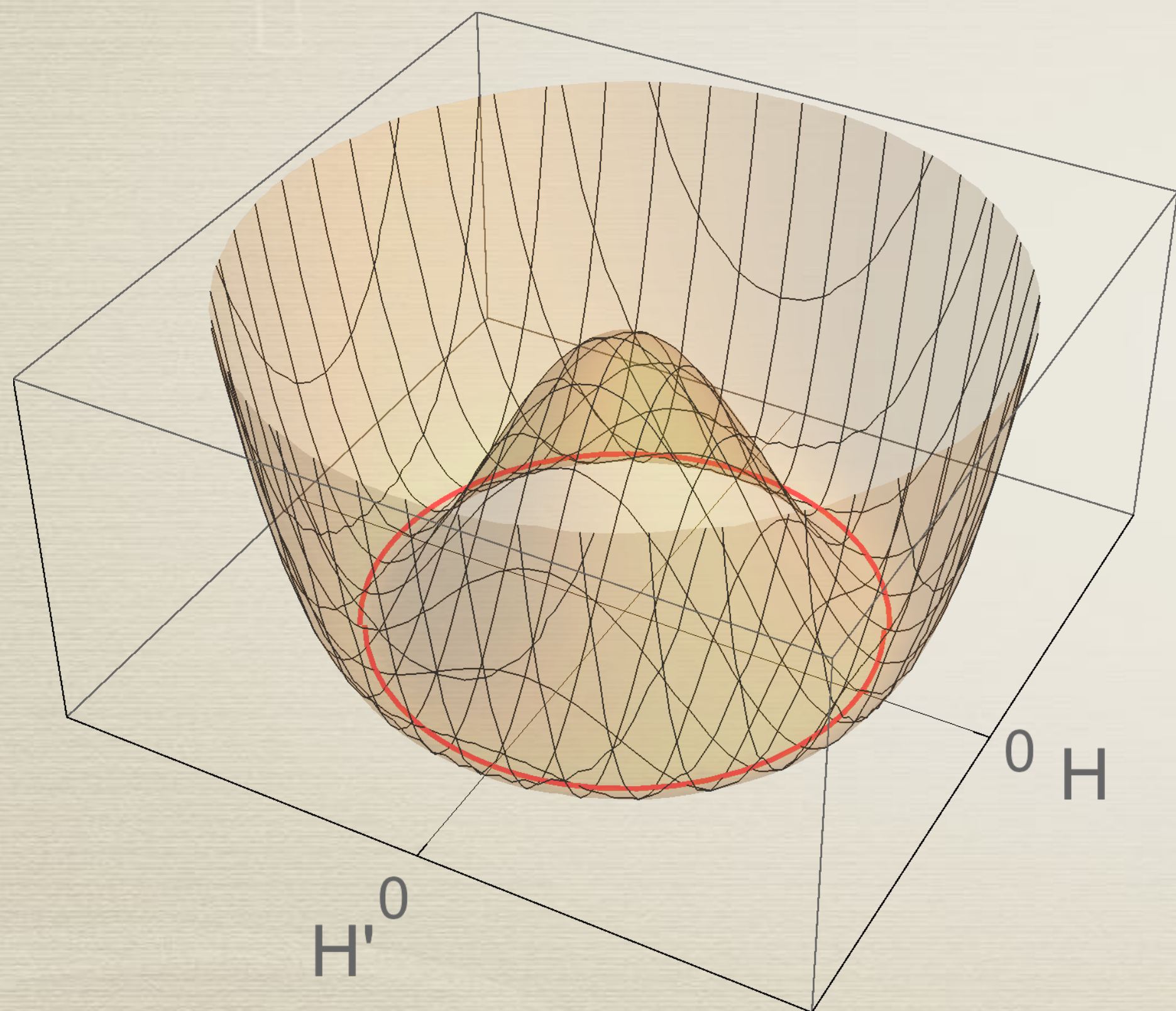
$$V = \lambda(|H|^2 + |H'|^2 - v'^2)^2 + y|H|^2|H'|^2$$



$$\langle H \rangle \ll \langle H' \rangle$$

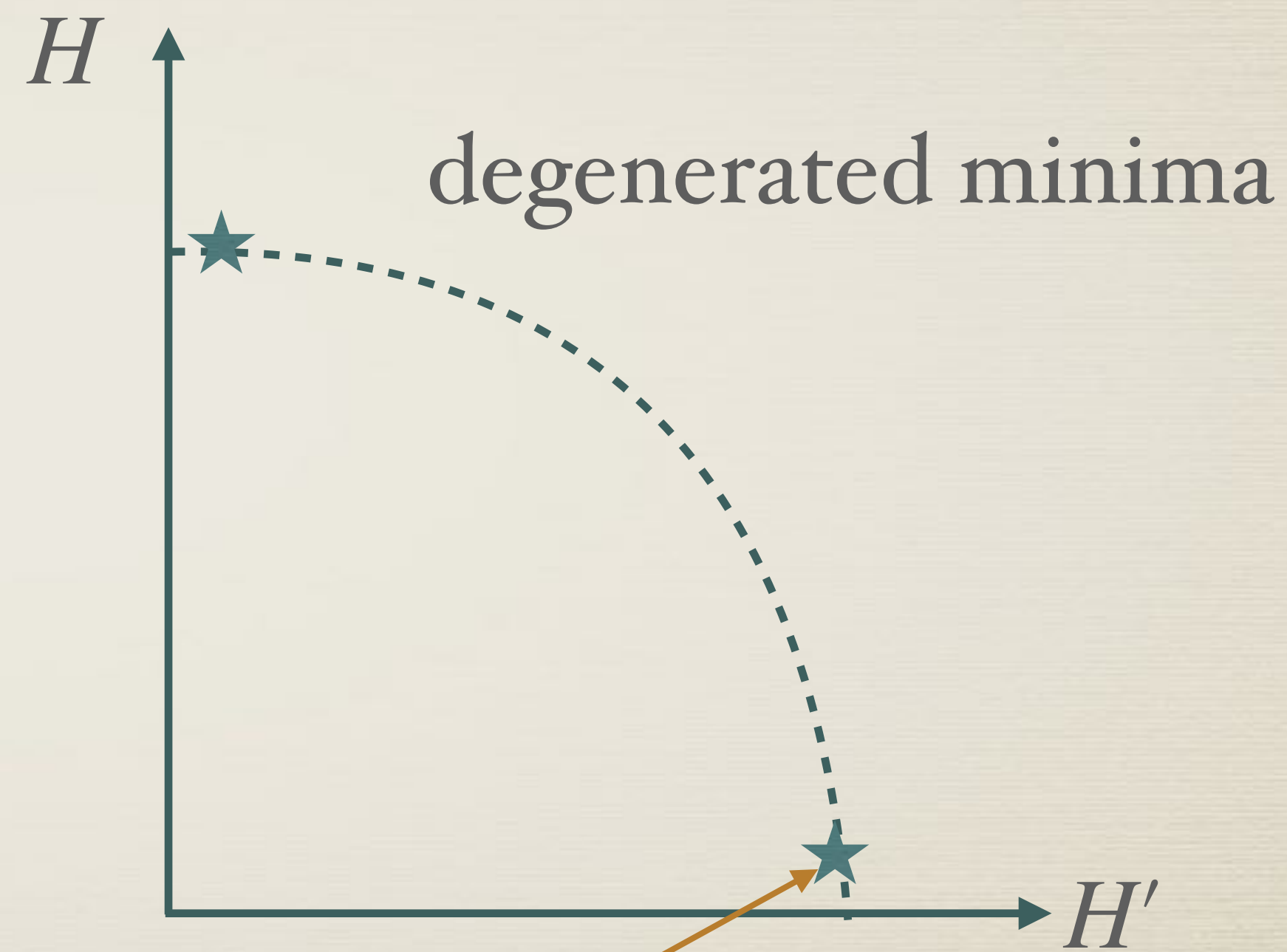
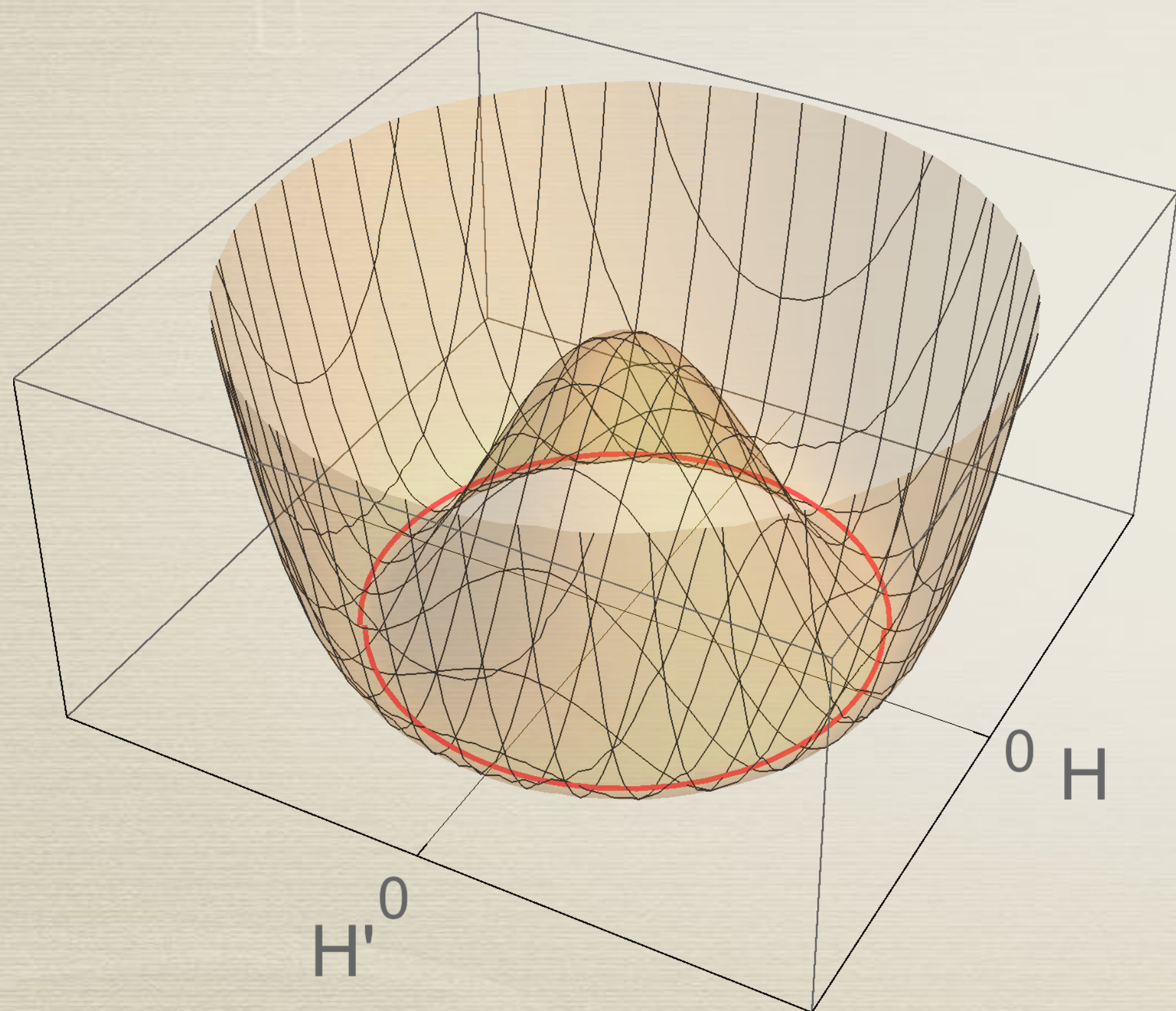
$$y \simeq 0$$

