Dark Photons Review

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290e Dark Matter Seminar

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Basic Assumptions

Suppose there is an additional U(1)' in nature

$$\mathcal{L}_{U(1)'} = -\frac{1}{4}X^{\mu\nu}X_{\mu\nu} + \frac{1}{2}m_{A'}^2A'_{\mu}A^{'\mu}$$

- Mass can come from Higgs or Stuckelberg mechanism
- Assume SM particles not charged under new gauge group
- ▶ However there would necessarily be kinetic mixing between $X^{\mu\nu}$ and $F^{\mu\nu}$ (satisfies gauge invariance) [Holdom '86]

$$\mathcal{L} \supset \frac{-1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{1}{4} X^{\mu\nu} X_{\mu\nu} + \frac{1}{2} m_{A'}^2 A'_{\mu} A^{\prime\mu} - \frac{\epsilon}{2} F^{\mu\nu} X_{\mu\nu} + J_{\mu} A^{\mu}$$

• This is renormalizable, dimension 4 operator \Rightarrow relevant at all energies

Kinetic Mixing

Can remove kinetic mixing term by diagonalizing, changes the mass eigenstates and interactions of the SM vector bosons.

In mass basis (A^µ → A^µ − εX^µ), A' couples directly to EM charged particles:

$$\mathcal{L} \supset \frac{-1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{1}{4} X^{\mu\nu} X_{\mu\nu} + \frac{1}{2} m_{A'}^2 A'_{\mu} A^{'\mu} + J^{\mu} (A_{\mu} - \epsilon A'_{\mu})$$

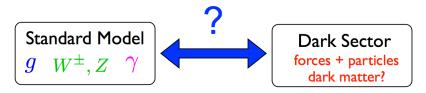
▶ In interaction basis $(X^{\mu} \mapsto X^{\mu} - \epsilon A^{\mu})$ there is $A' \leftrightarrow \gamma$ oscillation due to small mass-mixing:

$$\mathcal{L} \supset \frac{-1}{4} F^{\mu\nu} F_{\mu\nu} - \frac{1}{4} X^{\mu\nu} X_{\mu\nu} + \frac{1}{2} m_{A'}^2 A'_{\mu} A^{\prime\mu} + J^{\mu} A_{\mu} - \epsilon m_{A'}^2 A_{\mu} A^{\prime\mu}$$

▶ Two unknown parameters $\{m_{A'}, \epsilon\}$

Dark sector consists of particles that do not couple to known SM forces

- Can arise naturally from high energy physics including string theory constructions, low-scale supersymmetry models
- Dark sector need not lie at the weak-scale, and may harbor dark matter! (Although not motivated by just DM)



Dark sector could have rich structure, so what non-gravitational portals exist between dark sector and SM? Only a few possibilities

• Axion:
$$\frac{a}{f_a}F^{\mu\nu}\tilde{F_{\mu\nu}} \Rightarrow axion-like particles (ALPs)$$

Vector:
$$\frac{\epsilon}{2}F^{\mu\nu}X_{\mu\nu} \Rightarrow \text{dark photon } A^{\mu\nu}$$

• Higgs: $\epsilon_h |h|^2 |\phi|^2 \Rightarrow \text{dark scalar (exotic Higgs decays)}$

• Neutrino:
$$\epsilon_{\nu}(hL)\psi \Rightarrow$$
 sterile neutrino

A' could be messenger to the dark sector, could even constitute dark matter \Rightarrow strong theoretical arguments for searching entire $\{m_{A'}, \epsilon\}$ parameter space.

Generating $\epsilon, m_{A'}$

Consider loop diagrams of heavy fields charged under both photon and A'. Simple naturalness arguments $\Rightarrow \epsilon \sim \frac{eg_D}{6\pi^2} \log(\frac{m}{\Lambda}) \sim 10^{-8} - 10^{-2}$

[Essig, Schuster, Toro '09]



However, non-perturbative and large-volume effects in string theory constructions generate much smaller ϵ . No clear minimum, but generally $\epsilon\sim 10^{-12}-10^{-2}$ is predicted [Goodsell, Jaeckel, Redondo, Ringwald '09]

 $m_{A'}$ has a much larger parameter space to be explored, $m_{A'}\sim 2m_e\sim~{\rm MeV}$ is a natural dividing line

Very well-motivated portion of $m_{A'}$ parameter space with interesting phenomenology and experimental avenues.

