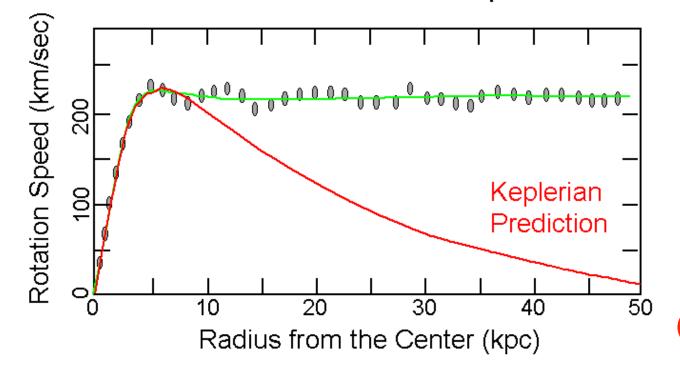


Bike-to-work Day, 2014 INPA Workshop

Astrophysical Results

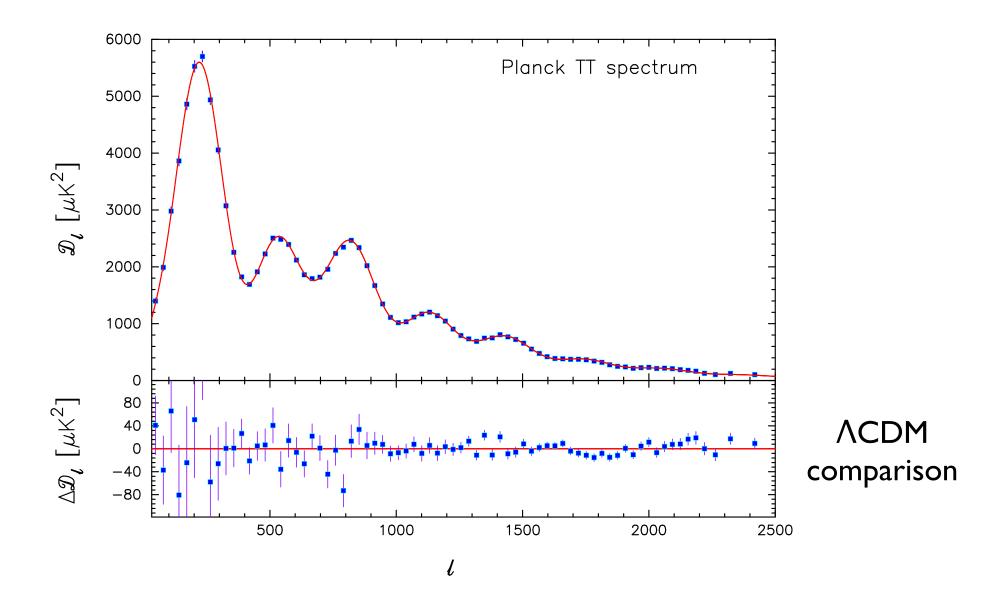


 $v \propto constant$ $\leftarrow m(r) \propto r$ $\rho(r) \propto 1/r^2$ (flat rotation curve)

 $\leftarrow v \propto I/\sqrt{r}$ (gravitating central mass)



NGC-6384 (from HST)



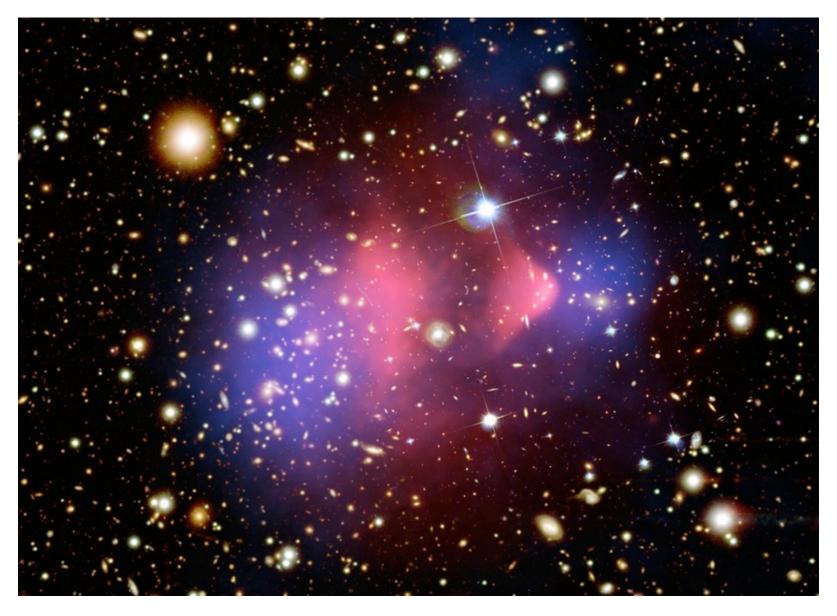
Planck XVI $\Omega_m \sim 0.315 \pm 0.017$

Required in simulations such as this (Bolshoi Collaboration) to reproduce observed cluster-cluster correlations*

Calculated with $\Omega_m = \Omega_{DM} + \Omega_b = 0.27 \sim \Omega_m^{\text{WMAP9}} \sim 0.2865 \pm 0.0088$

*Primack, Klypin, et al.

Bullet Cluster



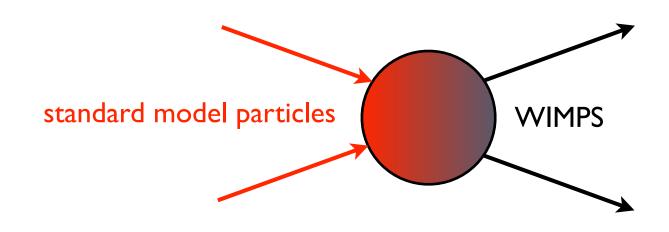
A collision between two clusters of galaxies, imaged by gravitational lensing, showing a separation of visible (pink) and dark (blue) matter

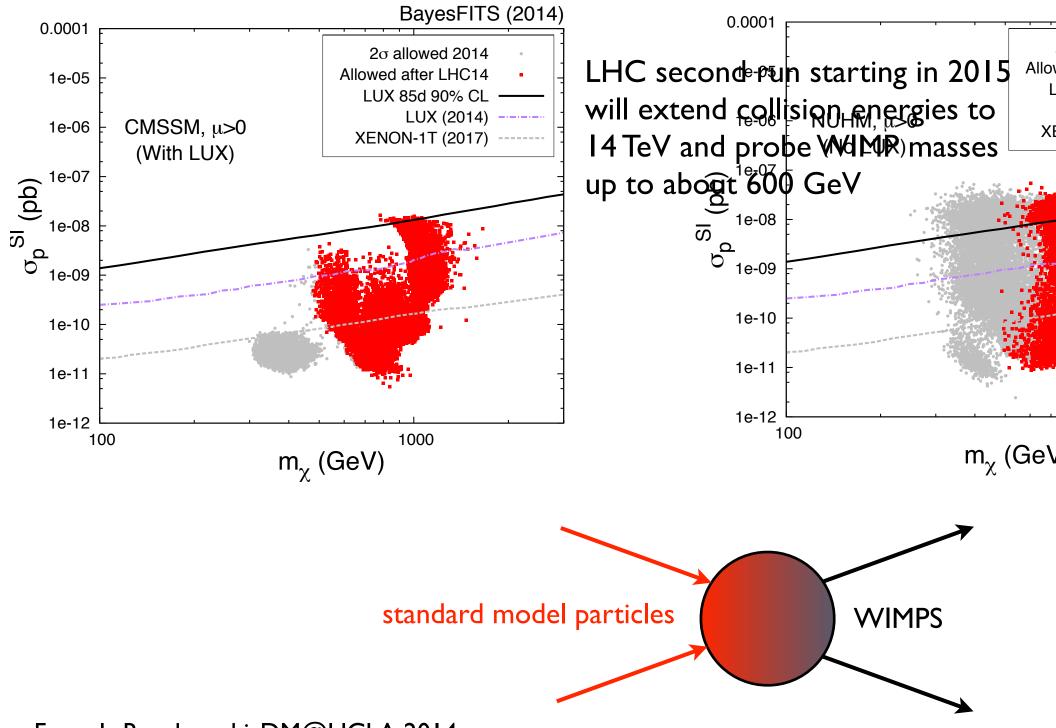
Nongravitational Evidence?



WIMP detection:

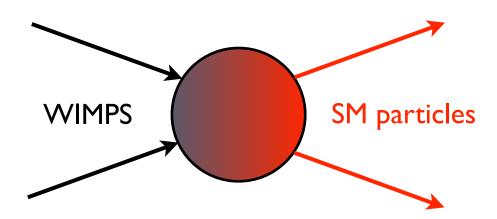
□ collider searches

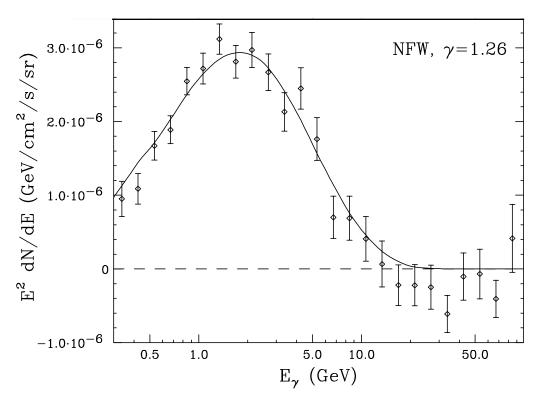




From L. Roszkowski, DM@UCLA 2014

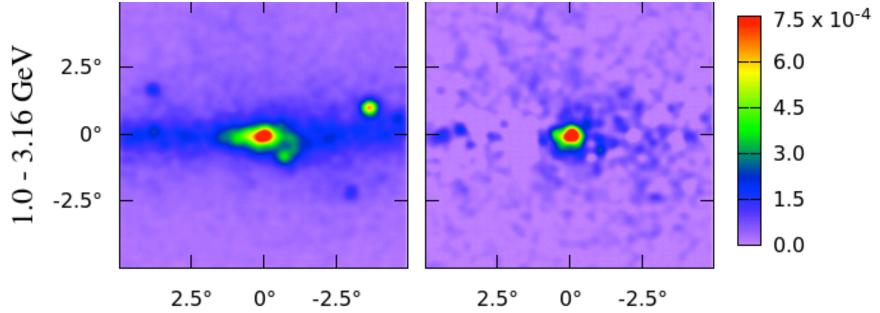
- □ collider searches
- □ indirect detection: astrophysical signals





