

Detecting rare events using optical bolometers with transition edge sensors.

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UC Berkeley

Outline

- Overview of bolometric detection principles
- CUORE experiment as state of the art large conventional bolometric detector
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 - energy resolution
 - backgrounds
 - limited capacity for active background rejection (anticoincidence)
- Transition Edge Sensors (TES's)
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- Optical bolometers
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 - overview
 - CRESST II as state of the art example
 - Cherenkov light bolometers
 - Neganov-Luke amplification for light detectors

Bolometers: general principles

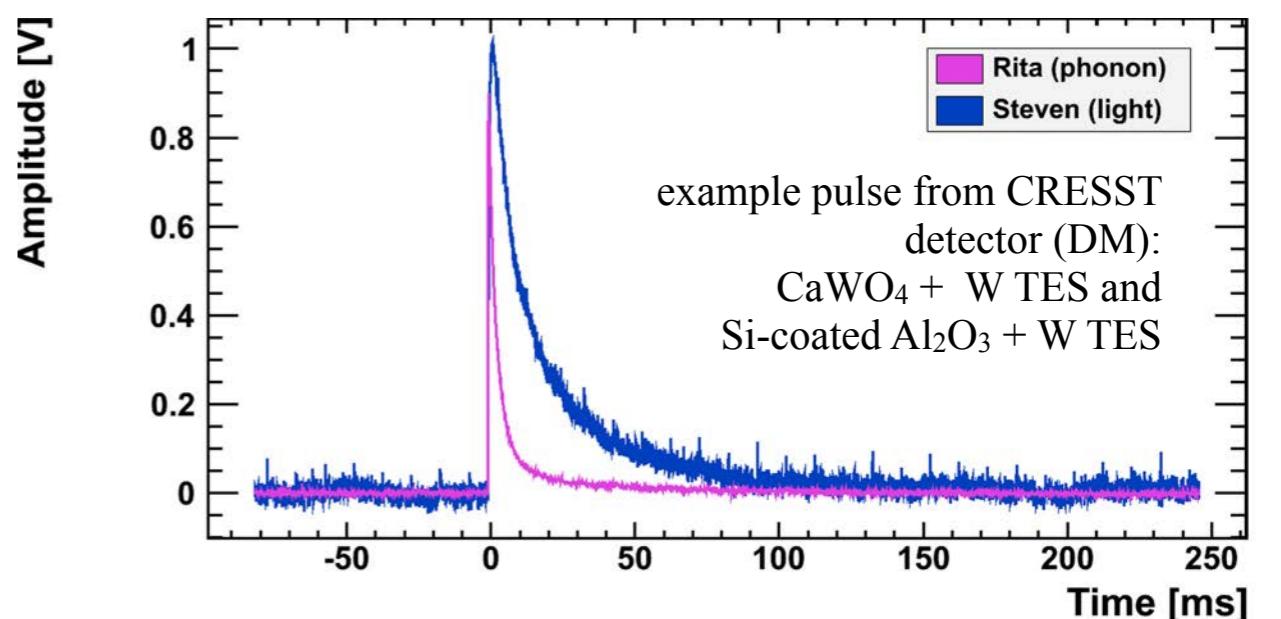
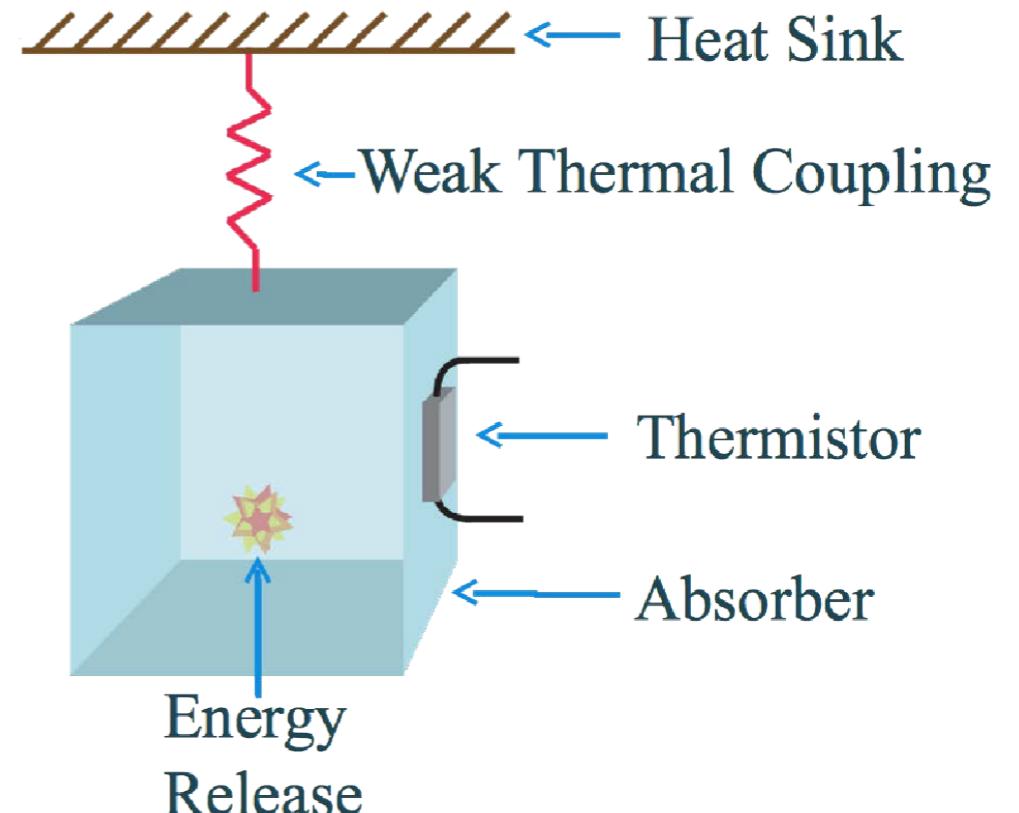
- Detection of thermal signatures of absorbed particles:

$$\Delta T = \frac{E_{event}}{C_{absorber}}$$

- Low T (10mK–100mK):
 - minimization of $C_{absorber}$
 - achieved using Dilution Refrigerator
- Fundamental limit on energy resolution very low:

$$\Delta E_{limit} \sim \sqrt{kCT^2}$$

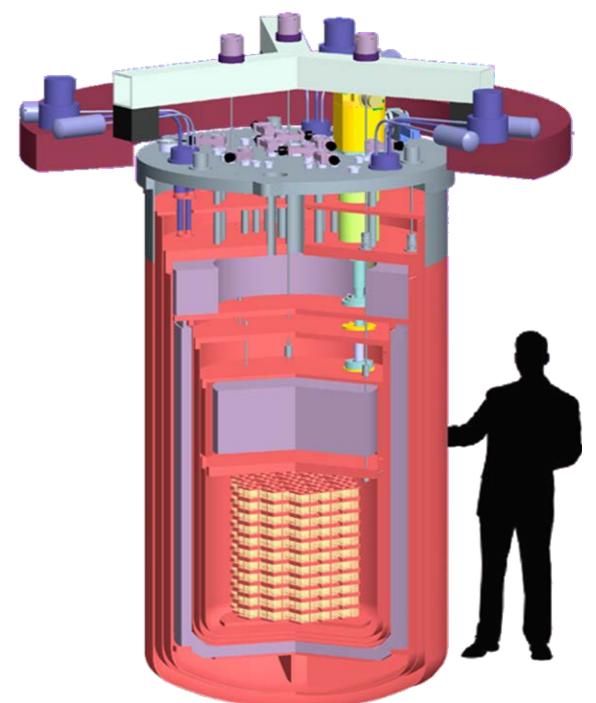
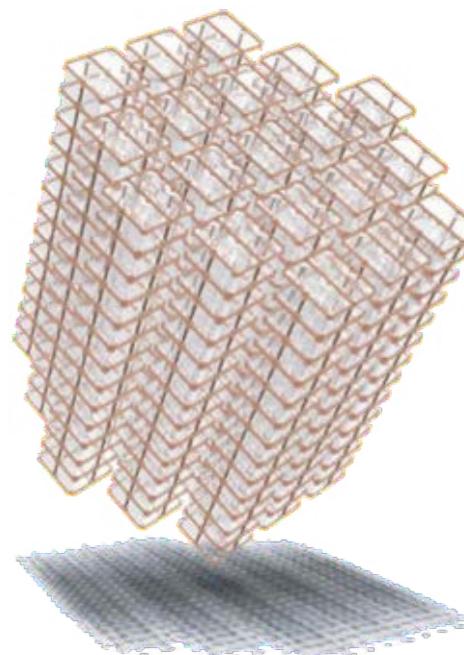
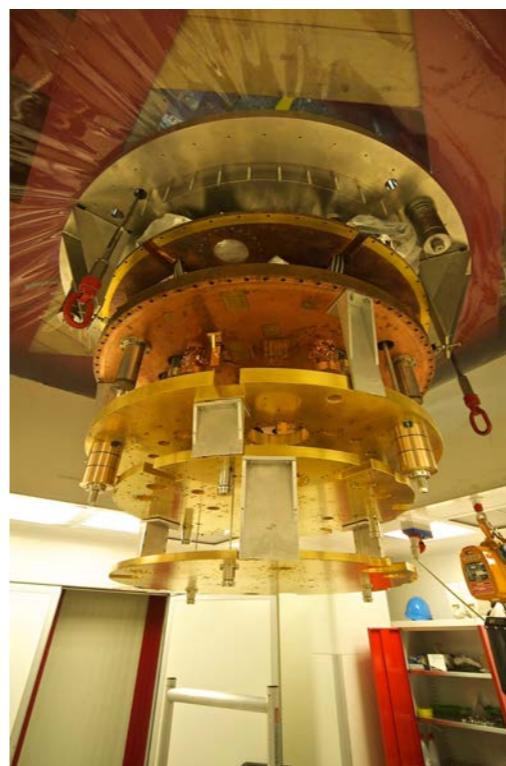
- Attractive for low-background, rare event searches like Dark Matter and $0\nu\beta\beta$.



CUORE: example of modern day
“traditional” bolometric detector
(w/o TES or optical channel)

CUORE

- Cryogenic Underground Observatory for Rare Events.
- Bolometric search for $0\nu\beta\beta$ in ^{130}Te (with secondary topics $2\nu\beta\beta$, axions, and others).
- A good practical example of currently-running/soon-to-run practical ton-scale bolometric detector with *only a thermal channel*.
- 988 TeO₂ 5x5x5cm³ crystals in 19 towers of 52: source and absorber, m_{isotope} = 204kg.
- T_{base} ≤ 10mK
- Commissioning ongoing, data in 2016.
- CUORE-0—single CUORE tower, turned off summer 2015, results spring 2015 (2y data).