$$V = \lambda(|H|^2 + |H'|^2 - v'^2)^2 + y|H|^2|H'|^2$$



$$y \simeq 0$$

$$V = \lambda(|H|^2 + |H'|^2 - v'^2)^2 + y|H|^2|H'|^2$$



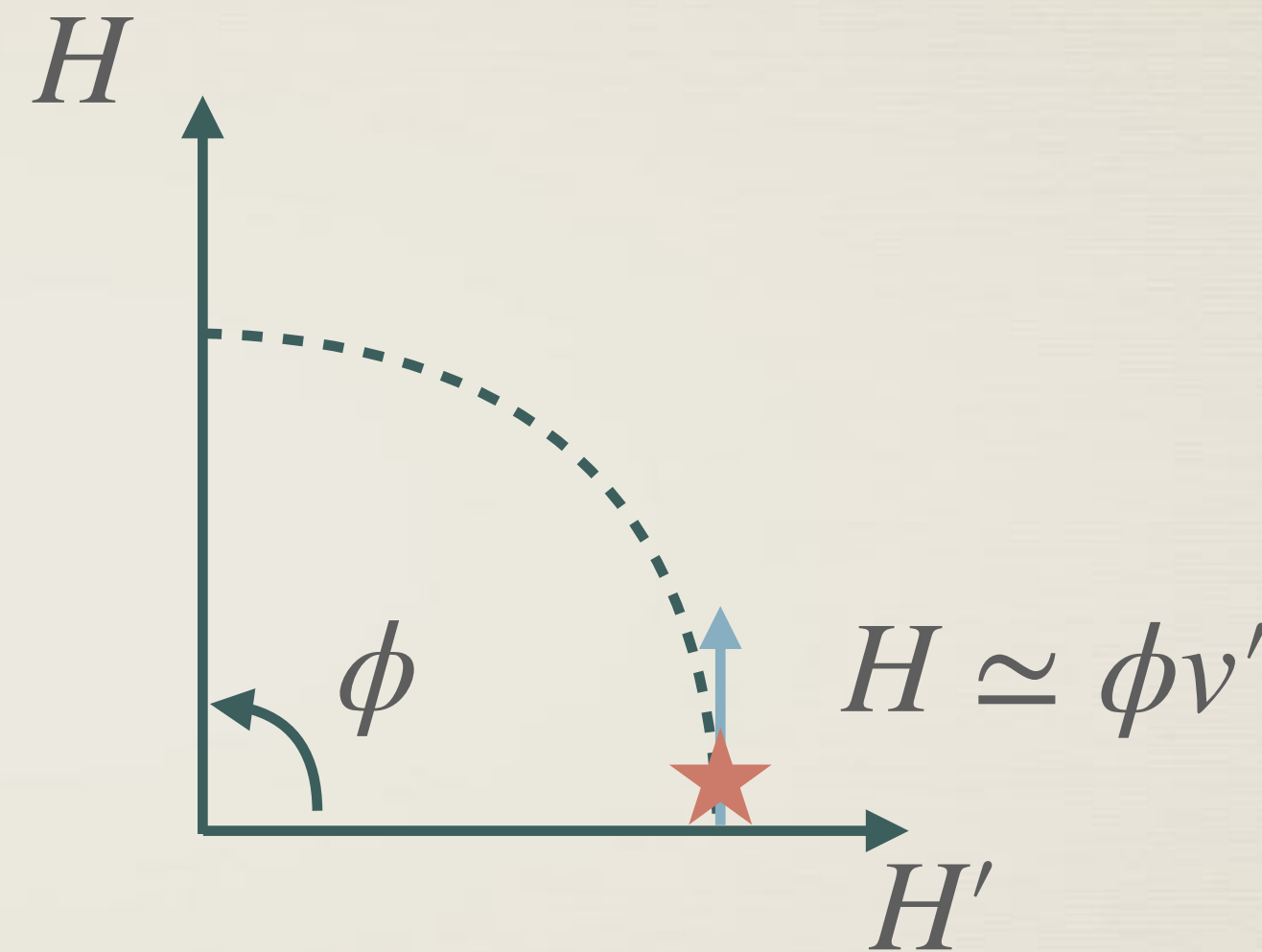
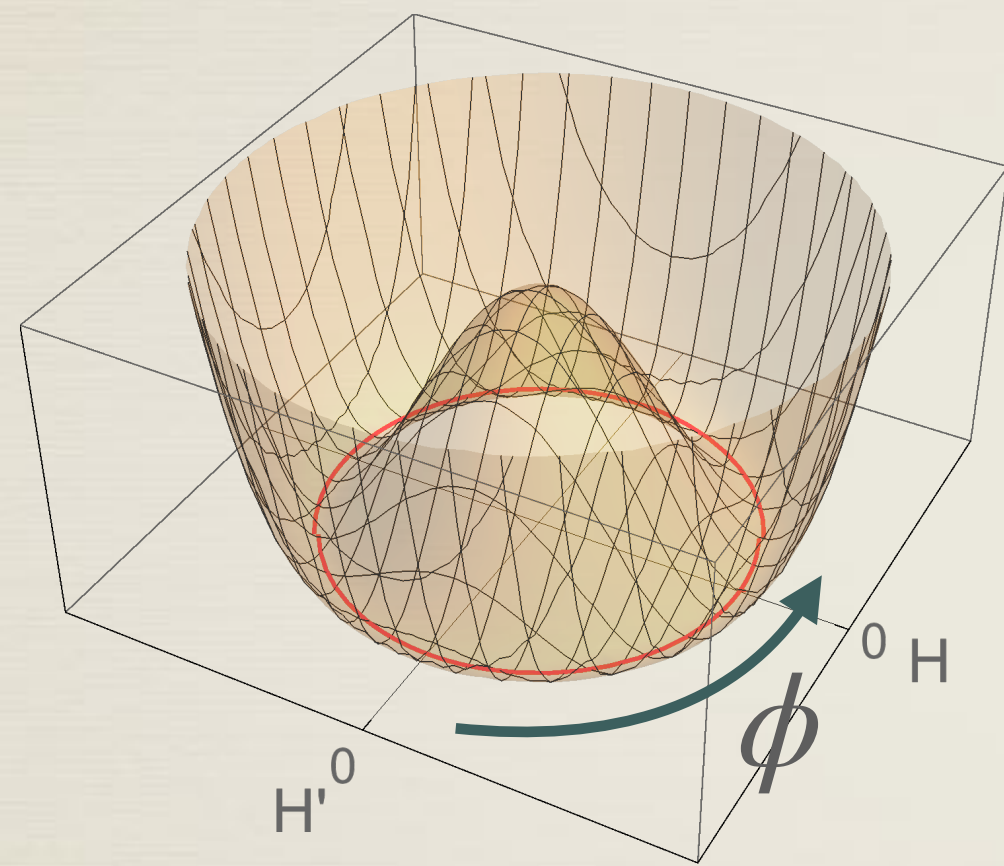
$$\langle H \rangle \ll \langle H' \rangle$$

This point can become a minimum by quantum corrections

Prediction on the quartic coupling

$$V \simeq \lambda(|H|^2 + |H'|^2 - v'^2)^2 + \text{small corrections}$$

Hall, KH (2018)



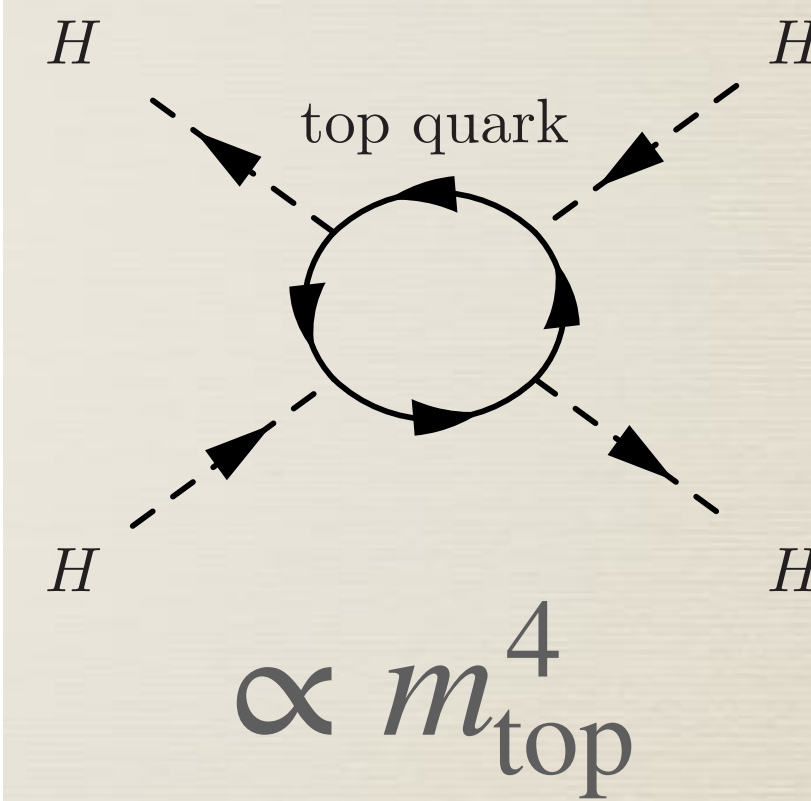
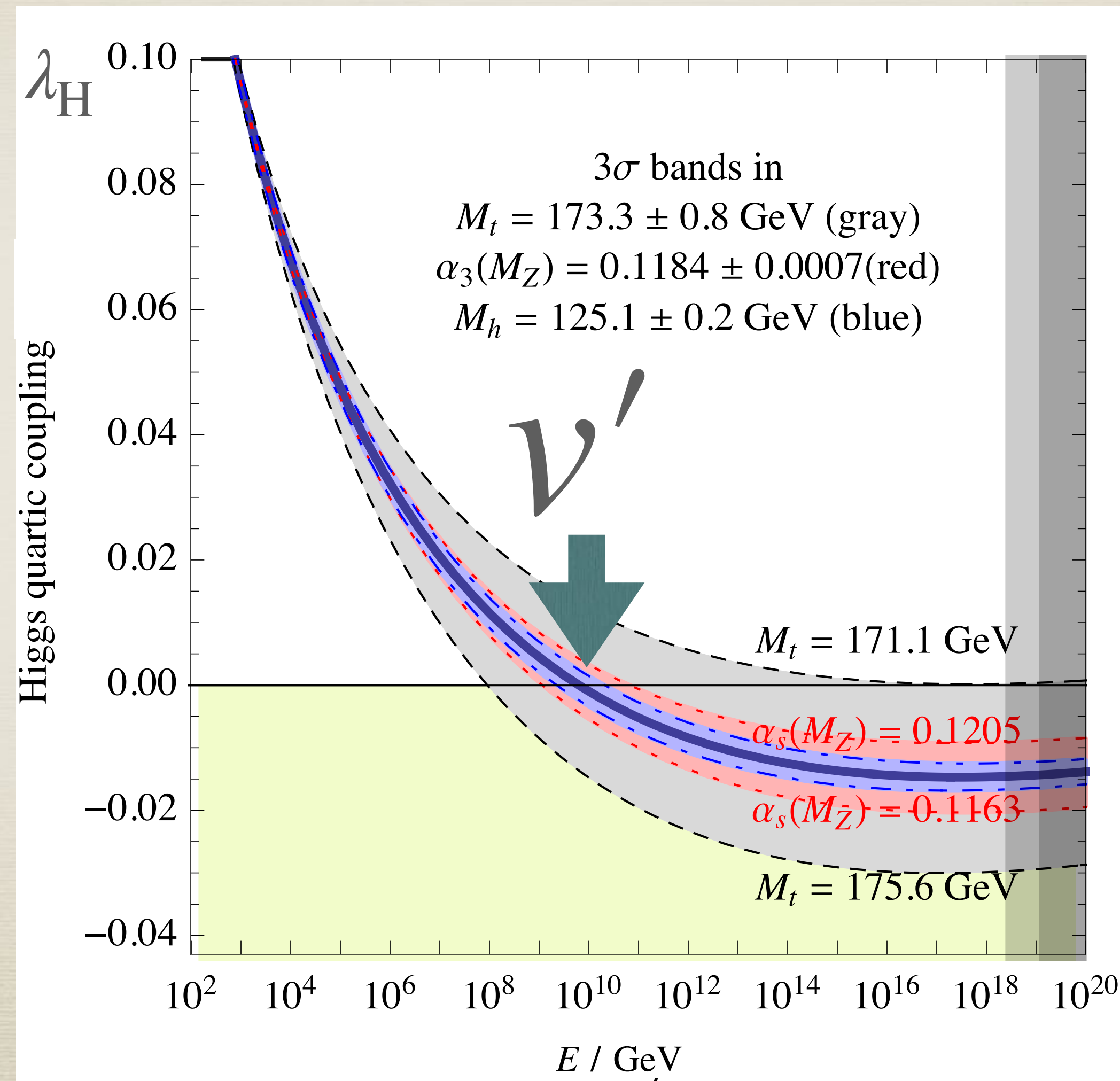
symmetry rotating the vector (H, H')