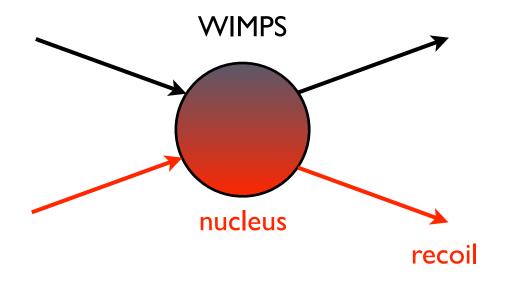
Some claims: From D. Hooper, UCLA DM Workshop

One interpretation: ~ 30-40 GeV WIMPs annihilating to b quarks, producing ~ 5 GeV gammas

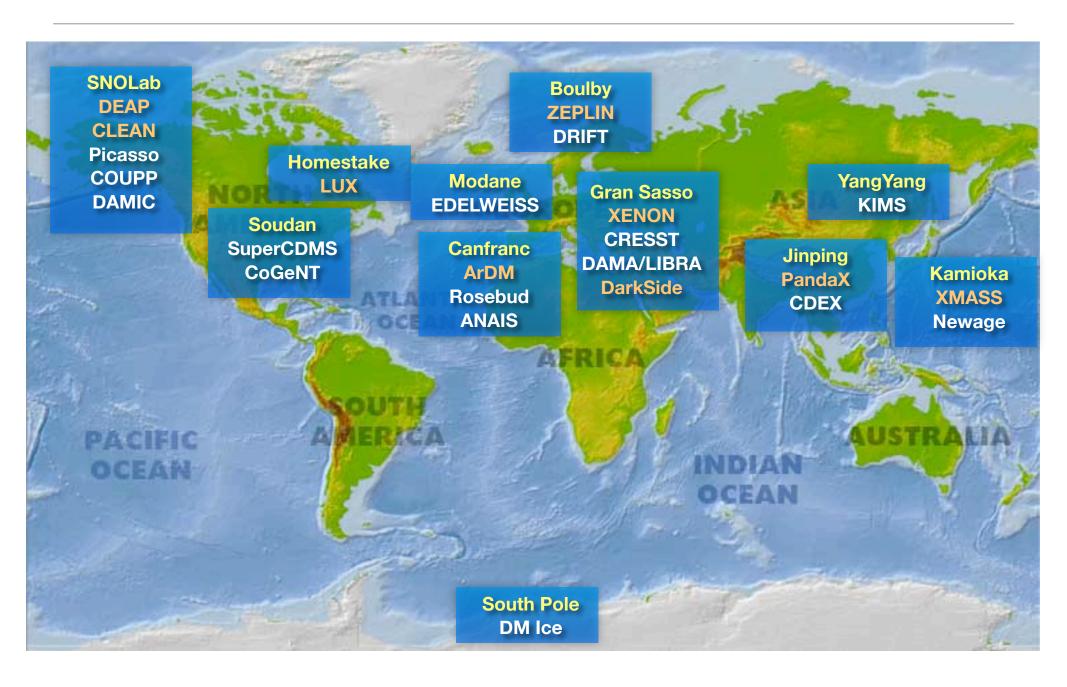


Detection: their detection channels include (other than large scale structure)

- □ collider searches
- □ indirect detection: astrophysical signals
- □ direct detection: large variety of experiments, some claims...



Essentially every underground lab has experiments



Xe: Xenon 100/1T; LUX/LZ; XMASS; Zeplin; NEXT

Si: CDMS; DAMIC

Ge: COGENT; Edelweiss; SuperCDMS; TEXONO; CDEX; GERDA; Majorana

Nal: DAMA/LIBRA; ANAIS; DM-ice; SABRE; KamLAND-PICO

CsI: KIMS

Ar: DEAP/CLEAN; ArDM; Darkside

Ne: CLEAN

C/F-based: PICO; DRIFT; DM-TPC

CF₃I: COUP

Cs2: DRIFT A large variety of nuclei with

TeO2: CUORE different spins, isospin, masses

CaWO4: CRESST

NOBLE GASSES

Single-phase detectors

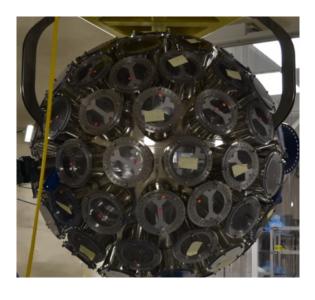
(SCINTILLATION LIGHT)

- Challenge: ultra-low absolute backgrounds
- LAr: pulse shape discrimination, factor 109-1010 for gammas/betas



XMASS-RFB at Kamioka:

835 kg LXe (100 kg fiducial), single-phase, 642 PMTs unexpected background found detector refurbished (RFB) new run this fall -> 2013



CLEAN at SNOLab:

500 kg LAr (150 kg fiducial) single-phase open volume under construction to run in 2014



DEAP at SNOLab:

3600 kg LAr (1t fiducial) single-phase detector under construction to run in 2014

Time projection chambers

(SCINTILLATION & IONIZATION)











XENON100 at LNGS:

161 kg LXe (~50 kg fiducial)

242 1-inch PMTs taking new science data

LUX at SURF:

350 kg LXe (100 kg fiducial)

122 2-inch PMTs physics run since spring 2013

PandaX at CJPL:

125 kg LXe (25 kg fiducial)

143 1-inch PMTs 37 3-inch PMTs started in 2013

ArDM at Canfranc:

850 kg LAr (100 kg fiducial)

28 3-inch PMTs in commissioning to run 2014

DarkSide at LNGS

50 kg LAr (dep in ³⁹Ar) (33 kg fiducial)

38 3-inch PMTs in commissioning since May 2013 to run in fall 2013

CRYSTALS, BUBBLE CHAMBERS, ...



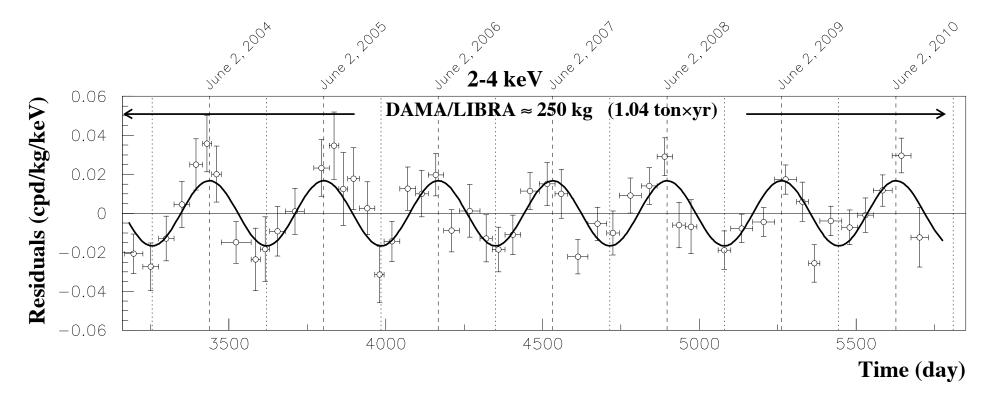




DAMA/LIBRA NAI

CDMS SI, GE COGENT GE

COUP CF₃I





DAMA/LIBRA: 9.30 variation of the signal over the year, and the expected over the expected over a country of the expected over a country of the signal on the expected over a country of the Sun of a DM signal on the expected over a country of the Sun over the year, and the expected over the year.

Time (day)

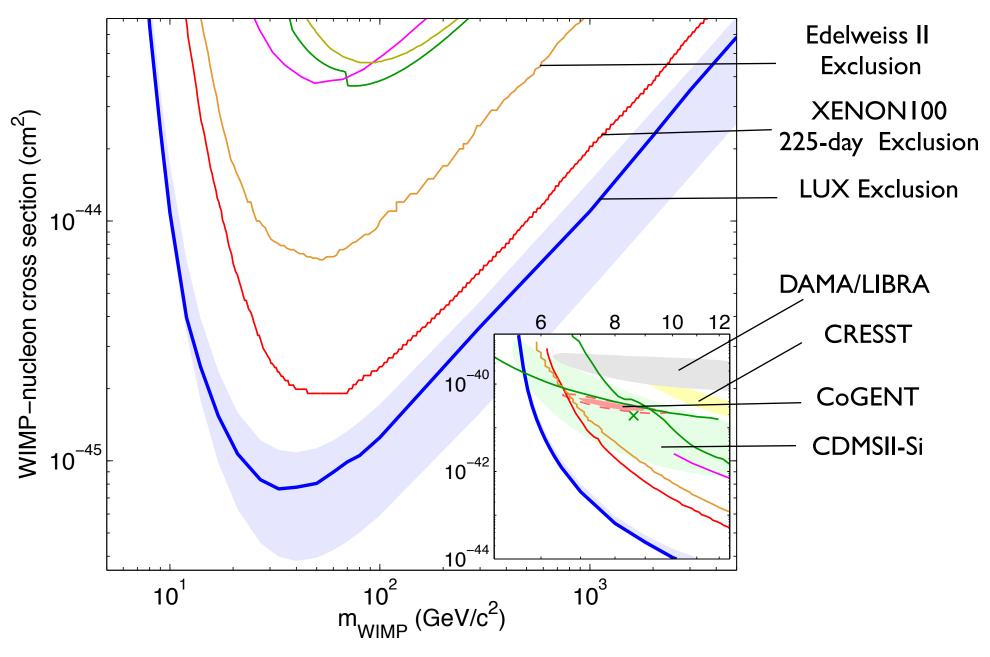
CoGENT: Ge detector in which a similar seasonal variation was seen at 2.8 σ , consistent with a light 7 GeV WIMP