CUORE: measurements

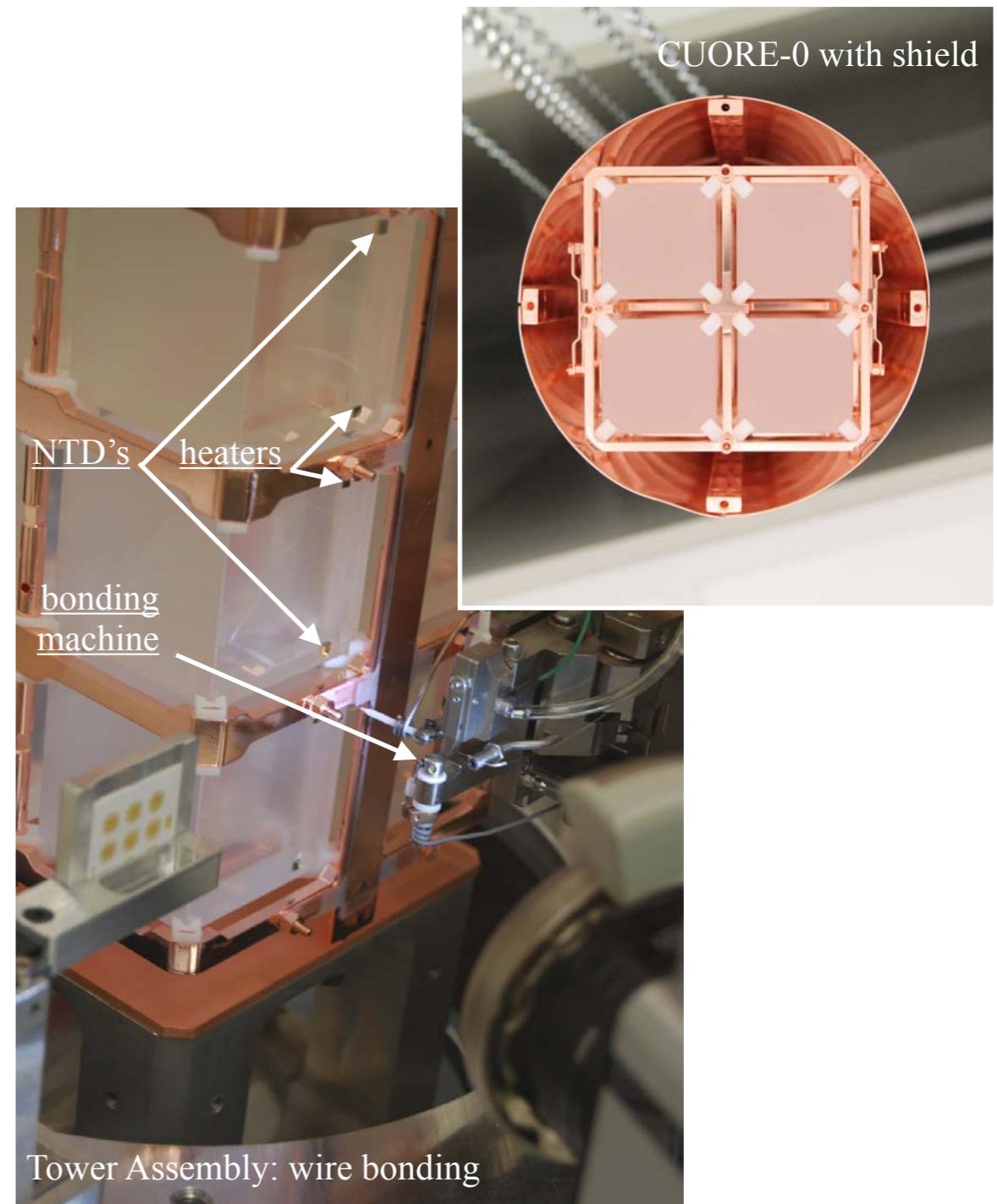
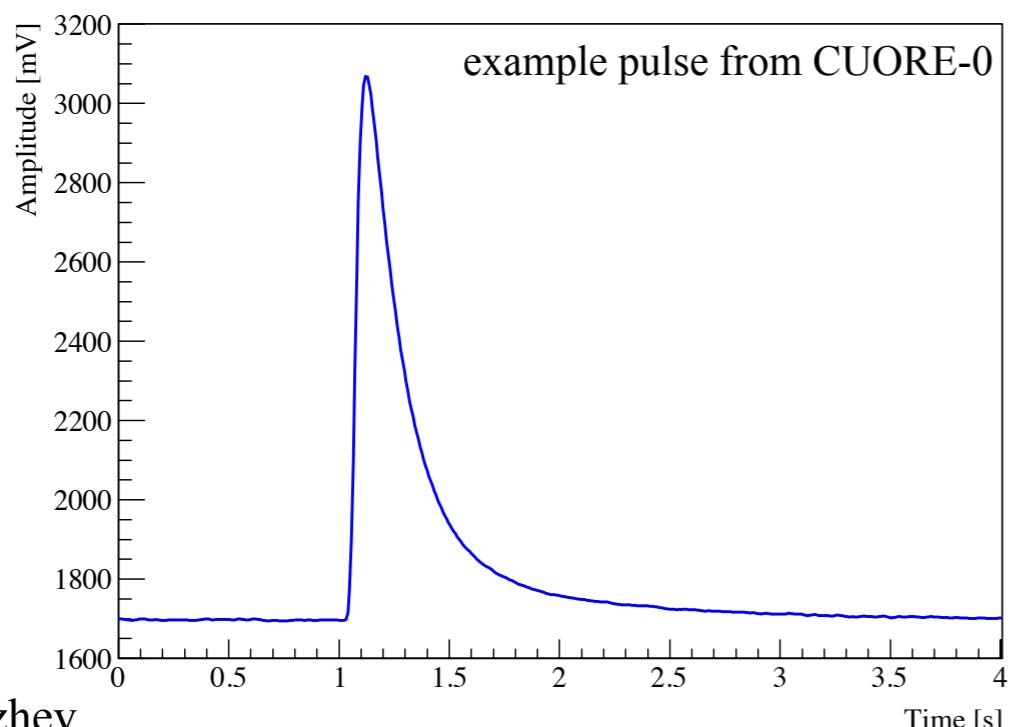
- $5 \times 5 \times 5 \text{ cm}^3 \text{ TeO}_2$ at 10 mK :

$$\Delta T_{event} = \frac{E_{event}}{C_{crystal}}$$

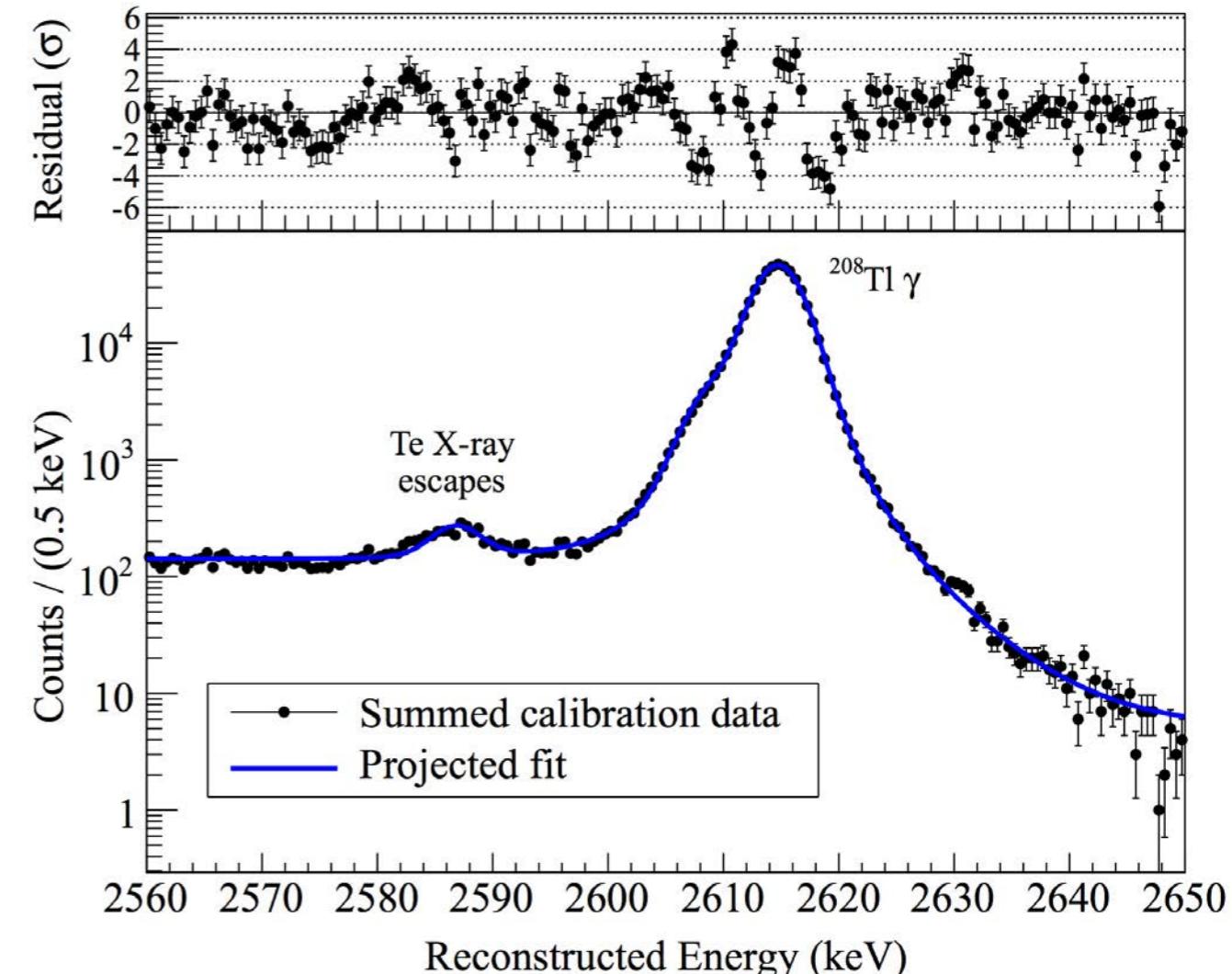
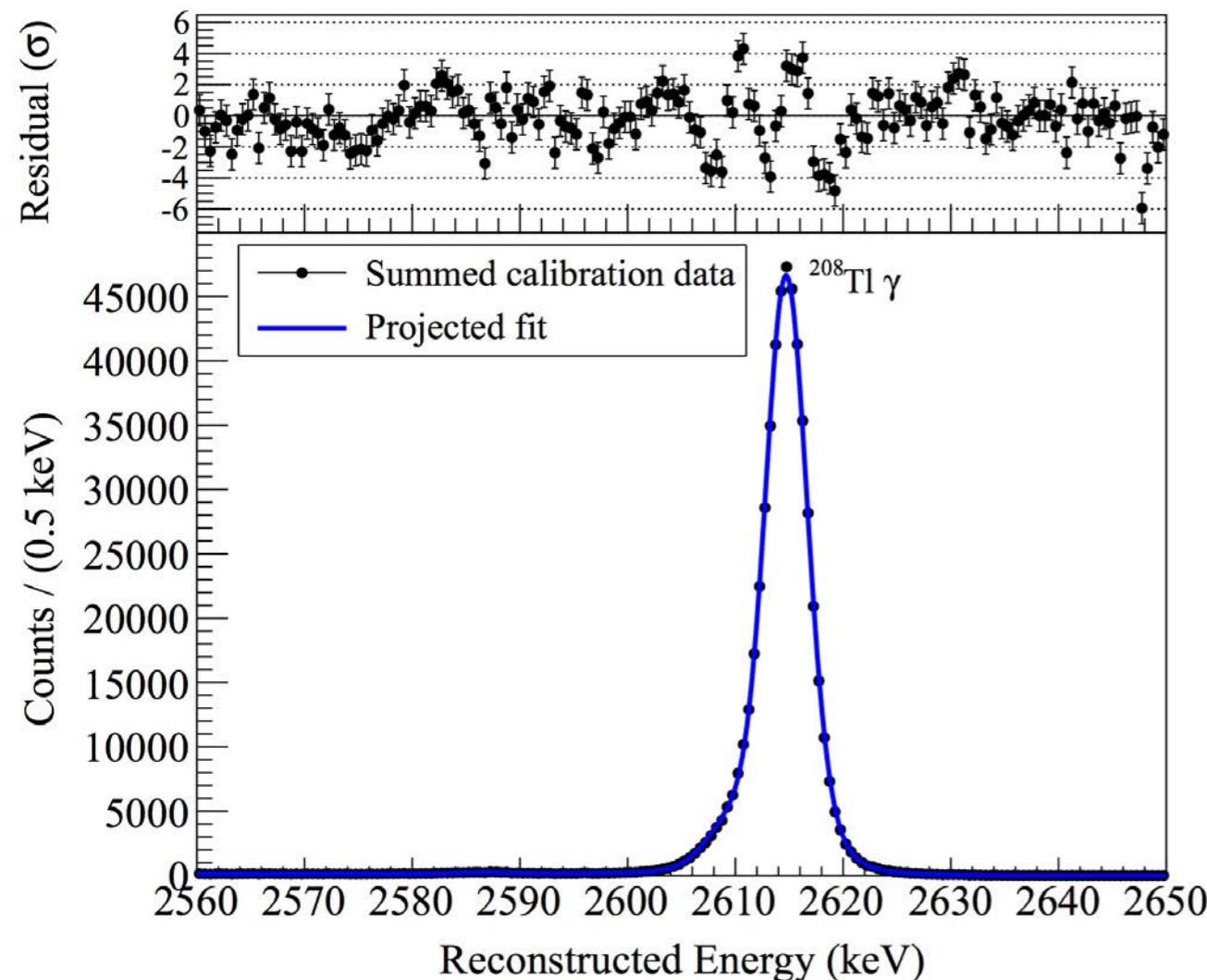
with $C_{crystal}^{-1} \approx 100 \mu\text{K}/\text{MeV}$