Standard Model Higgs is a (pseudo) Nambu-Goldstone boson associated with symmetry breaking by $\langle H' \rangle = v'$

$$\lambda_{\text{SM}}(v') \simeq 0$$


(up to calculable threshold correction)

Vanishing quartic coupling



Fine-tuning

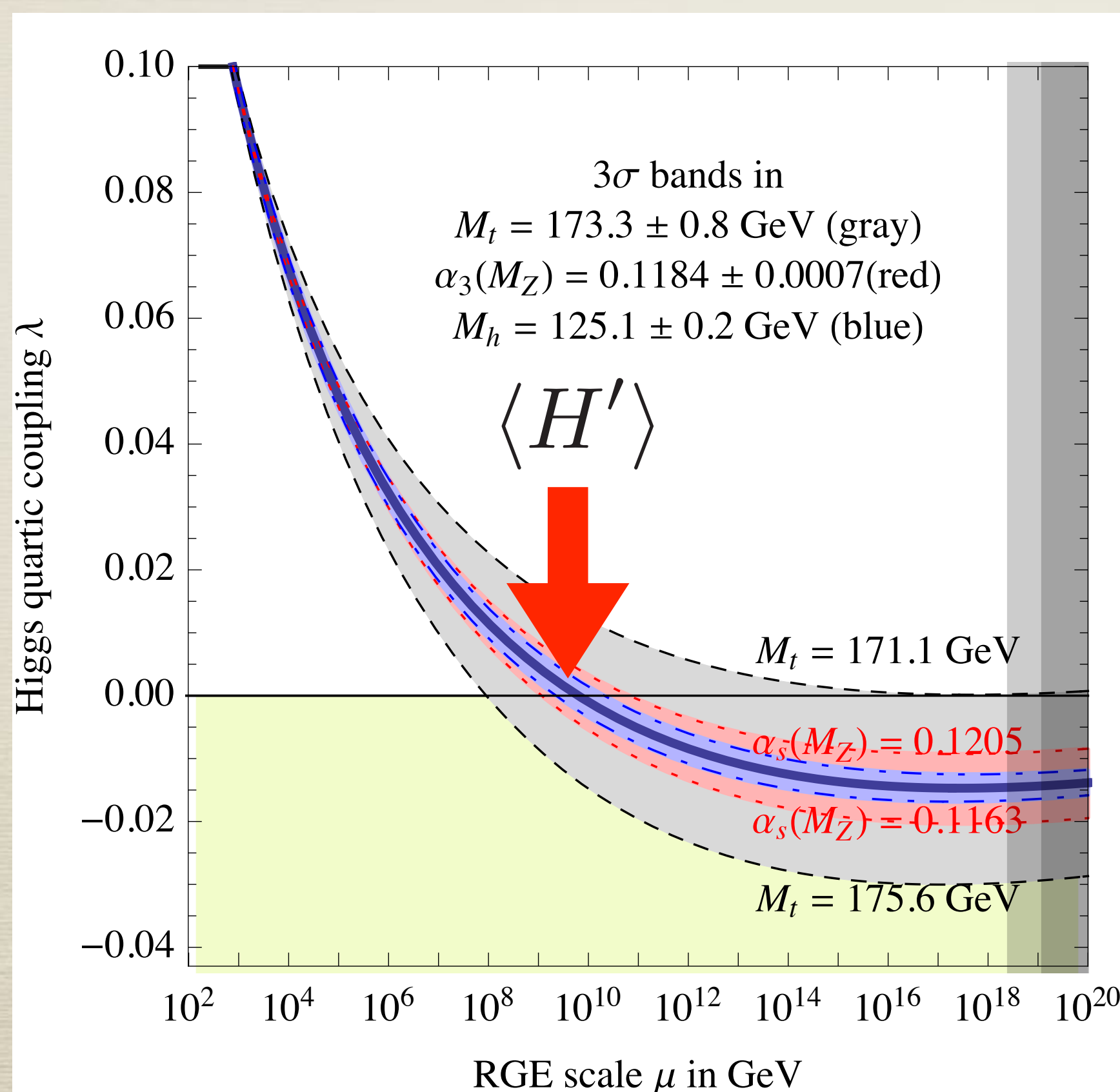
$$V = \lambda \left(|H|^2 + |H'|^2 - v'^2 \right)^2 + y |H|^2 |H'|^2$$


$$\frac{v'^2}{\Lambda_{\text{cut}}^2} \times \frac{v^2}{v'^2} = \frac{v^2}{\Lambda_{\text{cut}}^2}$$

Despite the intermediate scale v' ,
same as that of standard model

Generic Z_2

The scheme is applicable to generic models with the SM Higgs H and its Z_2 partner H'



Higgs Parity

Mirror Z_2 , Left-Right symmetry with or without space-time parity,

Precise measurement and Z_2

Hall and KH (2018, 2019)

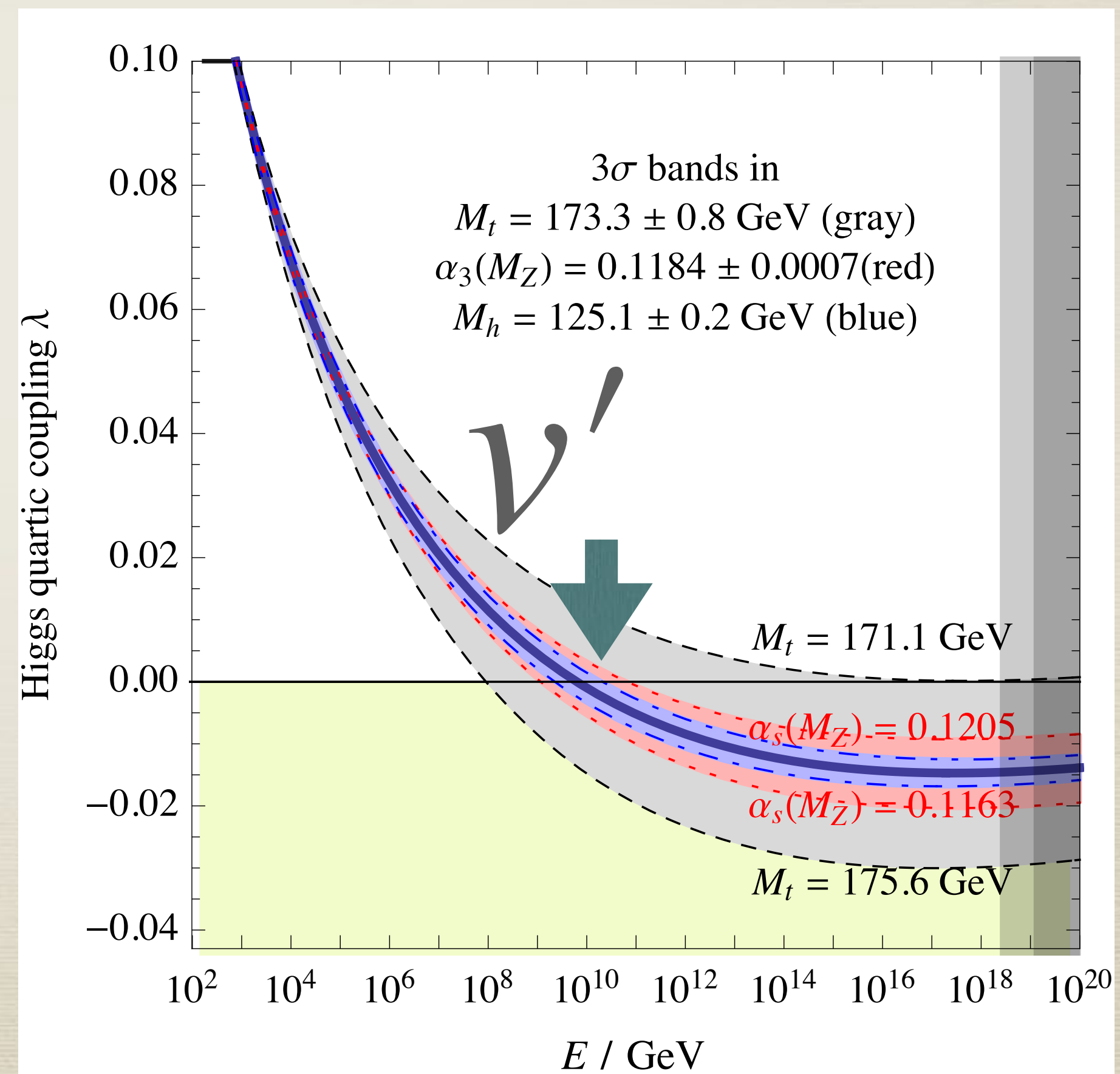
Dunsky, Hall and KH (2019, 2020, 2022)

Carrasco-Martinez, Dunsky, Hall and KH (2023)

top quark mass
Higgs mass
strong coupling constant



Z_2 symmetry
breaking scale



Precise measurement and new physics

Hall and KH (2018, 2019)

Dunsky, Hall and KH (2019, 2020, 2022)

Carrasco-Martinez, Dunsky, Hall and KH (2023)

top quark mass

Higgs mass

strong coupling constant



Z_2 symmetry
breaking scale



Dark matter and its detection,
gauge coupling unification
and proton decay, baryogenesis, ...

Precise measurement and new physics

Hall and KH (2018, 2019)

Dunsky, Hall and KH (2019, 2020, 2022)

Carrasco-Martinez, Dunsky, Hall and KH (2023)



Dark matter and its detection,
gauge coupling unification
and proton decay, baryogenesis, ...