No such signal found by the MALBEK Ge detector group

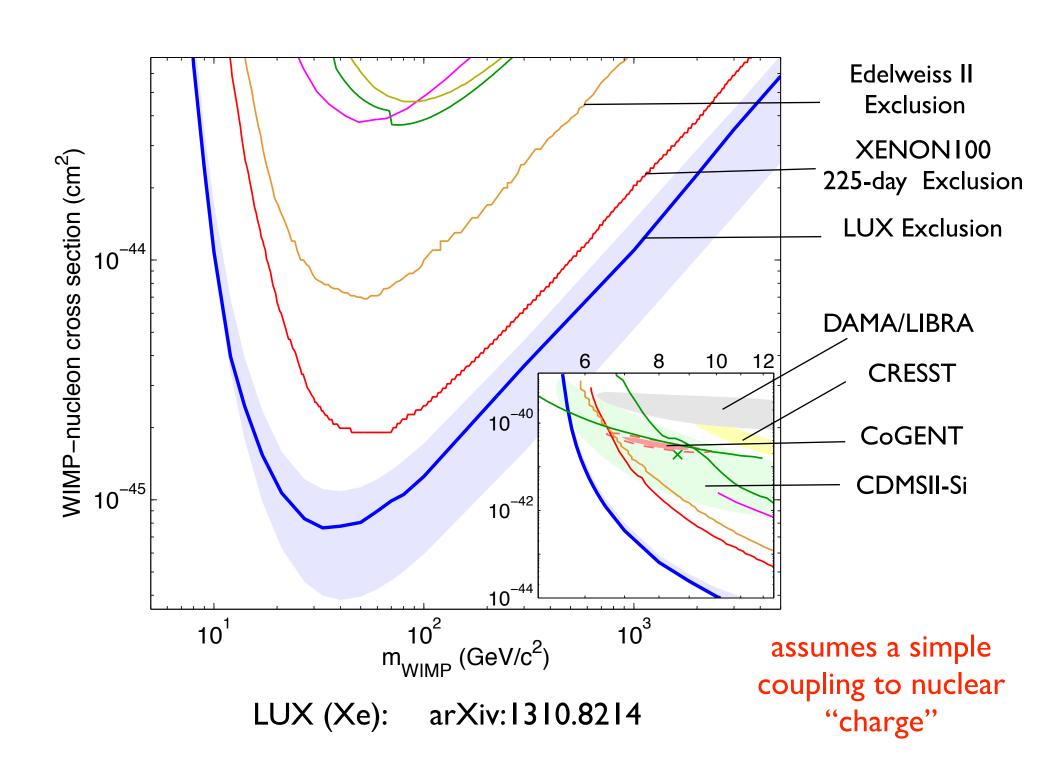
CDSM II-Si: upper bound established, but found three low-mass events vs. an expected background signal of ~ 0.41 events. If interpreted as DM, implies M_{WIMP} ~ 10 GeV







LUX (Xe): arXiv:1310.8214



How are these comparisons among experiments done?

We know some basic parameters

- WIMP velocity relative to our rest frame $\sim 10^{-3}$
- if mass is on the weak scale, WIMP momentum transfers in elastic scattering can range to $q_{\rm max} \sim 2v_{\rm WIMP}\mu_T \sim 200~{\rm MeV/c}$
- WIMP kinetic energy ~ 30 keV: nuclear excitation (in most cases) not posible
- $R_{NUC} \sim 1.2 \, A^{1/3} \, f \implies q_{max} \, R \sim 3.2 \Leftrightarrow 6.0$ for $F \Leftrightarrow Xe$: the WIMP can "see" the structure of the nucleus

plane yz, the sudden peak at $R \simeq 13$ kpc is due to the 8.b. Earth restframe (Summer) Halo restframe Our motion through the WIMP "wind" 3 can be in ordered *b*) 10 GHALO $_{
m tal}\sim 0.3$ (Kpwim) GHALO_s -10400 0 200 600 800 200 400 600 800 v [km/s] v [km/s] -15 -15 M₅Kuhlen et al, JCAP02 (2010) 030

-10

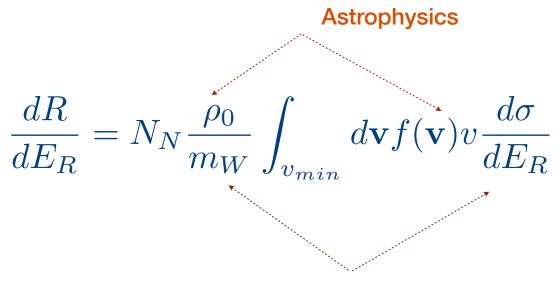
15

-5

y (kpc)

The galactic plane (blue), while it drops to a value ~ 0.9

An expression can be written for the rate as a function of nuclear recoil energy E_R



Particle+nuclear physics

$$N_N=$$
 number of target nuclei in detector $ho_0=$ Milky Way dark matter density $v_{min}=\sqrt{\frac{m_N E_{th}}{2\mu^2}}$ WIMP velocity distribution, Earth frame $m_W=$ WIMP mass

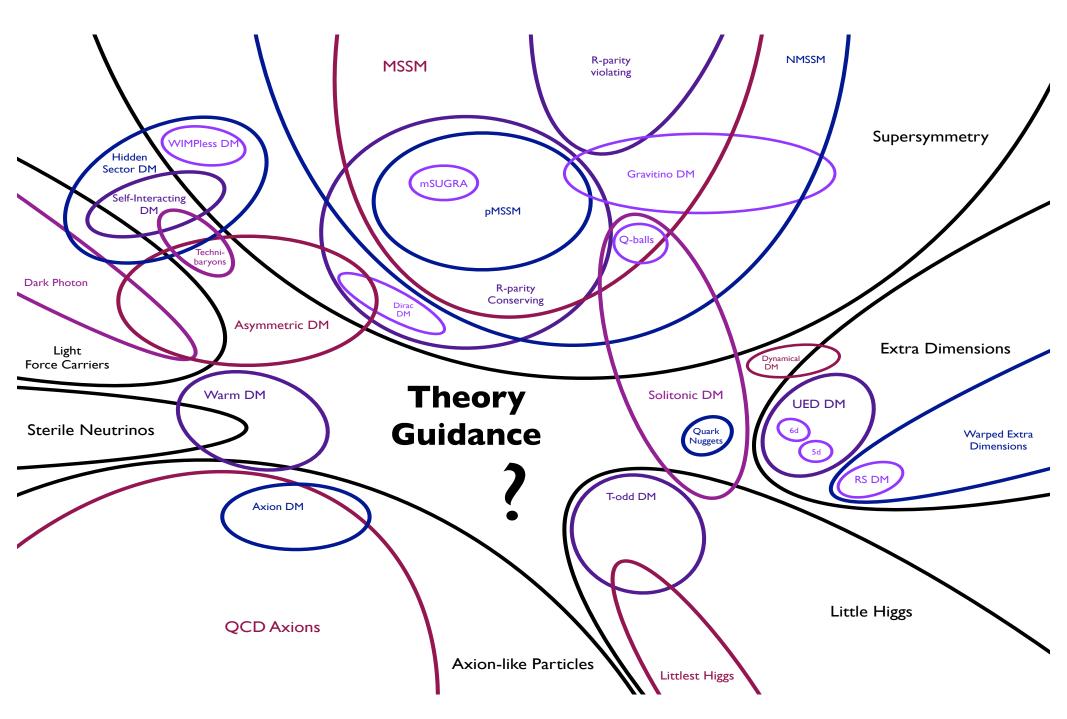
WIMP – nucleus elastic scattering cross section

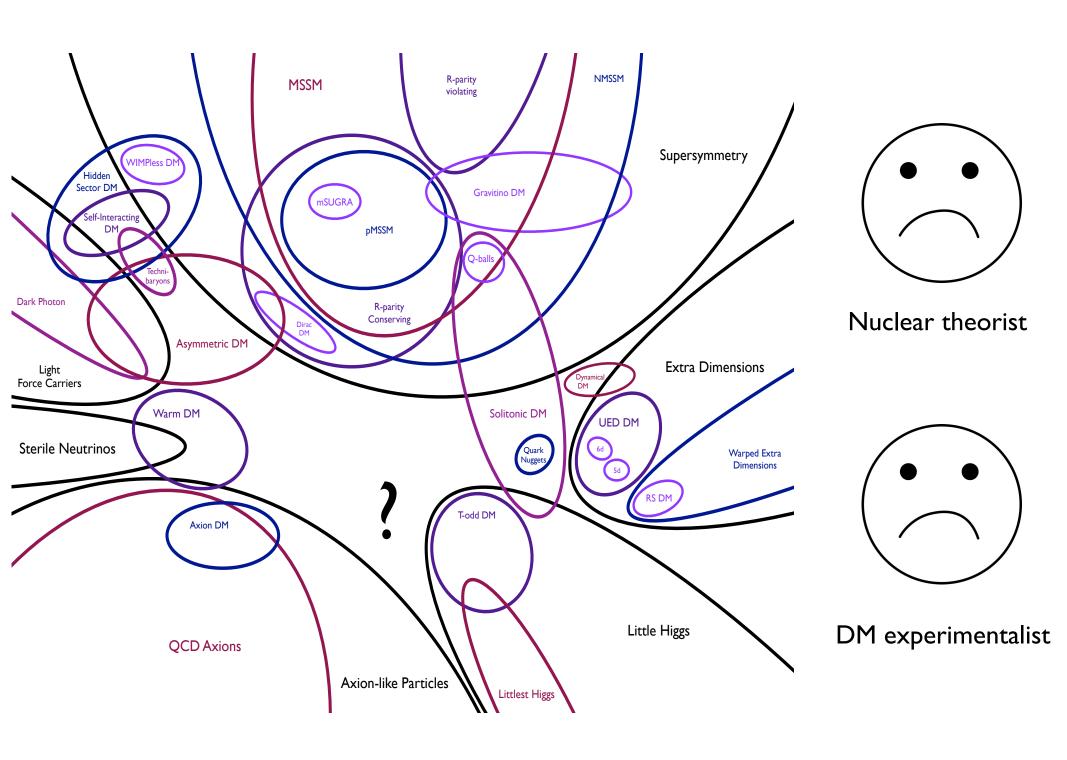
 $\sigma =$

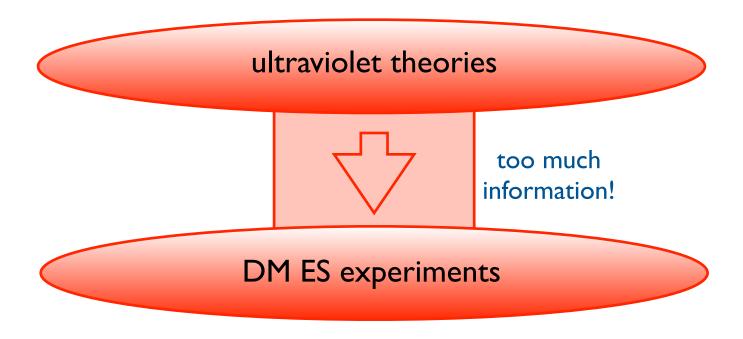
But where do we get the cross section -- the WIMP-nucleus interaction?