- Readout with Ge NTD thermistor:

$$R_{NTD} = R_0 e^{\sqrt{T_0/T}}$$



CUORE: energy resolution performance



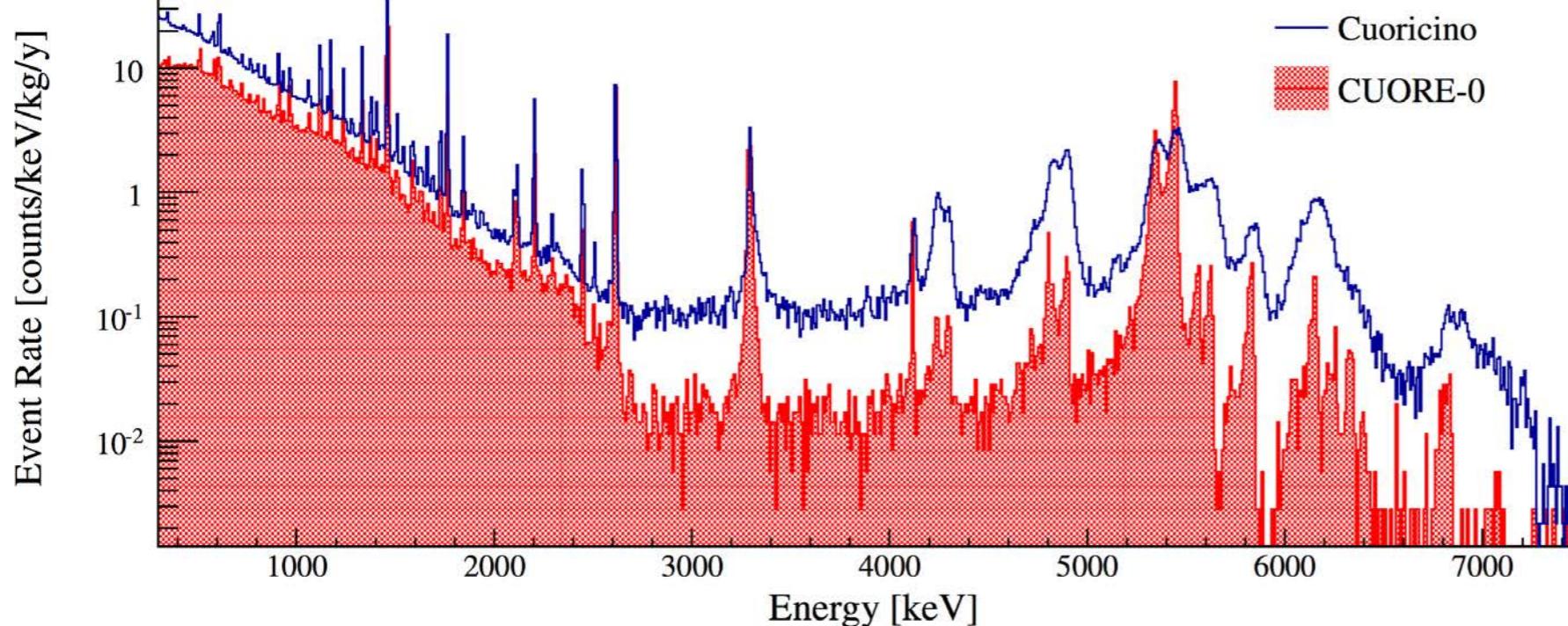
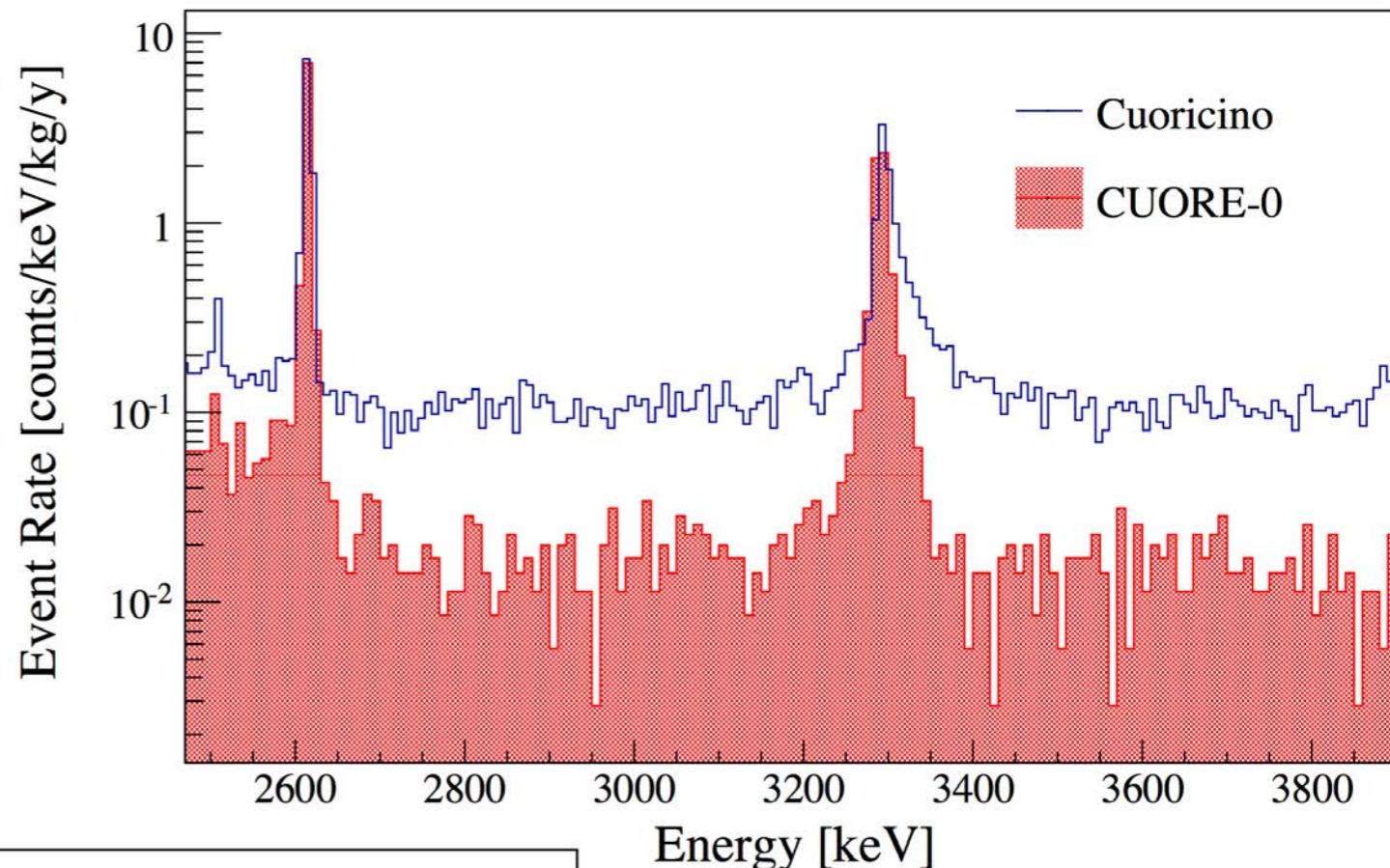
$$\Delta E_{limit} \sim \sqrt{kCT^2} \sim 10\text{eV} \quad \Delta E_{CUORE-0} \sim 5\text{keV FWHM}$$

K. Alfonso et al. arxiv: 1504.02454

at $\sim 2.6\text{MeV}$ ($^{208}\text{Tl} \gamma$ line)

CUORE: background performance

- Passive suppression to 0.02 counts/keV/kg/y (CUORE-0), expected 0.01 counts/keV/kg/y for CUORE.
- $Q=2527$ keV—high energy edge of γ region, but away from major peaks due to good energy resolution.
- Primary background is degraded α 's.