- ► Can decay to quarks, charged leptons (e⁺e⁻, μ⁺μ⁻, π⁺π⁻) either directly or indirectly (light dark sector states?)
- ▶ Generated by Higgs mechanism in context of supersymmetry, various models for communicating weak-scale SUSY-breaking to dark sector give rise to $m_{A'} \sim \sqrt{\epsilon}m_W$

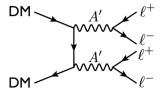
[Arkani-Hamed, Weiner '08; Cheung, Ruderman, Wang, Yavin '09; Morrissey, Poland, Zurek '09]

 $m_{A'} \sim \text{MeV} - \text{GeV}$

 \blacktriangleright May explain $(g-2)_{\mu}$ anomaly $\sim 3\sigma, 10^{-9}$ deviation from SM

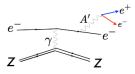
$$(g-2)^{A'}_{\mu} \sim \frac{\alpha}{2\pi} \times \epsilon^2, \quad m_{A'} < m_{\mu}$$

▶ If DM (~ TeV) coupled to A' in this range, would produce ~ GeV cosmic-ray electrons, positrons. Implications for AMS, PAMELA, Fermi?



Array of fixed-target experiments underway (APEX, HPS, DarkLight, ...)

- ▶ Intense beam of charged particles (e^-, p) shot onto a block or a foil
- A' produced forward and carries most of E_{beam} , $\sigma \sim \frac{\epsilon^2 Z^2}{m_{ef}^2}$
- ▶ Invariant mass distribution of e^+, e^- pairs has clear peak at $m_{A'}$
- ▶ Small ϵ , long-lived $A' \Rightarrow$ displaced vertices

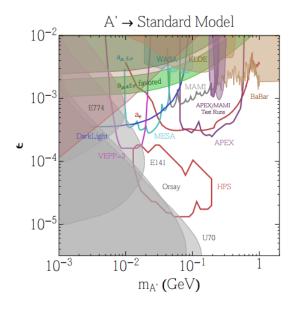


[Bjorken, Essig, Schuster, Toro '09]

Can also use existing e^+e^- collider experiments (BaBar, BELLE,...)

▶ Need high intensity, low-energy ($\sim 1 - 10 \, {
m GeV}$)

Current Reach



[Essig, et al. '13]

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Equally interesting parameter space, dark photons are a valid DM candidate! (Recall very light DM need to be produced non-thermally)

 Possibilities for generating correct relic density: Misalignment mechanism (works well for axions), inflationary fluctuations

[Nelson, Scholtz; Graham, Mardon, Rajendran]

$$\Omega \sim 0.3 \sqrt{\frac{m_{A'}}{\mathrm{keV}}} \left(\frac{H_i}{10^{12}\,\mathrm{GeV}}\right)$$

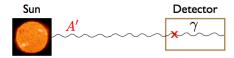
- Stability on cosmological scales $(A' \rightarrow 3\gamma)$ and current observations place upper limit on ϵ
- ► m_{A'} generated by Stuckelberg mechanism (no phase transition), occurs naturally in large volume string compactifications

Some ways to detect low-mass ${\cal A}^\prime$

"Light-Shining-Through-Walls" experiments

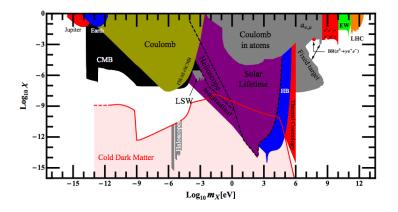
$$\gamma \sim A' \sim \gamma$$

▶ Helioscope: "Staring at the Sun"



Full A' Parameter Space

Evidently a large parameter space needs to be explored!



[Jaeckel '13]

- DM may be part of a more complicated dark sector
- Dark photons can provide a portal to DM (kinetic mixing), may even be the DM (kinetic mixing)
- Lots of parameter space still out there, and new/existing experiments are continuing the search!