In fact, what can and cannot be learned about the WIMP-matter interaction from these low-energy elastic scattering experiments?

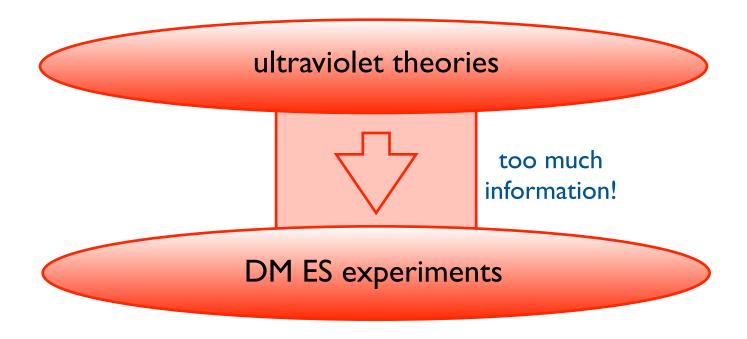
so just ask a particle theorist (or several)...







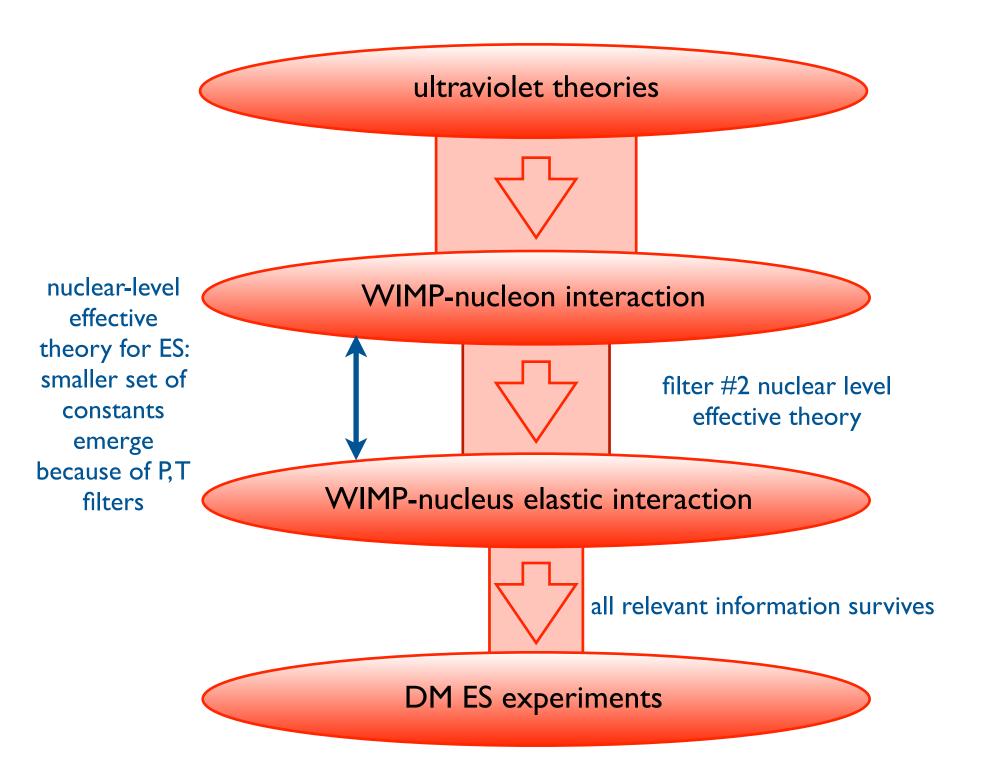
This is a very difficult step, and a tedious one as it must be taken for each candidate ultraviolet theory

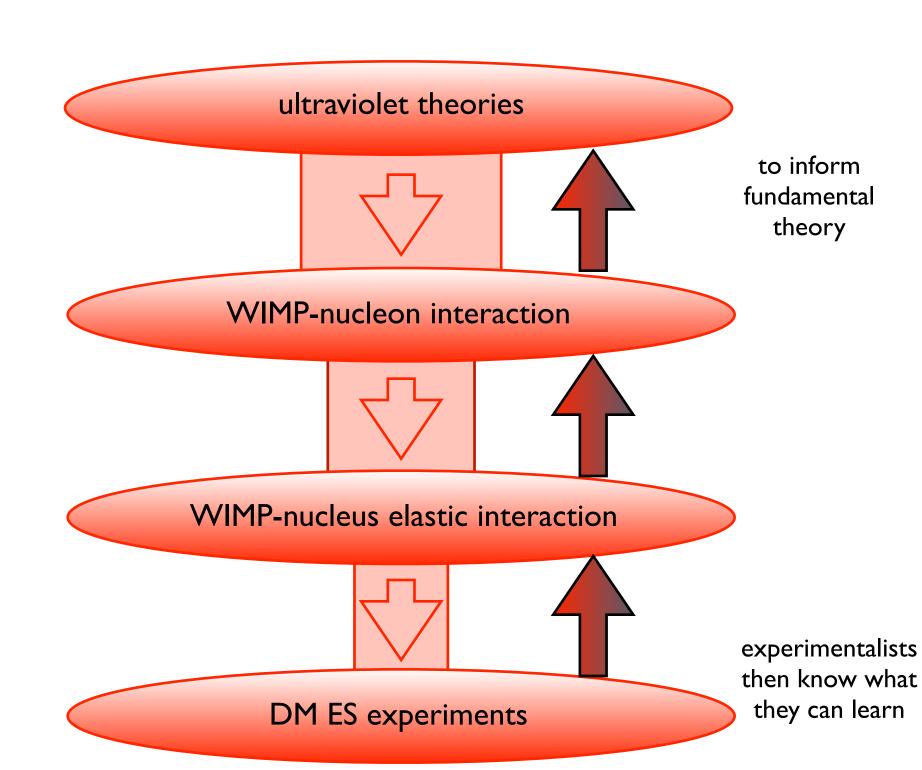


This is a very difficult step, and a tedious one as it must be taken for each candidate ultraviolet theory

An alternative is provided by effective field theory

ultraviolet theories ultraviolet physics encoded in a filter #I nucleon-level finite set of effective theory low-energy WIMP-nucleon coupling WIMP-nucleon interaction constants WIMP-nucleus elastic interaction DM ES experiments



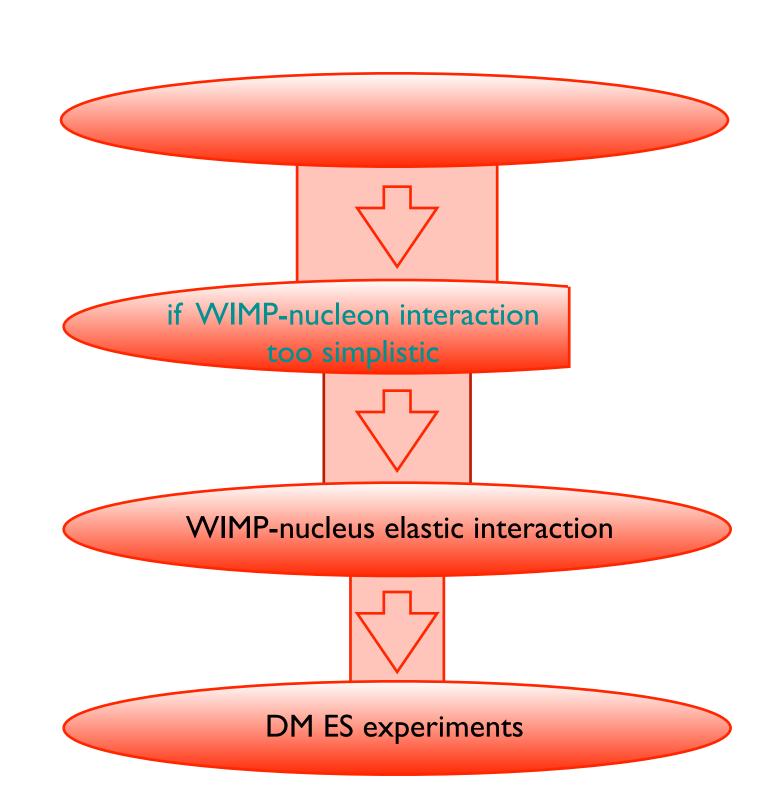


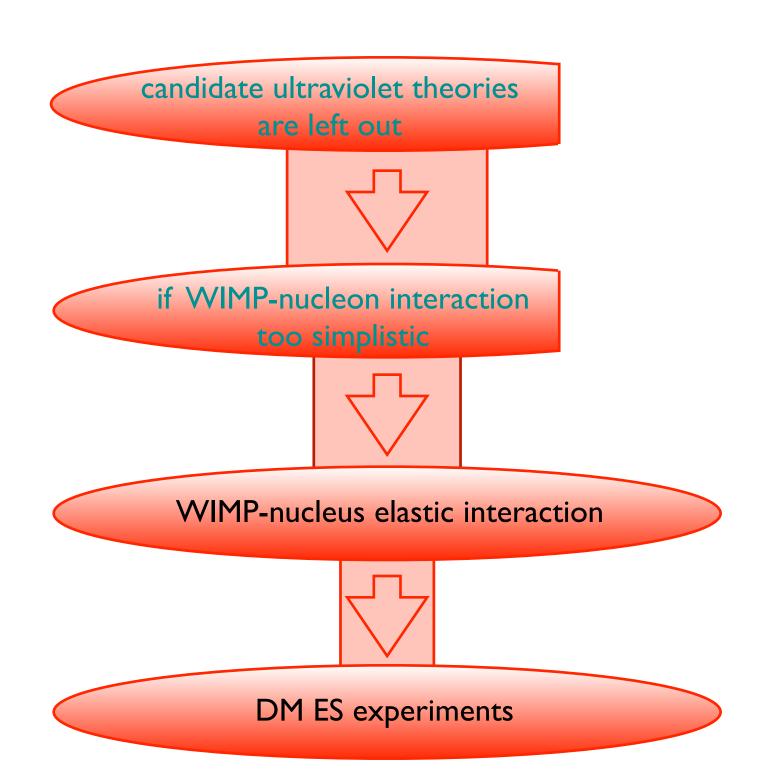
the effective theory process works only if each step is executed properly

