

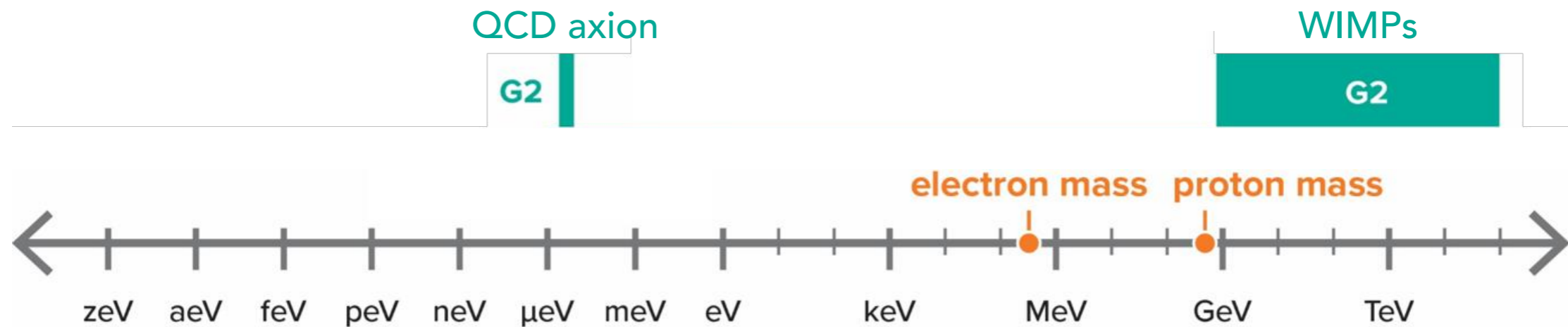
# Emerging Concepts in Dark Matter Detection

Tongyan Lin  
UCSD

February 22, 2023  
P5 Town Hall

# Dark Matter Landscape

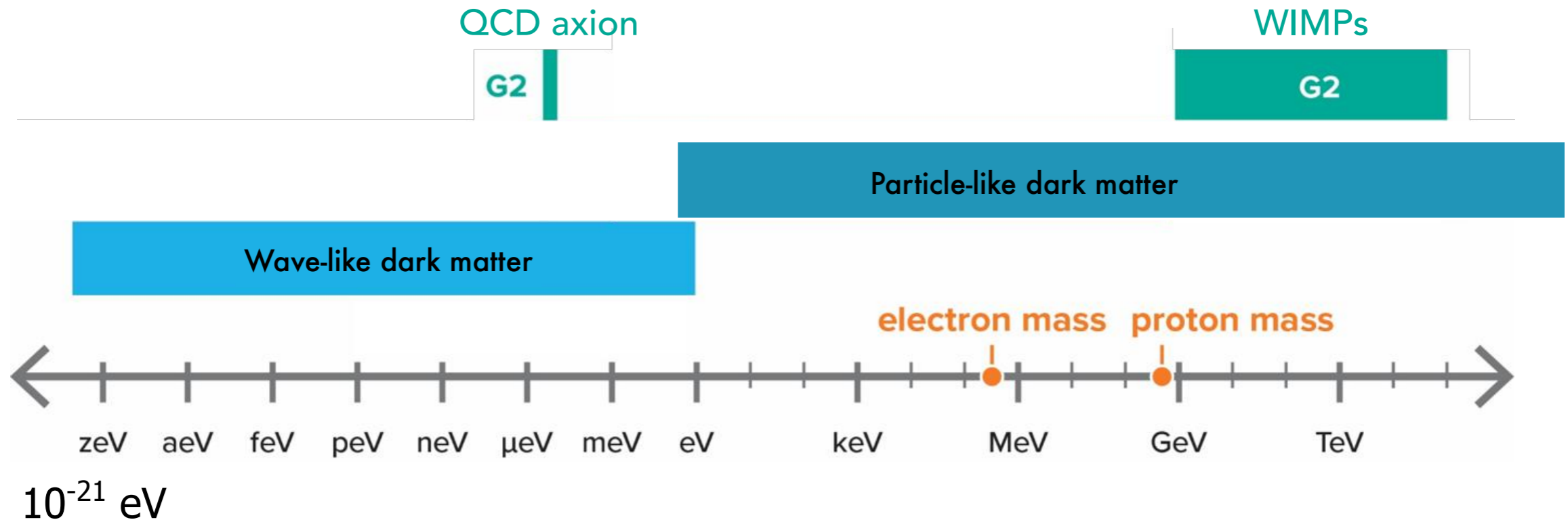
## 2014 P5 recommendation



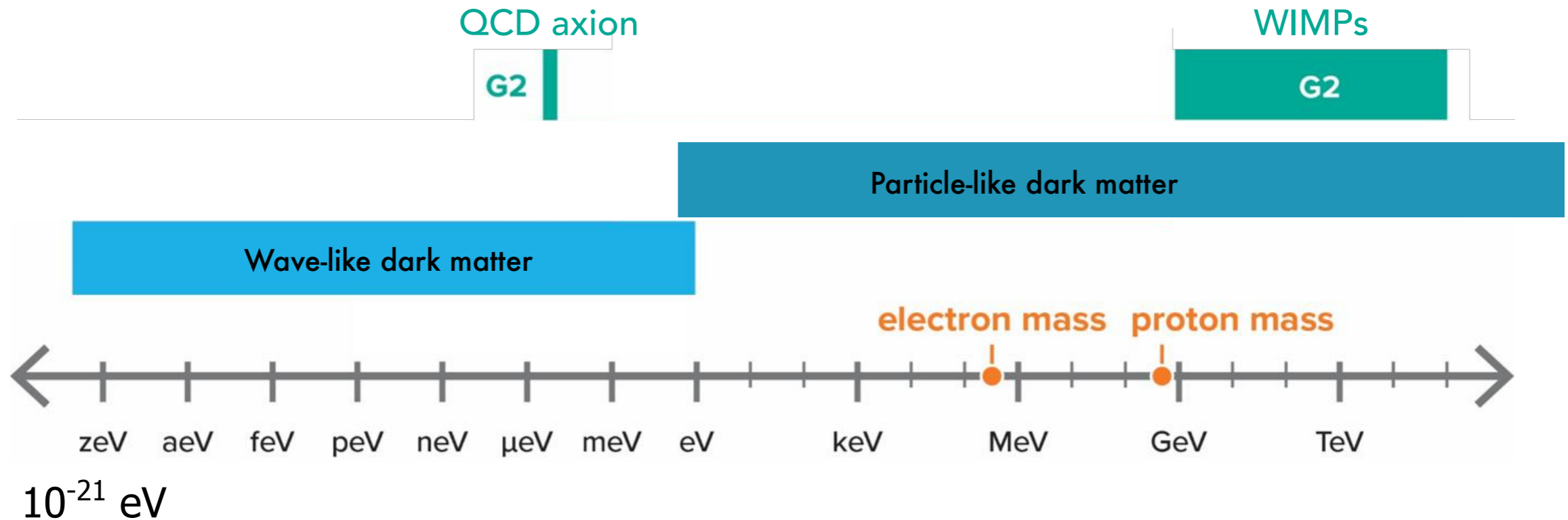
Excellent reasons to target DM in these mass ranges  
Well-honed tools, familiar theory motivations

Additional P5 recommendation: portfolio of smaller projects

# A wider landscape of dark matter candidates

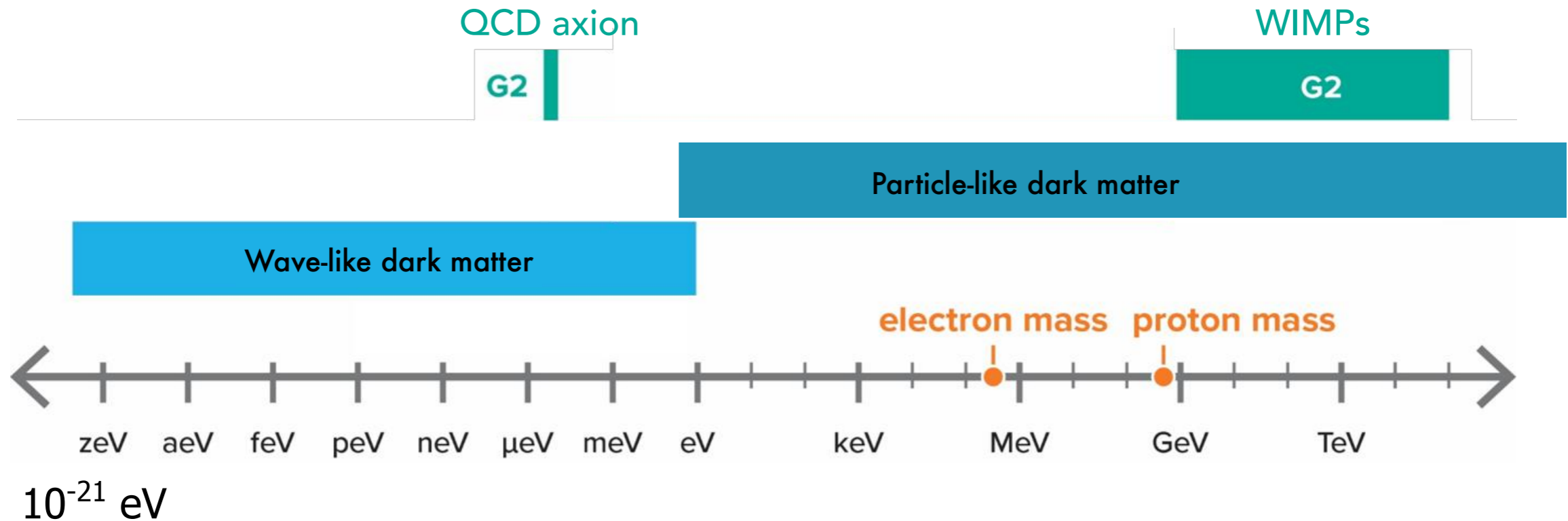


# A wider landscape of dark matter candidates



**New theory targets and tools to probe them**

# A wider landscape of dark matter candidates

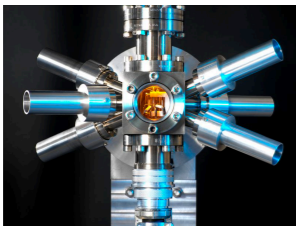


## New theory targets and tools to probe them

Previous P5 recommendation of small-scale projects supported  
Dark Matter New Initiatives (DMNI) effort

# New concepts and technologies can open new windows into dark sector physics

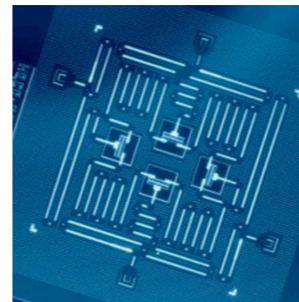
Atomic clocks



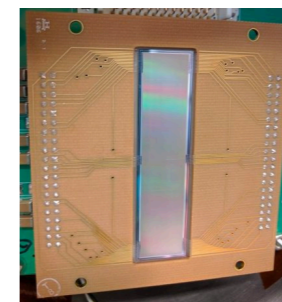
SRF Cavity



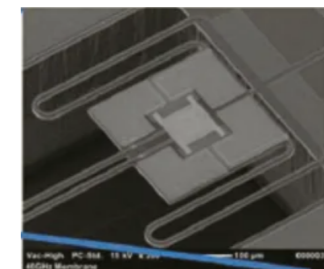
Qubits



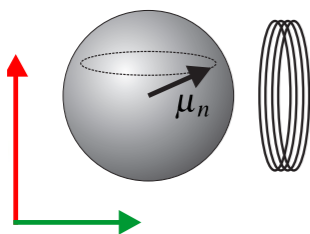
Skipper CCD



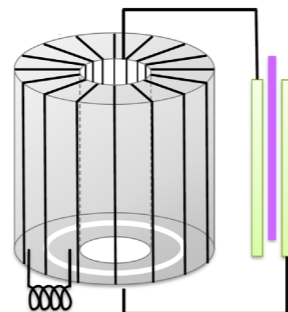
TES



NMR



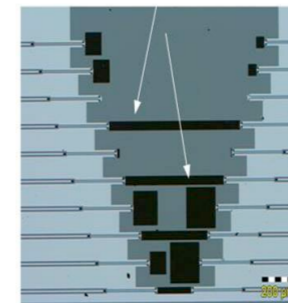
LC Resonators



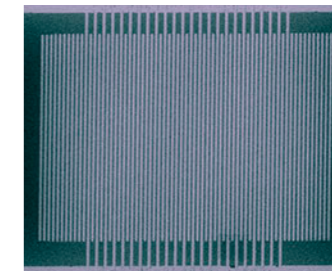
Metamaterials



KIDs



SNSPDs



and many more

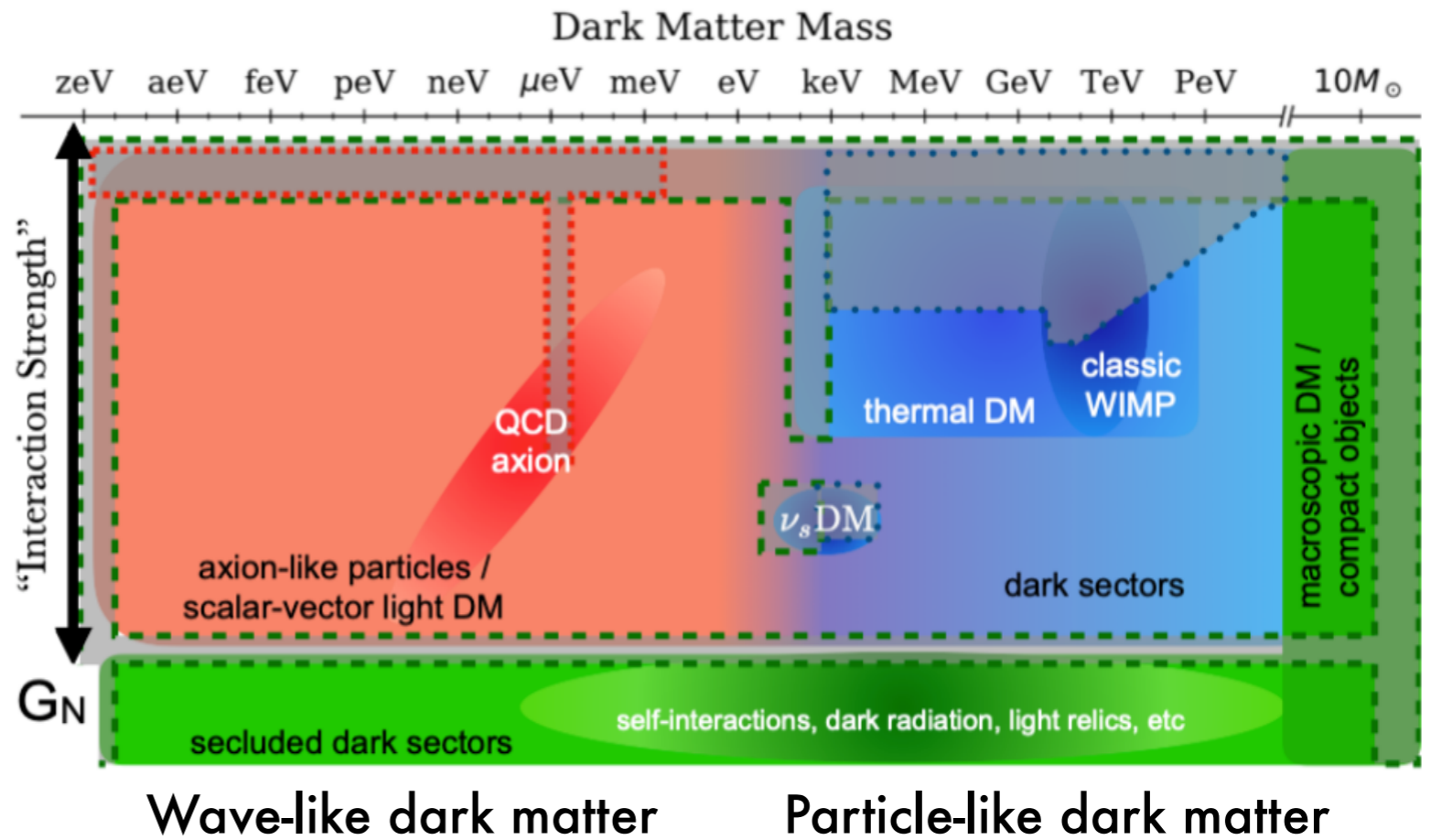
## Small-scale experiments can have a huge impact!