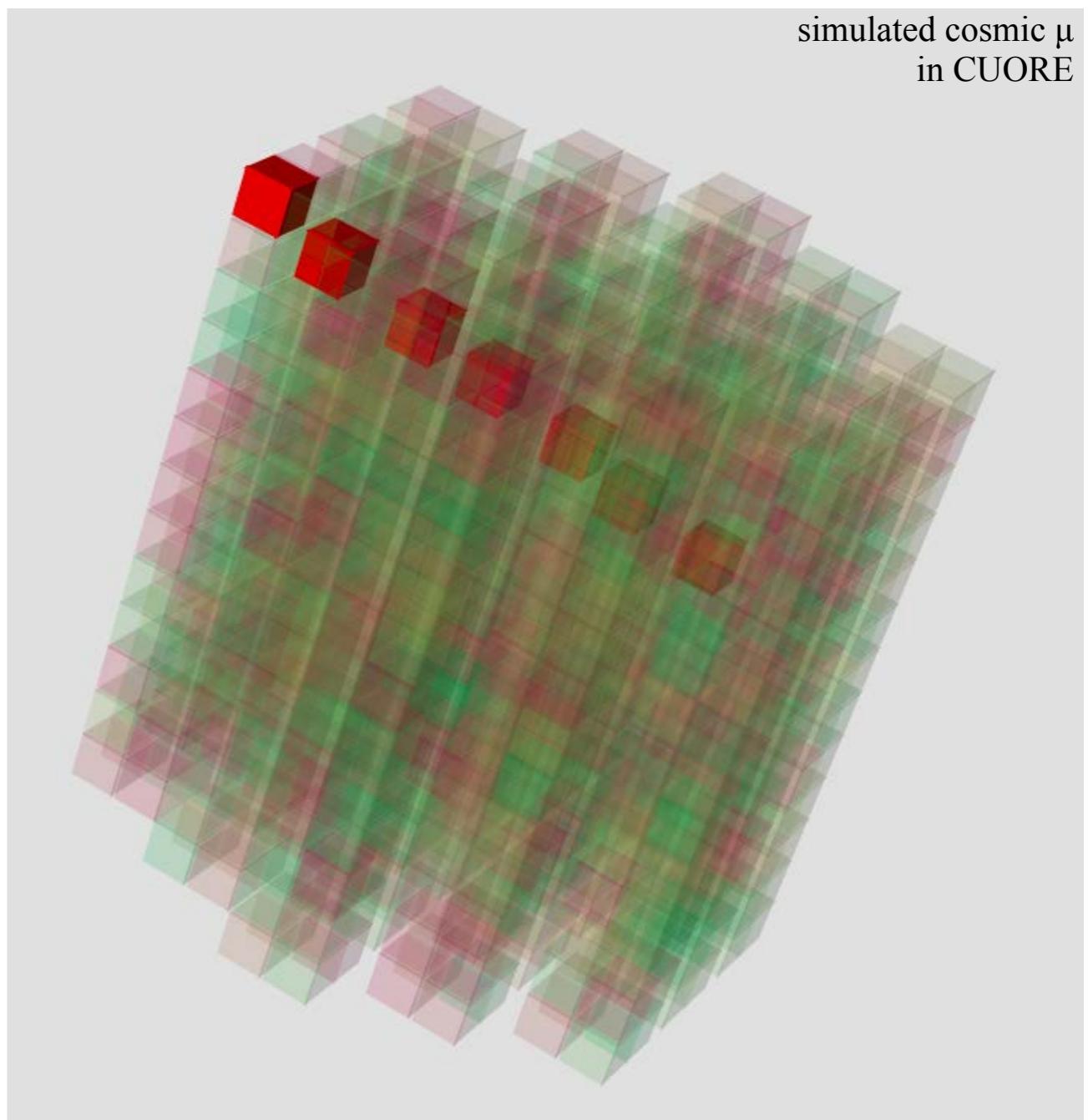


↑ α region

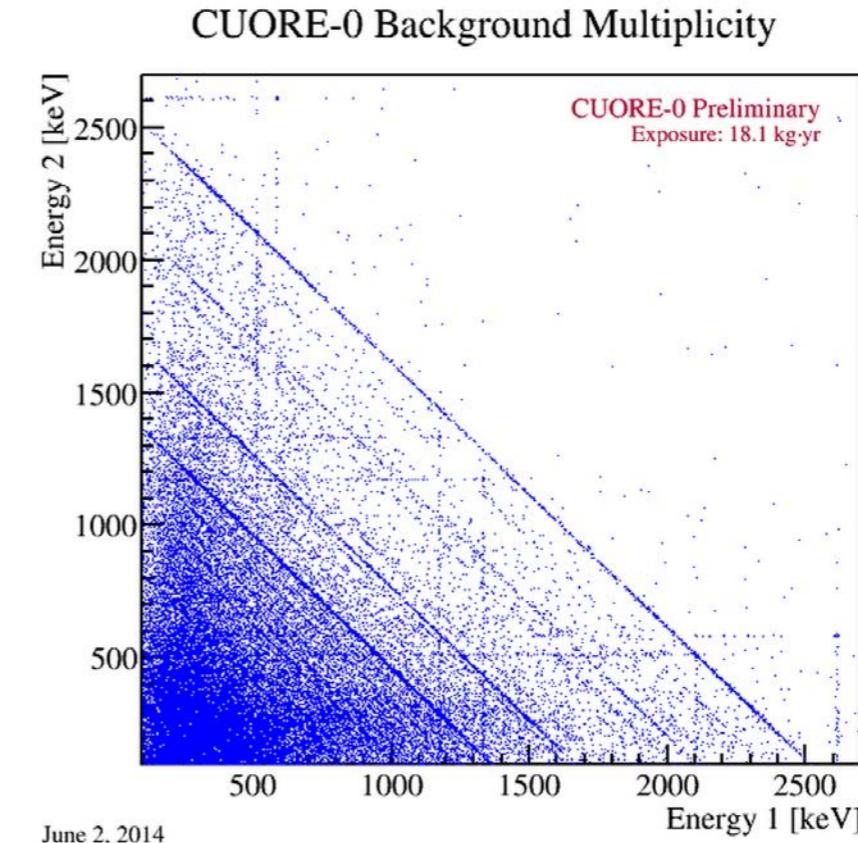
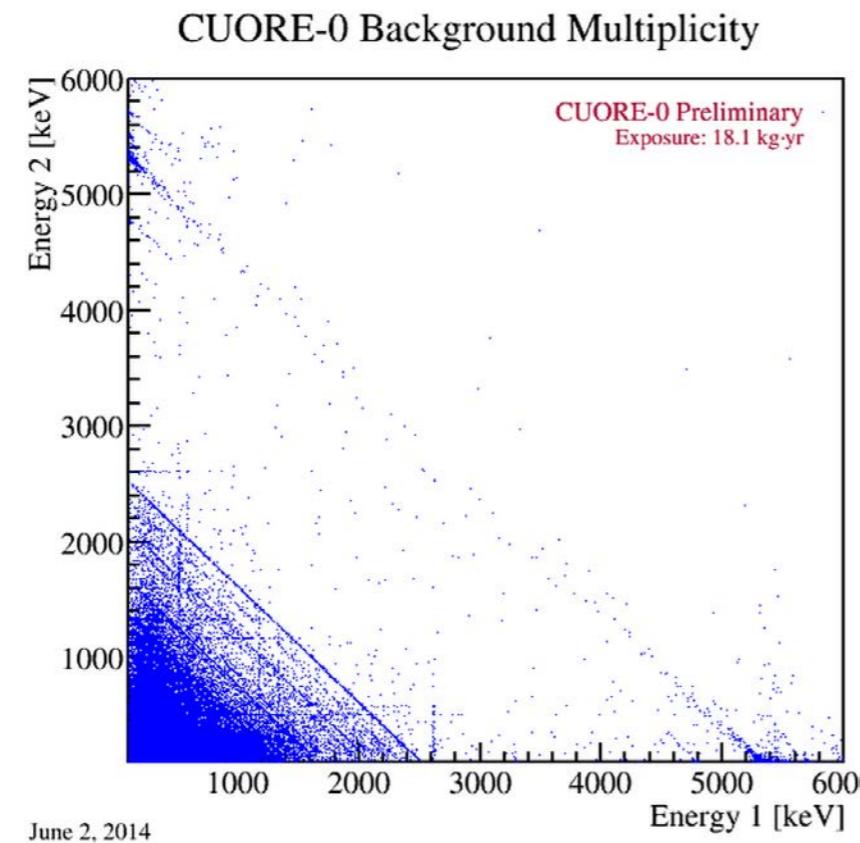
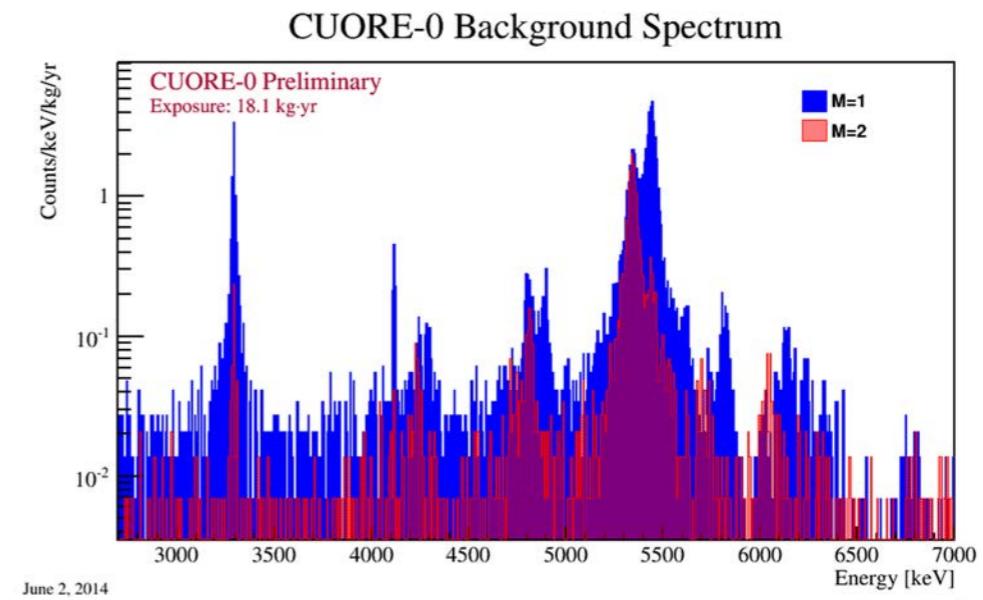
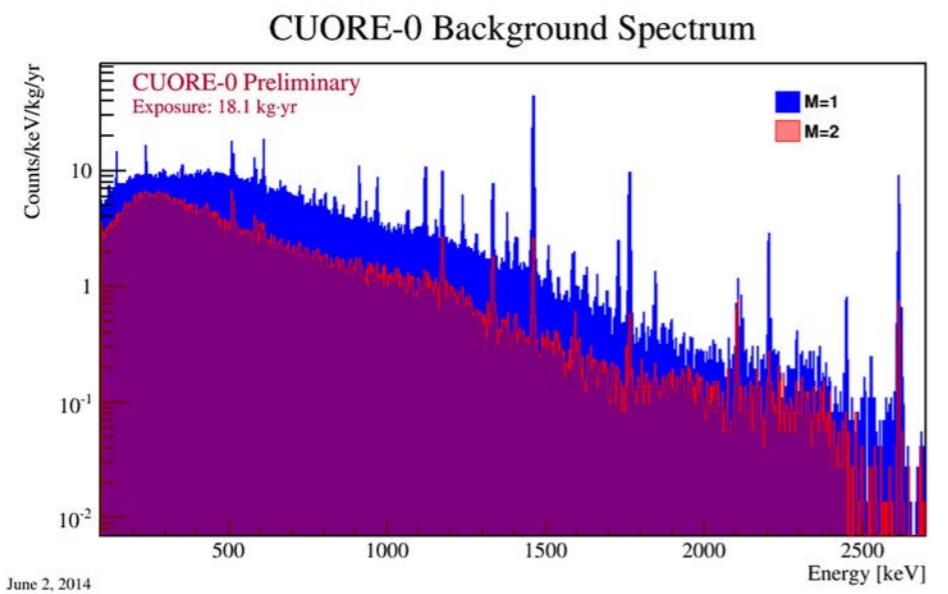
← full spectrum

CUORE: anti-coincidence background rejection

- Only thermal channel, NTD readout
⇒ only energy measurement ⇒ not possible to distinguish α , β , γ , μ , etc.
- Anti-coincidence:
 - $2\nu\beta\beta$, $0\nu\beta\beta$ events — 1 crystal.
 - Some surface α 's — 2 crystals— nuclear recoil back into crystal, α absorbed on adjacent crystal.
 - Some γ 's — multiple crystals (Compton length 10-20cm).
 - Cosmic μ 's— multiple crystals.



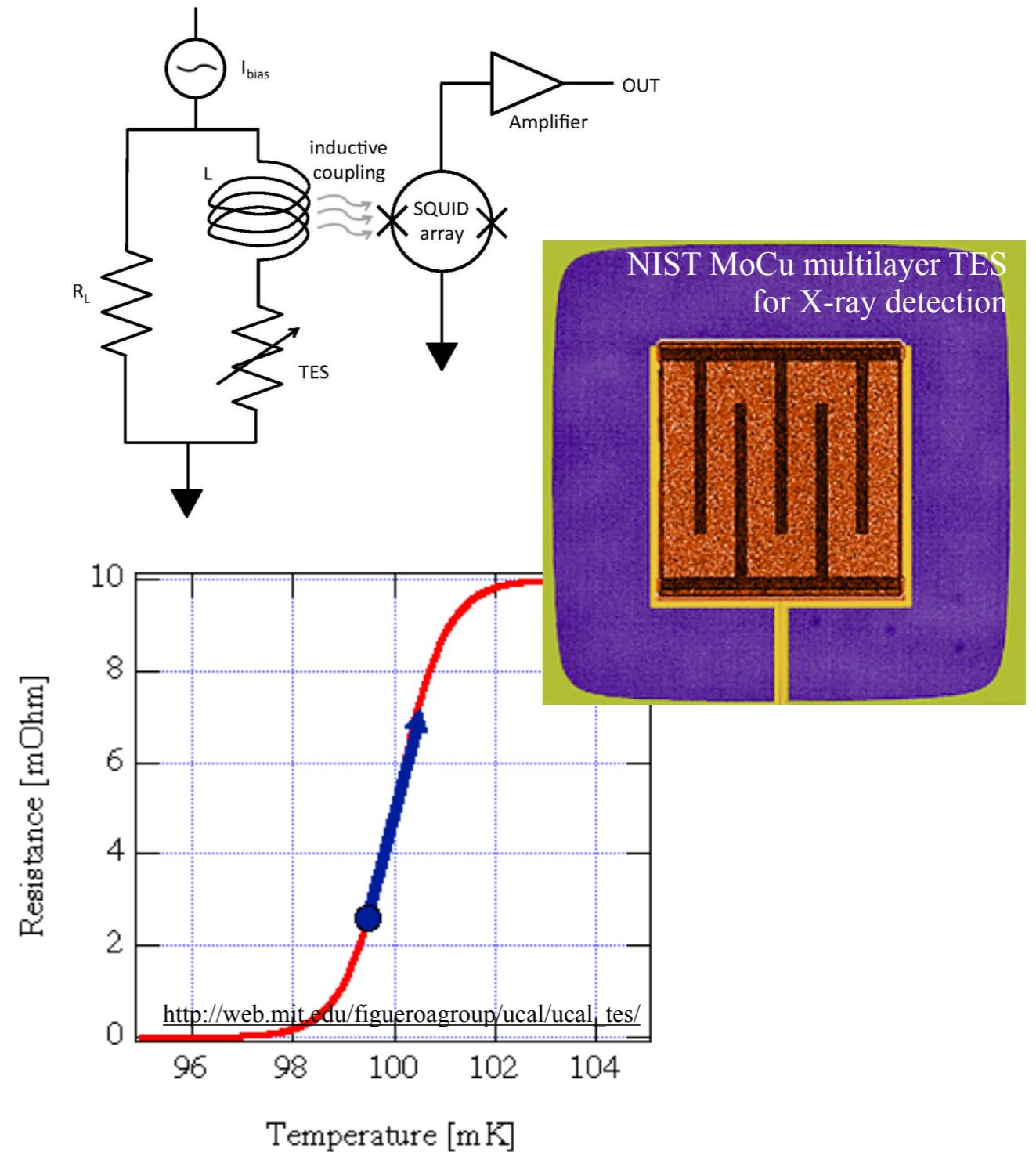
CUORE: anti-coincidence background rejection



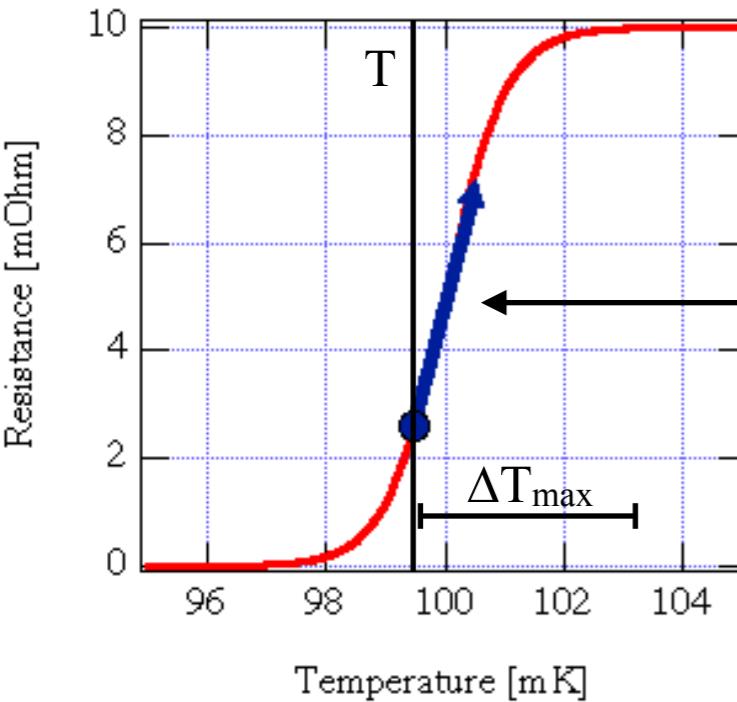
Transition Edge Sensors (TES's): improved phonon signal readout

Transition Edge Sensors: overview

- Temperature sensor based on superconducting transition of a metal or alloy.
- In narrow region around T_C , $R(T)$ much more sensitive than exponential.
- Typically read out by SQUID \Rightarrow low-T preamp \Rightarrow lower noise (than room-T preamp).
- 2 primary types in use:
 - W (tungsten) (α and β phases)
 - Bilayer or multilayer sample of superconducting and non-superconducting metals (e.g. Ir/Au).
- Main challenge/compromise for rare event bolometers: sufficiently low T_C for low temperature operation.
- Long history of use outside rare event searches for photon detection.