2023 at Flatiron Institute

Z2 symmetries

* Mirror Z₂

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times (SU(3)_c \times SU(2)_L \times U(1)_Y)'$$

e' dark matter, dark radiation,
GWs from mirror QCD PT

* Left-Right symmetry

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

SO(10) unification, leptogenesis,
right-handed neutrino DM

* Left-Right Parity

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

above + solving strong CP problem

* Mirror Parity

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times SU(2)'_L \times U(1)'_Y$$

e' dark matter, solving strong CP problem

Z2 symmetries

* Mirror Z₂

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times (SU(3)_c \times SU(2)_L \times U(1)_Y)'$$

e' dark matter, dark radiation,
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SO(10) unification, leptogenesis,
right-handed neutrino DM

* **Left-Right Parity**

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

above + **solving strong CP problem**

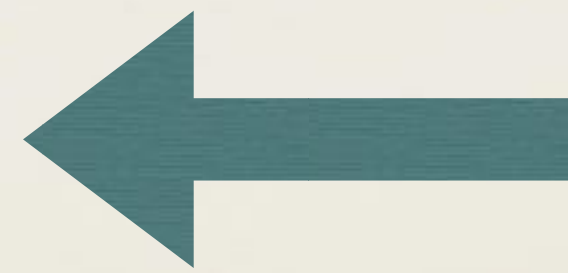
* Mirror Parity

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times SU(2)'_L \times U(1)'_Y$$

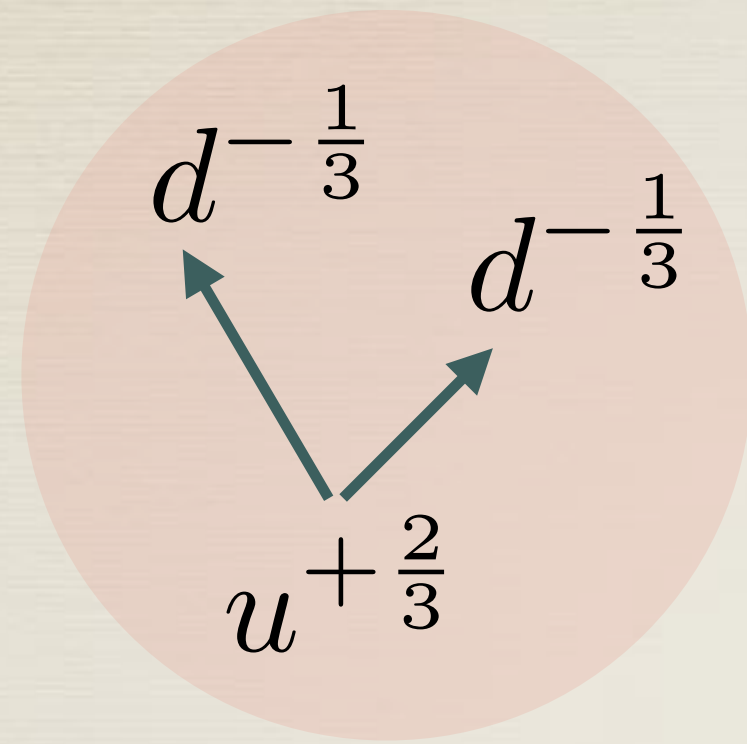
e' dark matter, solving strong CP problem

Outline

- * Z_2 symmetry of Higgs and its spontaneous breaking
- * Parity, $SO(10)$, and dark matter



The strong CP problem



Neutron Electric Dipole Moment

$$H = d_n \vec{E} \cdot \vec{S}$$

$$d_n/e \sim 0.1 \text{ fm} \sim 10^{-14} \text{ cm} ?$$

$$d_n/e < 2.9 \times 10^{-26} \text{ cm} \quad \text{Baker et.al (2006)}$$

Suggests CP symmetry

$$d_n \vec{E} \cdot \vec{S} \xrightarrow{\text{CP}} -d_n \vec{E} \cdot \vec{S} \quad \rightarrow \quad d_n = 0$$

CP is not preserved in SM

$$\mathcal{L} = y_{ij}^u H^\dagger q_i \bar{u}_j + y_{ij}^d H q_i \bar{d}_j + \frac{\theta_{\text{QCD}}}{32\pi^2} G\tilde{G}$$

CP violation

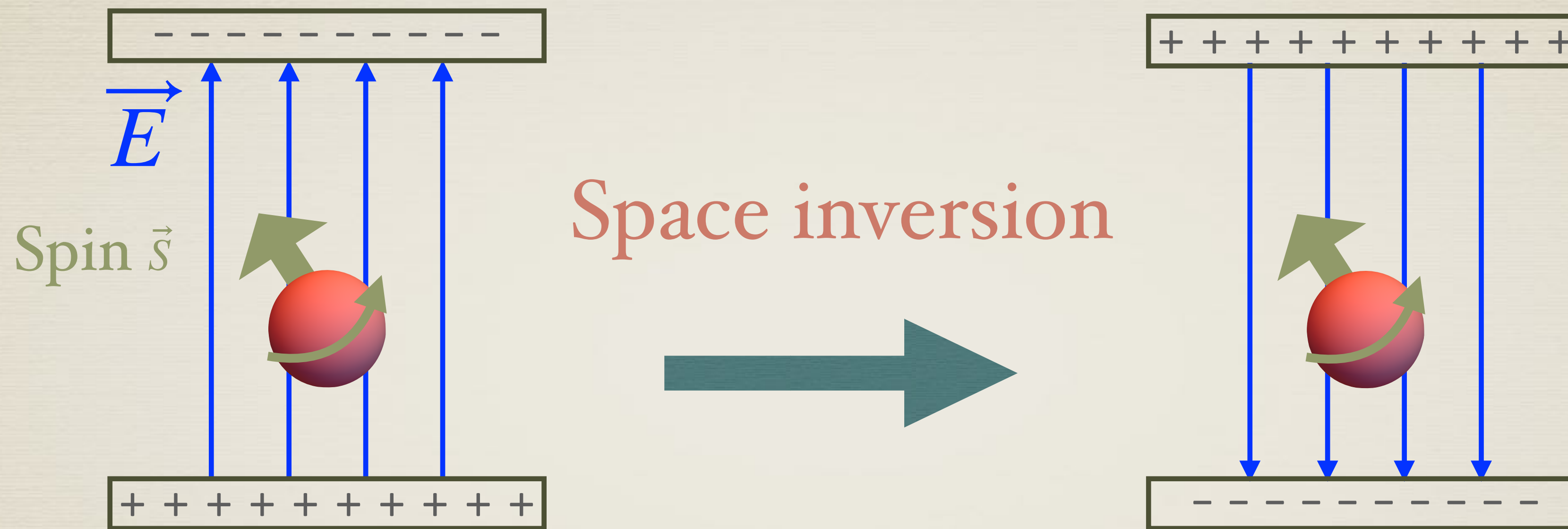
observed $\theta_{\text{CKM}} = O(1)$ \Rightarrow $y^{u,d}$ must have $O(1)$ complex phases
 \Rightarrow CP is violated

$$d_n/e \simeq 5 \times 10^{-16} \bar{\theta} \text{ cm} \quad \text{Crewther, Vecchia and Witten (1979)}$$

$$\bar{\theta} = \text{argdet}(m^u m^d) + \theta_{\text{QCD}} < 10^{-10}$$

The strong CP problem

Parity solution



$$~~H = d_n \vec{E} \cdot \vec{S}~~$$

Parity symmetry can forbid the neutron electric dipole moment

Left-Right Parity

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X \times P_{LR}$$

$q = (u, d)$	Parity	$\bar{q} = (\bar{u}, \bar{d})$
$(3, 2_L, 1, 1/6)$		$(\bar{3}, 1, 2_R, -1/6)$
$\ell = (\nu, e)$		$\bar{\ell} = (\bar{N}, \bar{e})$
H_L		H_R
$(1, 2_L, 1, -1/2)$		$(1, 1, 2_R, 1/2)$

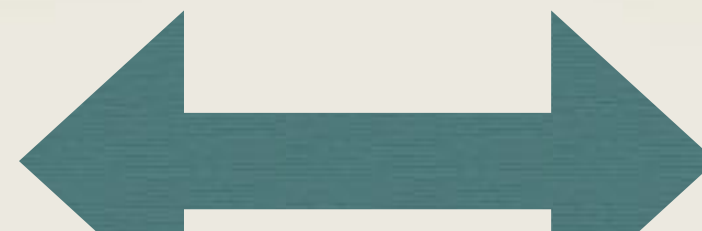
$$q(t, x) \leftrightarrow i\sigma_2 \bar{q}^*(t, -x)$$

Left-Right Parity

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X \times P_{LR}$$

$$q = (u, d)$$
$$(3, 2_L, 1, 1/6)$$

Parity



$$\bar{q} = (\bar{u}, \bar{d})$$
$$(\bar{3}, 1, 2_R, -1/6)$$

$$\bar{\ell} = (\bar{N}, \bar{e})$$
$$H_R$$

$$(1, 1, 2_R, 1/2)$$

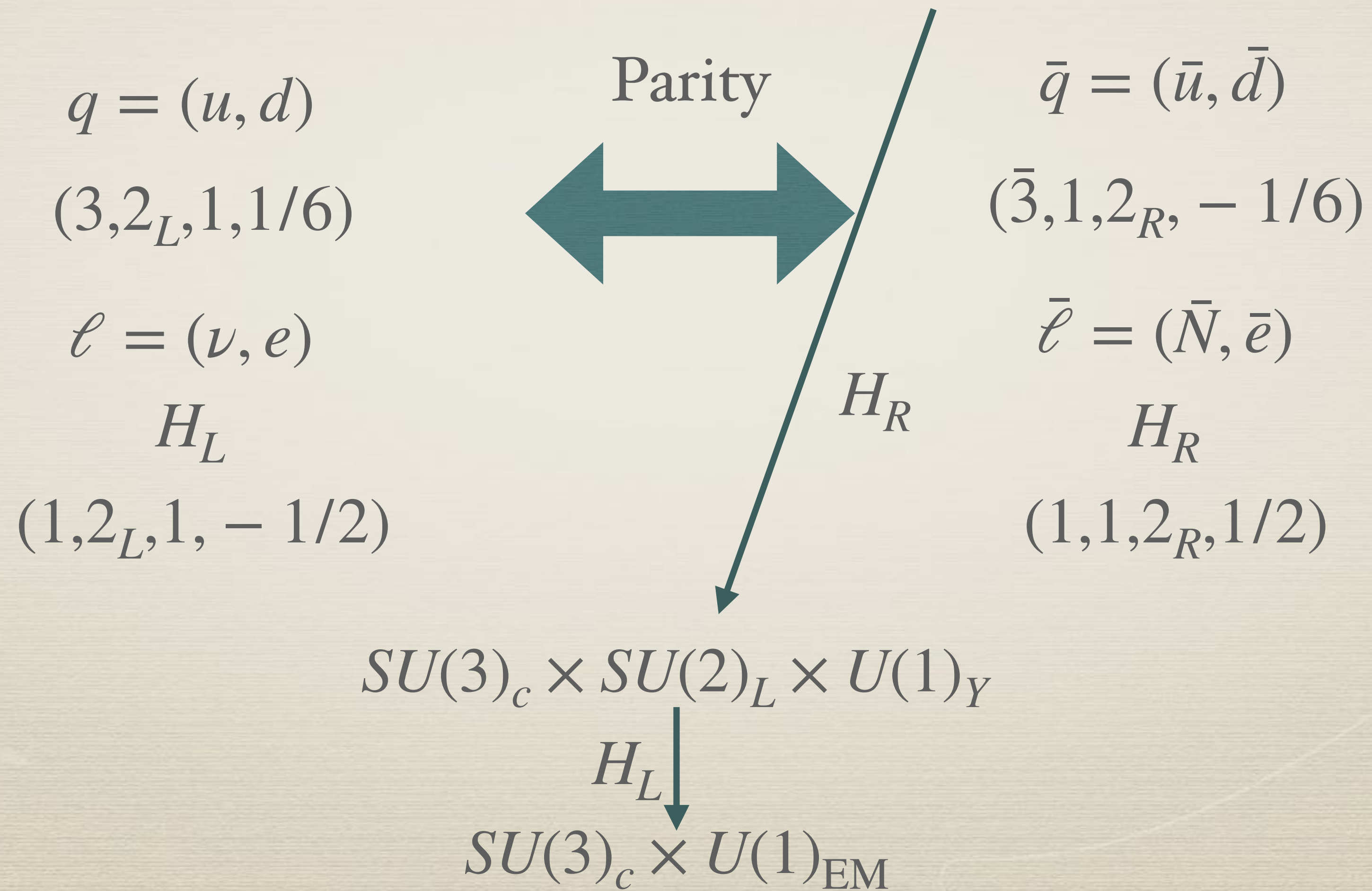
Mystery of missing
antimatter (2008)