# Cosmic Frontier Summary Report

The HEP community has identified potentially transformative opportunities to address fundamental physics questions via Cosmic Frontier programs. We aspire to **Aim High, Search Wide, and Delve Deep**:

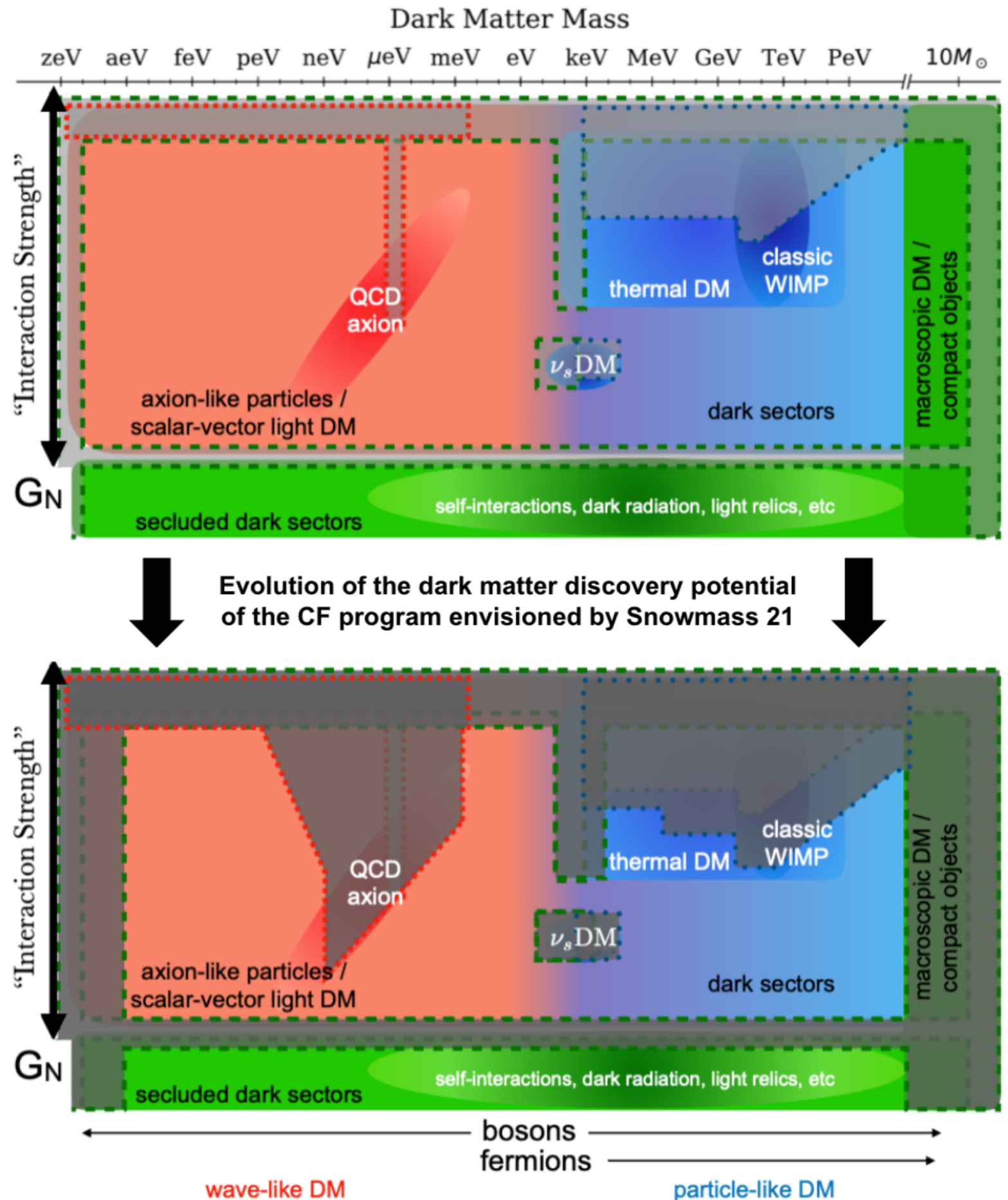
- Complete the CMB-S4 cosmic microwave background experiment and build a large spectroscopic facility (Spec-S5) to study physics including inflation, dark energy, light relics, modifications to general relativity, and dark matter.
- Pursue a broad program investigating the full landscape of dark matter candidates, including implementation of the existing Dark Matter New Initiatives (DMNI) portfolio and development of future DMNI-like programs focused on small projects; investment in new quantum technologies; and engagement of the HEP community in the development and execution of cosmic and indirect searches for dark matter, to take full advantage of the unique opportunities provided by cosmological and astrophysical probes (e.g., Rubin LSST and AugerPrime).
- Scale up mature technologies for weakly-interacting massive particle (WIMP) direct detection, fully exploring the parameter space down to the neutrino fog, and support high-energy gamma-ray telescopes (e.g., SWGO and CTA) to probe thermal WIMPs up to tens-of-TeV mass scales. Move ahead with new, construction-ready DMNI experiments with the capacity to probe the QCD axion over most of its viable mass range.

Opportunity for orders of magnitude coverage in sensitivity from portfolio of small projects





Opportunity for orders of magnitude coverage in sensitivity from portfolio of small projects



# Relic abundance from dark sectors

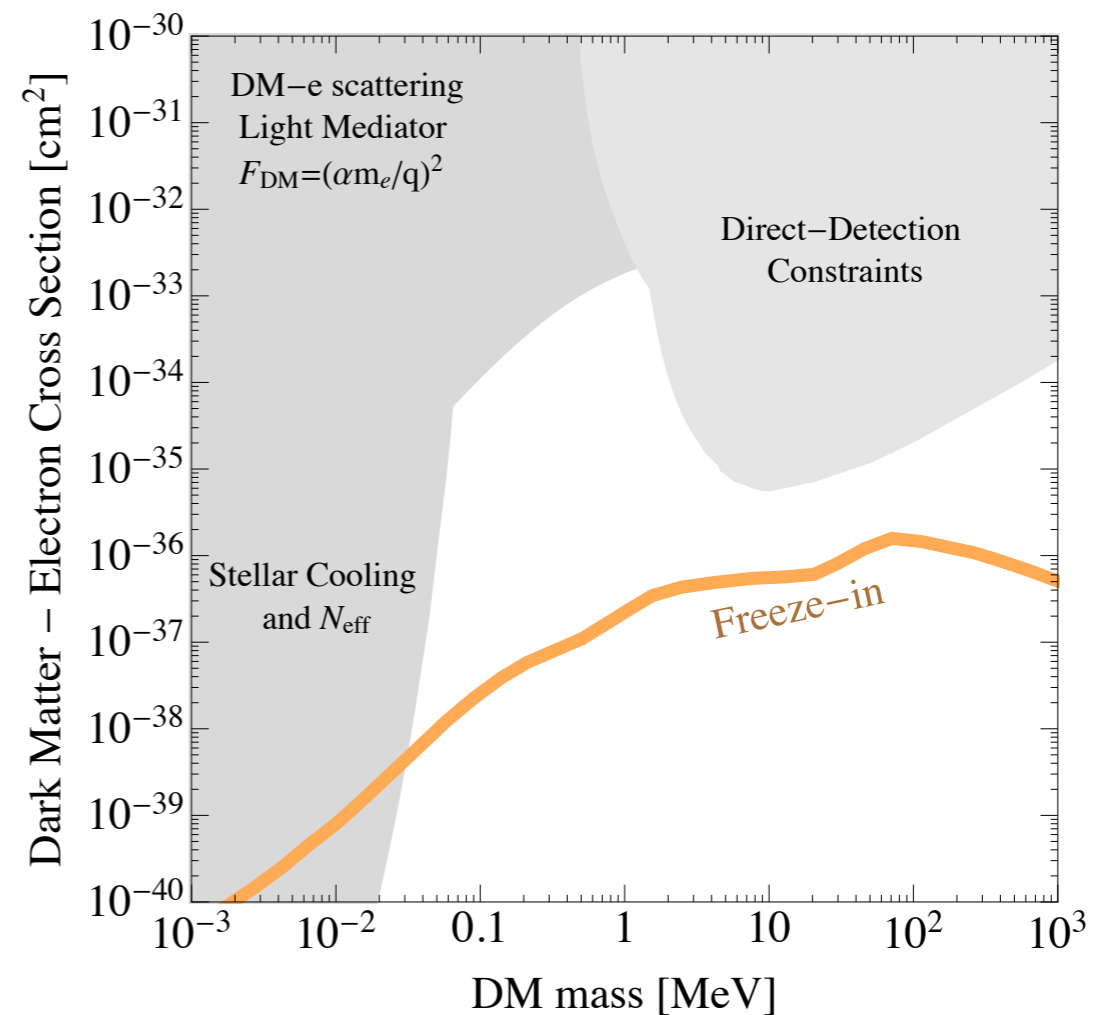
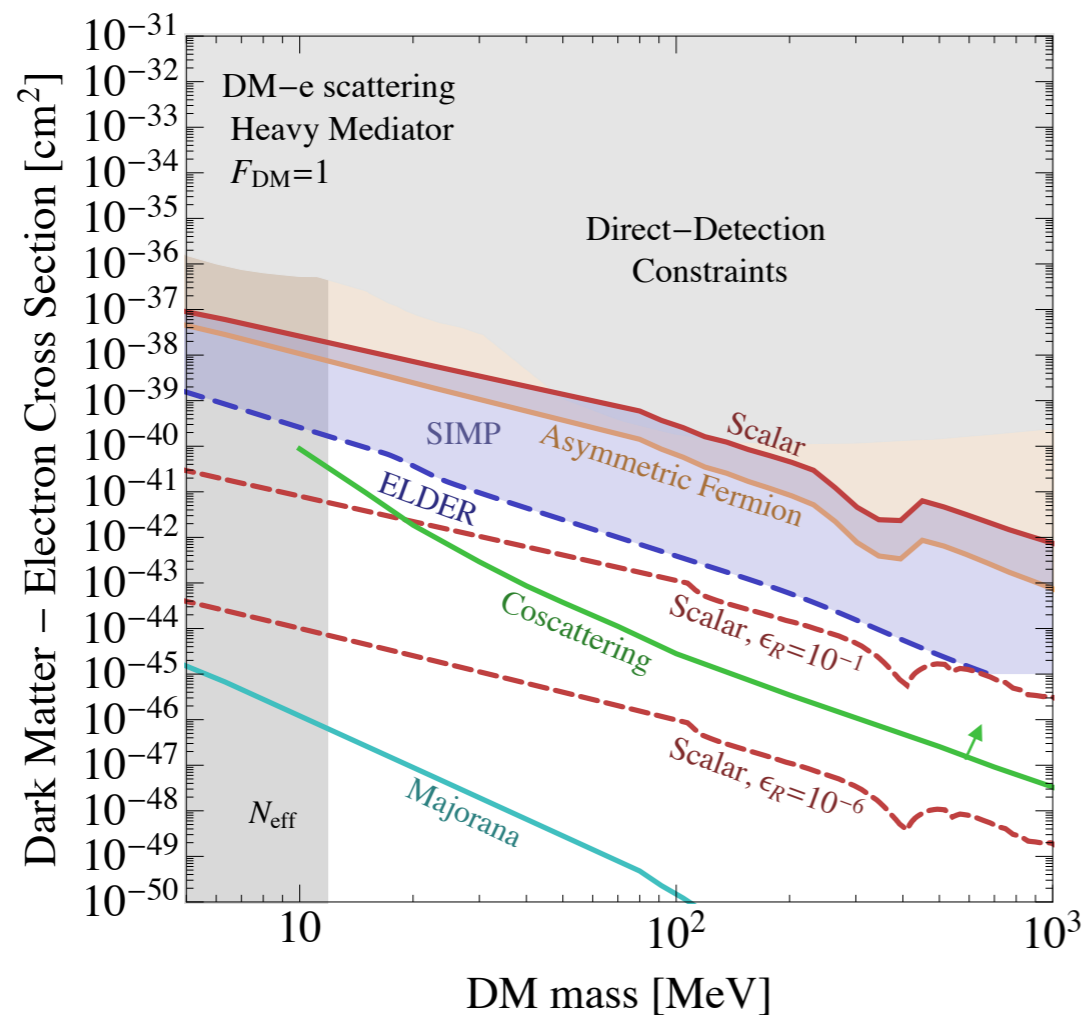
Significant theory progress in mechanisms for DM genesis.



# Relic abundance from dark sectors

Significant theory progress in mechanisms for DM genesis.

For  $\text{keV} \lesssim m_\chi \lesssim \text{GeV}$ , targets in DM-electron scattering:



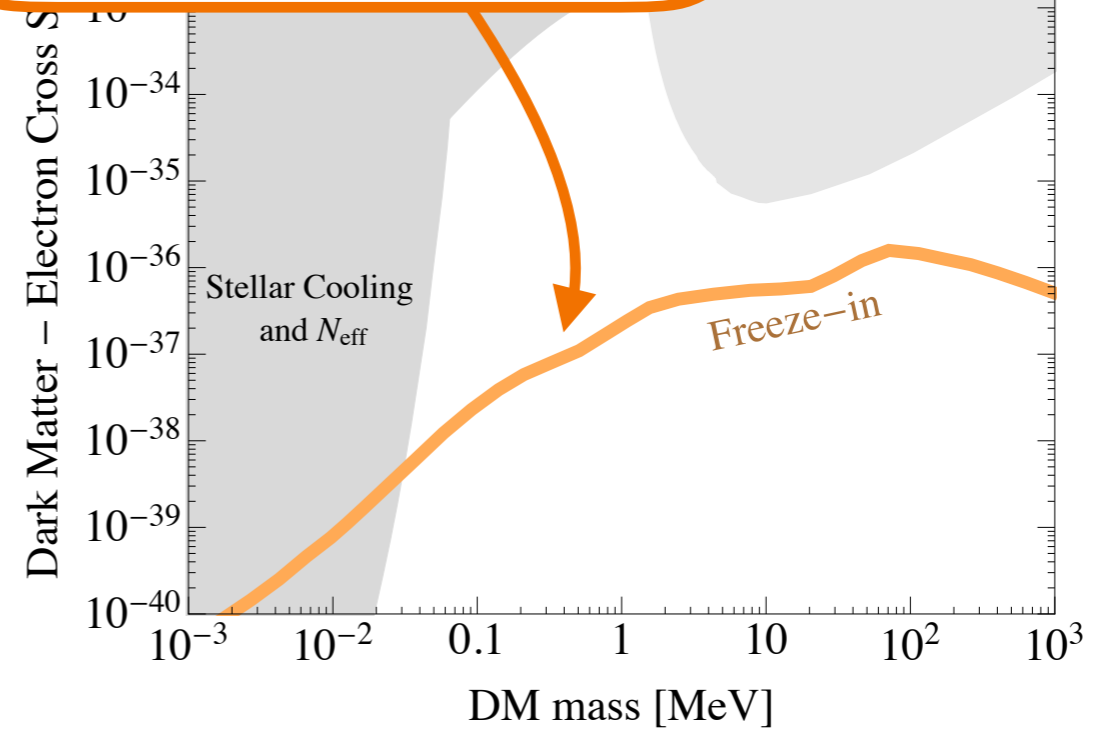
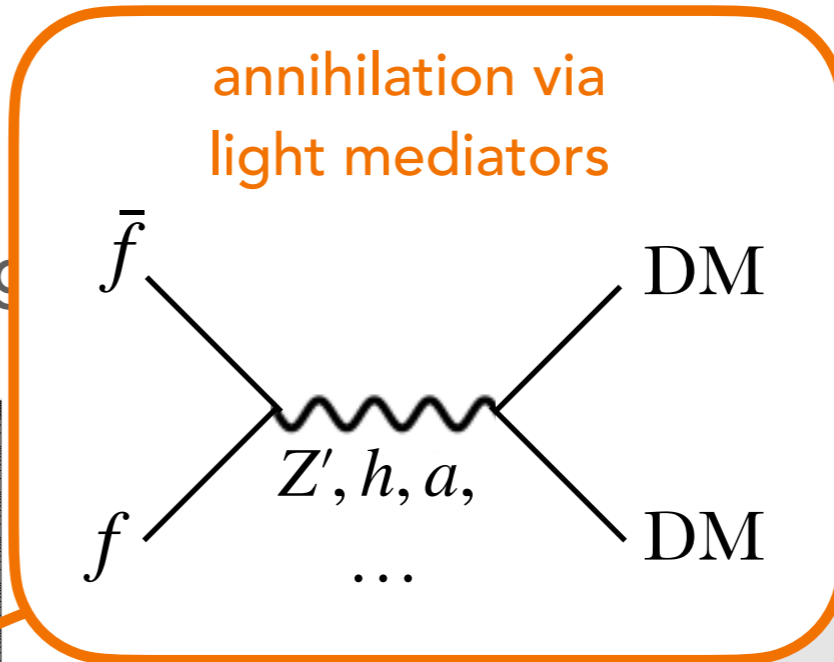
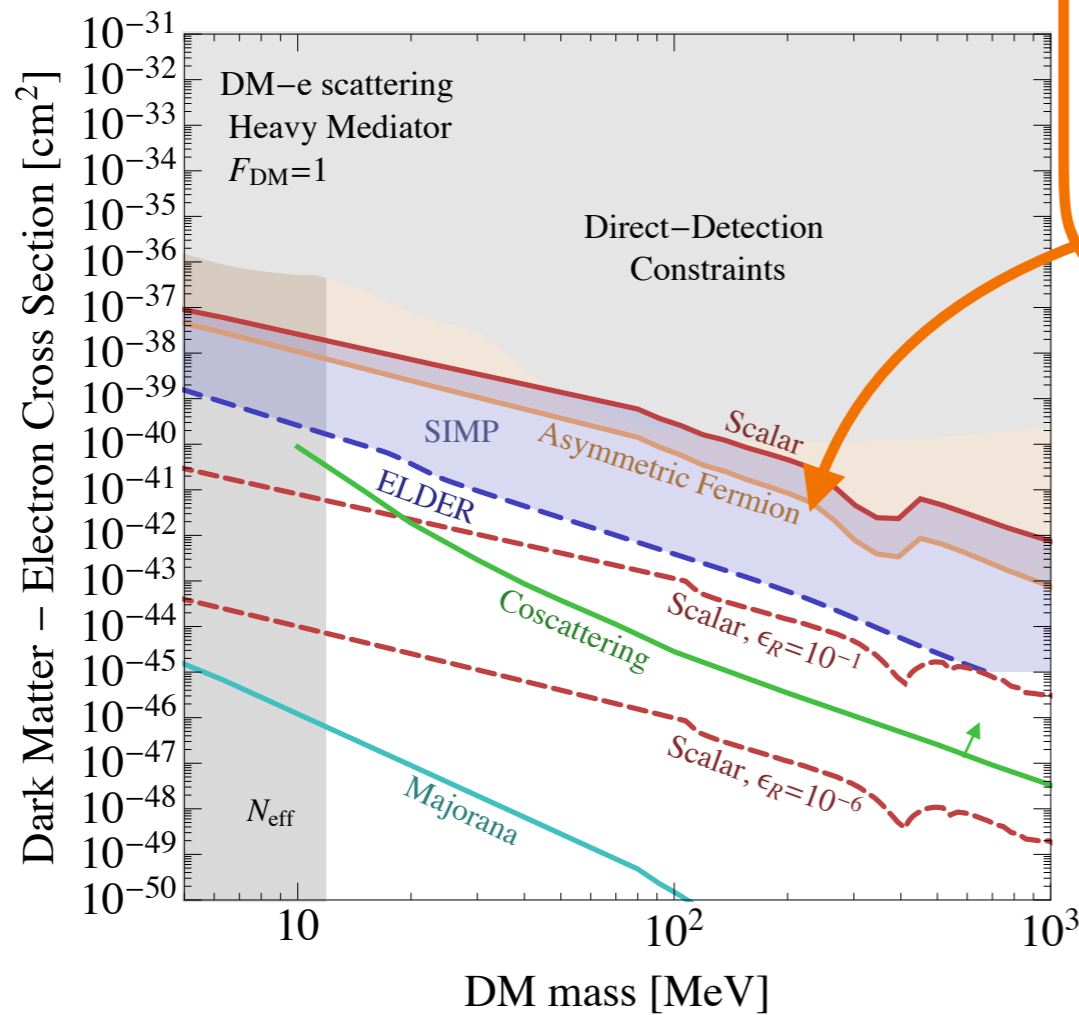
# Relic abundance from dark sectors