Transition Edge Sensors: operation theory



$$\alpha = \frac{T}{R} \frac{dR}{dT}$$

$$\Delta E = 2.35 \sqrt{4kT^2 \frac{C}{\alpha}}$$

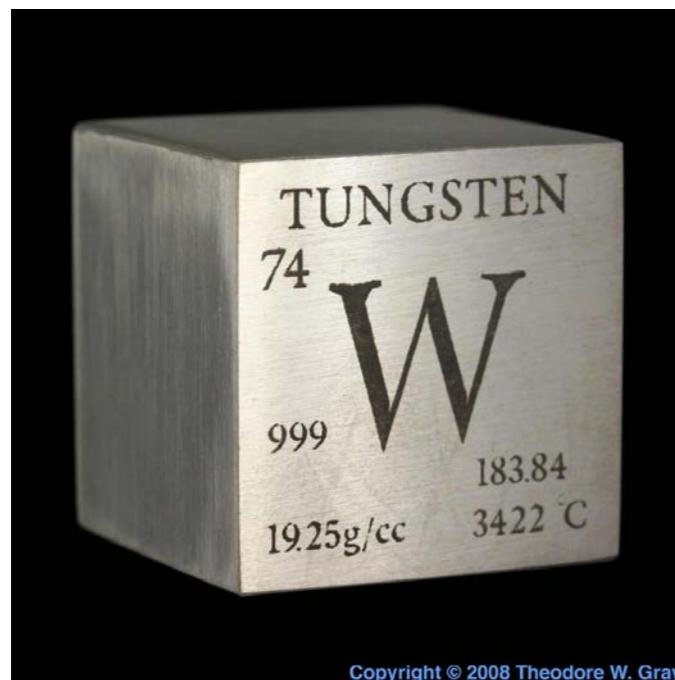
$$E_{max} = C \Delta T_{max} \simeq \frac{C}{\alpha} T$$

$$\Delta E \simeq 2.35 \sqrt{4kT^2 E_{max}}$$

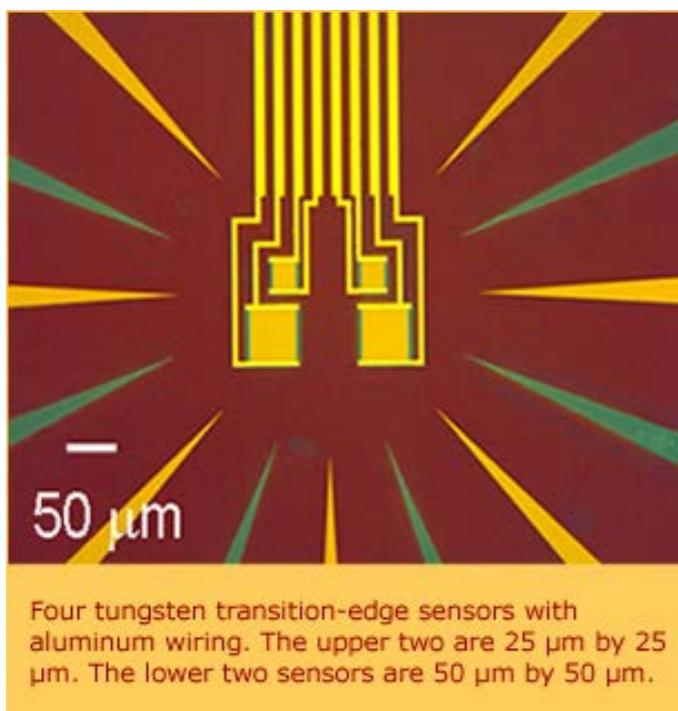
- figure of merit α —steepness of transition
- low operating T , low C of absorber (interconnected), large α improve energy resolution ΔE .
- low C and large α also lower bandpass
⇒ best performing TES must be very well tuned for precise application.
- Example: typical X-ray TES @ 100mK has bandpass $\sim 0.1\text{keV} - \sim 10\text{keV}$ w/ $\Delta E \sim 1.5\text{eV}$ FWHM (plot)
- low R (+ electrothermal feedback) ⇒ possibility of larger signal bandwidth (than NTD) ⇒ possibility of pulse shape discrimination

http://web.mit.edu/figueroagroup/ucal/ucal_tes/

Tungsten (W) TES's



- 2 crystalline forms:
 1. α : BCC structure, isometric grains. $T_C \sim 10\text{mK} - 15\text{mK}$
 2. β : A15 cubic, columnar habit. $T_C \sim 1\text{K} - 4\text{K}$
- 2 forms can be mixed to achieve intermediate T_C 's.
- Pure α (or β) sample very hard to fabricate:
 - Most examples' $T_C > 15\text{mK}$ necessitates higher operating temperature.
 - Irregularity/high rejection rate
 - \Rightarrow O.K. for small detectors like CRESST II; problematic for ton-scale experiments like CUPID (planned CUORE successor) or EURECA (planned CRESST successor).
- Examples of use:
 - CRESST II (close to pure α -W; main and light bolometers)
 - CDMS ($T_C \approx 90\text{mK}$, mixed phases; in athermal phonon collector of Ge detector)



Multilayer TES's

- 2+ $O(100\mu m)$ layers of superconducting and normal metals sputtered or atomic-deposited on substrate.
- Ir/Au, Ir/Pt, Ir/IrMn/Au, Ir/Pd/Au, Mo/Au, Mo/Cu, Mo/Ti, Al/Au, Al/Ti/Au, others...
- T_C of superconducting metal is suppressed by leakage of carriers from normal metal layer.
- Usadel theory (right) to estimate T_C of bilayer—impossible to get transmission parameter, empirical methods are necessary.
- Motivations for developing multilayer TES's :
 - Theoretically possible to set T_C precisely, including very cold $<15mK$ transitions (no need to compromise like with W).
 - Dedicated clean, high vacuum, high precision fabrication facility should allow for cheaper and more robust production than W TES's for large detectors.
- Most practical examples now: T_C 's $\sim 100mK$; used for X-rays, cosmology, nonproliferation R&D, etc.
- No workable examples with $T_C < 20mK - 30mK$

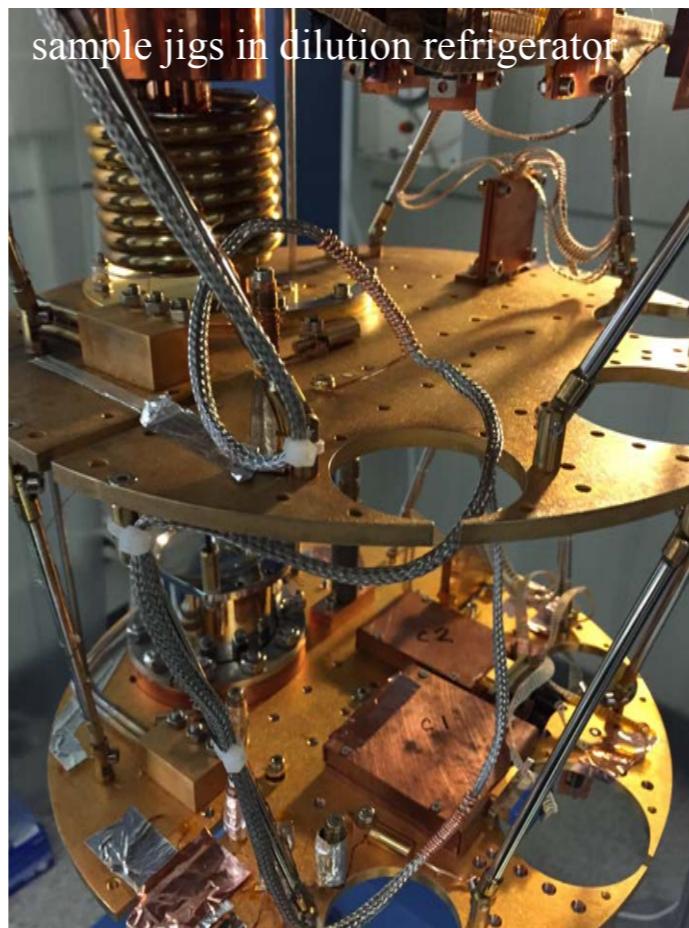
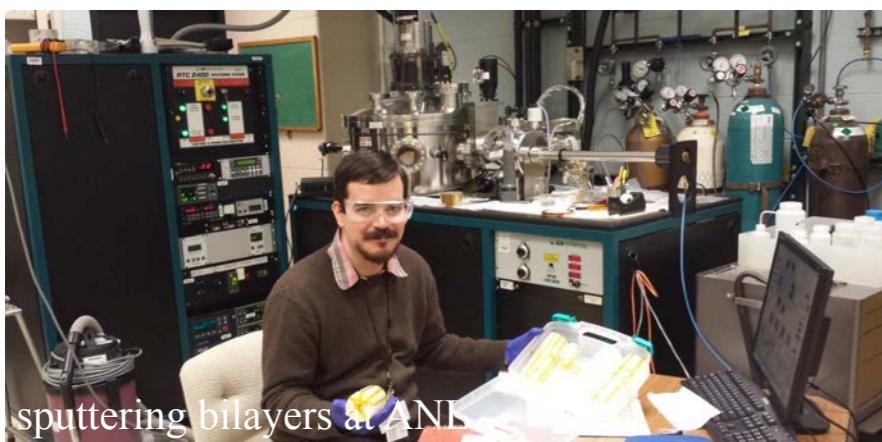
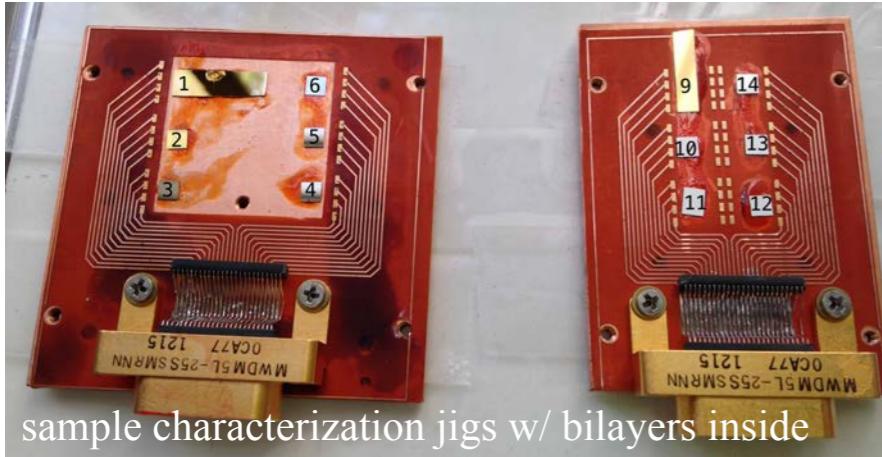
$$T_C = T_{C_0} \left[\frac{d_s}{d_0} \frac{1}{1.13(1 + 1/\alpha)} \frac{1}{t} \right]^\alpha$$

$$\frac{1}{d_0} = \frac{\pi}{2} k_B T_{C_0} \lambda_f^2 n_s,$$

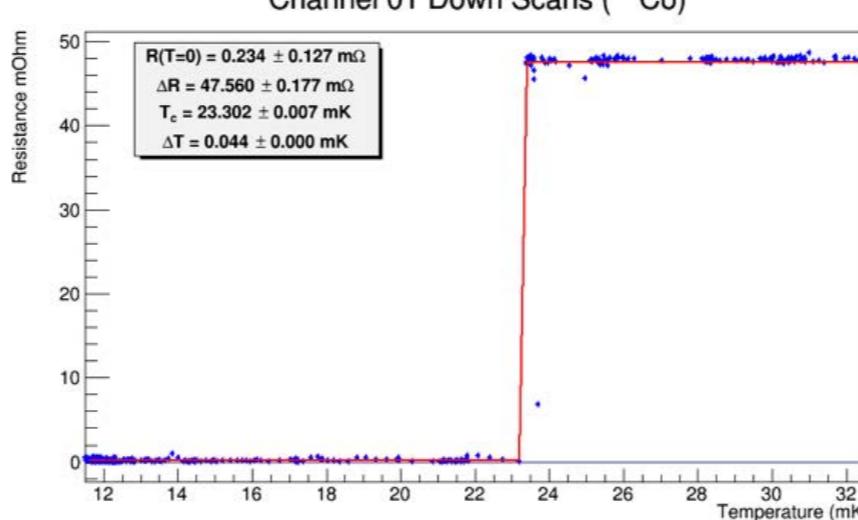
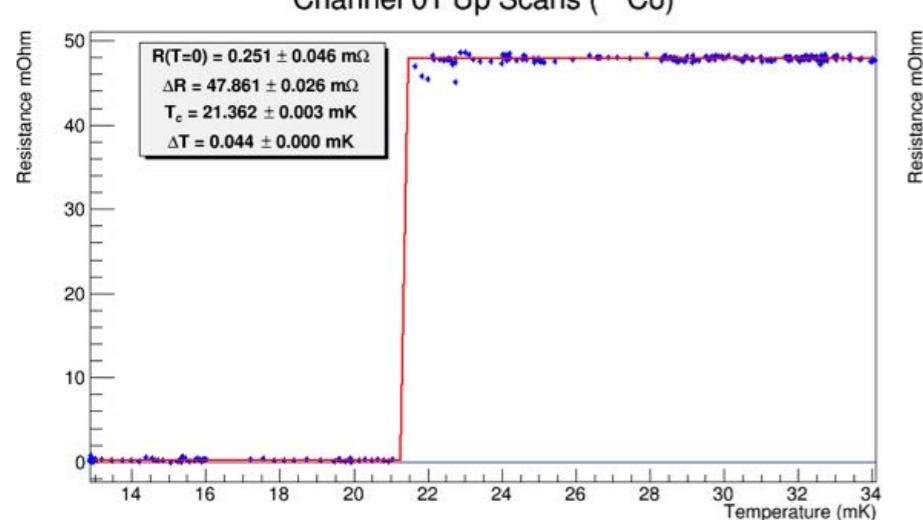
$$\alpha = d_n n_n / d_n n_s.$$

n_n, n_s — densities of states
 d_n, d_s — film thicknesses
 T_{C_0} — natural s.c. layer T_C
 λ_f — Fermi λ in normal metal
 t — unitless interface transmission parameter $O(1)$

Low-T_c bilayer TES R&D at Berkeley



- Dilution refrigerator facility:
 - Oxford Instruments cryogen-free DR:
 - $P_{\text{cooling}}(100\text{mK}) \geq 400\mu\text{W}$ (record $\sim 450\mu\text{W}$)
 - $T_{\text{base}} \leq 10\text{mK}$ (record $\sim 6\text{mK}$)
 - Readout:
 - AC370 resistance bridges
 - new Magnicon SQUID electronics
 - Thermometry:
 - RuO₂ resistance thermometers down to $20\text{mK} - 30\text{mK}$
 - ⁶⁰Co nuclear orientation thermometer down to base temperature
 - new Magnicon noise thermometer down to base temperature.
 - Ir/Au, Ir/Pt, Ir/IrMn/Au, Ir/Pd/Au bilayers being fabricated and tested in collaboration with ANL.
 - Next step: light detector w/ meandered bilayer TES with semiconductor absorber.