Carrasco-Martinez, Dunsky, Hall and KH (2023)

Left-Right Parity

$$SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X \times P_{LR}$$



Left-Right in $SO(10)$

$SO(10)$ contains left-right symmetry without space-time inversion

$$\psi_{16} = (\ell, \bar{\ell}, q, \bar{q})$$

$$C_{LR} : q \leftrightarrow \bar{q}, \ell \leftrightarrow \bar{\ell}$$

$$H_{16} = (H_L, H_R, H_L^c, H_R^c)$$

$$SO(10) \xrightarrow{H_{45}} SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X$$

C_{LR} is spontaneously broken

Left-Right Parity in $SO(10)$

$$\text{Add CP: } \psi_{16}(t, x) \rightarrow i\sigma_2\psi_{16}^*(t, -x)$$

$$C_{LR} : q \leftrightarrow \bar{q}, \ell \leftrightarrow \bar{\ell}$$

$$(\text{CP and } C_{LR}) = P_{LR} : q(t, x) \leftrightarrow i\sigma_2\bar{q}^*(t, -x), \ell(t, x) \leftrightarrow i\sigma_2\bar{\ell}^*(t, -x)$$

$$SO(10) \times CP \quad \longrightarrow \quad SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_X \times P_{LR}$$

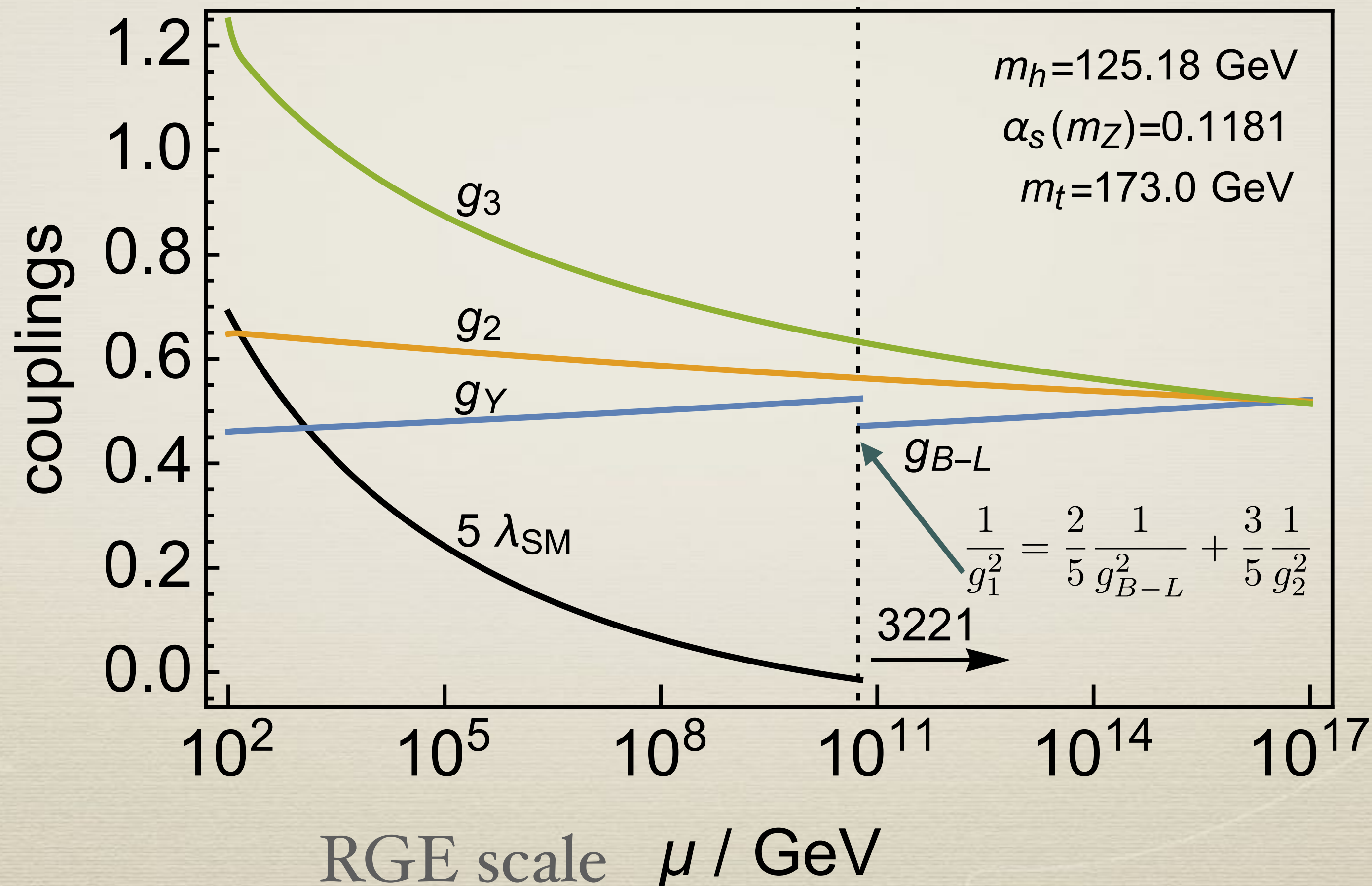
H_{45}

assuming H_{45} is CP odd

Coupling unification

Hall, KH (2018, 2019)

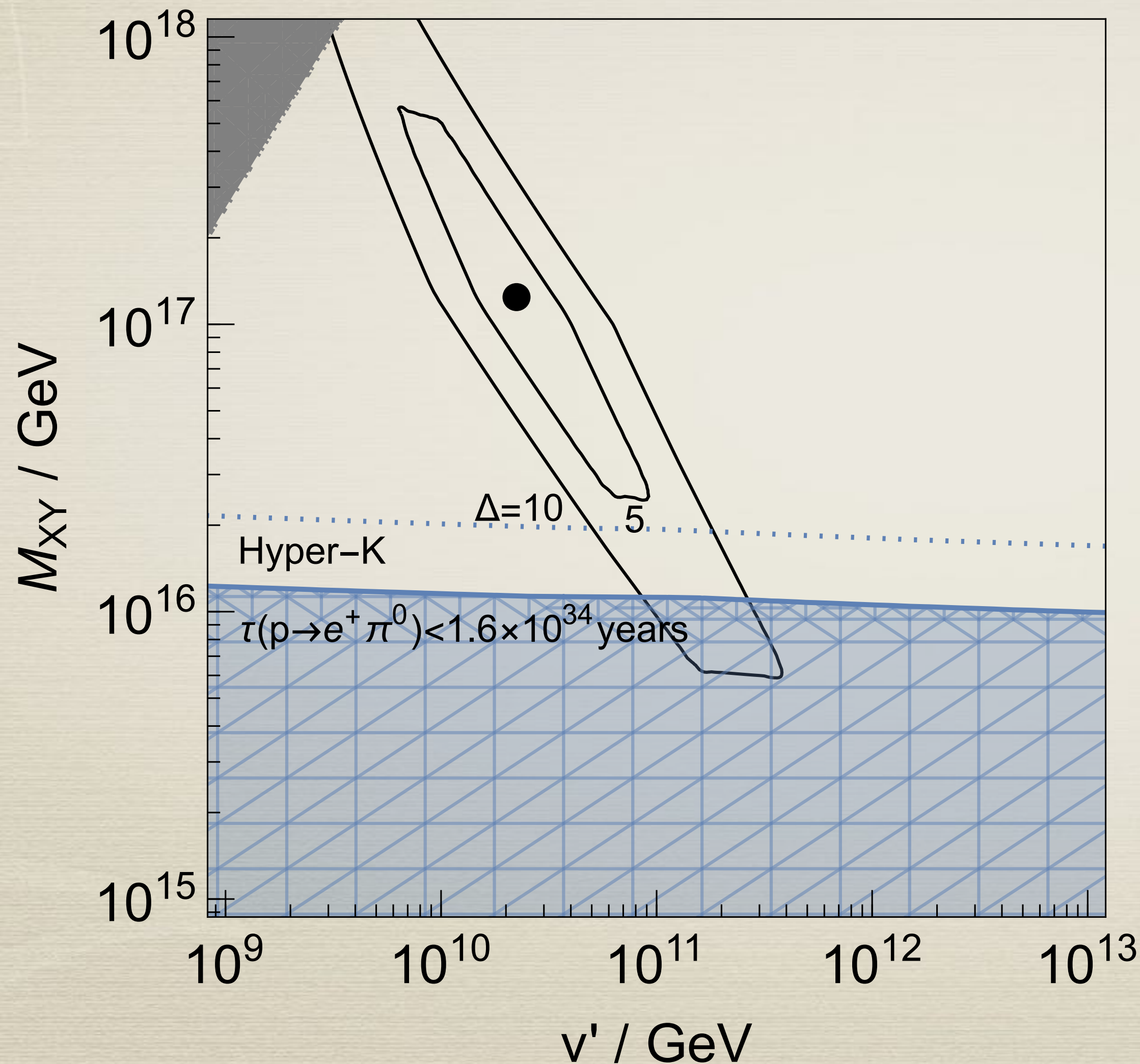
energy-dependent couplings



Quantify unification

Hall, KH (2019)

mass of new gauge boson
mediating proton decay



There can be quantum corrections from heavy particles around the GUT scale

$$\Delta = \max_{i,j} \left| \frac{8\pi^2}{g_i^2} - \frac{8\pi^2}{g_j^2} \right|$$