Significant theory progress

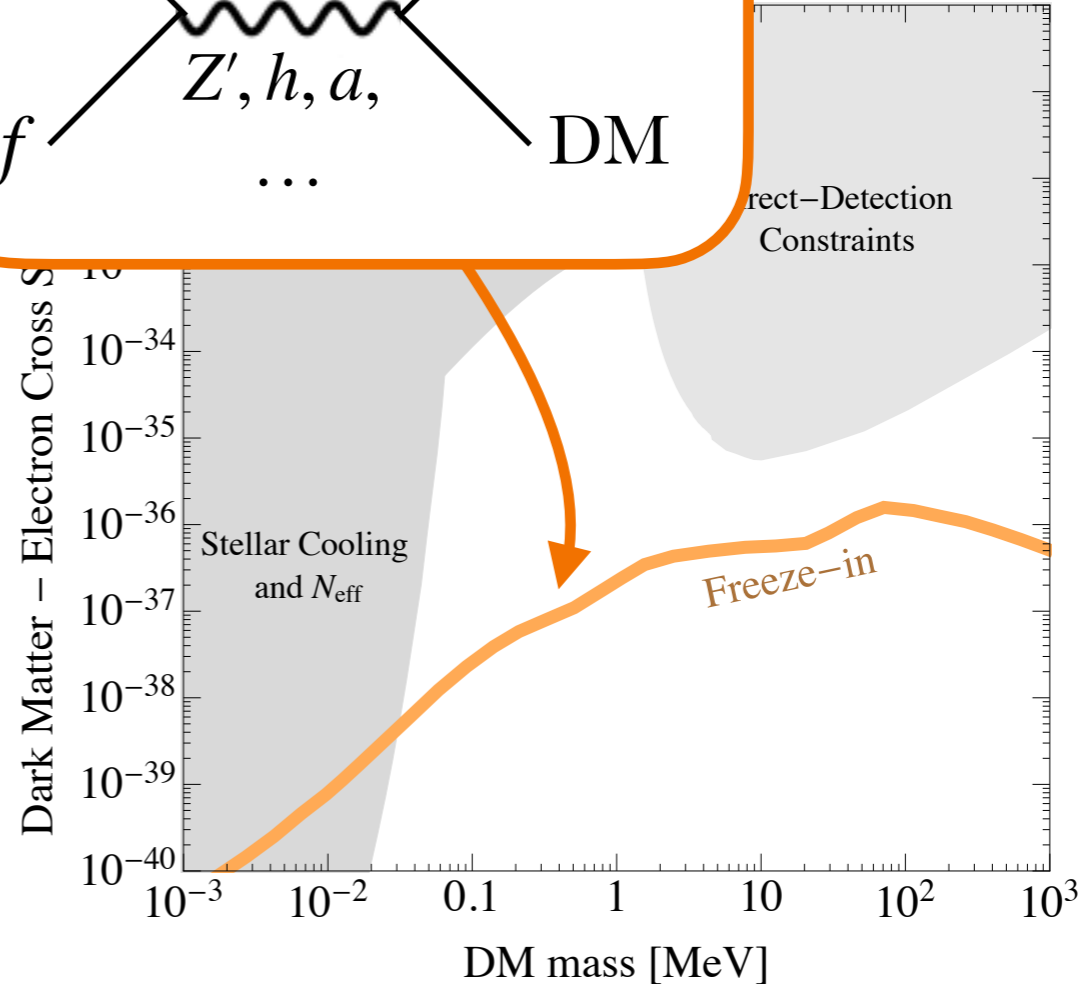
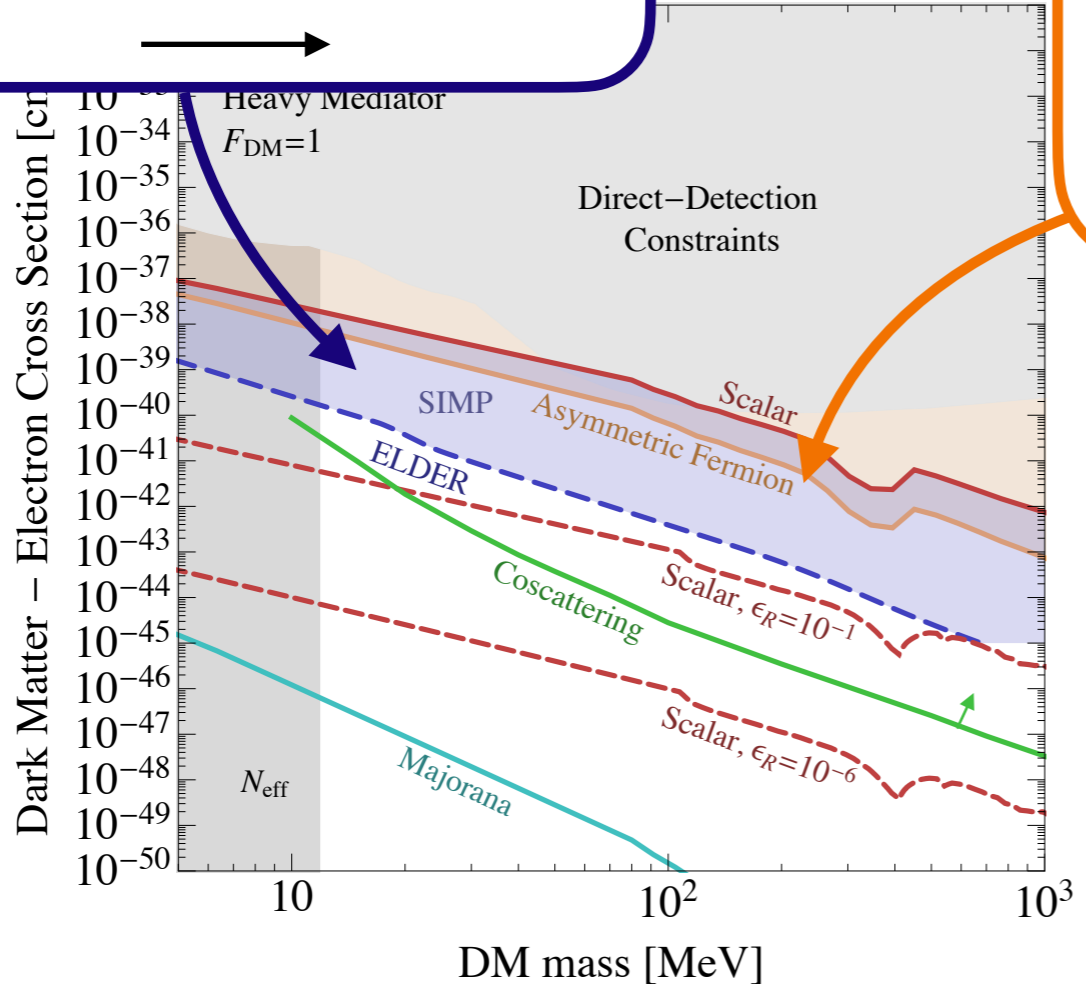
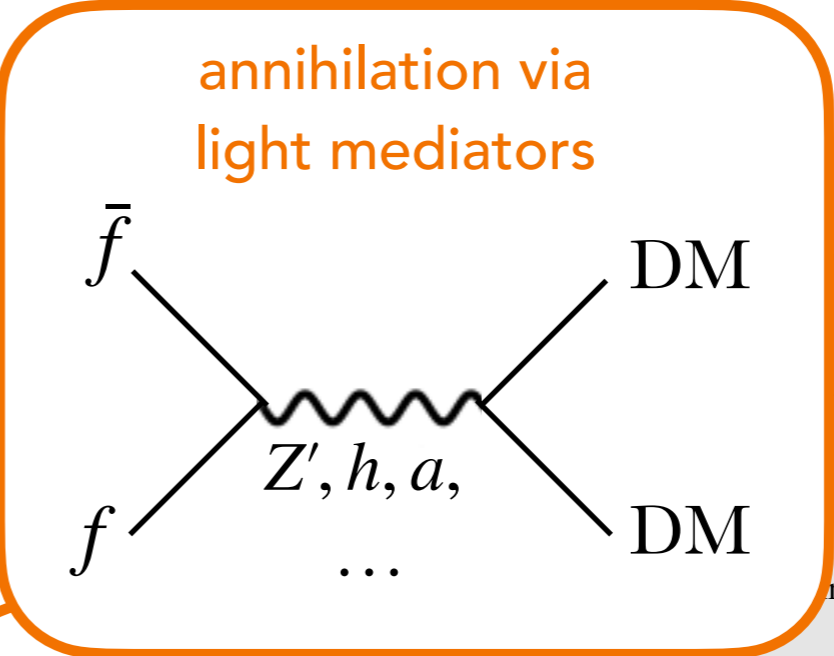
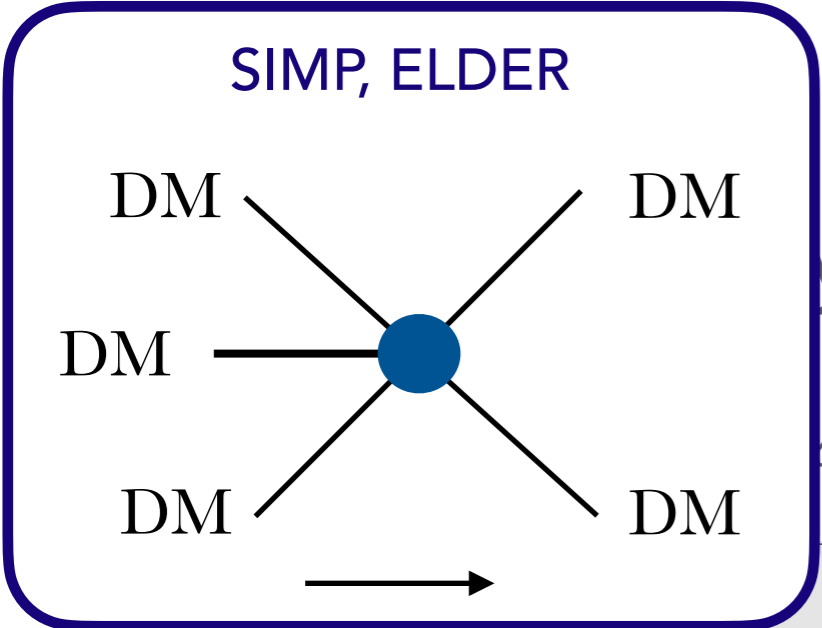
For  $\text{keV} \lesssim m_\chi \lesssim \text{GeV}$ , target

genesis.

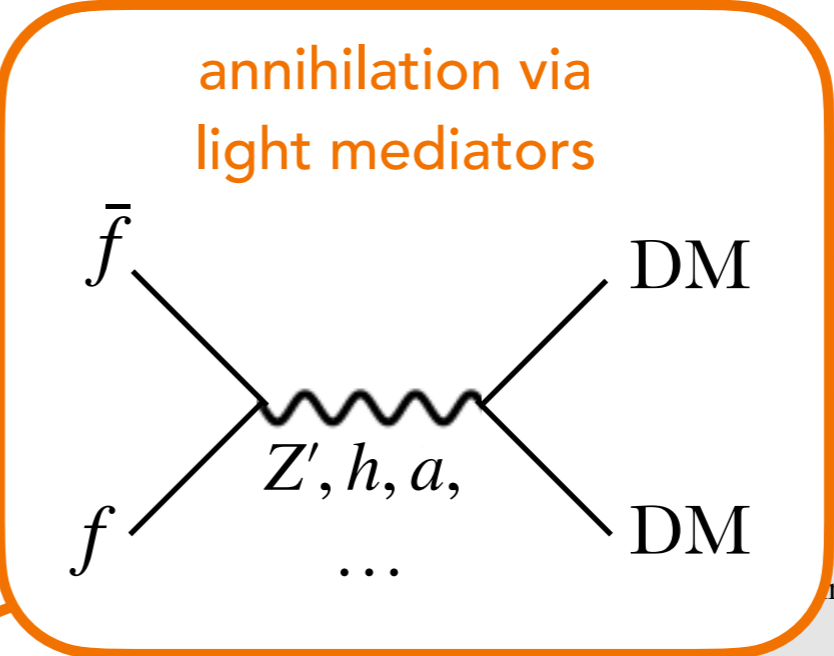
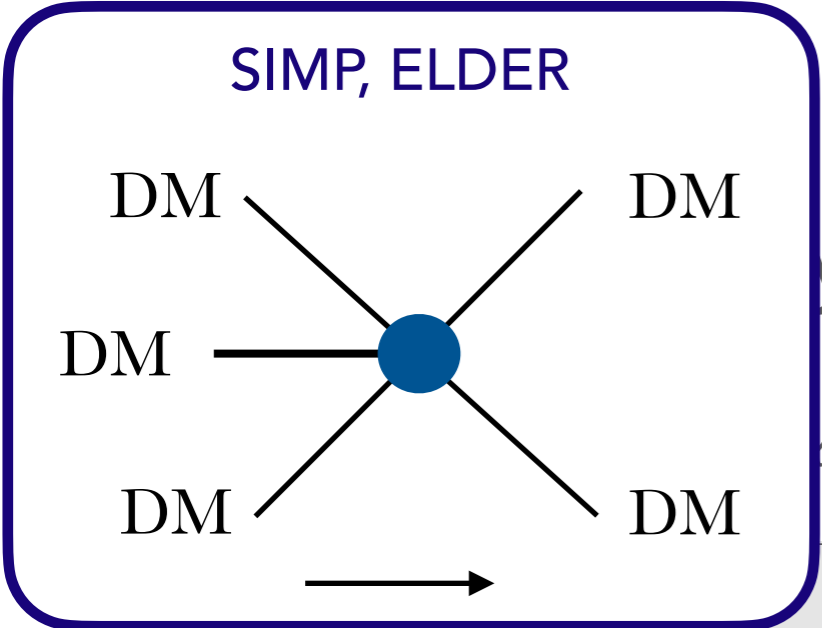
ing:



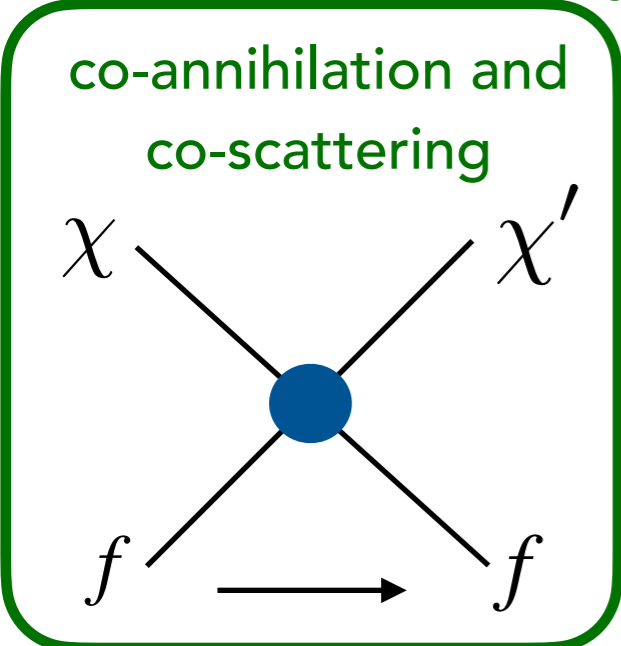
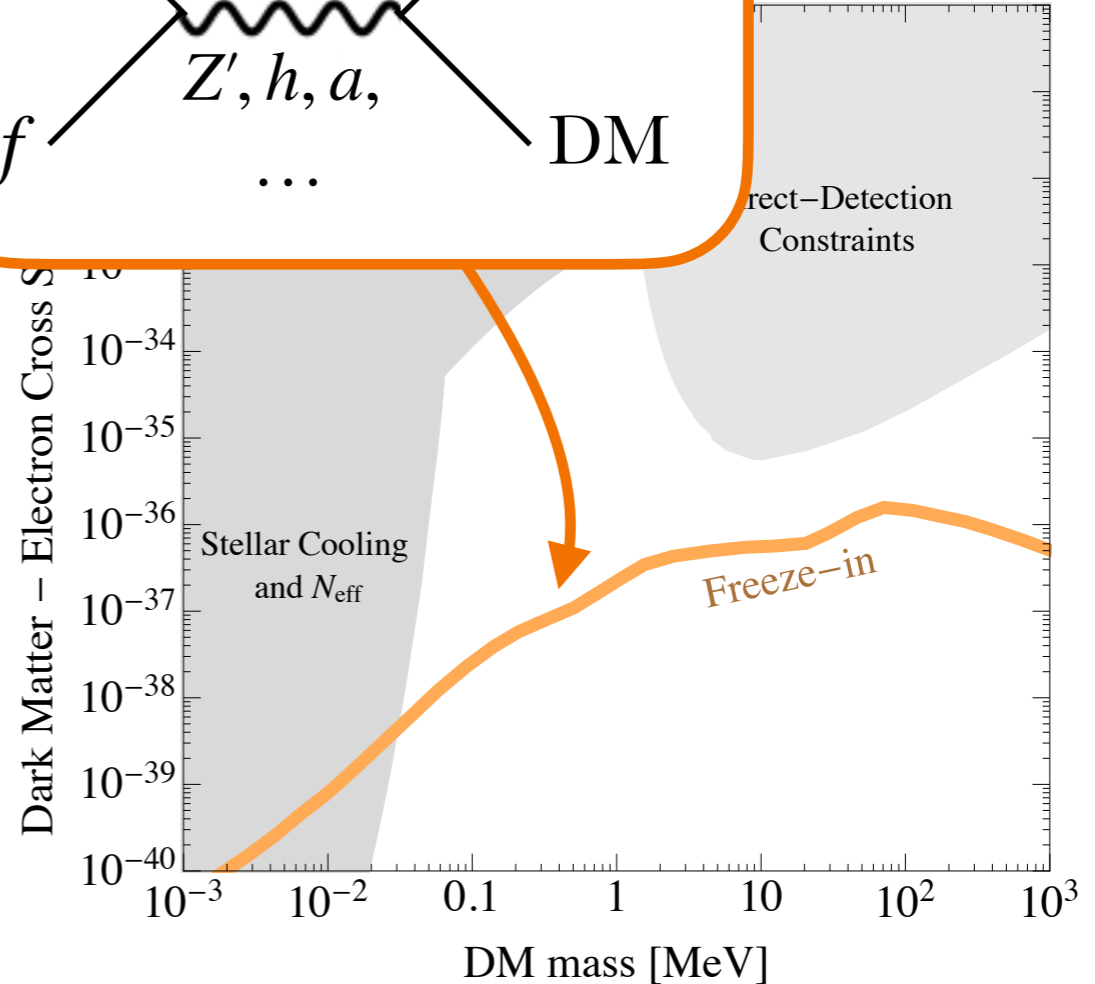
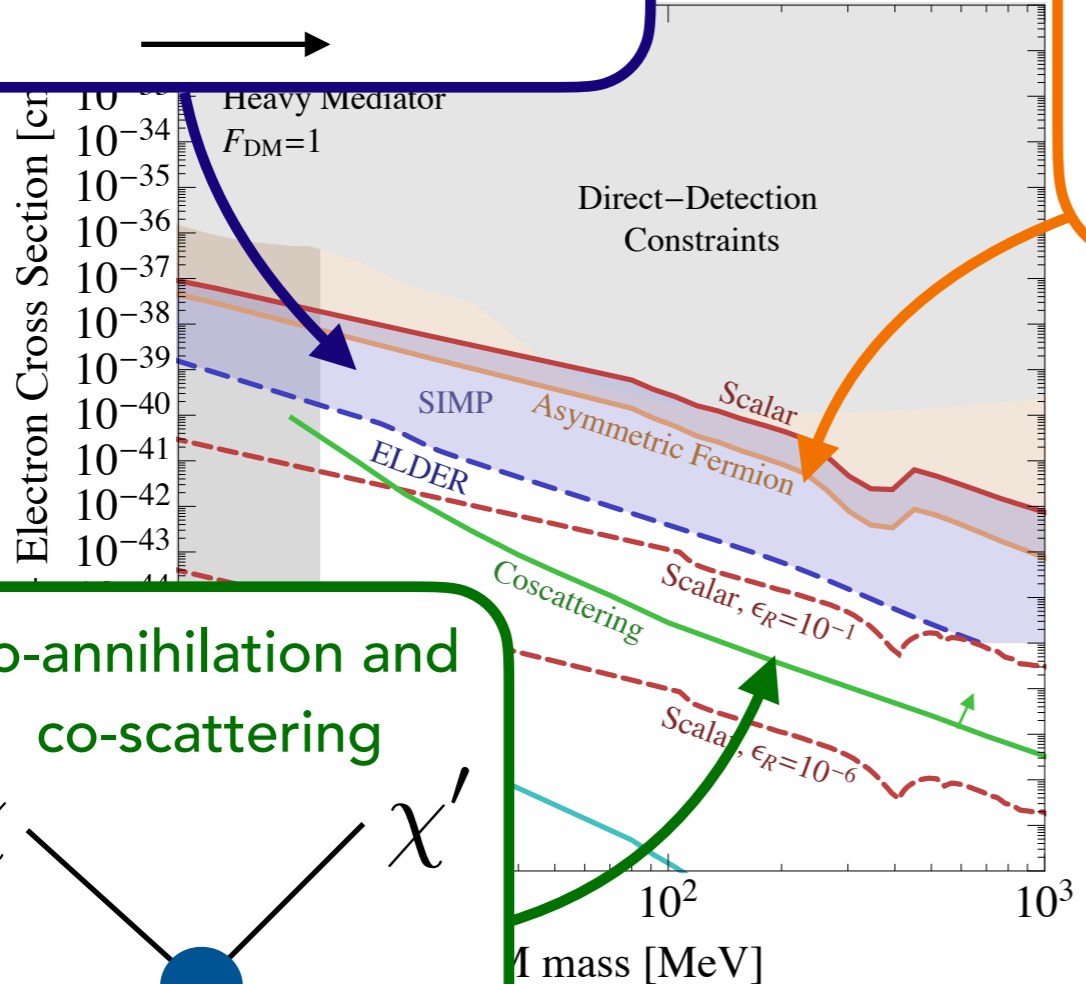
# Dark Matter production from dark sectors



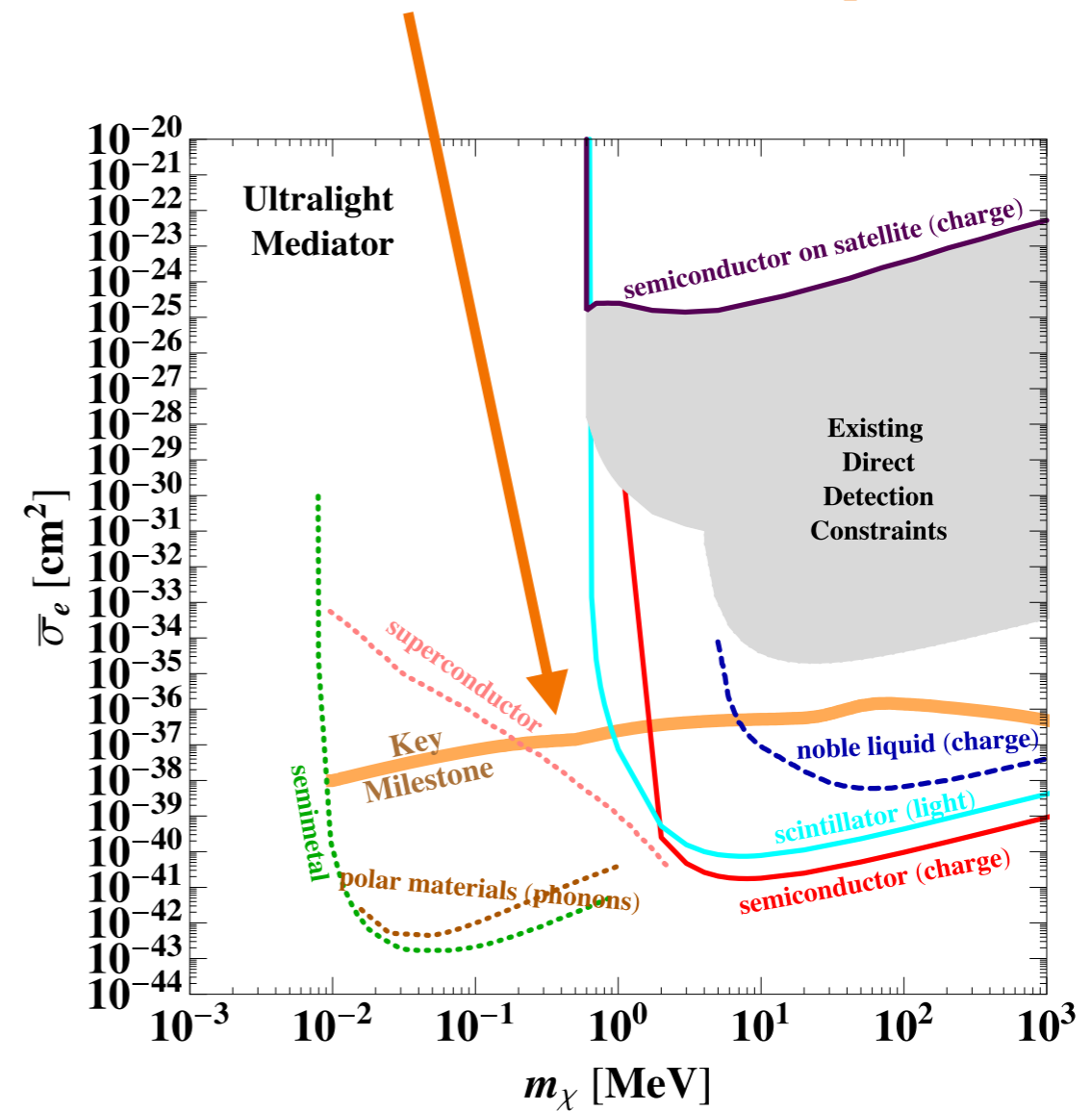
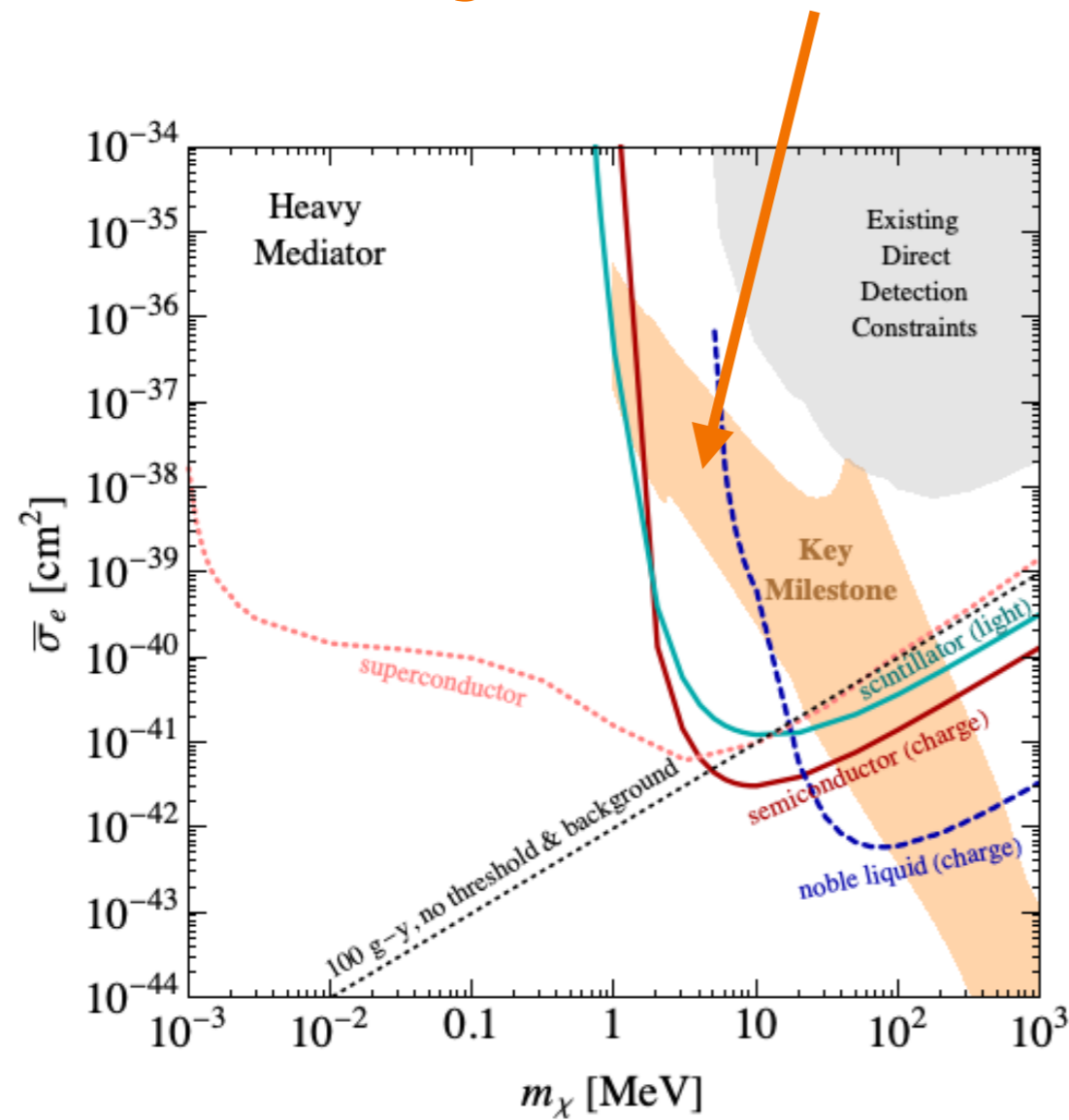
# Dark Matter production from dark sectors



genesis.  
ing:



# Opportunity to probe DM targets with a variety of small-scale detector concepts



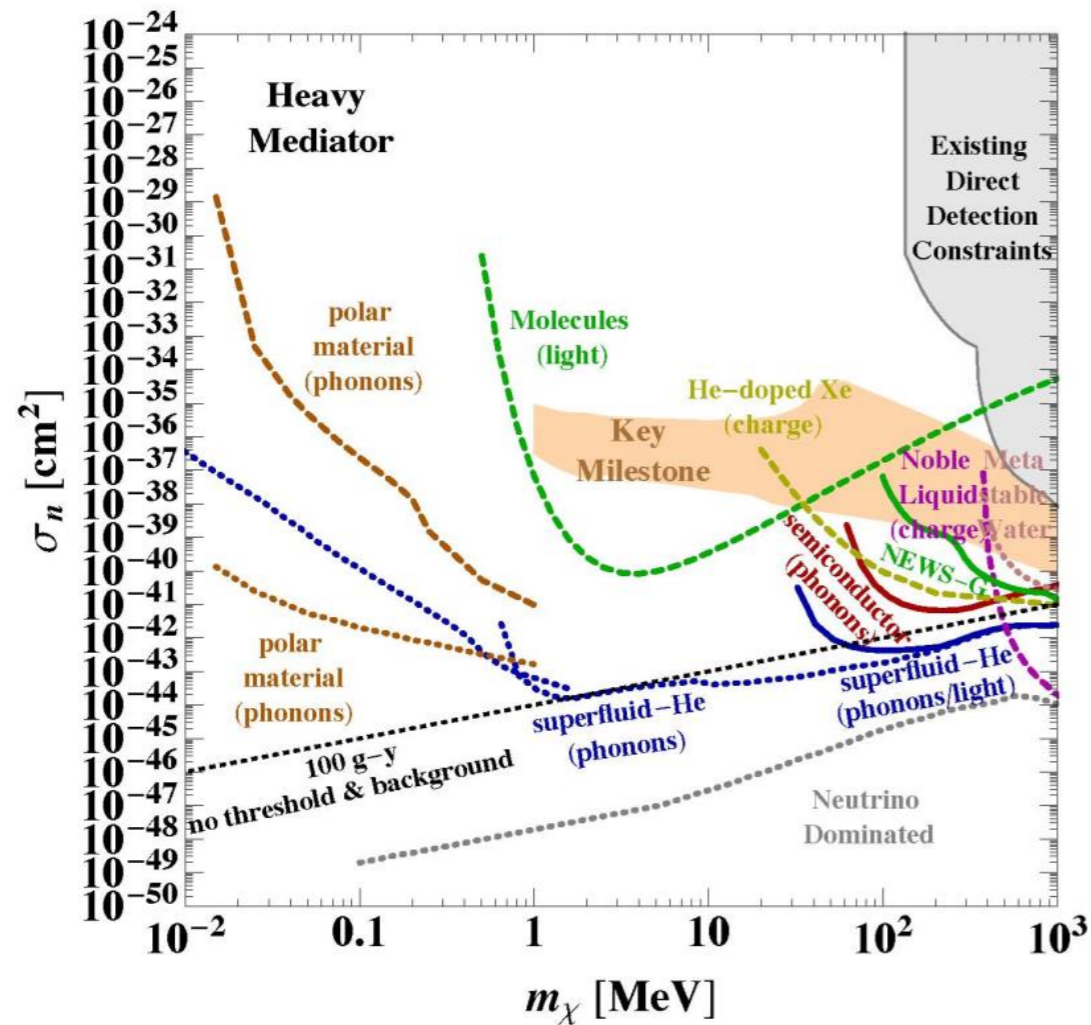
DM-electron interactions for  $m_{\text{DM}} \gtrsim \text{keV}$