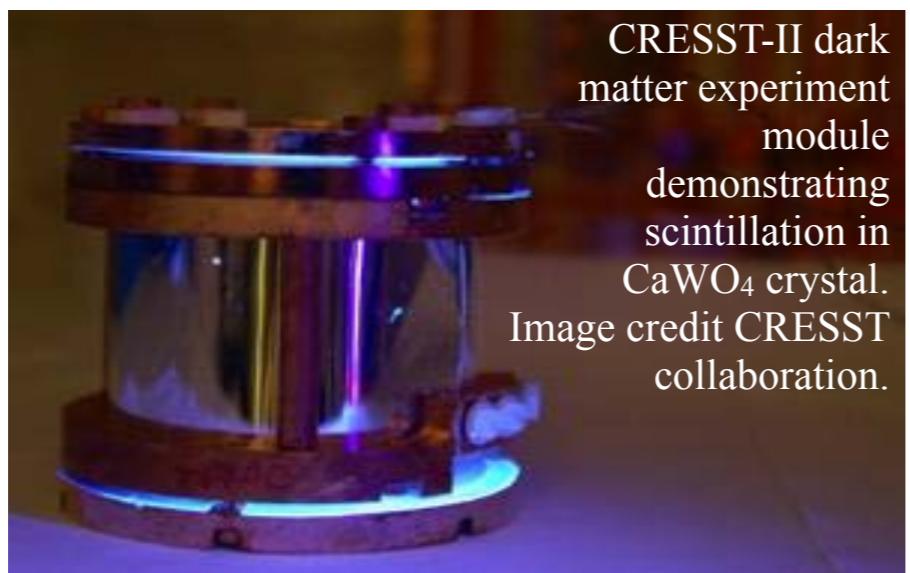
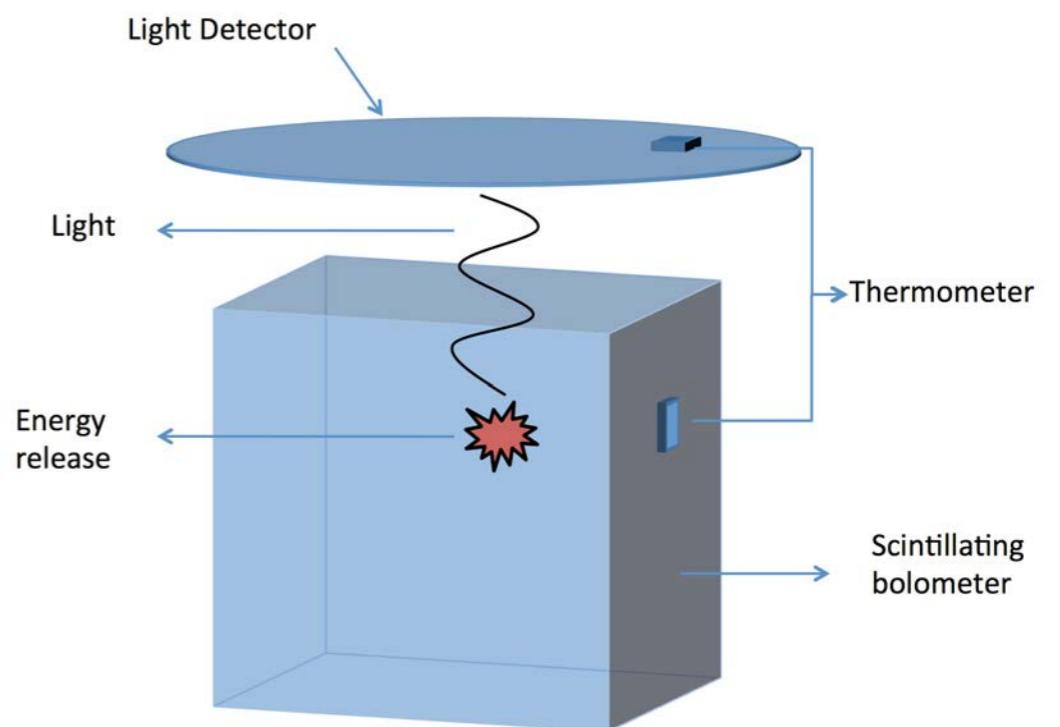


Ir/Au bilayer w/ $T_c \approx 22\text{mK}$ (hysteresis effect in plots from AC bridge readout)

Optical bolometers: active background discrimination

Scintillating bolometers

- Secondary bolometer w/ thin wafer as absorber to detect scintillation light from main bolometer crystal.
- Heat signature + light collection:
 - phonons: best E resolution
 - photons: separation of nuclear recoil (α , n or WIMP) from electron recoil (γ or β) via light yield, quenching.

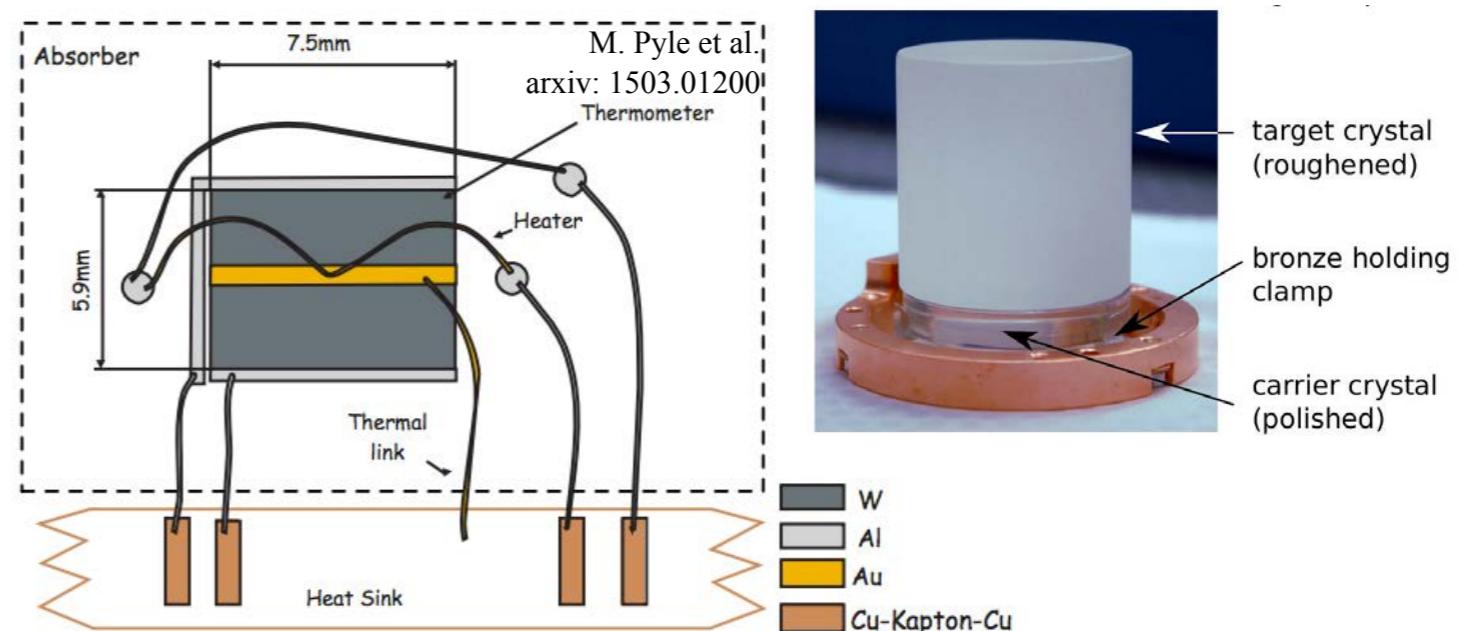


Scintillating bolometers: CRESST II

- Cryogenic Rare Event Search with Superconducting Thermometers II
- European Dark Matter experiment at Gran Sasso.
- Main strength: low WIMP mass sensitivity.
- 18 modules:
 - clean Cu structure
 - scintillating 3M reflector
 - phonon detector: 300g CaWO₄ scintillating crystal (total 5.4kg)
 - photon detector: Si-coated Al₂O₃ or Si absorber.
 - readout: W TES w/ $T_C \sim 10\text{mK}$ (closer to $\sim 13\text{mK}$, pure α W)
 - Variations in crystal attachment module to module.



<http://www.cresst.de/material.html>

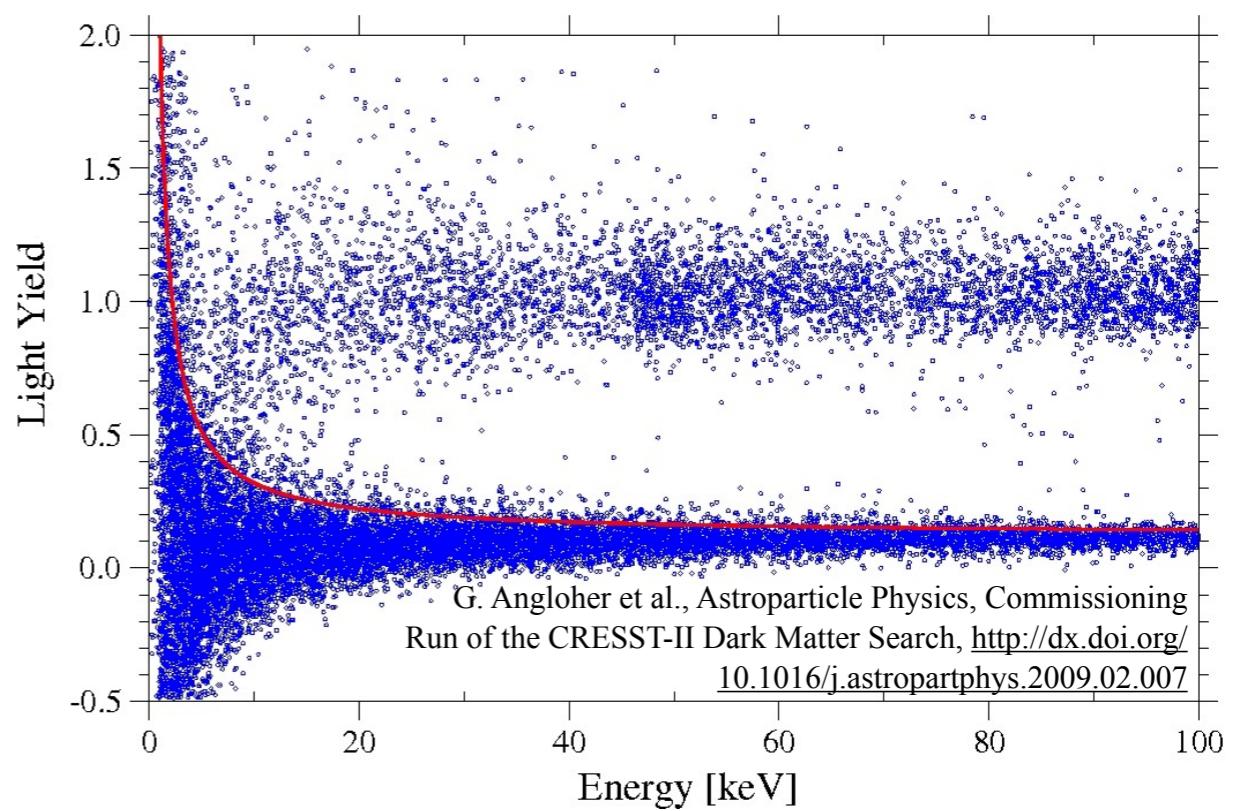
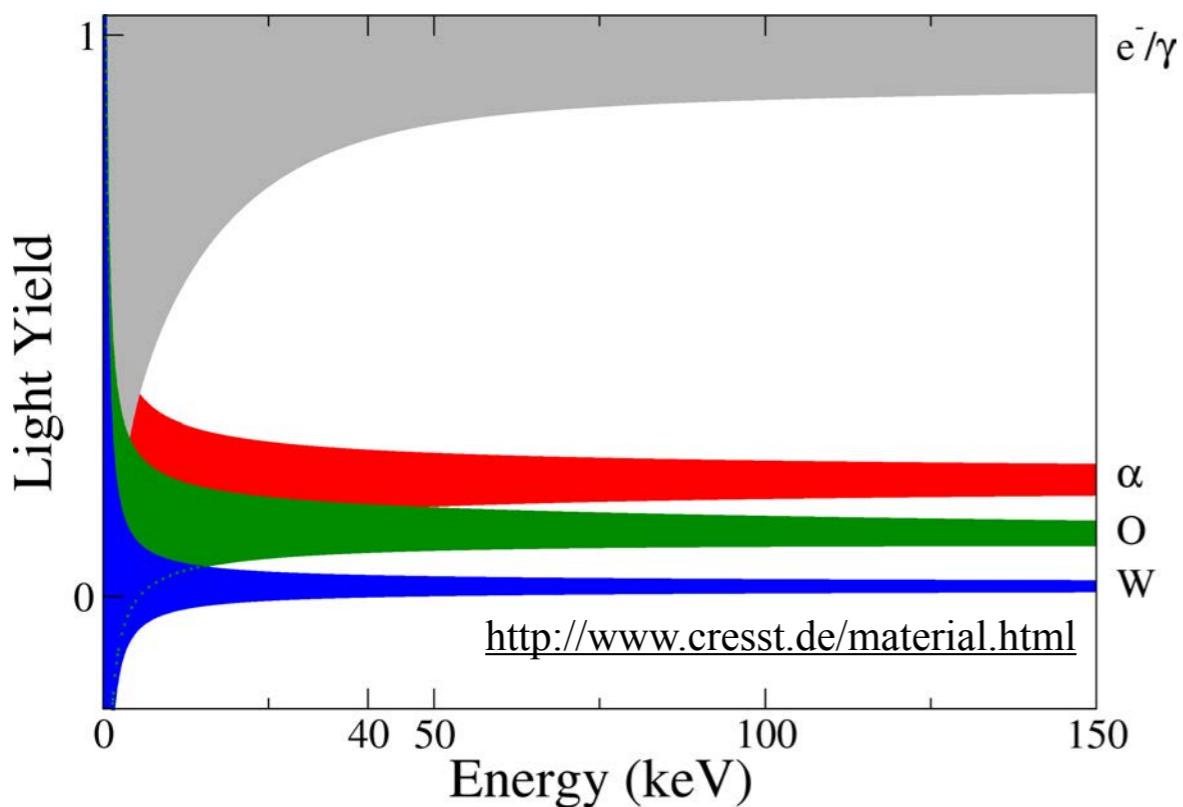


Schematic of the TES and its connections on the original CRESST-II phonon detector

Scintillating bolometers: CRESST II

- Example from CRESST II : less light from nuclear recoil (from WIMP, α , n) than from electron recoil (from γ or β) in CaWO_4 crystals (readout w/ W TES; light detection w/ Si-coated sapphire absorber and W TES).