typically

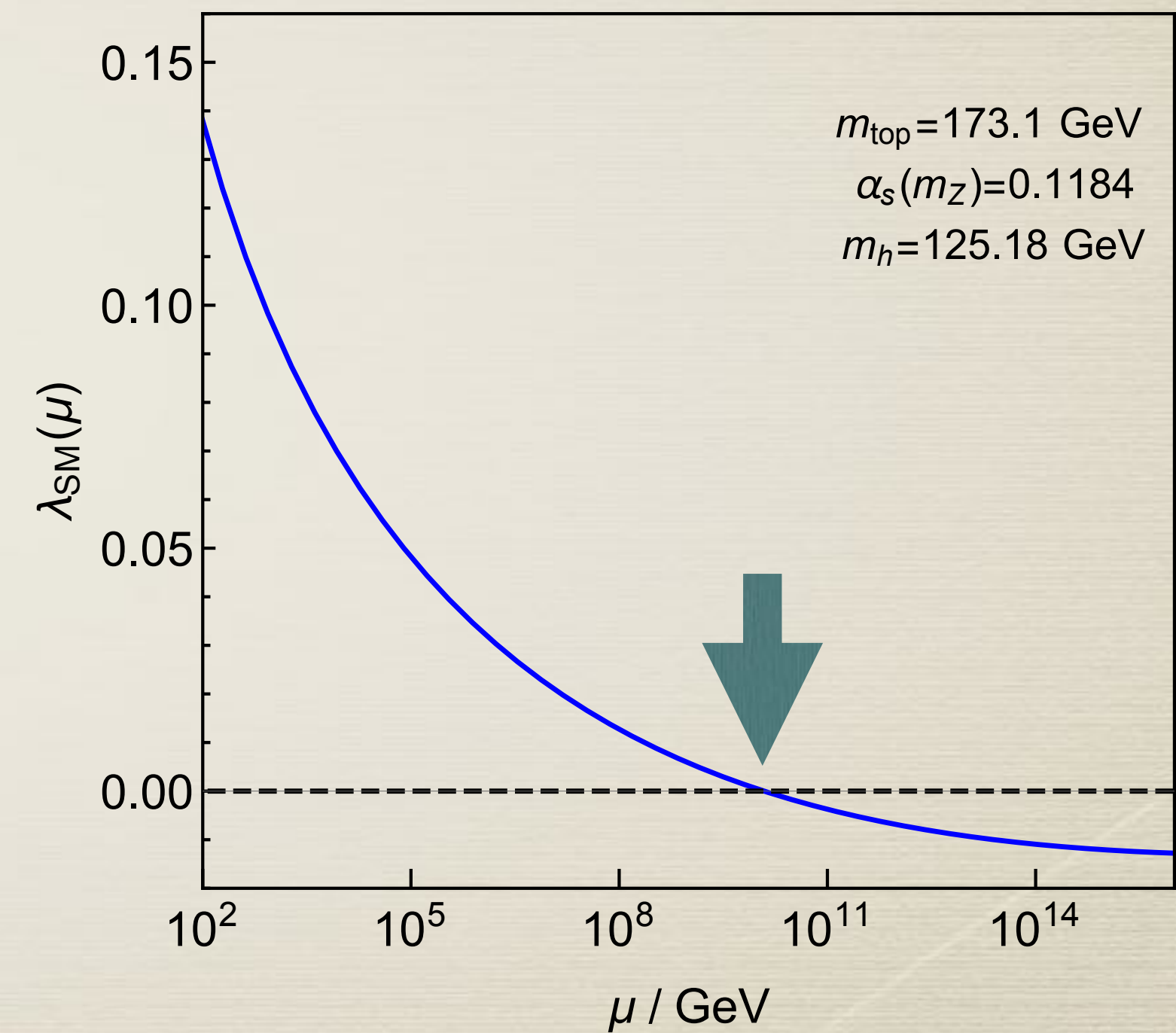
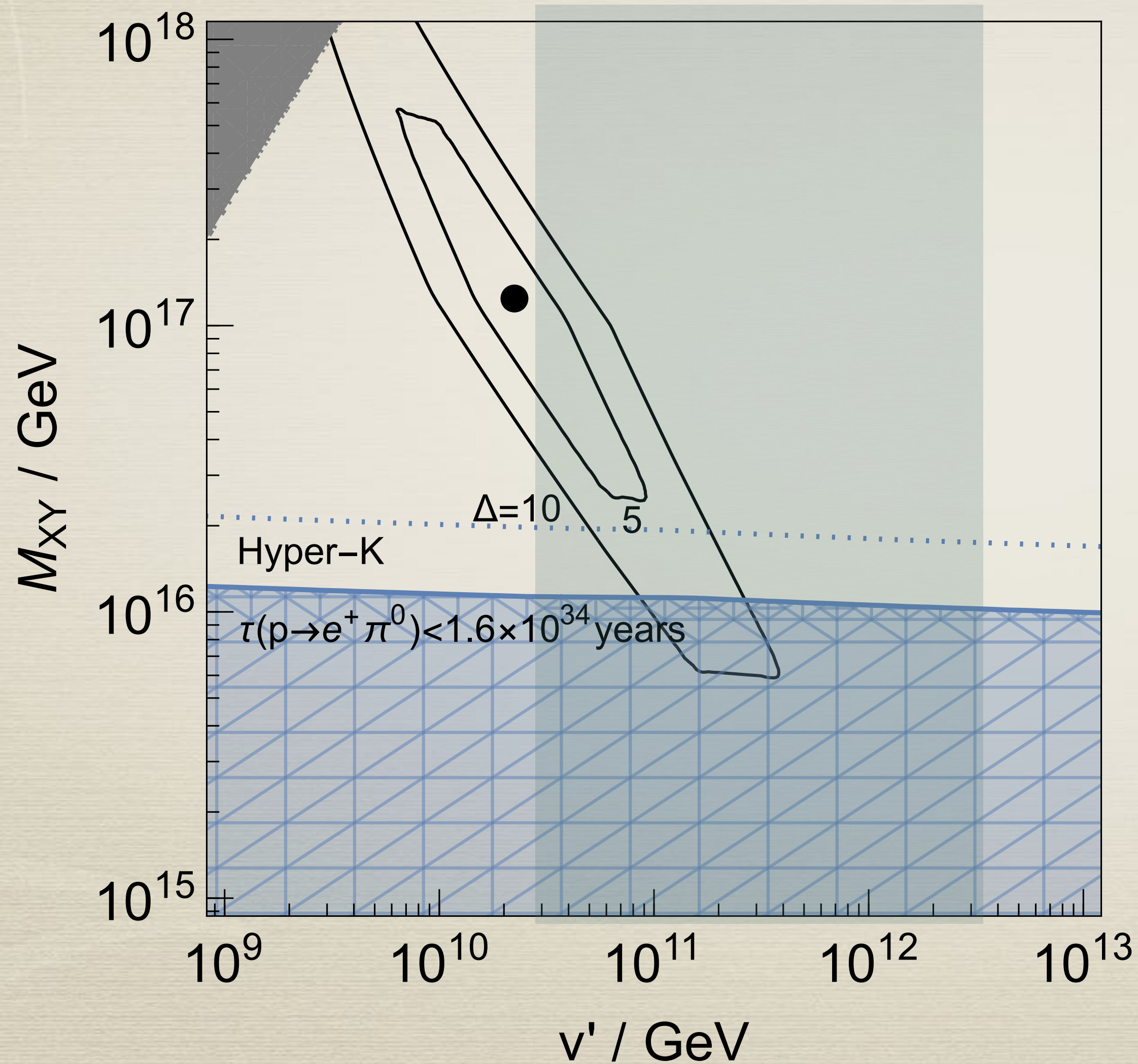
$$\Delta = \text{few} - 10$$

(smaller than SUSY GUT)

Quantify unification

Hall, KH (2019)

mass of new gauge boson
mediating proton decay



Dark Matter?

- * One of the right-handed neutrinos?

Difficult because of the up-neutrino unification

- * Dark sector?

- * Free massive scalar?

- * A new particle with non-trivial electroweak charge?

DM in 10

Let us add a fermion χ in 10 of $SO(10)$

	χ_D	$\chi_{\bar{D}}$	χ_L	$\chi_{\bar{L}}$
$SU(3)$	3	$\bar{3}$		
$SU(2)$			2	2
$U(1)$	$-1/3$	$1/3$	$-1/2$	$1/2$