Lines assume O(kg-yr) exposure

# Dark sector interactions

DM-nucleon interactions

$$m_{\text{DM}} \gtrsim \text{keV}$$

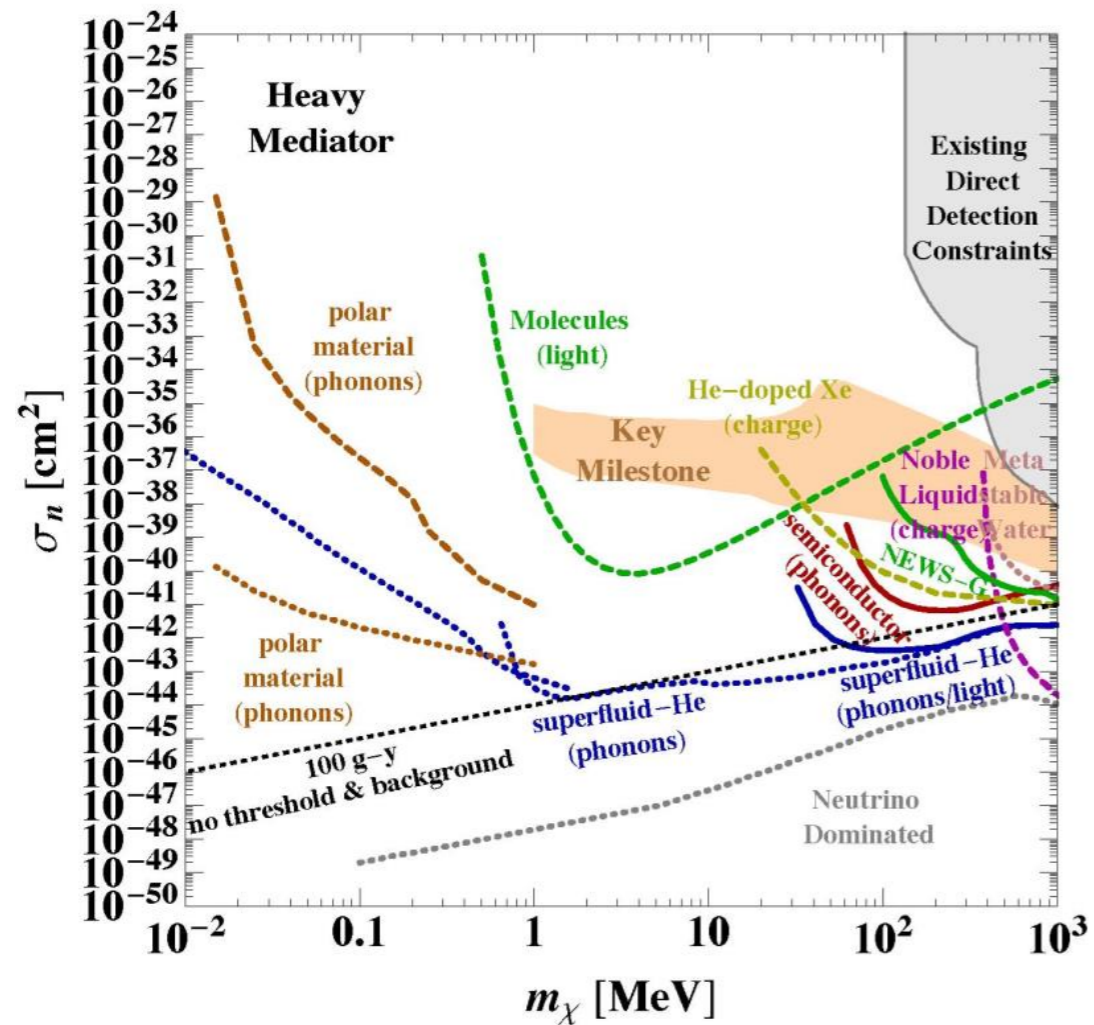




# Dark sector interactions

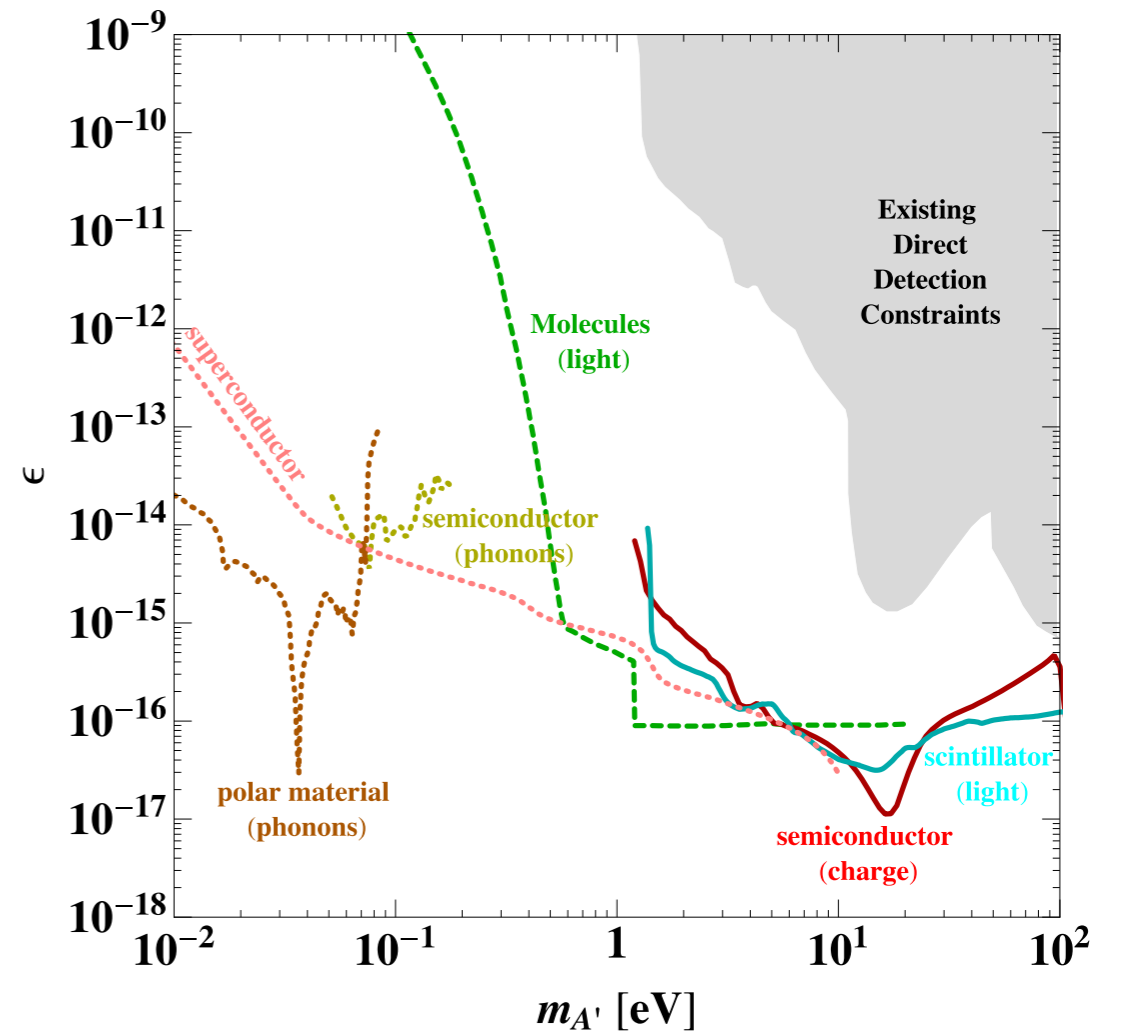
DM-nucleon interactions

$$m_{\text{DM}} \gtrsim \text{keV}$$



Dark photon DM absorption

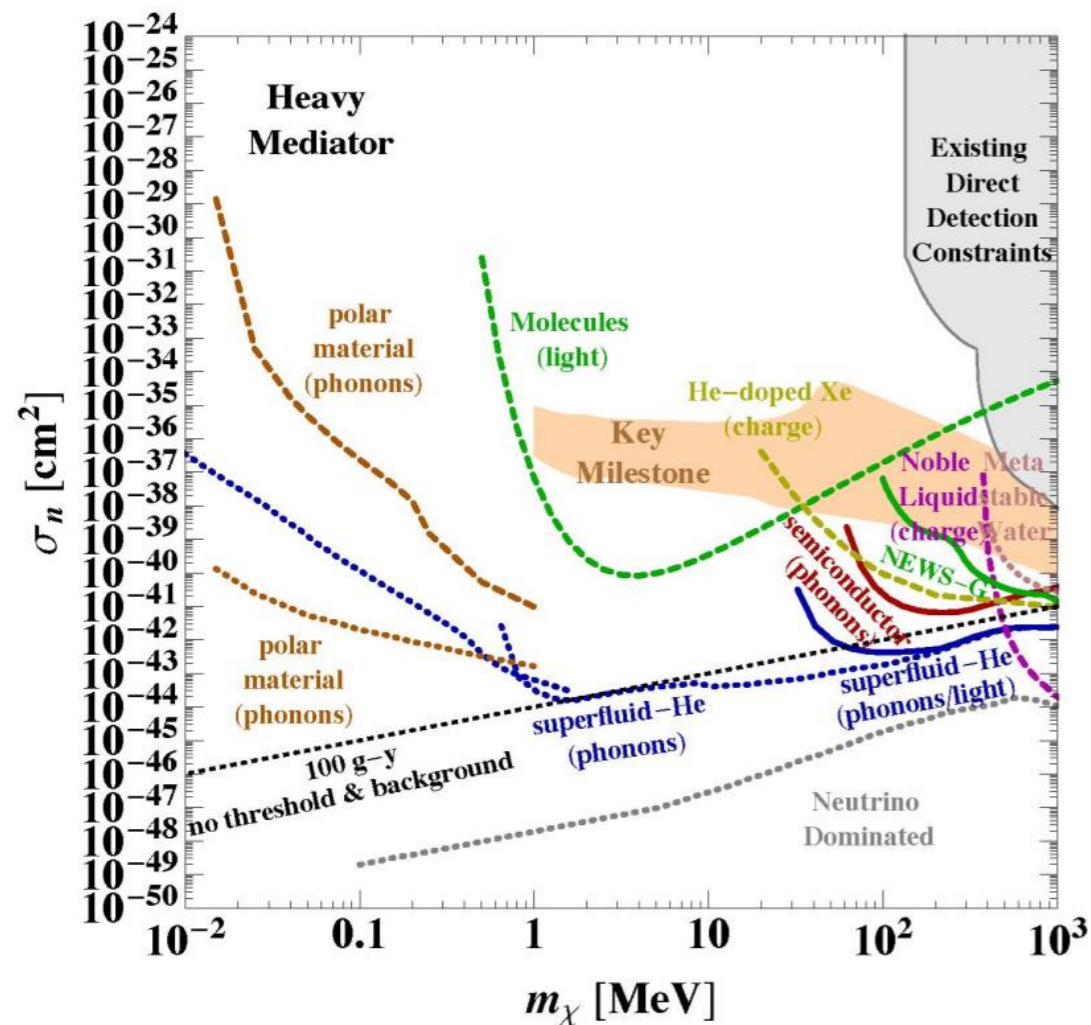
$$\text{meV} \lesssim m_{\text{DM}} \lesssim \text{keV}$$



# Dark sector interactions

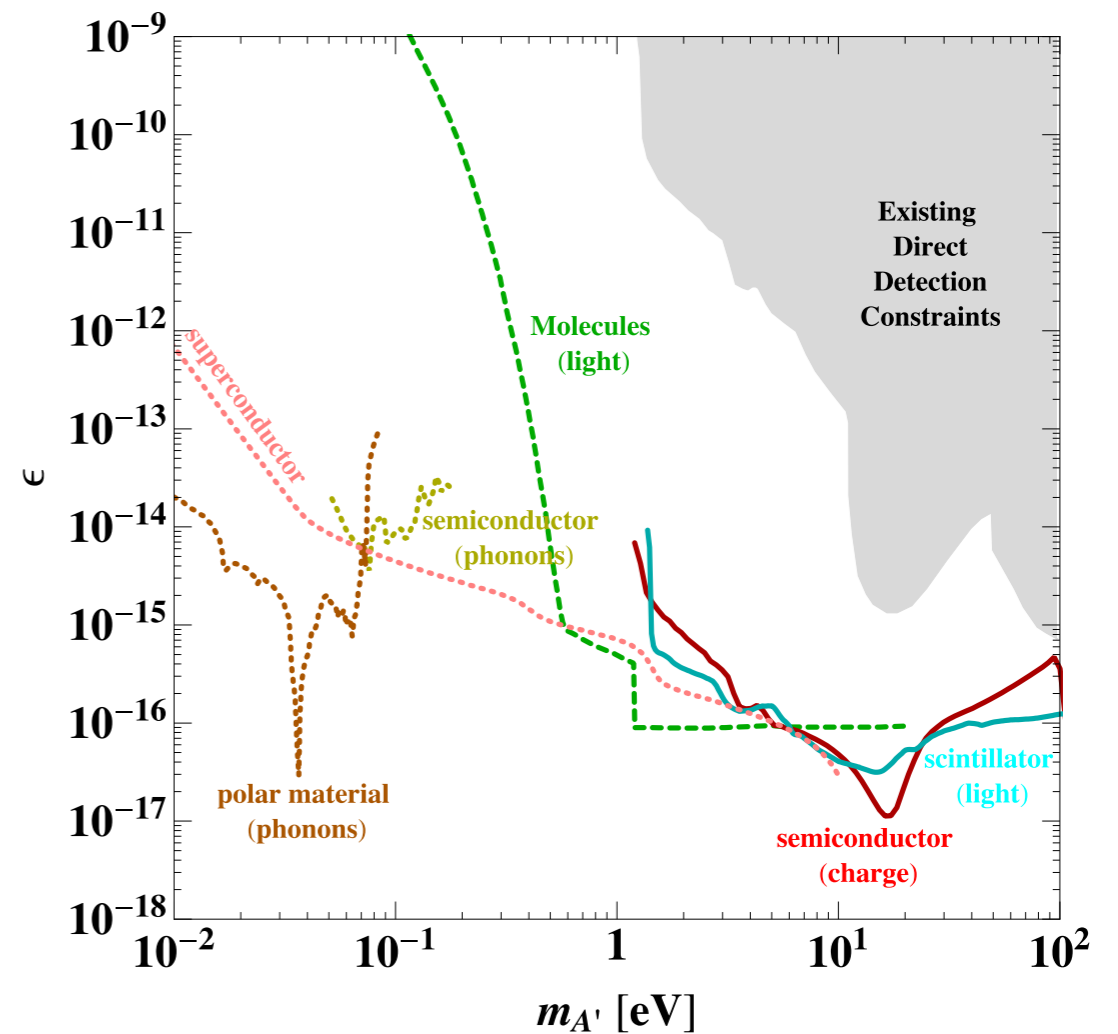
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Dark photon DM absorption

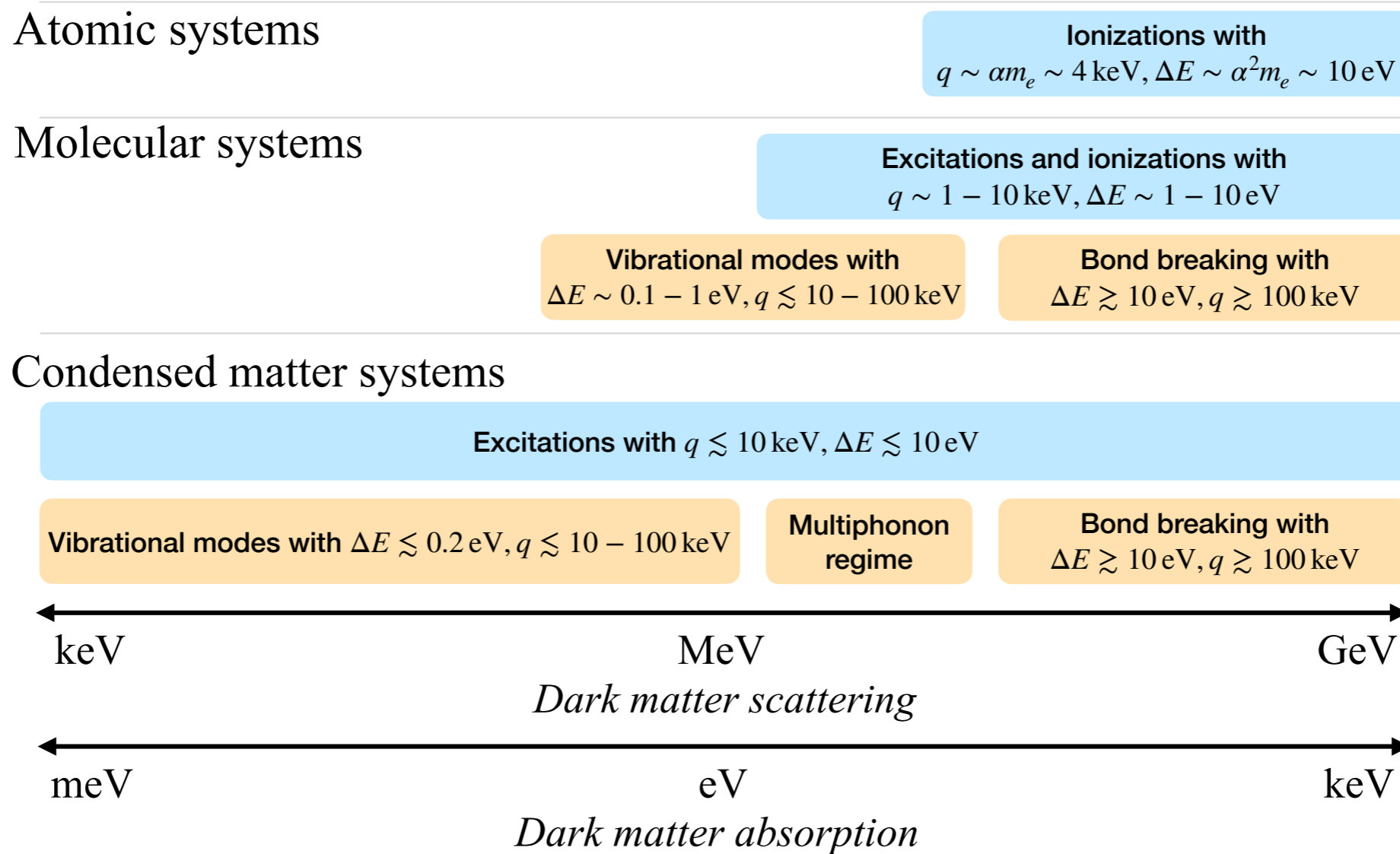
$$\text{meV} \lesssim m_{\text{DM}} \lesssim \text{keV}$$