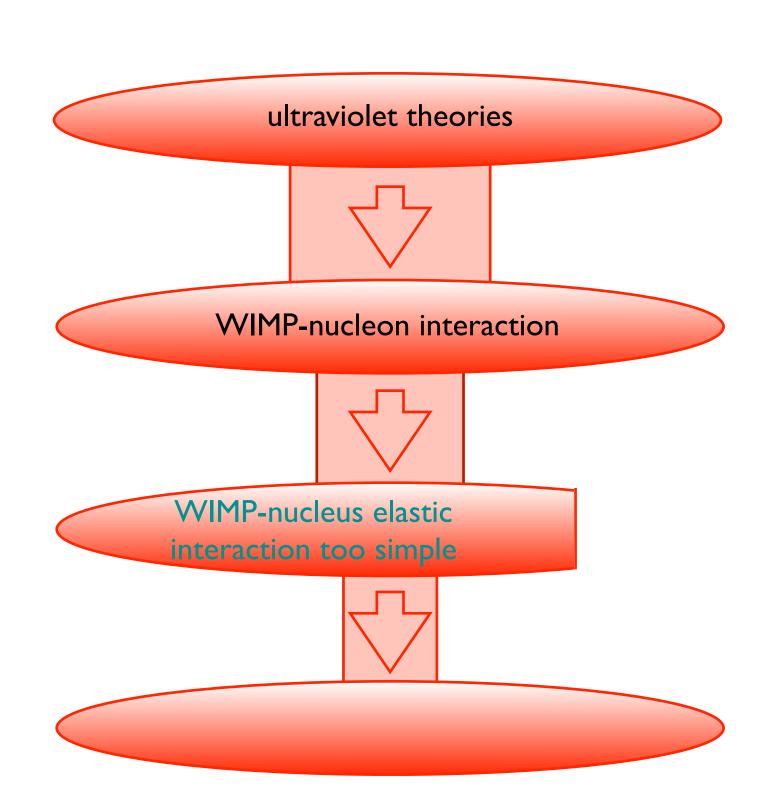




this not this











WIMP-nucleon interaction



WIMP-nucleus elastic interaction too simple



Too few experiments done, too little learned

 Experiments are frequently analyzed and compared in a formalism in which the nucleus is treated as a point particle

S.I.
$$\Rightarrow \langle g.s. | \sum_{i=1}^{A} (a_0^F + a_1^F \tau_3(i)) | g.s. \rangle$$

S.D. $\Rightarrow \langle g.s. | \sum_{i=1}^{A} \vec{\sigma}(i) (a_0^{GT} + a_1^{GT} \tau_3(i)) | g.s. \rangle$

Is this treatment sufficiently general, to ensure a discovery strategy that will lead to the right result? Is it some lowest order ET? Reminiscent of an old weak interactions problem: we knew a lot more -- four fermions, contact interaction,... Constructed the low energy effective theory

$$S+V+A+P+T$$

Yet took a lot of effort to determine V-A, including eliminating wrong experiments in favor of correct ones We can attack this problem with similar techniques: What is the form of the elastic response for a nonrelativistic theory with charges and current at most linear in velocities?

		even	odd
charges:	vector axial	$\begin{bmatrix} C_0 \\ C_0^5 \end{bmatrix}$	$C_1 \\ C_1^5$

currents:

	even	odd	even	odd	even	odd
axial spin vector velocity vector spin — velocity	$egin{array}{c} L_0^5 \ L_0 \ L_0 \end{array}$	$L_1^5 \ L_1 \ L_1$	$T_2^{ m 5el} \ T_2^{ m el} \ T_2^{ m el}$	$T_1^{ m 5el} \ T_1^{ m el} \ T_1^{ m el}$	$T_2^{5\mathrm{mag}} \ T_2^{\mathrm{mag}} \ T_2^{\mathrm{mag}}$	$T_1^{5\mathrm{mag}} \ T_1^{\mathrm{mag}} \ T_1^{\mathrm{mag}}$

(where we list only the leading multipoles in J above)

We can attack this problem with similar techniques: What is the form of the elastic response for a nonrelativistic theory with charges and current at most linear in velocities?

		even	odd	
charges:	vector axial		$C_1 \ C_1^5$	symmetry arguments ⇒ form of the nuclear ET

currents:

	even	odd	even	odd	even	odd
axial spin vector velocity vector spin – velocity	$\begin{array}{c c} L_0^5 \\ L_0 \\ L_0 \end{array}$	$L_1^5 \ L_1 \ L_1$	$T_2^{ m 5el} \ T_2^{ m el} \ T_2^{ m el}$	$T_1^{ m 5el} \ T_1^{ m el} \ T_1^{ m el}$	$T_2^{5\mathrm{mag}} \ T_2^{\mathrm{mag}} \ T_2^{\mathrm{mag}}$	$T_1^{5\mathrm{mag}} \ T_1^{\mathrm{mag}} \ T_1^{\mathrm{mag}}$

(where we list only the leading multipoles in J above)

Response constrained by good parity and time reversal of nuclear g.s.

	even	odd
vector axial	C_0	C_1^5

	even	odd	even	odd	even	odd
axial spin vector velocity vector spin — velocity	$egin{array}{c} L_0 \ L_0 \end{array}$	L_1^5	$T_2^{ m el} \ T_2^{ m el}$	$T_1^{ m 5el}$	$T_2^{5\mathrm{mag}}$	$T_1^{ m mag} \ T_1^{ m mag}$

Response constrained by good parity and time reversal of nuclear g.s.

	even	odd
vector axial	C_0	

	even	odd	even	odd	even	odd
axial spin vector velocity vector spin — velocity	L_0	L_1^5	$T_2^{ m el}$	$T_{1}^{ m 5el}$	_	$T_1^{ m mag}$

The resulting table of allowed responses has six entries (not two): just determined by the symmetries, not yet related to WIMP couplings to the nucleon/nucleus

One of the union rules for theorists:

Interactions allow by symmetries must be (and will be) included in a proper effective theory

- This suggests more can be learned about ultraviolet theories from ES than is generally assumed - that's good
- But what quantum mechanics are we missing? How do these responses from interactions?
 - to answer this in full, we need to construct the Galilean-invariant ET at the nucleon level (the ladder to HE)

Helpful to first understand the quantum mechanics that is missing in the standard treatment -- which assumes the nucleus behaves as a point particle, and is thus characterized by just its charge and spin...

(this helps us define the ET construction)

The new responses are connected with velocity-dependent interactions that is, with theories that have derivative couplings

Let's take an example: consider

$$\sum_{i=1}^{A} \vec{S}_{\chi} \cdot \vec{v}^{\perp}(i)$$

the velocity is defined by Galilean invariance $ec{v}^{\perp}(i) = ec{v}_{\scriptscriptstyle Y} - ec{v}_N(i)$

$$\vec{v}^{\perp}(i) = \vec{v}_{\chi} - \vec{v}_{N}(i)$$

 ${\color{blue} extbf{D}}$ In the point-nucleus limit $\ \vec{S}_{\chi} \cdot \vec{v}_{ ext{WIMP}} \sum 1(i)$

 $\vec{v}_{\mathrm{WIMP}} \sim 10^{-3}$.



$$\{\vec{v}^{\perp}(i), i = 1, ...A\} \rightarrow \{\vec{v}_{\text{WIMP}}; \ \vec{\dot{v}}(i), i = 1, ..., A - 1\}$$

and $\vec{v}(i) \sim 10^{-1}$: SI/SD retains the least important term

Parameter counting in the effective theory

- $\hfill\Box$ These velocities hide: the $\vec{\dot{v}}(i)$ carry odd parity and cannot contribute by themselves to elastic nuclear matrix elements.
- But in elastic scattering, momentum transfers are significant. The full velocity operator is

$$e^{i\vec{q}\cdot\vec{r}(i)}\vec{v}(i)$$
 where $\vec{q}\cdot\vec{r}(i)\sim 1$

flux We can combine the two vector nuclear operators $ec{r}(i),\ ec{\dot{v}}$ to form a scalar, vector, and tensor. To first order in $ec{q}$ for the new "SD" case

$$-\frac{1}{i}q\vec{r}\times\vec{\dot{v}} = -\frac{1}{i}\frac{q}{m_N}\vec{r}\times\vec{\dot{p}} = -\frac{q}{m_N}\vec{\ell}(i)$$

 $\vec{\ell}(i)$ is a new dimensionless operator. And we deduce an instruction for the ET that is not obvious. Internal nucleon velocities are encoded

$$\dot{v} \sim 10^{-1} \sim \frac{q}{m_N}$$

Galilean invariant effective theory

 The most general Hermitian WIMP-nucleon interaction can be constructed from the for variables

$$ec{S}_{\chi}$$
 $ec{S}_{N}$ $ec{v}^{\perp}$ $\dfrac{q}{m_{N}}$

This interaction (filter #1) can be constructed to 2nd in velocities

$$H_{ET} = \begin{bmatrix} a_1 + a_2 \ \vec{v}^{\perp} \cdot \vec{v}^{\perp} + a_5 \ i \vec{S}_{\chi} \cdot \left(\frac{\vec{q}}{m_N} \times \vec{v}^{\perp} \right) \end{bmatrix} + \vec{S}_N \cdot \left[a_3 \ i \frac{\vec{q}}{m_N} \times \vec{v}^{\perp} + a_4 \ \vec{S}_{\chi} + a_6 \ \frac{\vec{q}}{m_N} \vec{S}_{\chi} \cdot \frac{\vec{q}}{m_N} \right]$$

$$+ \left[a_8 \ \vec{S}_{\chi} \cdot \vec{v}^{\perp} \right] + \vec{S}_N \cdot \left[a_7 \ \vec{v}^{\perp} + a_9 \ i \frac{\vec{q}}{m_N} \times \vec{S}_{\chi} \right]$$
 (parity odd)
$$+ \left[a_{11} \ i \vec{S}_{\chi} \cdot \frac{\vec{q}}{m_N} \right] + \vec{S}_N \cdot \left[a_{10} \ i \frac{\vec{q}}{m_N} + a_{12} \ \vec{v}^{\perp} \times \vec{S}_{\chi} \right]$$
 (time and parity odd)
$$+ \vec{S}_N \cdot \left[a_{13} \ i \frac{\vec{q}}{m_N} \vec{S}_{\chi} \cdot \vec{v}^{\perp} + a_{14} \ i \vec{v}^{\perp} \ \vec{S}_{\chi} \cdot \frac{\vec{q}}{m_N} \right]$$
 (time odd)