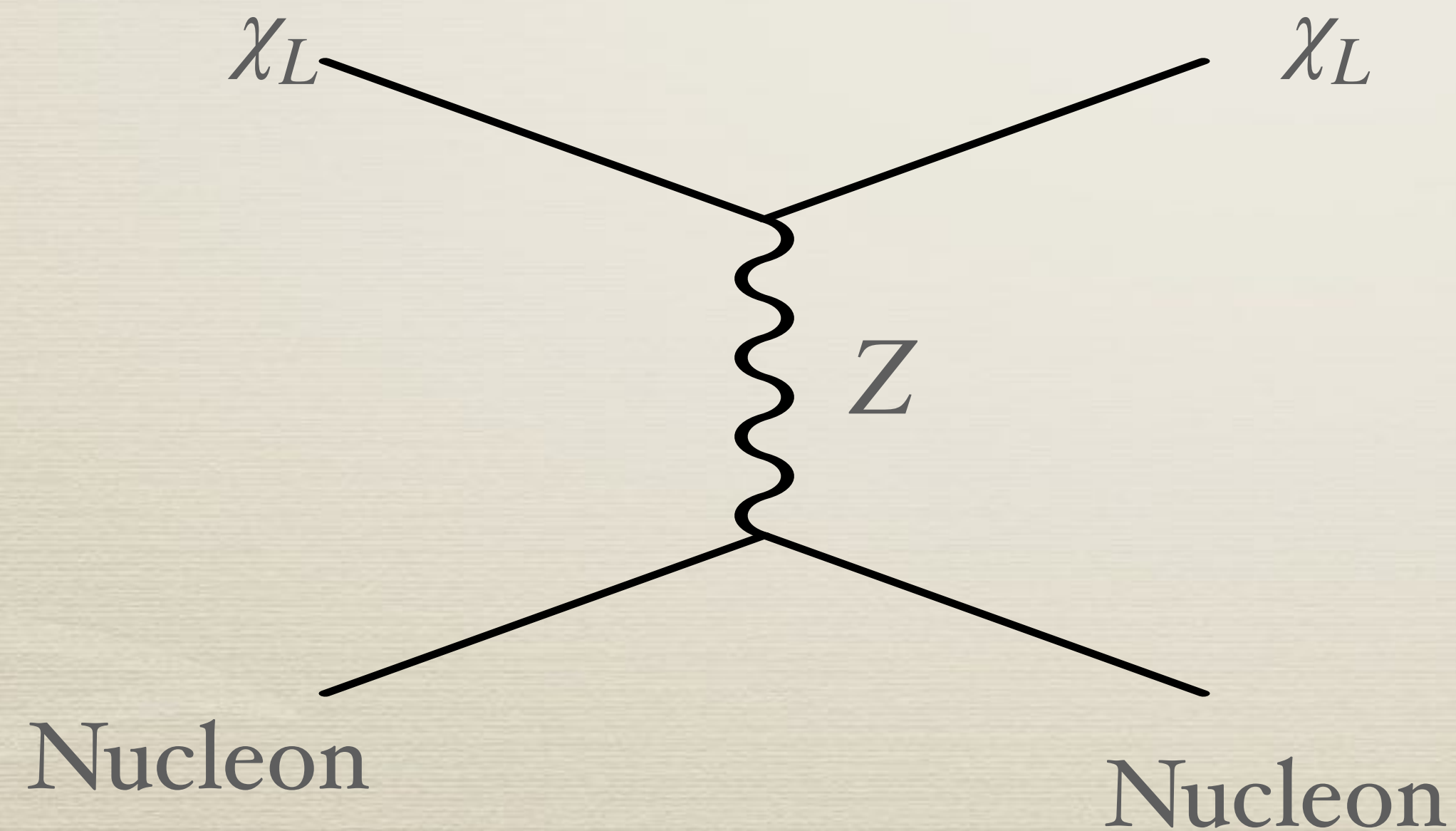


Higgsino-like DM

Higgsino-like DM

Freeze-out mechanism : $m_L \simeq 1 \text{ TeV}$

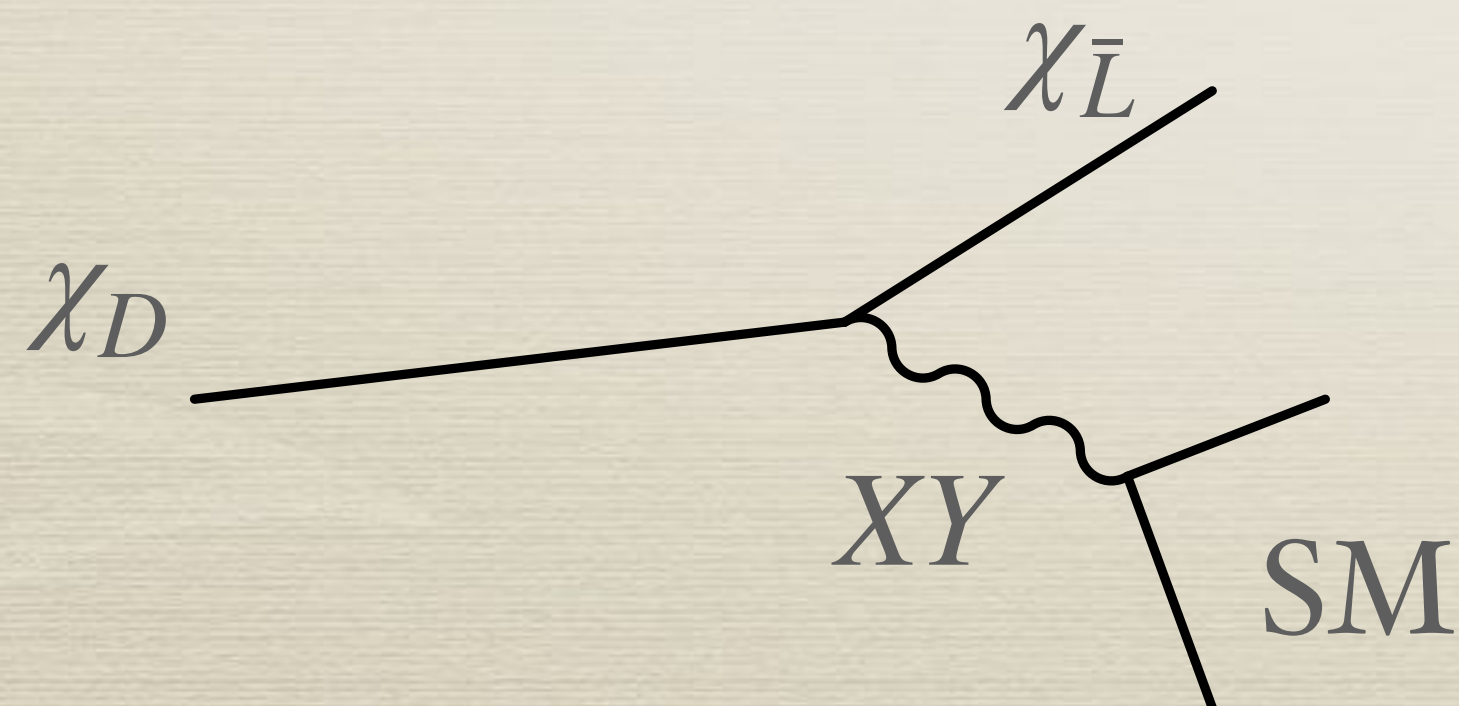
It should not remain a Dirac fermion
in order to avoid direct detection bound



Add a mixing with a singlet
that has a Majorana mass

Colored partner

	χ_D	$\chi_{\bar{D}}$	χ_L	$\chi_{\bar{L}}$
$SU(3)$	3	$\bar{3}$		
$SU(2)$			2	2
$U(1)$	$-1/3$	$1/3$	$-1/2$	$1/2$



The lifetime is longer
than the age of the universe
if $m_D \sim \text{TeV}$

Exotic stable relic

χ_D annihilation freezes-out around $T \sim 0.1m_D$

They re-annihilate around the QCD PT

$$\chi_D \bar{u} + \chi_{\bar{D}} u \rightarrow SM$$

$$\chi_D \chi_D u + \chi_D \bar{u} \rightarrow \chi_D \chi_D \chi_D + SM$$

Majority of them end up in $\chi_D \chi_D \chi_D$, as $\chi_D \bar{u}$ etc has large radius

$$\Omega_{\chi_D \chi_D \chi_D} h^2 \sim 0.1 \left(\frac{m_D}{10 \text{ TeV}} \right)^2 \quad \text{De Luca et.al (2018)}$$

Charged particles with such large abundance is excluded by heavy isotope searches

Splitting by missing VEV

$$\langle H_{45} \rangle = -iv_{45} \times \begin{pmatrix} \sigma_2 & 0 & 0 & \\ 0 & \sigma_2 & 0 & 0_{4 \times 6} \\ 0 & 0 & \sigma_2 & \\ & 0_{6 \times 4} & & 0_{4 \times 4} \end{pmatrix}$$

$SU(2)_R$ symmetry

$$\chi^{1-6} : \chi_{D, \bar{D}}, \chi^{7-10} : \chi_{L, \bar{L}}$$

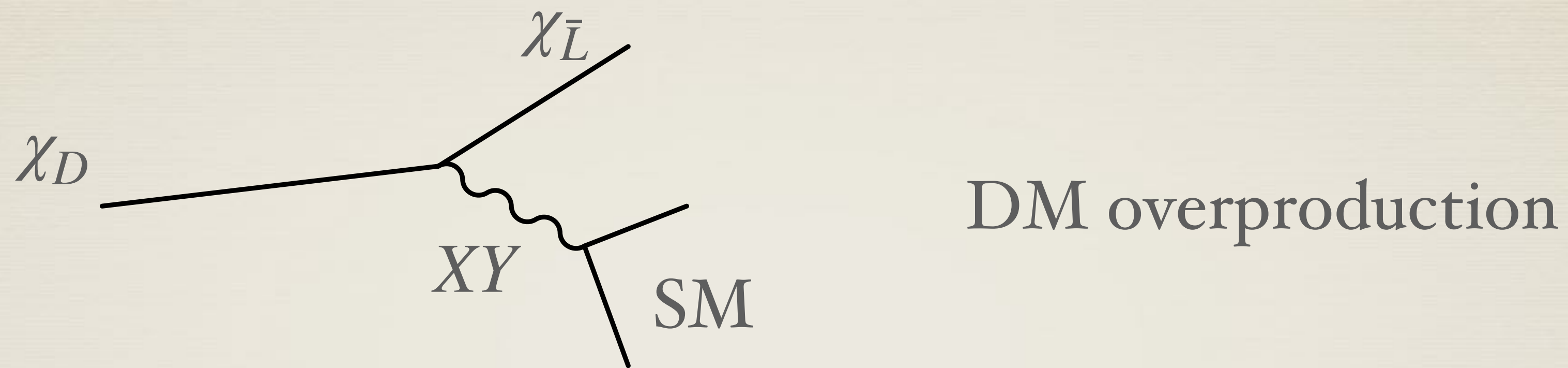
See Dimopoulos and Wilczek (1981)
for models with SUSY protection

By giving mass to χ by H_{45} maybe large mass splitting can be achieved

Baldwin and KH (2024)

(Models with small enough quantum corrections need one more 10-plet)

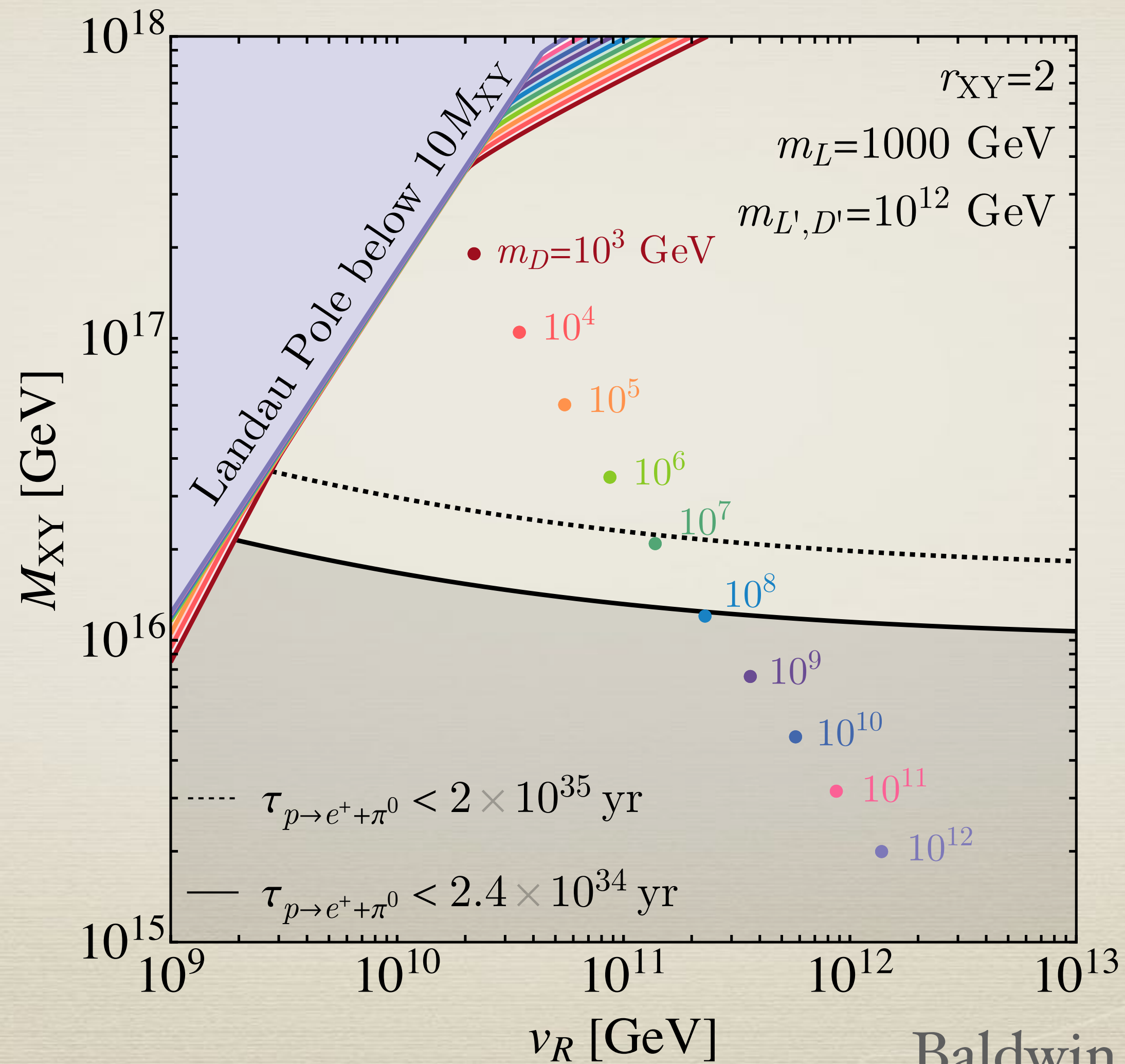
Lower bound on the mass



$$\frac{\rho_{\text{DM}}}{s} \simeq 0.4 \text{ eV} \left(\frac{m_L}{100 \text{ GeV}} \right)^3 \frac{0.05 \text{ GeV}}{T_{\text{dec}}}$$

