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Non-perturbative techniques for constraining the cosmological collider

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Upcoming large-scale structure (LSS) surveys will provide significant insight into the physics of the early Universe by searching for primordial non-Gaussianity (PNG). Whereas traditional approaches towards constraining PNG with LSS typically focus on phenomenological amplitudes, e.g., $f_{\rm NL}^{\rm loc.}$, which characterize a given "shape" of non-Gaussianity, there exist a range of inflationary scenarios that can produce unique signatures of PNG. A particularly compelling example is the cosmological collider scenario, wherein the presence of massive spinning particles during inflation can lead to poles or oscillations in the squeezed bispectrum. In this talk, I will present a novel approach towards running and analyzing N-body simulations with initial conditions generated for cosmological collider type models. Using these simulations, I will validate a model for the non-linear squeezed matter bispectrum and collapsed trispectrum for the collider physics scenario based on the LSS consistency relations – non-perturbative statements about the structure of LSS correlation functions and non-perturbative models can be used to develop pipelines to directly constrain the masses and spins of particles present during inflation.

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