Many concepts from theorists *and* experimentalists

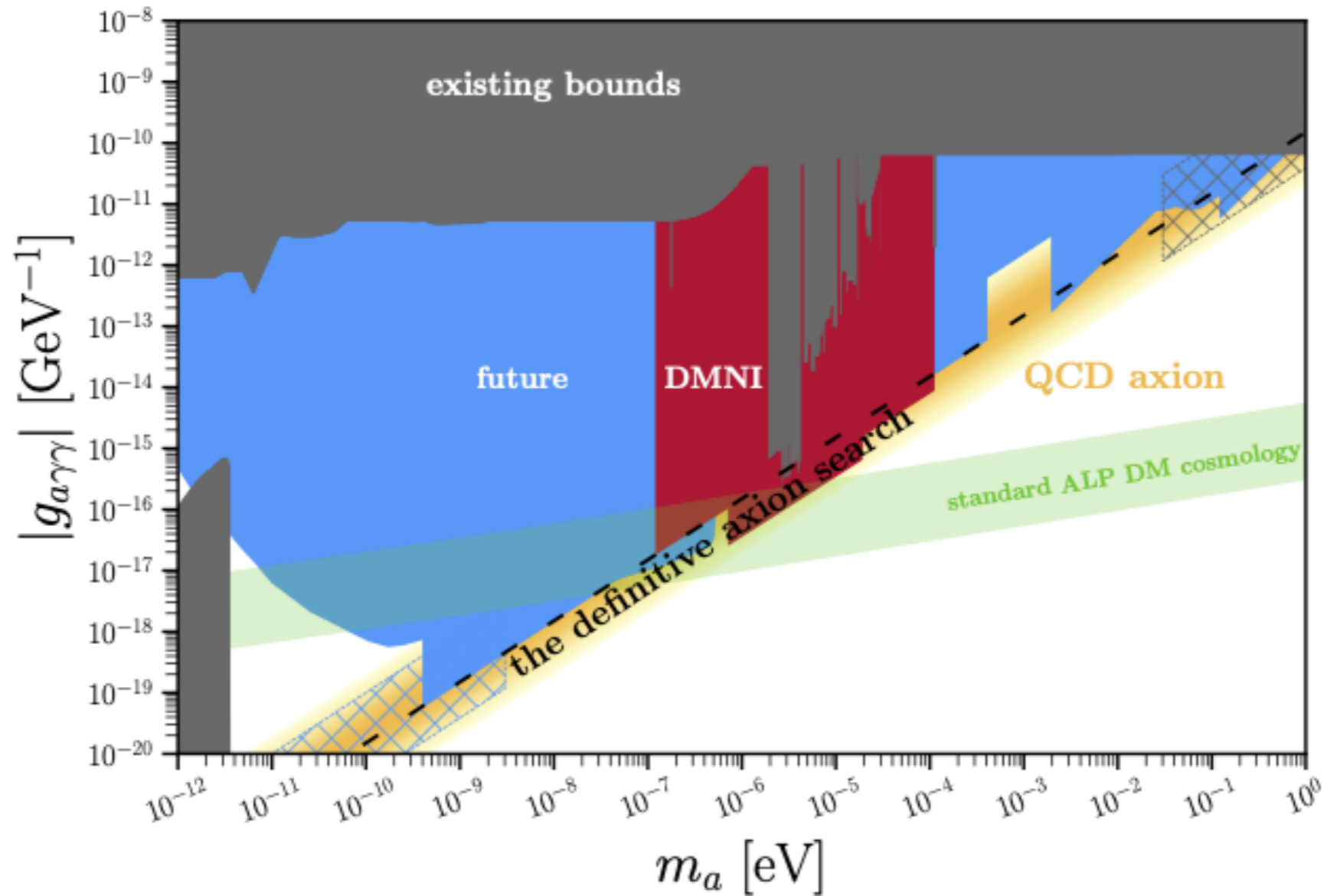
# Detection concepts for particle-like DM

Utilize diversity of excitation modes in nature and advances in neighboring fields including materials science, quantum sensing.

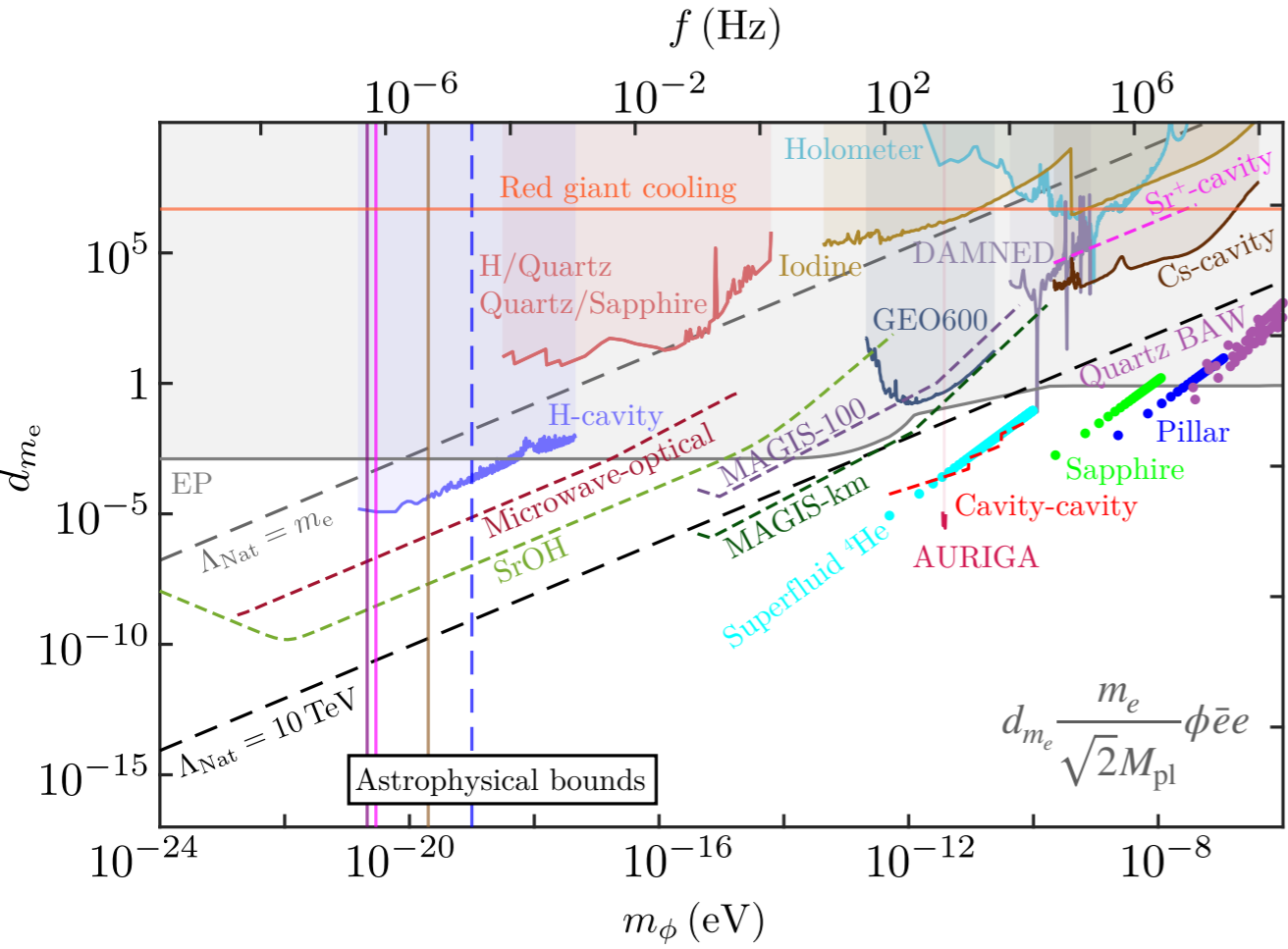


# Bosonic (wave-like) dark matter

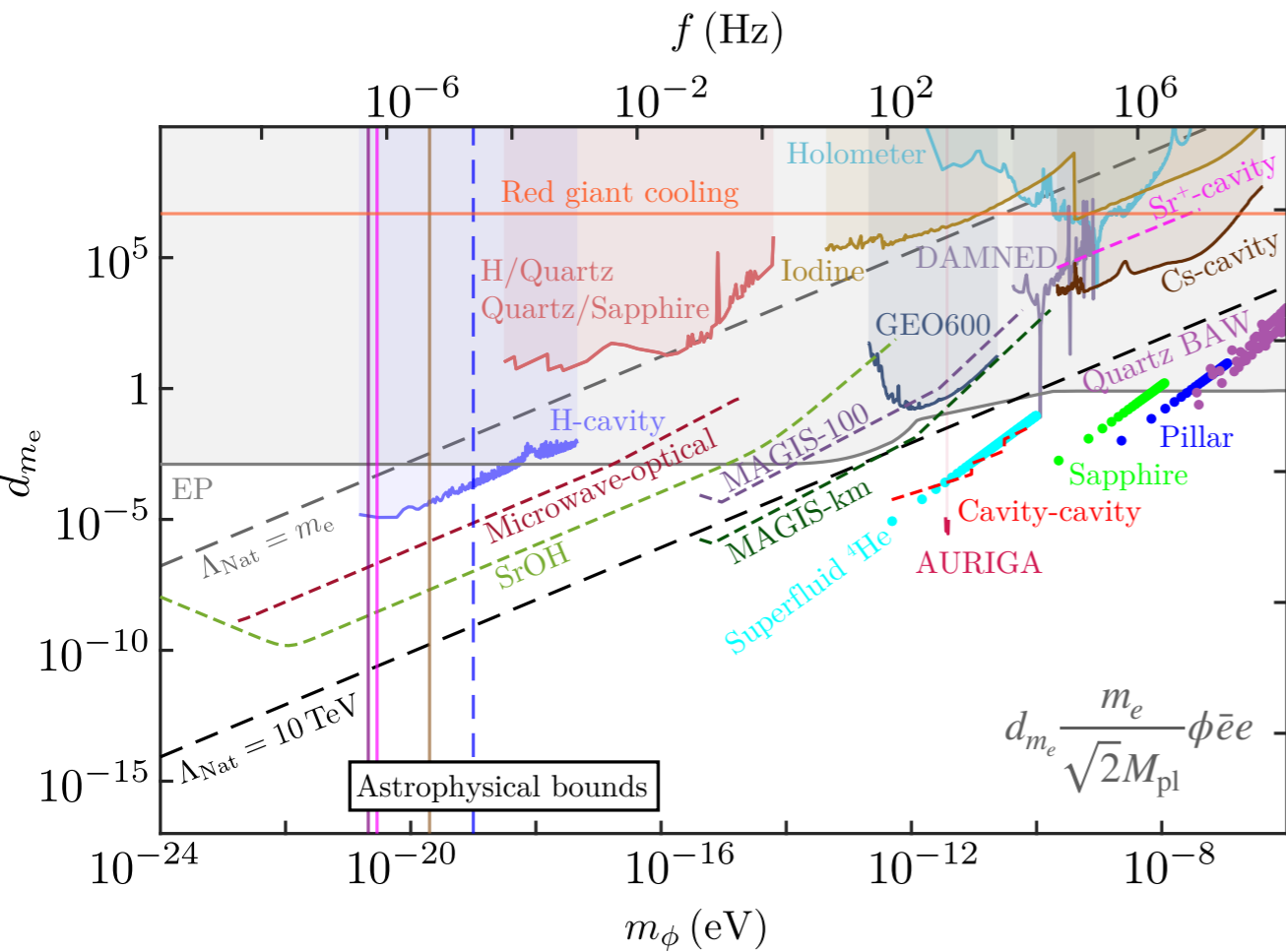
QCD axion DM and axion-like particles (ALPs)



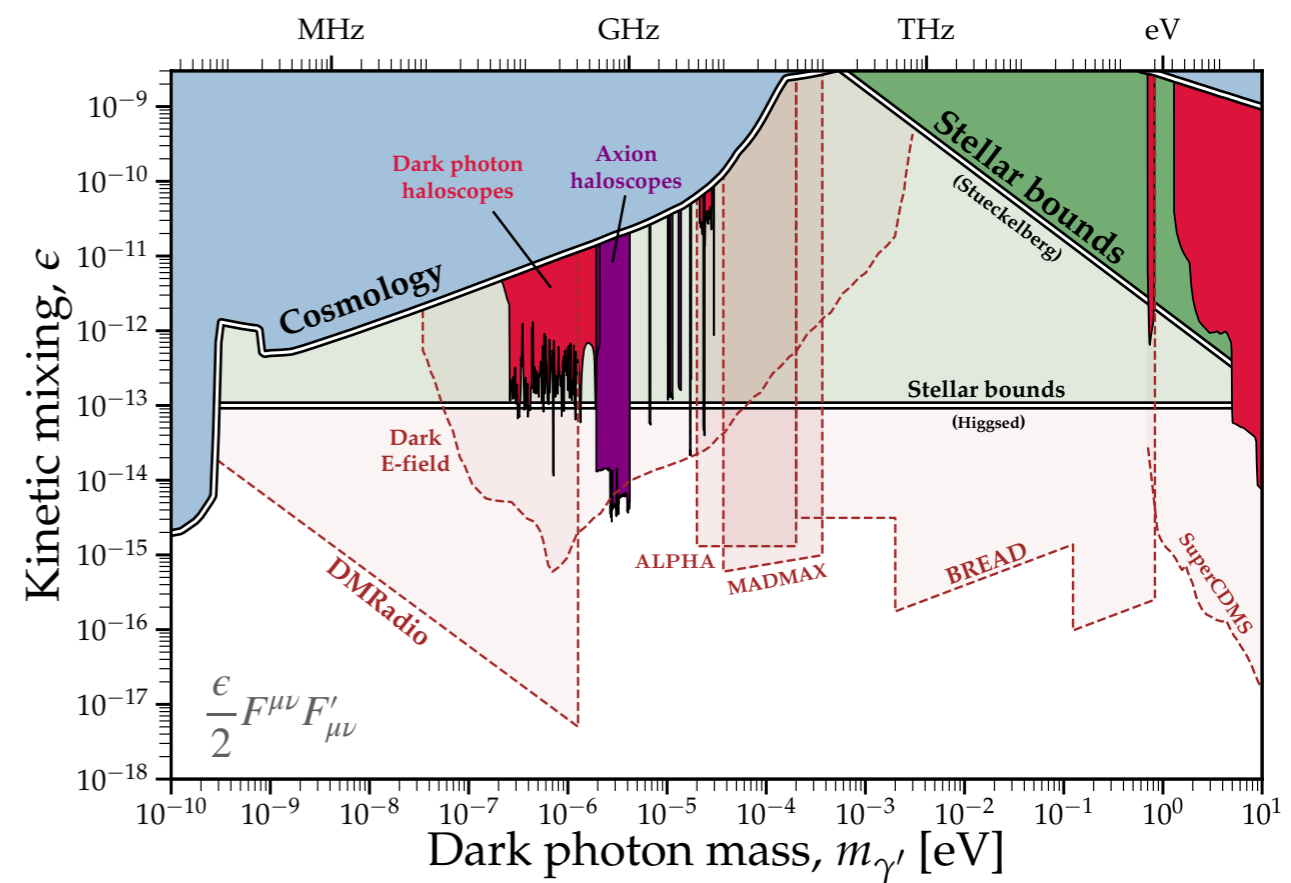
# Scalar DM with $m_{\text{DM}} \gtrsim 10^{-21} \text{eV}$



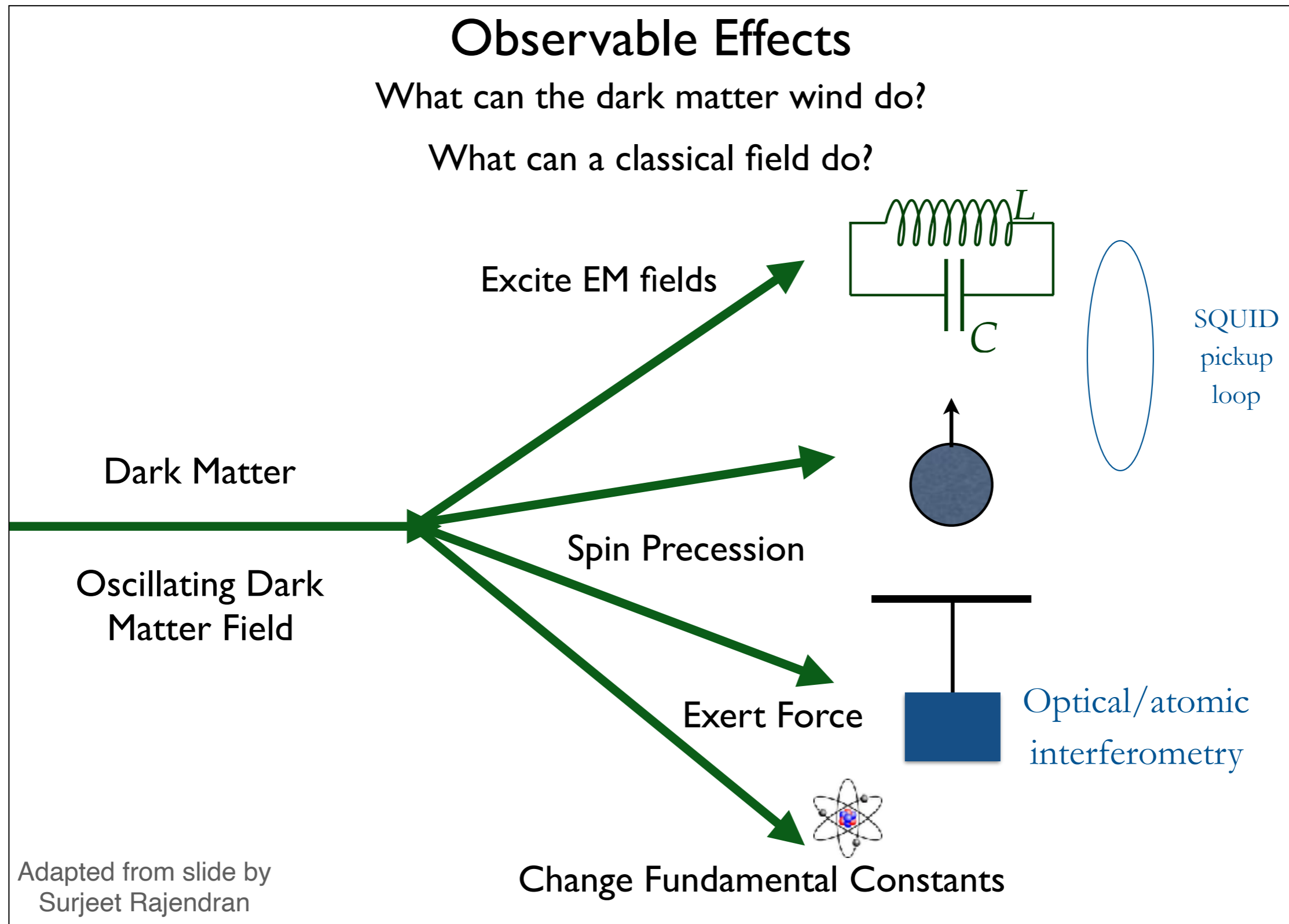
# Scalar DM with $m_{\text{DM}} \gtrsim 10^{-21} \text{eV}$