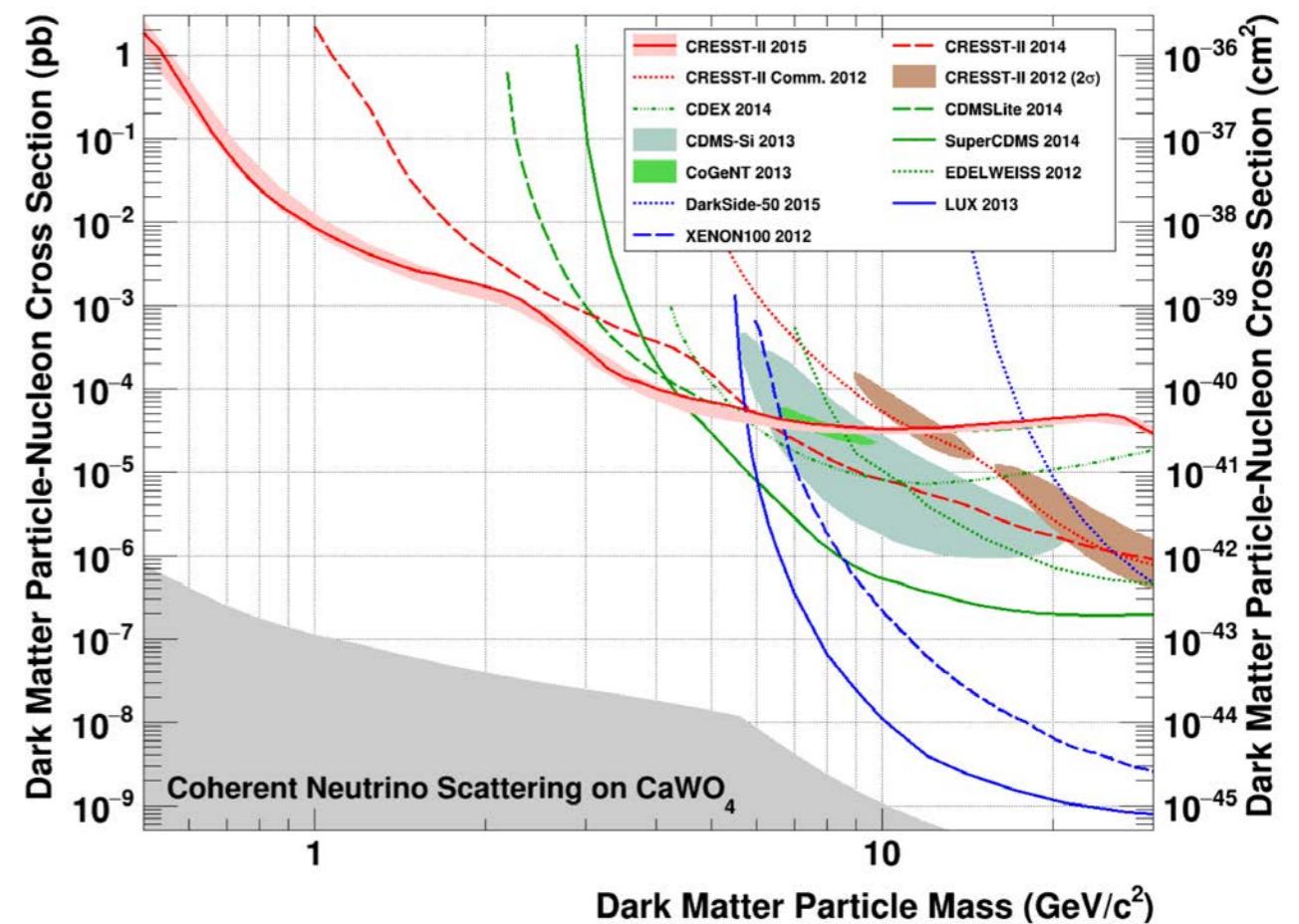
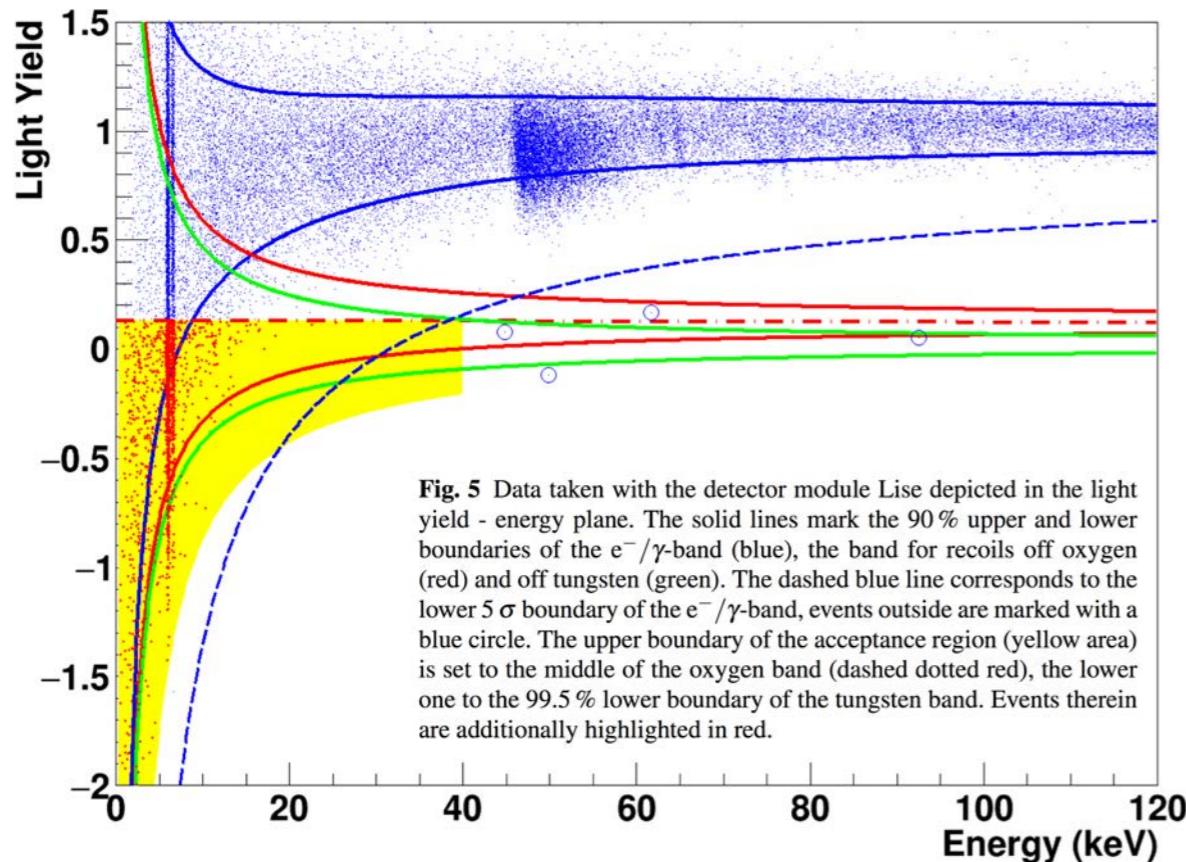
Left: schematic of bands. Right: data from calibration with neutrons.



* negative light yield values stem from amplitude fitting procedure that allows for negative amplitudes for unbiased treatment of noise. Phenomenon not unique to CRESST II.

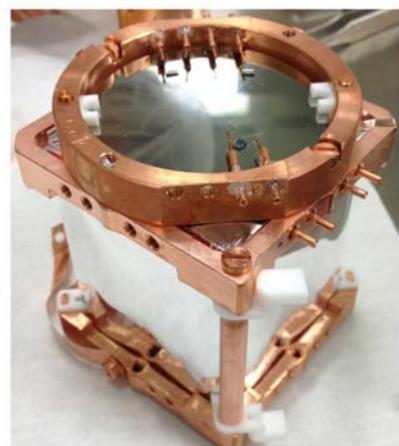
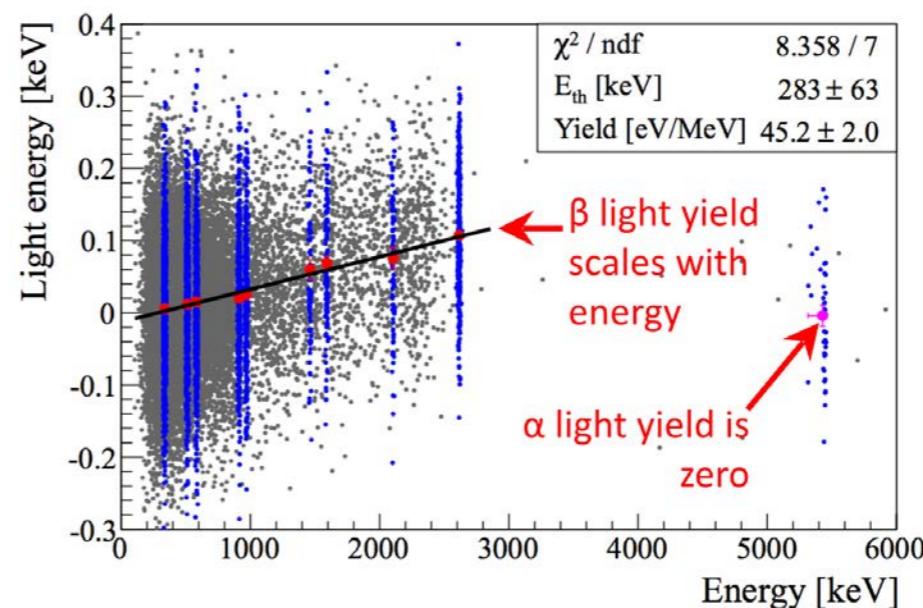
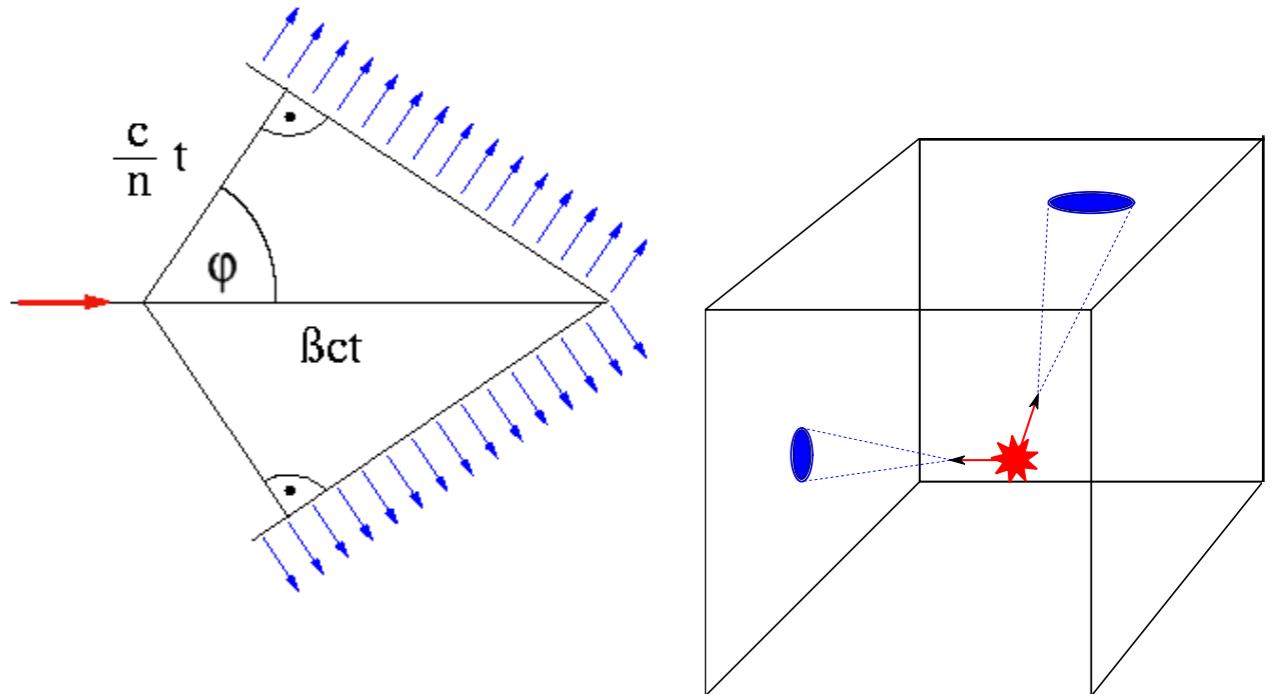
Scintillating bolometers: CRESST II

- Example from CRESST II : dark matter data.