The coefficients represent the information that survive at low energy from a semi-infinite set of high-energy theories

Galilean invariant effective theory

 The most general Hermitian WIMP-nucleon interaction can be constructed from the for variables (previous physics)

$$\vec{S}_{\chi}$$
 \vec{S}_{N} \vec{v}^{\perp} $\frac{q}{m_{N}}$ argument defines this dimensionless parameter)

This interaction (filter #1) can be constructed to 2nd in velocities

$$\begin{split} H_{ET} = & \left[a_1 + a_2 \ \vec{v}^\perp \cdot \vec{v}^\perp + a_5 \ i \vec{S}_\chi \cdot \left(\frac{\vec{q}}{m_N} \times \vec{v}^\perp \right) \right] + \vec{S}_N \cdot \left[a_3 \ i \frac{\vec{q}}{m_N} \times \vec{v}^\perp + a_4 \ \vec{S}_\chi + a_6 \ \frac{\vec{q}}{m_N} \vec{S}_\chi \cdot \frac{\vec{q}}{m_N} \right] \\ + & \left[a_8 \ \vec{S}_\chi \cdot \vec{v}^\perp \right] + \vec{S}_N \cdot \left[a_7 \ \vec{v}^\perp + a_9 \ i \frac{\vec{q}}{m_N} \times \vec{S}_\chi \right] \\ + & \left[a_{11} \ i \vec{S}_\chi \cdot \frac{\vec{q}}{m_N} \right] + \vec{S}_N \cdot \left[a_{10} \ i \frac{\vec{q}}{m_N} + a_{12} \ \vec{v}^\perp \times \vec{S}_\chi \right] \\ + & \vec{S}_N \cdot \left[a_{13} \ i \frac{\vec{q}}{m_N} \vec{S}_\chi \cdot \vec{v}^\perp + a_{14} \ i \vec{v}^\perp \ \vec{S}_\chi \cdot \frac{\vec{q}}{m_N} \right] \end{aligned} \tag{time and parity odd)}$$

The coefficients represent the information that survive at low energy from a semi-infinite set of high-energy theories

We can then embed this in the nucleus (filter #2) to find what information survives, accessible to experiment.

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) \ W_i(q^2 b^2)$$

We can then embed this in the nucleus (filter #2) to find what information survives, accessible to experiment.

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_{i} R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2 b^2)$$

WIMP tensor: contains all of the DM particle physics

depends on two "velocities"

$$\vec{v}^{\perp 2} \sim 10^{-6}$$
 $\frac{\vec{q}}{m_N}^2 \sim \langle v_{\text{internucleon}} \rangle^2 \sim 10^{-2}$

We can then embed this in the nucleus (filter #2) to find what information survives, accessible to experiment.

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2b^2)$$

Nuclear tensor:

"nuclear knob" that can be turned by the experimentalists to deconstruct dark matter

Game - vary the W_i to determine the R_i : change the nuclear charge, spin, isospin, and any other relevant nuclear properties that can help

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2b^2)$$

$$W_1 \sim \langle J | \sum_{i=1}^A 1(i) | J \rangle^2$$
.

take q→0, suppress isospin

the S.I. response

contributes for J=0 nuclear targets

take $q \rightarrow 0$,

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i (\vec{v}^{\perp\,2}, \frac{\vec{q}^{\,2}}{m_N^2}) \underbrace{W_i (q^2 b^2)}_{W_i (q^2 b^2)}$$
 take q \rightarrow 0, suppress isospin
$$W_2 \sim \langle J | \sum_{i=1}^A \hat{q} \cdot \vec{\sigma}(i) \; |J\rangle^2$$

the S.D. response (J>0) but split into two components, as the longitudinal and transverse responses are independent, coupled to different particle physics

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2b^2)$$

take q→0, suppress isospin

$$W_4 \sim \frac{q^2}{m_N^2} \langle J | \sum_{i=1}^A \vec{\ell}(i) | J \rangle^2$$

A second type of vector (requires J>0) response, with selection rules very different from the spin response

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2b^2)$$

take q→0, suppress isospin

$$W_5 \sim \frac{q^2}{m_N^2} \langle J | \sum_{i=1}^A \vec{\sigma}(i) \cdot \vec{\ell}(i) | J \rangle^2$$

A second type of scalar response, with coherence properties very different from the simple charge operator

$$\frac{d\sigma}{dE_R} \sim G_F^2 \sum_i R_i(\vec{v}^{\perp 2}, \frac{\vec{q}^2}{m_N^2}) W_i(q^2b^2)$$

$$W_6 \sim \frac{q^2}{m_N^2} \langle J | \sum_{i=1}^A \left[\vec{r}(i) \otimes \left(\vec{\sigma}(i) \times \frac{1}{i} \vec{\nabla}(i) \right)_1 \right]_2 |J\rangle^2$$

A exotic tensor response: in principle interactions can be constructed where no elastic scattering occurs unless J is at least I

Response $\times \left[\frac{4\pi}{2J_i+1}\right]^{-1}$	Leading Multipole	Long-wavelength Limit	Response Type
$\sum_{J=0,2,\dots}^{\infty} \langle J_i M_{JM} J_i \rangle ^2$	$M_{00}(q\vec{x}_i)$	$\frac{1}{\sqrt{4\pi}}1(i)$	M_{JM} : Charge
$\sum_{\substack{J=1,3,\dots\\\infty}}^{\infty} \langle J_i \Sigma_{JM}'' J_i\rangle ^2$	$\Sigma_{1M}^{\prime\prime}(q\vec{x}_i)$	$rac{1}{2\sqrt{3\pi}}\sigma_{1M}(i)$	$L_{JM}^5:$ Axial Longitudinal
$\sum_{\substack{J=1,3,\ldots\\\infty}}^{\infty} \langle J_i \Sigma'_{JM} J_i\rangle ^2$	$\Sigma'_{1M}(q\vec{x}_i)$	$rac{1}{\sqrt{6\pi}}\sigma_{1M}(i)$	$T_{JM}^{\mathrm{el5}}: \mathrm{Axial}$ Transverse Electric
$\sum_{J=1,3,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Delta_{JM} J_i \rangle ^2$	$\frac{q}{m_N} \Delta_{1M}(q\vec{x}_i)$	$-rac{q}{2m_N\sqrt{6 au}}\ell_{1M}(i)$	T_{JM}^{mag} : Transverse Magnetic
$\sum_{J=0,2,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Phi_{JM}'' J_i\rangle ^2$	$\frac{q}{m_N}\Phi_{00}''(q\vec{x}_i)$	$-rac{q}{3m_N\sqrt{4\pi}}ec{\sigma}(i)\cdotec{\ell}(i)$	L_{JM} : Longitudinal
$\sum_{J=2,4,\dots}^{\infty} \langle J_i \frac{q}{m_N} \tilde{\Phi}'_{JM} J_i\rangle ^2$	$\frac{\frac{q}{m_N}\Phi_{2M}^{\prime\prime}(q\vec{x}_i)}{\frac{q}{m_N}\tilde{\Phi}_{2M}^{\prime}(q\vec{x}_i)}$	$-\frac{q}{m_N\sqrt{30\pi}} \left[x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i}\vec{\nabla})_1 \right]_{2M}$ $-\frac{q}{m_N\sqrt{20\pi}} \left[x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i}\vec{\nabla})_1 \right]_{2M}$	$T_{JM}^{ m el}$: Transverse Electric

† Full responses: operators familiar from standard semileptonic weak interactions

Response $\times \left[\frac{4\pi}{2J_i+1}\right]^{-1}$	Leading Multipole	Long-wavelength Limit	Response Type
$\sum_{\substack{J=0,2,\dots\\\infty}}^{\infty} \langle J_i M_{JM} J_i\rangle ^2$	$M_{00}(qec{x}_i)$	$\frac{1}{\sqrt{4\pi}}1(i)$	M_{JM} : Charge
$\sum_{J=1,3,\dots}^{\infty} \langle J_i \Sigma_{JM}'' J_i \rangle ^2$	$\Sigma_{1M}''(q\vec{x}_i)$	$rac{1}{2\sqrt{3\pi}}\sigma_{1M}(i)$	$L_{JM}^5:$ Axial Longitudinal
$\sum_{\substack{J=1,3,\dots\\\infty}}^{\infty} \langle J_i \Sigma'_{JM} J_i\rangle ^2$	$\Sigma'_{1M}(q\vec{x}_i)$	$rac{1}{\sqrt{6\pi}}\sigma_{1M}(i)$	$T_{JM}^{ m el5}:$ Axial Transverse Electric
$\sum_{J=1,3,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Delta_{JM} J_i \rangle ^2$	$ \frac{q}{m_N} \Delta_{1M}(q\vec{x}_i) $	$-rac{q}{2m_N\sqrt{6 au}}\ell_{1M}(i)$	T_{JM}^{mag} : Transverse Magnetic
$\sum_{J=0,2,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Phi_{JM}'' J_i \rangle ^2$	$\frac{q}{m_N}\Phi_{00}''(q\vec{x}_i)$	$-rac{q}{3m_N\sqrt{4\pi}}ec{\sigma}(i)\cdotec{\ell}(i)$	L_{JM} : Longitudinal
$\sum_{J=2,4,\dots}^{\infty} \langle J_i \frac{q}{m_N} \tilde{\Phi}'_{JM} J_i \rangle ^2$		$-\frac{q}{m_N\sqrt{30\pi}} \left[x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i} \vec{\nabla})_1 \right]_{2M}$ $-\frac{q}{m_N\sqrt{20\pi}} \left[x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i} \vec{\nabla})_1 \right]_{2M}$	$T_{JM}^{ m el}$: Transverse Electric