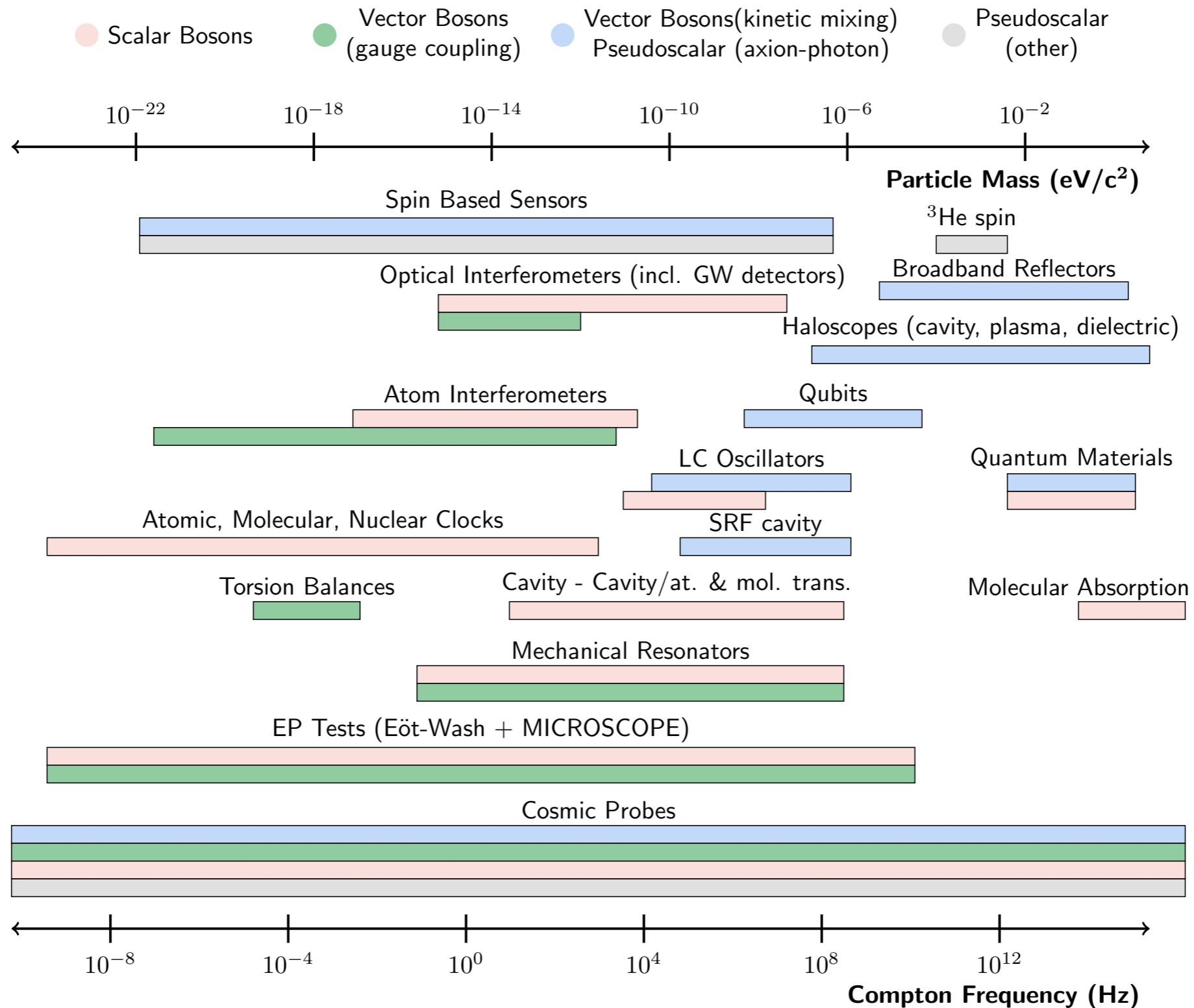
# Vector DM with $m_{\text{DM}} \gtrsim 10^{-10} \text{eV}$



# Detection concepts for wave-like DM



# Detection concepts for wave-like DM





# Progress in the past decade

## DM-electron scattering

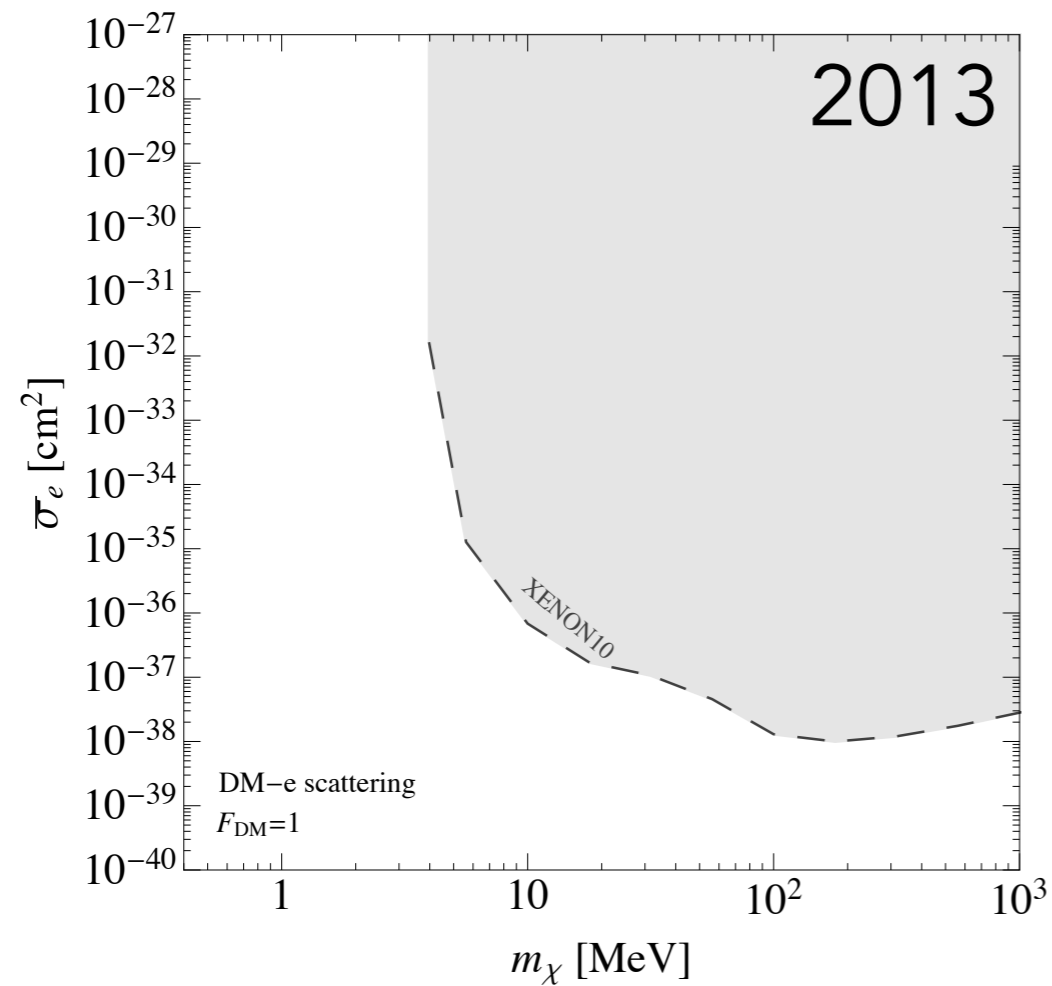


Fig from R. Essig and CF1 white paper  
“Landscape of low-threshold detection”

# Progress in the past decade

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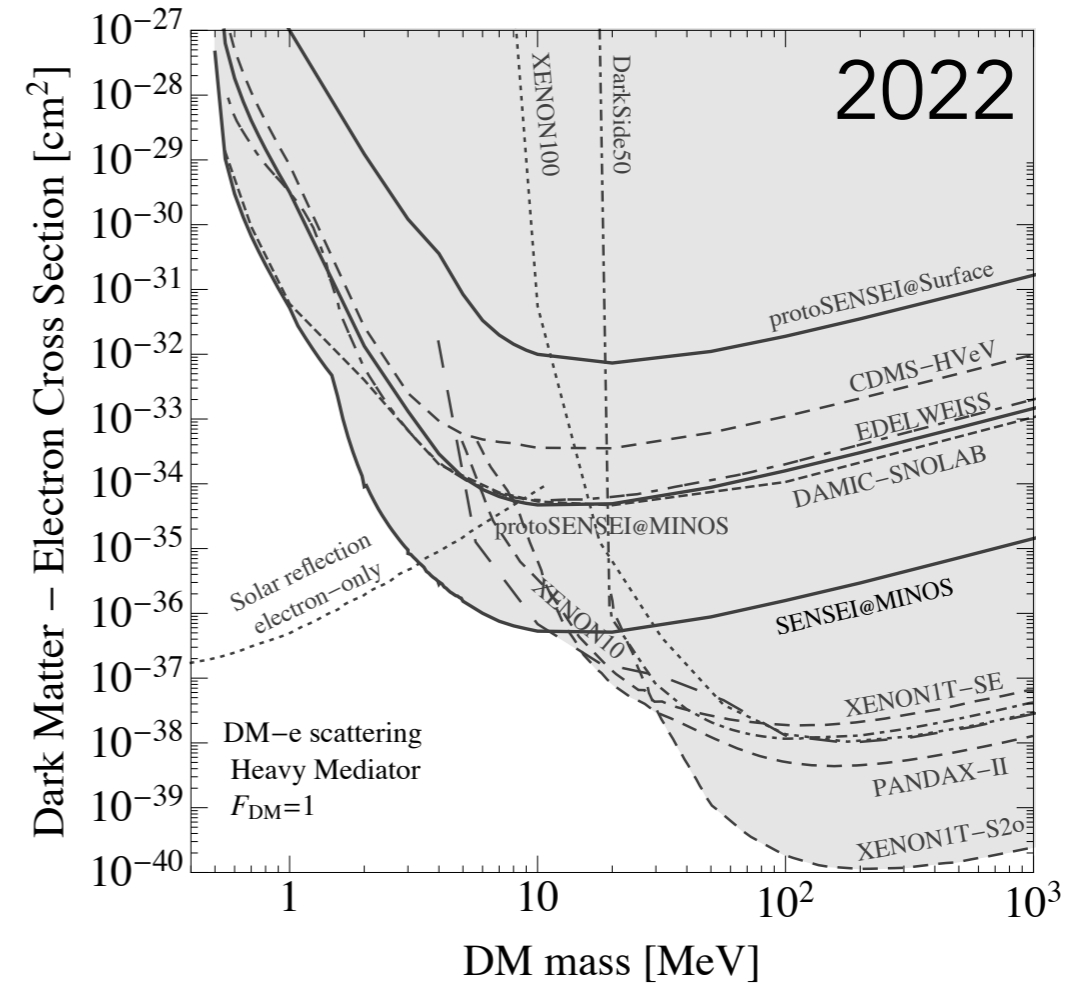
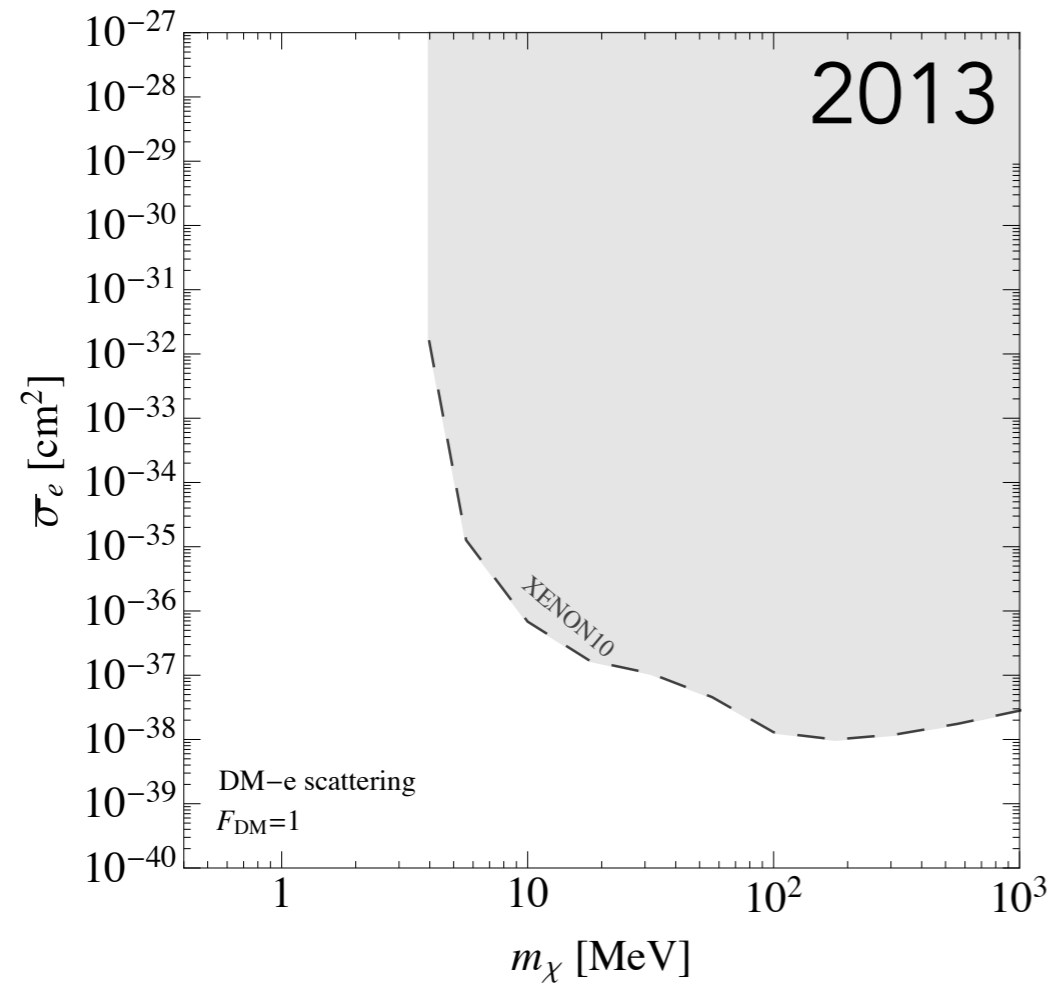
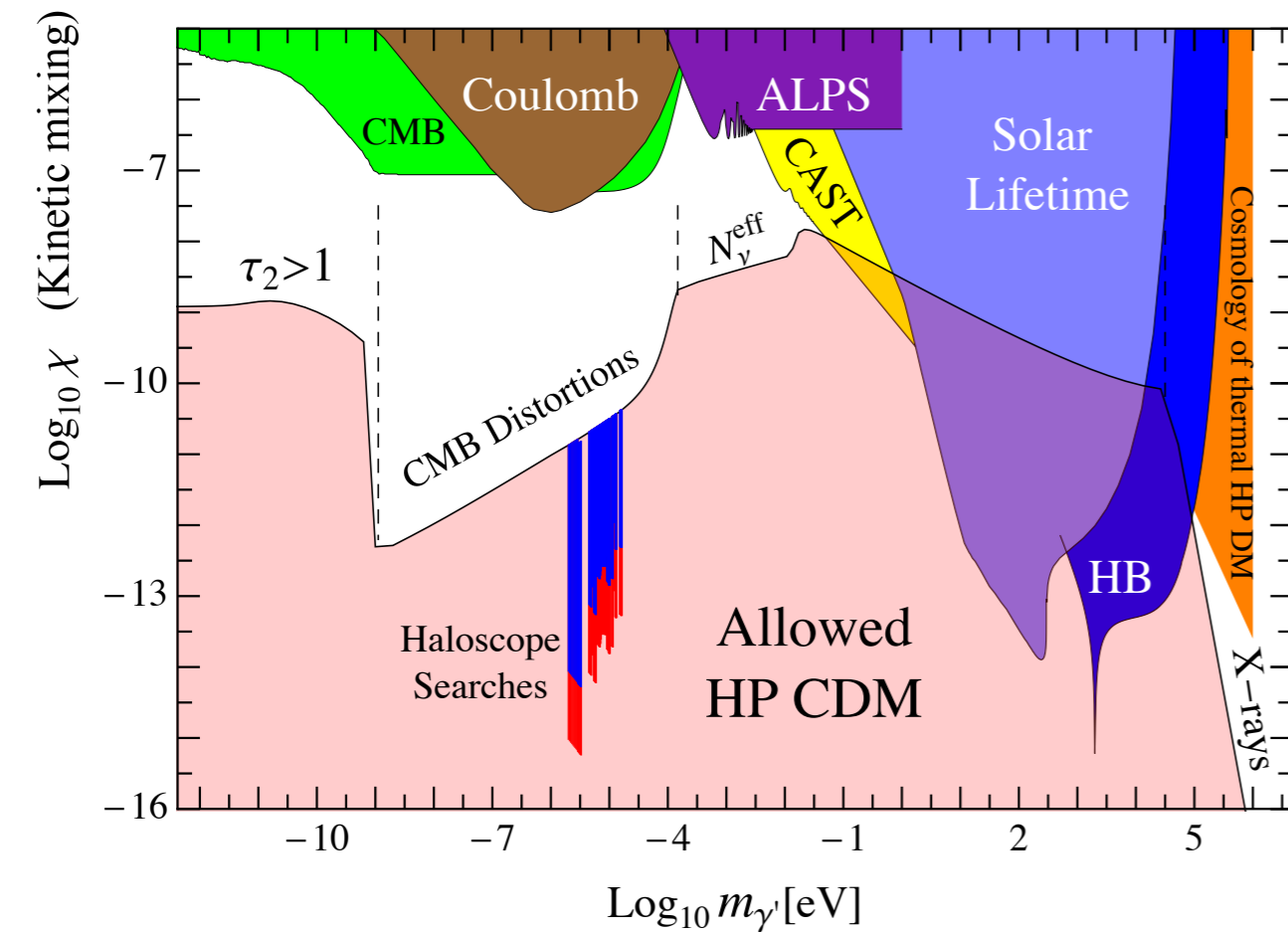