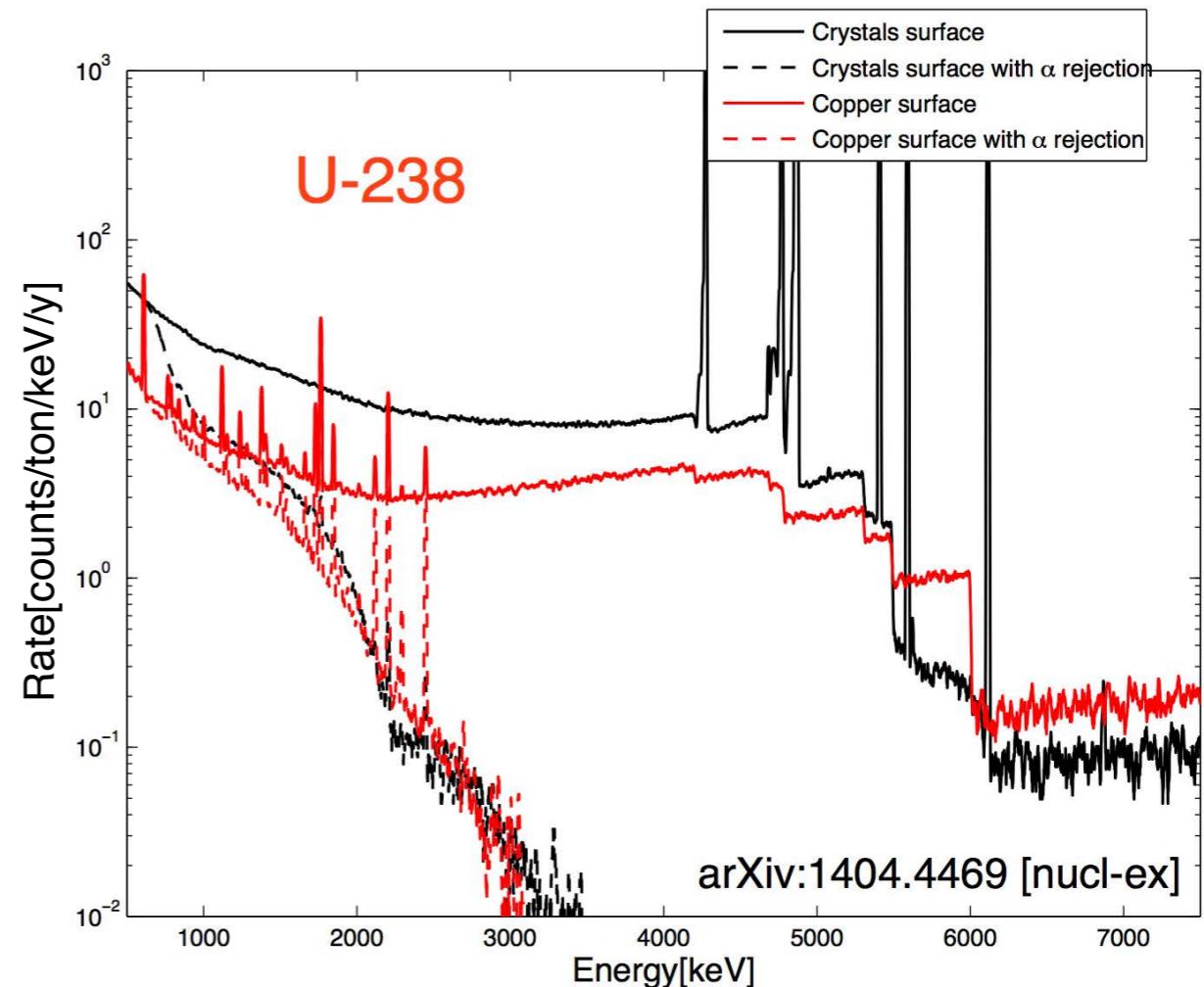
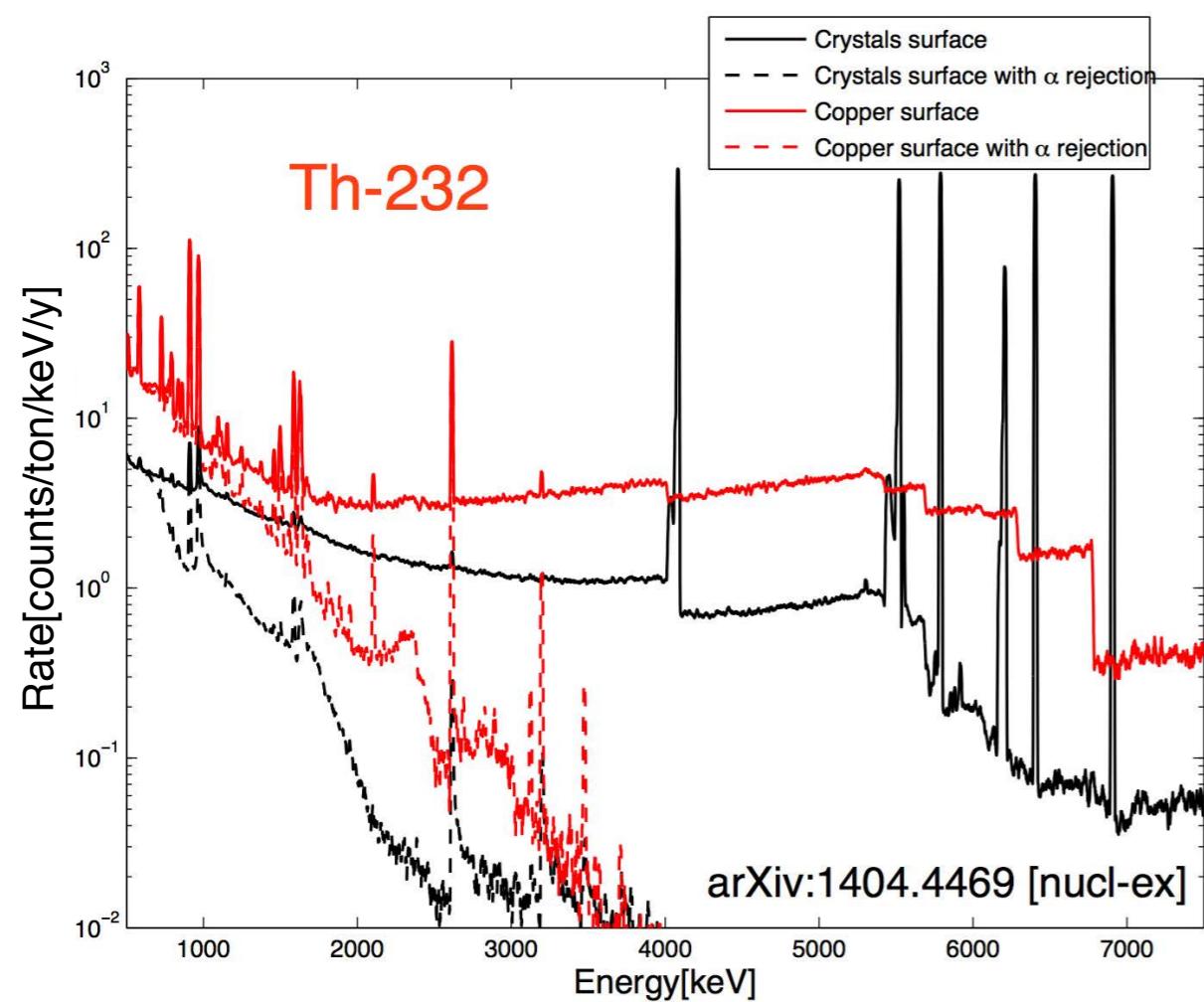


Cherenkov bolometers

- Analogous hardware configuration to scintillating bolometers
- Instead of scintillation photons, Cherenkov photons are detected:
 - β 's radiate; slow α 's and many n 's do not.
 - \Rightarrow WIMPS would not radiate
- degraded α 's a major background in cryogenic detectors \Rightarrow Cherenkov detectors not very useful for DM searches.
- Active interest from $0\nu\beta\beta$ researchers due to necessity of including isotope of interest in crystal (difficult w/ scintillators) and simple binary discrimination
 - challenging weak signals (10's of photons, thresholds as low as 10eV).



Improvement of CUORE performance with complete rejection of α background (simulation)



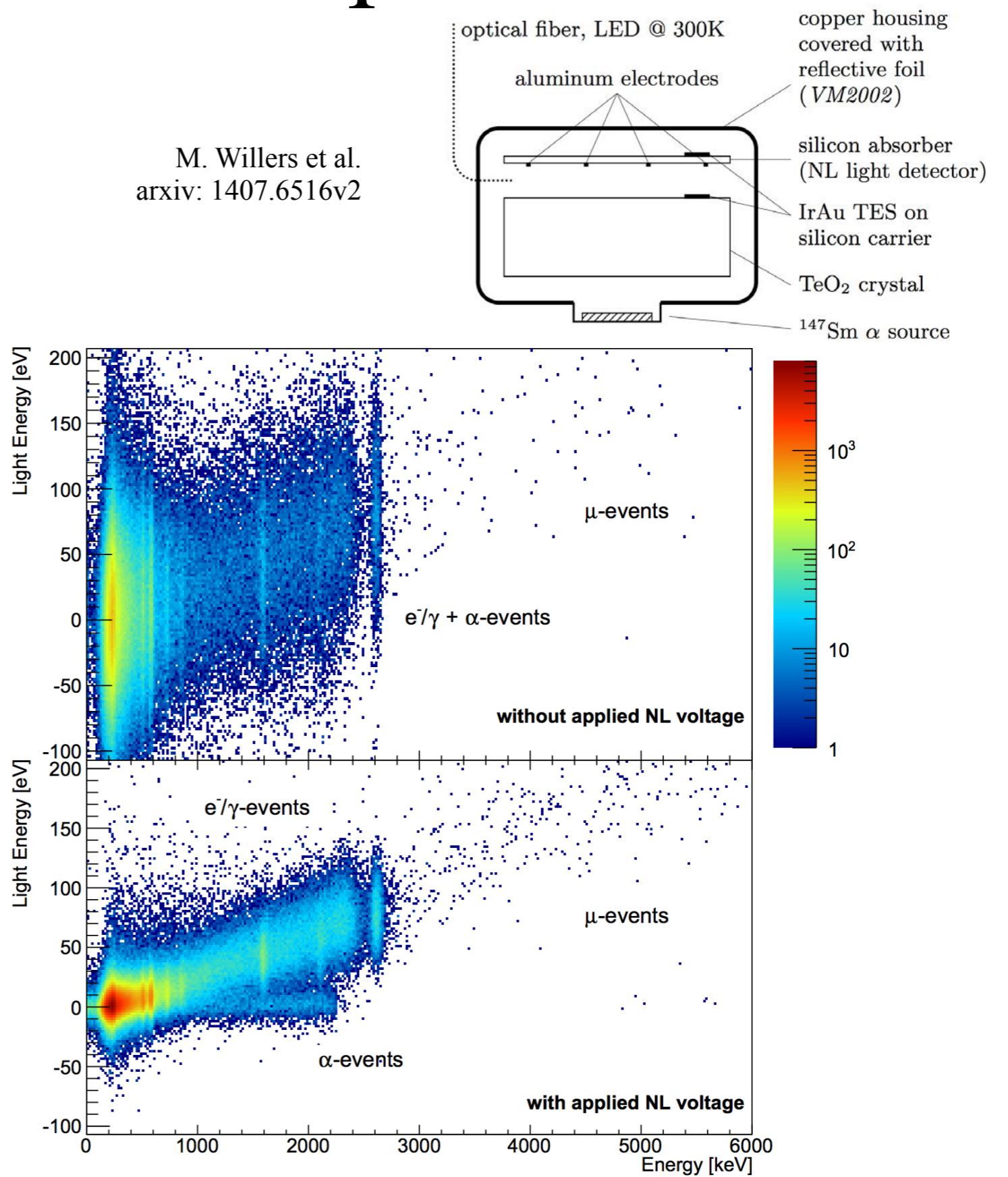
Neganov-Luke amplification

- Devices being developed in framework of both direct DM searches (EURECA) and $0\nu\beta\beta$ searches (CUPID).
- Semiconductor absorber equipped with electrodes, biased.
- Created electrons & holes drifted to create stronger thermal signal:

$$G = 1 + \frac{e \cdot V_{NL}}{E_{ph}/\eta}$$

η —quantum efficiency of electron-hole pair production, E_{ph} —photon energy, V_{NL} —applied voltage

- Likely essential for Cherenkov bolometers for $0\nu\beta\beta$.
- Also seriously improves CaWO_4 scintillating bolometer performance.



thank you for your attention