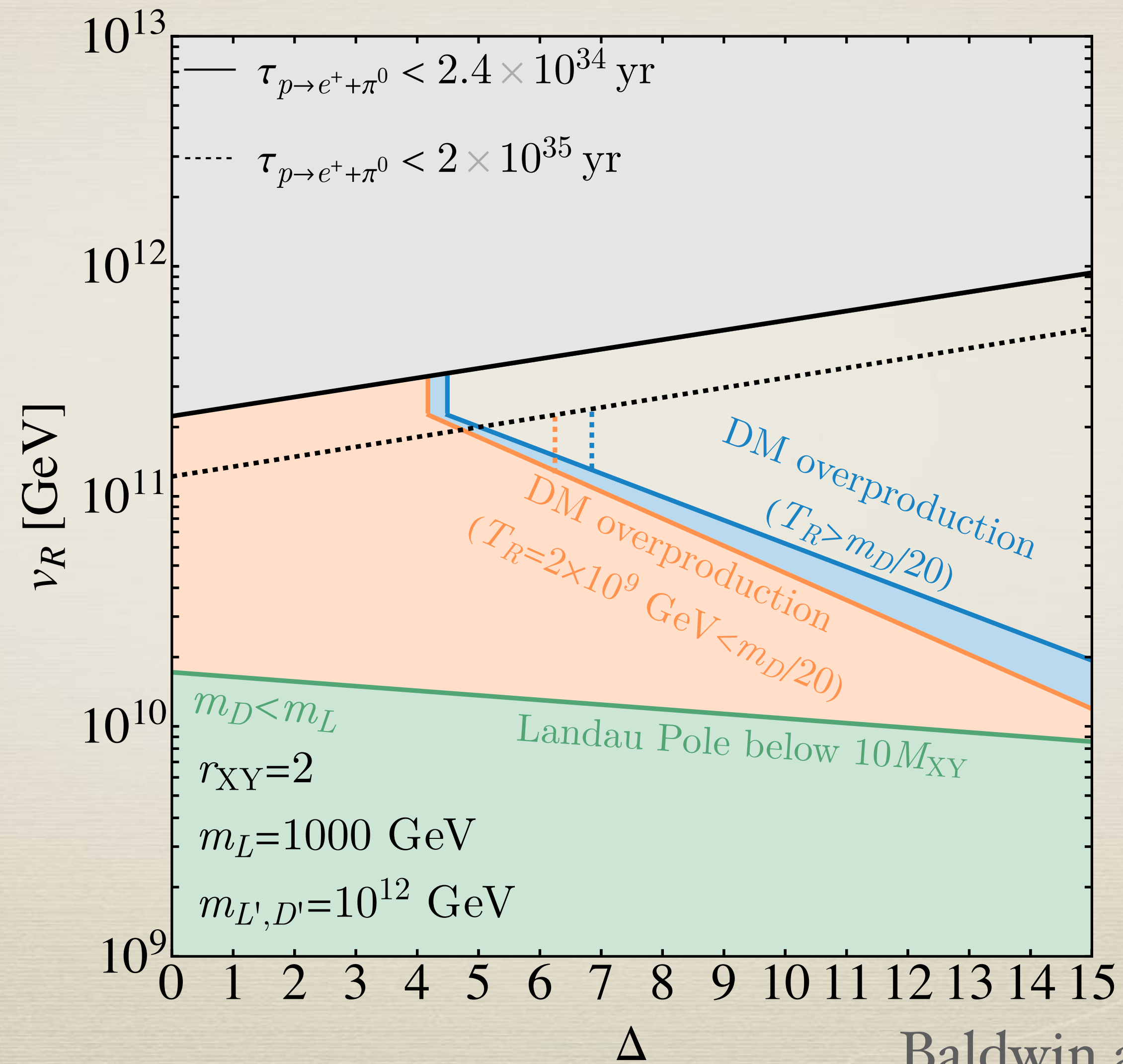
$$m_D > 3 \times 10^9 \text{ GeV} \times \left(\frac{m_{XY}}{10^{16} \text{ GeV}} \right)^{4/5} \left(\frac{m_L}{100 \text{ GeV}} \right)^{6/5}$$

Gauge coupling unification

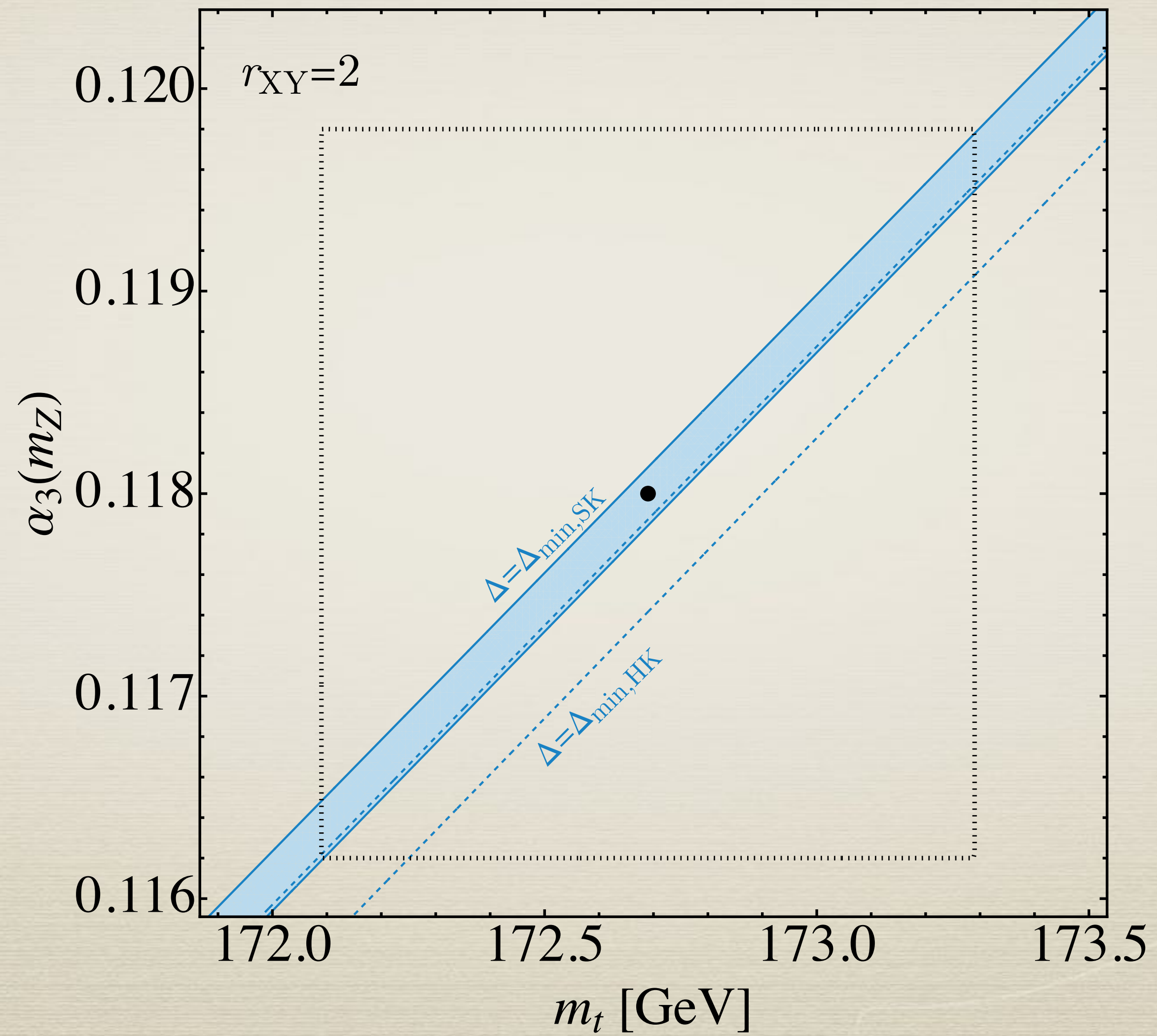


Baldwin and KH (2024)

Constraints



SM parameters



Summary

- * Z_2 symmetry of Higgs is phenomenologically interesting
- * The breaking scale of a class of Z_2 symmetry, Higgs Parity, is correlated with the SM parameters
- * Combined with space-time parity, the strong CP problem can be solved
- * Implications to $SO(10)$ unification, dark matter detections, etc.

Back up

Fermion masses

$$xH_{16}\psi_{16}X + \frac{1}{2}m_X X^2$$

X : 10 for down and electron, 45 or 54 for up and neutrino

CKM phase

$$xH_{16}\psi_{16}X + \frac{1}{2}m_X X^2$$

CP requires that x and m_X are real

$$i\frac{H_{45}}{M}H_{16}\psi_{16}X + \frac{1}{2}i\lambda H_{45}X^2$$

Setup

$\chi^a \chi^b H_{45}^{ab}$ identically vanishes

$$\chi^a \chi^b H_{45}^{ac} H_{45}^{bc} \longrightarrow \chi^a \chi^a H_{45}^{bc} H_{45}^{bc}$$

quantum
correction

So let us introduce two fermions : χ_1 and χ_2

Setup

$$\lambda \chi_1^a \chi_2^b H_{45}^{ab} + \frac{1}{2} M_2 \chi_2^a \chi_2^a$$

For example, for $\lambda v_{45} < M_2$,

$$\chi_2 \text{ ————— } M_2$$

$$\chi_{D_1} \text{ ————— } \lambda^2 v_{45}^2 / M_2$$

$$\chi_{L_1} \text{ ————— } 0$$

Corrections

$$\lambda \chi_1^a \chi_2^b H_{45}^{ab} + \frac{1}{2} M_2 \chi_2^a \chi_2^a$$

The first term preserve Z_4 under which $\chi_{1,2}(1), H_{45}(2)$

Ignoring the second terms,
corrections always involve odd number of H_{45}

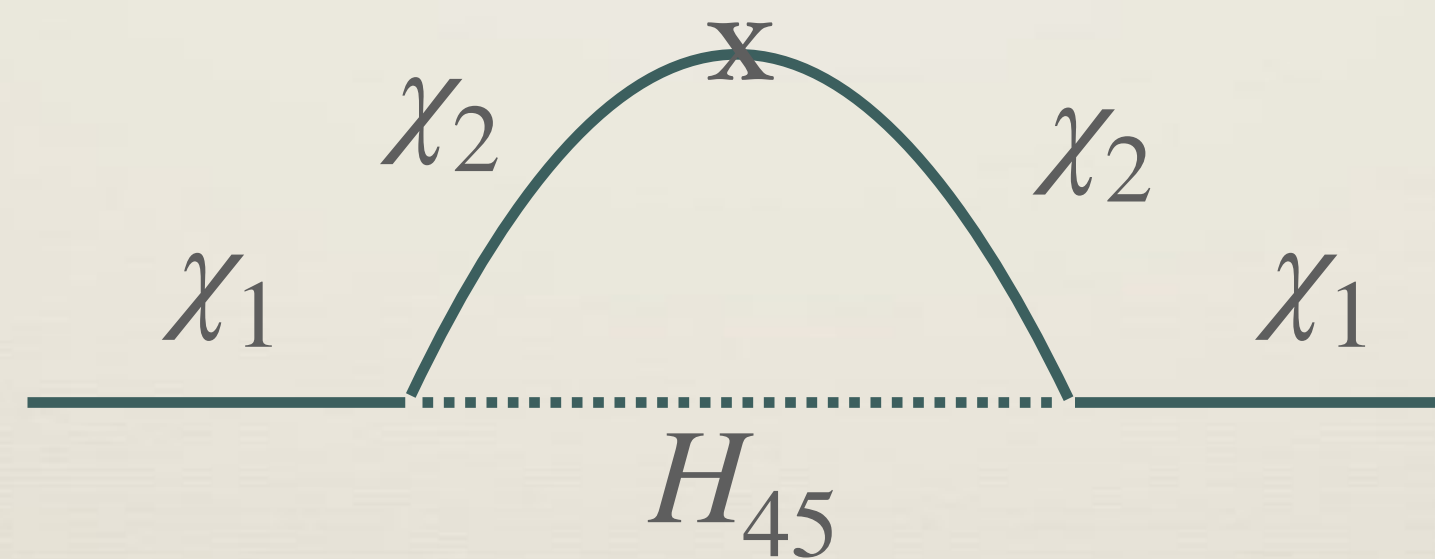
Two of the indices of H_{45} s must be contracted with χ

Correction to χ_L mass vanishes

Corrections

$$\lambda \chi_1^a \chi_2^b H_{45}^{ab} + \frac{1}{2} M_2 \chi_2^a \chi_2^a$$

Using the second term,



$$\delta m \sim \frac{1}{16\pi^2} \lambda^2 m_2 = 100 \text{ GeV} \left(\frac{\lambda}{10^{-4}} \right)^2 \frac{M_2}{10^{12} \text{ GeV}}$$

Corrections

$$\langle H_{45} \rangle = -iv_{45} \times \begin{pmatrix} \sigma_2 & 0 & 0 & \\ 0 & \sigma_2 & 0 & 0_{4 \times 6} \\ 0 & 0 & \sigma_2 & \\ & 0_{6 \times 4} & & 0_{4 \times 4} \end{pmatrix}$$

$SU(2)_R$ symmetry

The vev should be non-zero after $SU(2)_R$ symmetry breaking

Indeed, $H_{45}H_{45}H_{16}H_{16}^*$ gives

$$\frac{\delta v}{v_{45}} \sim \frac{v_R^2}{v_{45}^2}$$