consistent with our symmetry arguments

Response $\times \left[\frac{4\pi}{2J_i+1}\right]^{-1}$	Leading Multipole	Long-wavelength Limit	Response Type
$\sum_{J=0,2,\dots}^{\infty} \langle J_i M_{JM} J_i \rangle ^2$	$M_{00}(q\vec{x}_i)$	$rac{1}{\sqrt{4\pi}}1(i)$	M_{JM} : Charge
$\sum_{J=1,3,\dots}^{\infty} \langle J_i \Sigma_{JM}'' J_i \rangle ^2$	$\Sigma_{1M}^{\prime\prime}(q\vec{x}_i)$	$rac{1}{2\sqrt{3\pi}}\sigma_{1M}(i)$	$L_{JM}^5: Axial$ Longitudinal
$\sum_{\substack{J=1,3,\ldots\\\infty}}^{\infty} \langle J_i \Sigma'_{JM} J_i \rangle ^2$	$\Sigma'_{1M}(q\vec{x}_i)$	$rac{1}{\sqrt{6\pi}}\sigma_{1M}(i)$	$T_{JM}^{\mathrm{el5}}: \mathrm{Axial}$ Transverse Electric
$\sum_{J=1,3,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Delta_{JM} J_i \rangle ^2$	$\frac{q}{m_N} \Delta_{1M}(q\vec{x}_i)$	$-rac{q}{2m_N\sqrt{6 au}}\ell_{1M}(i)$	T_{JM}^{mag} : Transverse Magnetic
$\sum_{J=0,2,\dots}^{\infty} \langle J_i \frac{q}{m_N} \Phi_{JM}'' J_i\rangle ^2$	$\frac{q}{m_N}\Phi_{00}''(q\vec{x}_i)$	$-rac{q}{3m_N\sqrt{4\pi}}ec{\sigma}(i)\cdotec{\ell}(i)$	L_{JM} : Longitudinal
$\sum_{J=2,4,\dots}^{\infty} \langle J_i \frac{q}{m_N} \tilde{\Phi}'_{JM} J_i\rangle ^2$	$\frac{\frac{q}{m_N}\Phi_{2M}^{"}(q\vec{x}_i)}{\frac{q}{m_N}\tilde{\Phi}_{2M}^{'}(q\vec{x}_i)}$	$-\frac{q}{m_N\sqrt{30\pi}} [x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i}\vec{\nabla})_1]_{2M}$ $-\frac{q}{m_N\sqrt{20\pi}} [x_i \otimes (\vec{\sigma}(i) \times \frac{1}{i}\vec{\nabla})_1]_{2M}$	$T_{JM}^{ m el}$: Transverse Electric

Two interference terms: M - L

Tel5 - Tmag just as in neutrino scattering

$$\begin{split} \frac{d\sigma}{d\Omega} &\sim \frac{4\pi}{2J_i + 1} \sum_{\tau = 0, 1} \sum_{\tau' = 0, 1} \left\{ R_C^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 0, 2, \dots}^{\infty} \langle J_i || \; M_{J;\tau}(q) \; || J_i \rangle \langle J_i || \; M_{J;\tau'}(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_L^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 0, 2, \dots}^{\infty} \langle J_i || \; \Phi_{J;\tau}''(q) \; || J_i \rangle \langle J_i || \; \Phi_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_L^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 0, 2, \dots}^{\infty} \langle J_i || \; \Phi_{J;\tau}''(q) \; || J_i \rangle \langle J_i || \; \Phi_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Phi_{J;\tau}''(q) \; || J_i \rangle \langle J_i || \; \Phi_{J;\tau'}''(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}''(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}''(q) \; || J_i \rangle \\ &+ R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \\ &+ \frac{\vec{q}^{\; 2}}{m_N^2} \; R_T^{\tau\tau'}(\vec{v}_T^{\perp 2}, \frac{\vec{q}^{\; 2}}{m_N^2}) \sum_{J = 1, 3, \dots}^{\infty} \langle J_i || \; \Sigma_{J;\tau}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \langle J_i || \; \Sigma_{J;\tau'}'(q) \; || J_i \rangle \langle J_$$

experimentalists have all of these nuclear "knobs" to turn

$$\begin{split} \frac{d\sigma}{d\Omega} \sim \frac{4\pi}{2J_{i}+1} \sum_{\tau=0,1} \sum_{\tau'=0,1} \left\{ &R_{C}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=0,2,\dots}^{\infty} \langle J_{i} || \; M_{J;\tau}(q) \; || J_{i} \rangle \langle J_{i} || \; M_{J;\tau'}(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{L}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=0,2,\dots}^{\infty} \langle J_{i} || \; \Phi_{J;\tau}''(q) \; || J_{i} \rangle \langle J_{i} || \; \Phi_{J;\tau'}''(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{L/C}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=0,2,\dots}^{\infty} \langle J_{i} || \; \Phi_{J;\tau}''(q) \; || J_{i} \rangle \langle J_{i} || \; \Phi_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Phi_{J;\tau}''(q) \; || J_{i} \rangle \langle J_{i} || \; \Phi_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ R_{L^{5}}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Sigma_{J;\tau}''(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Sigma_{J;\tau}'(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Sigma_{J;\tau}'(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Delta_{J;\tau}(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Delta_{J;\tau}(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Delta_{J;\tau}(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Delta_{J;\tau}(q) \; || J_{i} \rangle \langle J_{i} || \; \Sigma_{J;\tau'}'(q) \; || J_{i} \rangle \\ &+ \frac{\vec{q}^{\;2}}{m_{N}^{2}} \; R_{T}^{\tau\tau'}(\vec{v}_{T}^{\perp 2}, \frac{\vec{q}^{\;2}}{m_{N}^{2}}) \sum_{J=1,3,\dots}^{\infty} \langle J_{i} || \; \Delta_{J;\tau}$$

to extract the low-energy DM information embedded in the DM responses

The coefficients are what one "measures." They define the particle physics that can be mapped back to high energies, to constrain models

$$\begin{array}{lll} R_{M}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{c_{1}^{\tau}c_{1}^{\tau'}} + \frac{j_{\chi}(j_{\chi}+1)}{3} \left[\frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{5}^{\tau} c_{5}^{\tau'} + \vec{v}_{T}^{\perp2} c_{8}^{\tau} c_{8}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{11}^{\tau} c_{11}^{\tau'} \right] \\ R_{\Phi''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{3}^{\tau} c_{3}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{12} \left(c_{12}^{\tau} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) \left(c_{12}^{\tau'} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) \\ R_{\Phi''M}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & c_{3}^{\tau}c_{1}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{3} \left(c_{12}^{\tau} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) c_{11}^{\tau'} \\ R_{\Phi''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{j_{\chi}(j_{\chi}+1)}{12} \left[c_{12}^{\tau}c_{12}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{13}^{\tau}c_{13}^{\tau'} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{12} \left[c_{4}^{\tau}c_{4}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{12}^{\tau^{2}} c_{13}^{\tau'} \vec{v}_{13}^{\tau'} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{\vec{q}^{4}}{m_{N}^{2}} c_{6}^{\tau}c_{6}^{\tau'} + \vec{v}_{T}^{\perp2} c_{12}^{\tau}c_{12}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{13}^{\tau}c_{13}^{\tau'} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{7}^{\tau}c_{7}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{m_{N}^{2}} \left[c_{4}^{\tau}c_{4}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{13}^{\tau}c_{13}^{\tau'} \right]} \\ R_{\Delta}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{2m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{3}^{\tau}c_{3}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{2} \left(c_{12}^{\tau} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) \left(c_{12}^{\tau'} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau'} \right) + \underbrace{\frac{\vec{q}^{2}}{2m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{14}^{\tau}c_{14}^{\tau'}} \right]} \\ R_{\Delta}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{2m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{13}^{\tau} c_{15}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{$$

The point-nucleus world is a very simple one

Generally any derivative coupling is seen most easily in the new responses

$$\begin{array}{lll} R_{M}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{c_{1}^{\tau}c_{1}^{\tau'}} + \frac{j_{\chi}(j_{\chi}+1)}{3} \left[\frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{5}^{\tau} c_{5}^{\tau'} + \vec{v}_{T}^{\perp2} c_{8}^{\tau} c_{8}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{11}^{\tau} c_{11}^{\tau} \right] \\ R_{\Phi''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{3}^{\tau} c_{3}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{12} \left(c_{12}^{\tau} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) \left(c_{12}^{\tau'} - \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) \\ R_{\Phi''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & c_{3}^{\tau}c_{1}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{3} \left(c_{12}^{\tau} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{15}^{\tau} \right) c_{11}^{\tau} \\ R_{\Phi''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{j_{\chi}(j_{\chi}+1)}{12} \left[c_{12}^{\tau}c_{12}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} c_{13}^{\tau}c_{13}^{\tau} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{j_{\chi}(j_{\chi}+1)}{12} \left[c_{4}^{\tau}c_{4}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{13}^{\tau}c_{13}^{\tau'} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{\vec{q}^{4}}{m_{N}^{2}} c_{5}^{\tau}c_{5}^{\tau'} + \vec{v}_{T}^{\perp2} c_{12}^{\tau}c_{12}^{\tau'} + \frac{\vec{q}^{2}}{m_{N}^{2}} \vec{v}_{T}^{\perp2} c_{13}^{\tau}c_{13}^{\tau'} \right]} \\ R_{\Sigma''}^{\tau\tau'}(\vec{v}_{T}^{\perp2},\frac{\vec{q}^{2}}{m_{N}^{2}}) & = & \underbrace{\frac{\vec{q}^{2}}{4m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{10}^{\tau}c_{10}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{10}^{\tau'}c_{10}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{15}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{13}^{\tau'}c_{13}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{13}^{\tau'}c_{13}^{\tau'} + \frac{\vec{v}_{T}^{\perp2}}{m_{N}^{2}} c_{10}^{\tau'}c_{10}^{\tau'} + \frac{\vec{v}_{T}^{\perp$$

Observations:

- The set of operators found here map on to the ones necessary in describing known SM electroweak interactions
- ES can in principle give us 8 constraints on DM interactions
- This argues for a variety of detectors or at least, continued development of a variety of detector technologies: G2???
- There are a significant number of relativistic operators that reduce in leading order to the new operators
- $\hfill\Box$ Power counting -- e.g., $1~{\rm VS}~q/m_N$ -- does not always work as the associated dimensionless operator matrix elements differ widely
 - examples can be given

- $^{\Box}$ As noted before velocity-dependent interactions will generate a SI or SD coupling, but proportional to $\vec{v}^{\perp 2}$ and misleading
 - ▶ the predicted strength is 10⁻⁴ the actual strength
 - the associated SI/SD operator will have the wrong rank, e.g., predicted small SI when the dominant contribution is "spin"-dependent (e.g., governed by $\vec{\ell}(i)$)

Could be really confusing!