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# Progress in the past decade

## Dark photon dark matter

2012

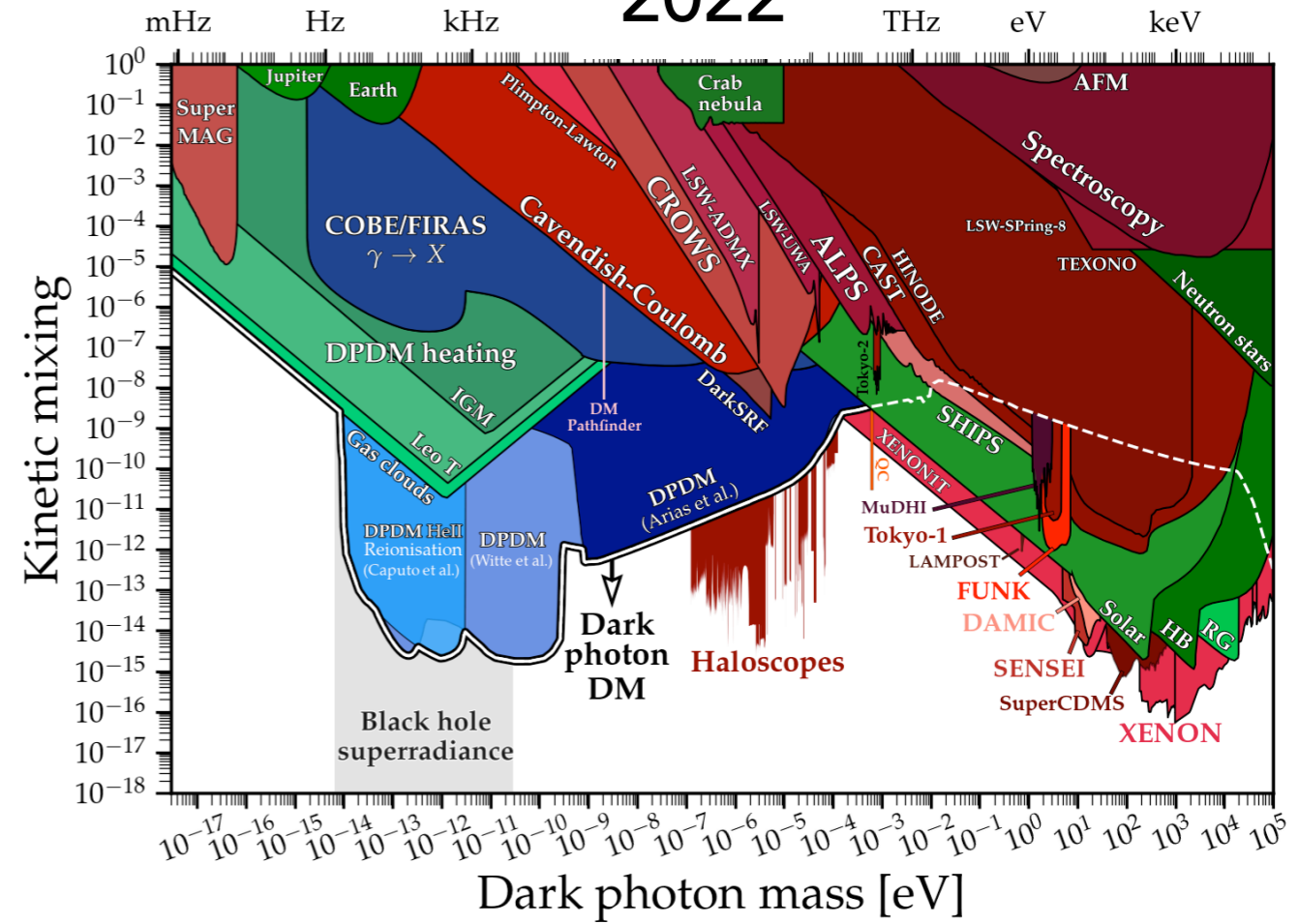
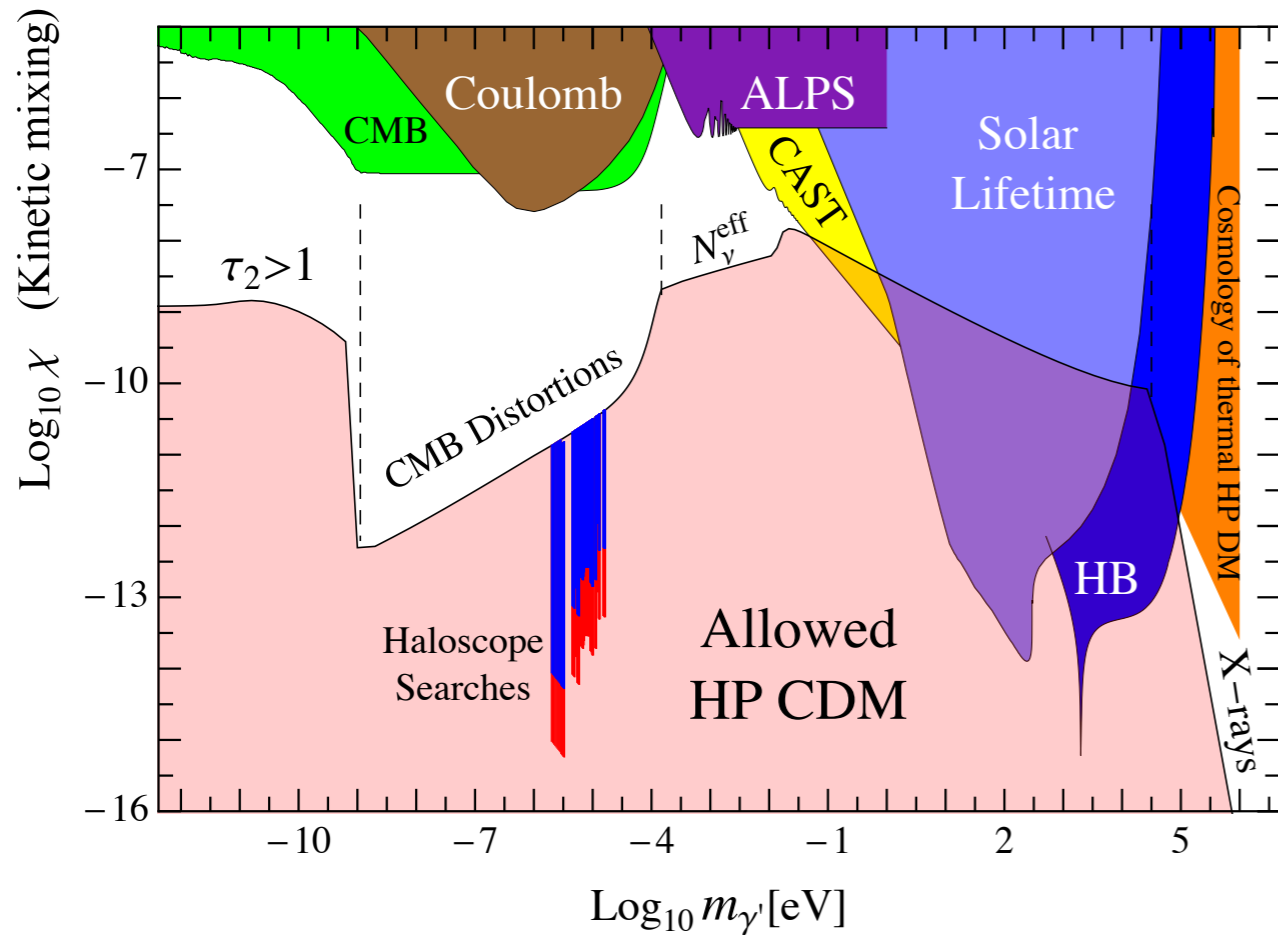


# Progress in the past decade

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2022

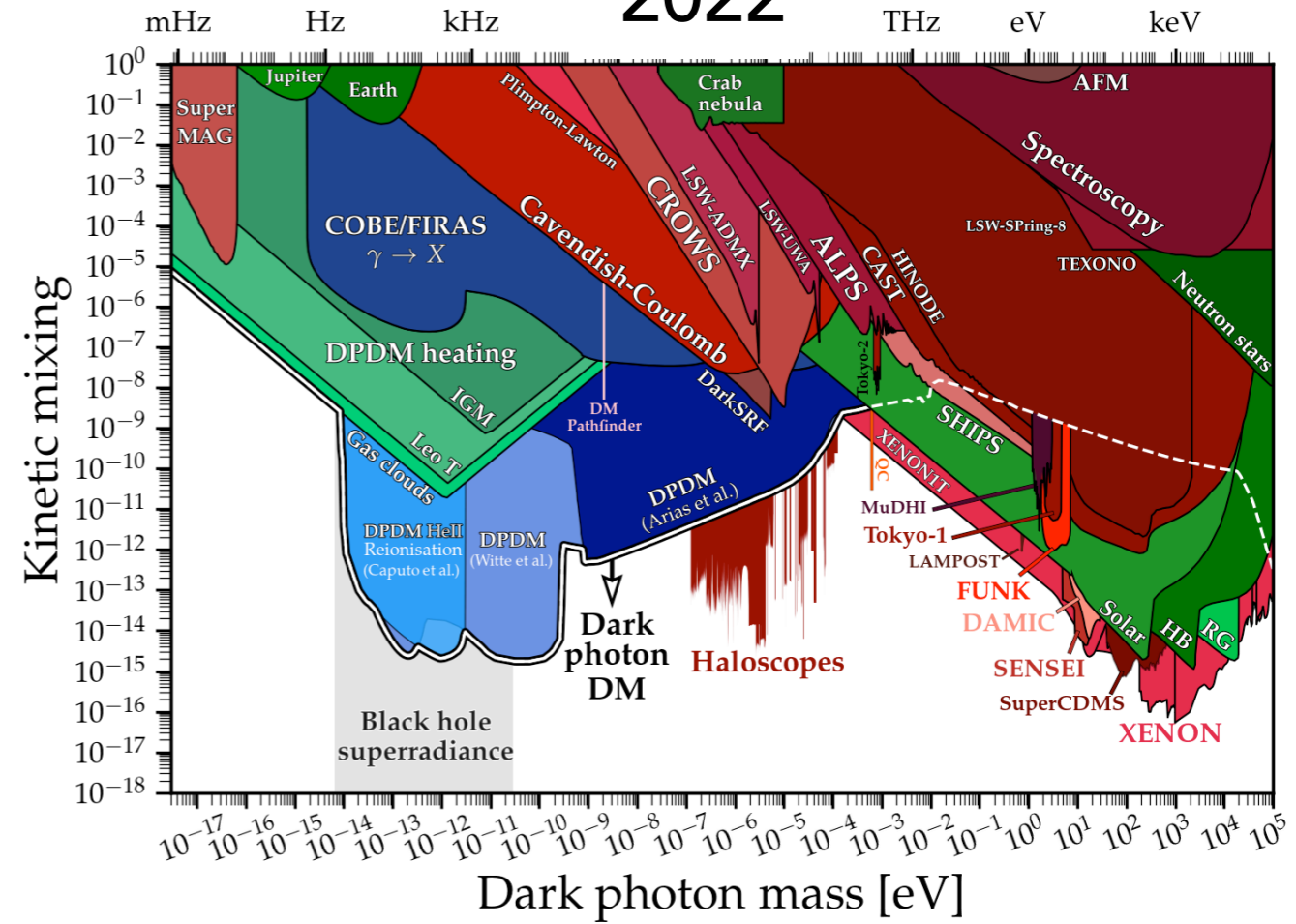
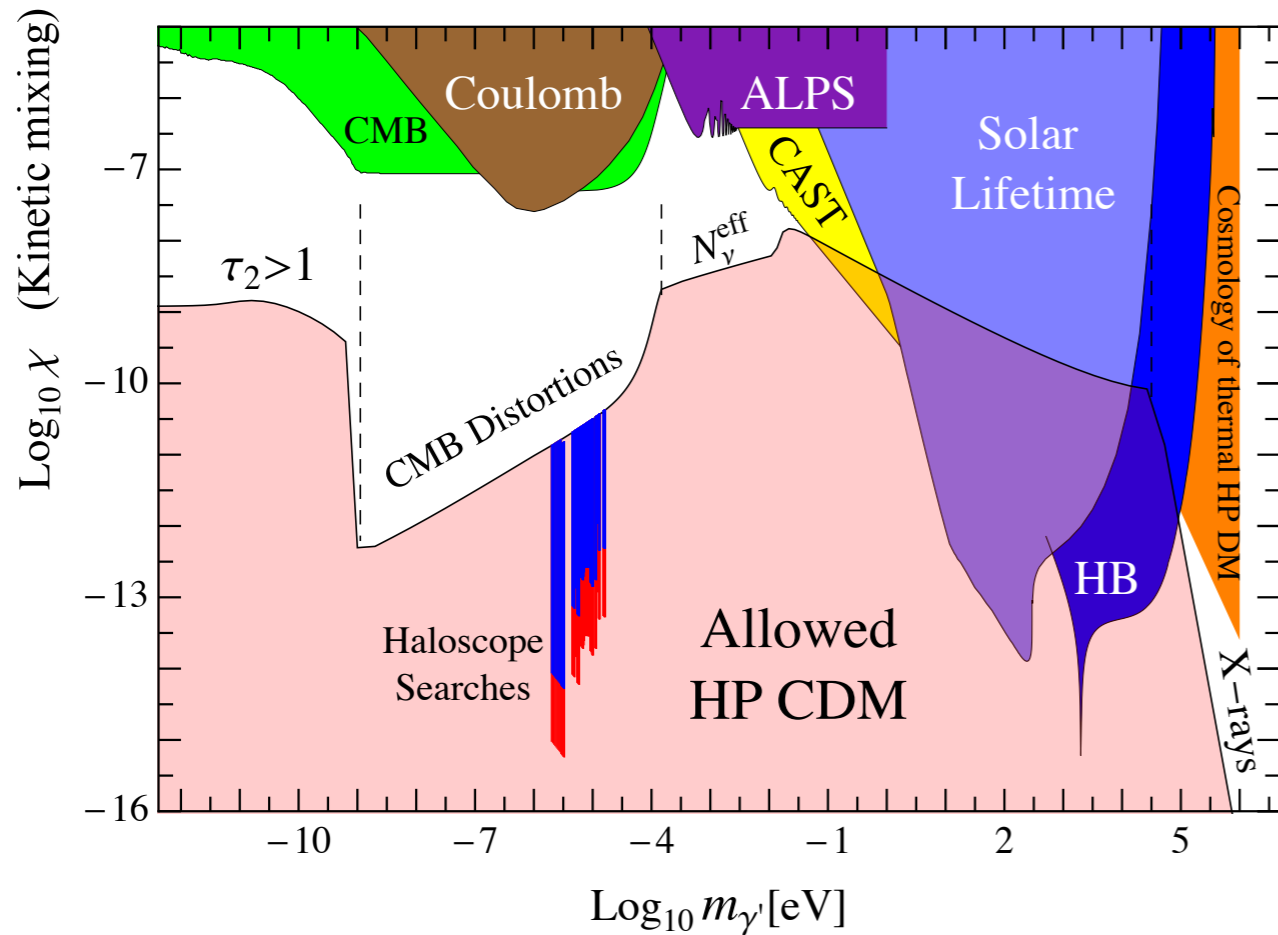


# Progress in the past decade

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2022



(and more!)

**How do we build on this progress?**

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- Support theory research: particle theory and interdisciplinary research (condensed matter, quantum sensing, etc).

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**We want to find the right keys to unlock a discovery!**

**Extra**

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## QCD axion DM and axion-like particles (ALPS)