For illustration purposes only!

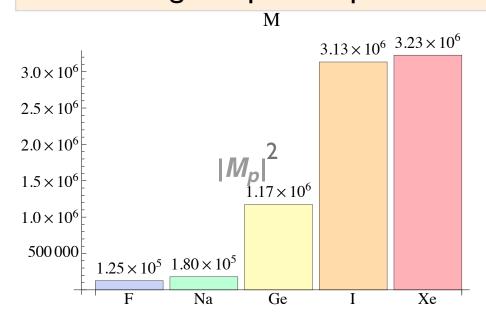
DAMA/LIBRA: Nal

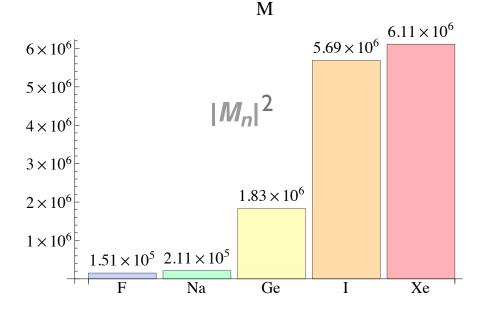
CoGENT: Ge

LUX: Xe

scalar charge responses: p vs. n S.I.

(normalized to natural abundance)

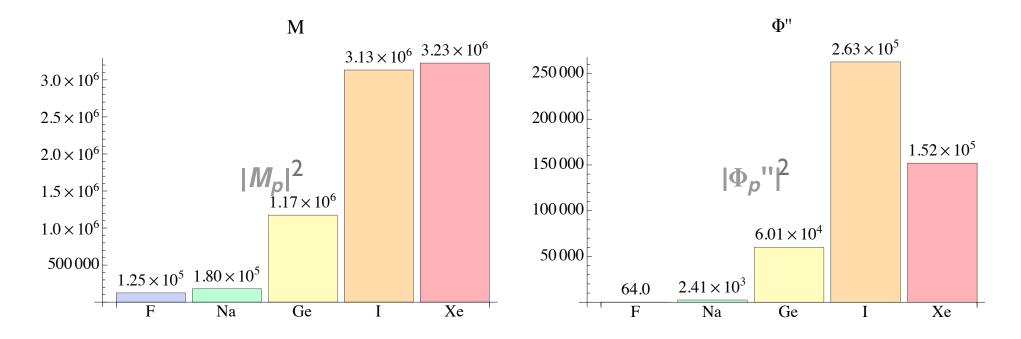




Standard SI sensitivities: LUX (Xe) > DAMA (NaI) > CDMS-Ge

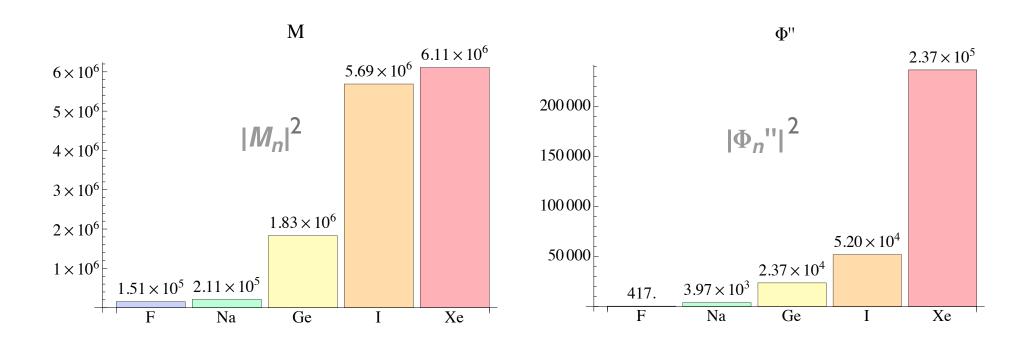
Little sensitivity to isospin (unless tuned)

Scalar operators, p: 1(i) vs $\vec{\sigma}(i) \cdot \vec{\ell}(i)$



LUX (Xe) ~ DAMA (NaI) ⇒ DAMA > LUX

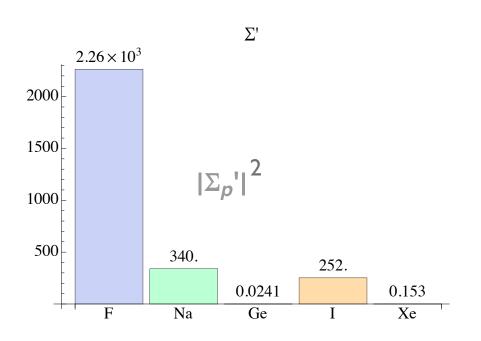
Scalar operators, n: $1(i) \text{ vs } \vec{\sigma}(i) \cdot \vec{\ell}(i)$

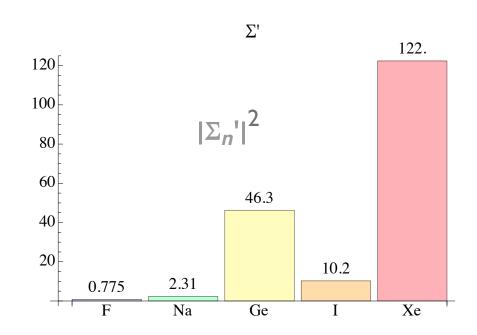


LUX (Xe) ~ DAMA (NaI) ⇒ DAMA < LUX

vector (transverse) spin response

(normalized to natural abundance)





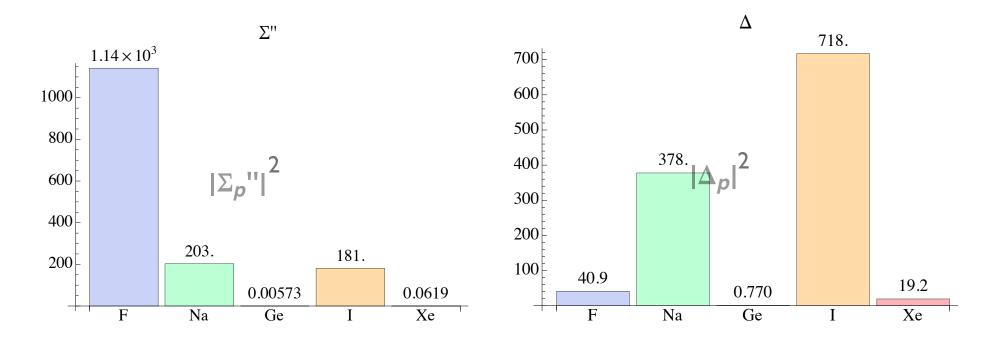
proton coupled: neutron coupled:

Picasso (F) > DAMA (Nal) » CDMS-Ge & LUX

LUX & CDMS-Ge » DAMA » Picasso

isospin

Vector, proton coupled: $\vec{\sigma}(i)$ vs. $\vec{\ell}(i)$

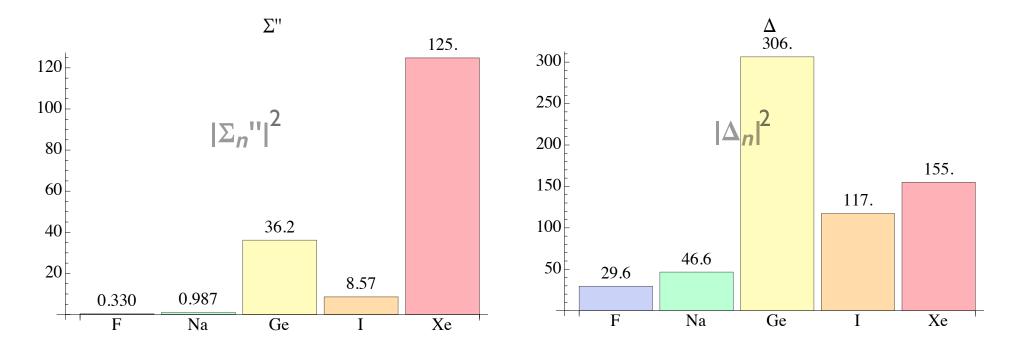


spin coupled: Picasso (F) > DAMA (Nal)

I-coupled coupled: DAMA (NaI) \gg Picasso (F) F: $2s_{1/2}$

orbital vs. spin ambiguity

Vector, neutron coupled: $\vec{\sigma}(i)$ vs. $\vec{\ell}(i)$



spin coupled: LUX > CDMS-Ge » DAMA

I-coupled coupled: CDMS-Ge > LUX ~ DAMA

orbital vs. spin ambiguity

Note on inelastic excitations:

Fortunate that key targets like Xe and Ge have unusual low-lying states: transitions can be inelastic

- □ Important because recoil energies for 100 GeV WIMP ≤ 250 keV
- Analyses again treated only in the SI/SD framework
- But we have seen that the elastic scattering "filter" is imposed by symmetry constraints of P and T: T does not constrain inelastic transitions
- There are familiar operators that effectively cannot contribute to elastic scattering, but can have large inelastic cross sections
 - most familiar of these is the axial charge operator $\vec{\sigma}(i) \cdot \vec{p}(i)$

<u>Summary</u>

- There is a lot of variability that can be introduced between detector responses by altering operators (and their isospins)
- Pairwise exclusion of experiments in general difficult
- But the bottom line is a favorable one: there is a lot more that can be learned from elastic scattering experiments than is apparent in conventional analysis
- This suggests we should do more experiments, not fewer
- When the first signals are seen, things will get very interesting: those nuclei that do not show a signal may be as important as those that do

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